

# A decade of short-period earthquake rupture histories from multi-array back-projection

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## Abstract

Teleseismic back-projection has emerged as a widely-used tool for understanding the rupture histories of large earthquakes. However, its application often suffers from artifacts related to the receiver array geometry, notably the ‘swimming’ artifact. We present a teleseismic back-projection method with multiple arrays and combined P and pP waveforms. The method is suitable for defining arrays ad-hoc in order to achieve a good azimuthal distribution for most earthquakes. We present a catalog of short-period rupture histories (0.5-2.0 Hz) including all 54 earthquakes from 2010 to 2021 with  $M_w$  [?] 7.5 and depth less than 200 km. The method provides semi-automatic estimates of rupture length, directivity, speed, and aspect ratio, which are related to the complexity of large ruptures. We determined short-period rupture length scaling relations that are in good agreement with previously published relations based on estimates of total slip. Rupture speeds were consistently in the sub-Rayleigh regime for thrust and normal earthquakes, whereas a tenth of strike-slip events propagated in the unstable supershear range. Many of the rupture histories exhibited complex behaviors such as rupture on conjugate faults, bilateral ruptures, and dynamic triggering by a P wave. For megathrust earthquakes, ruptures encircling asperities were frequently observed, with down-dip, up-dip, double encircling, and segmented patterns. Although there is a preference for short-period emissions to emanate from central and down-dip parts of the megathrust, emissions up-dip of the main asperities are more frequent than suggested by earlier results.

1     **A decade of short-period earthquake rupture histories**  
2                     **from multi-array back-projection**

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6   Germany.

7     **Key Points:**

- 8     • We provide a complete catalog of high-frequency rupture histories for  $M \geq 7.5$  events  
9         2010-2021.
- 10    • We develop a semi-automatic method for estimating rupture length, speed, direc-  
11         tivity, and aspect ratio.
- 12    • Asperity encircling ruptures and emissions up-dip of main asperity common in large  
13         megathrust earthquakes.

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16 rupture histories of large earthquakes. However, its application often suffers from artifacts  
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32 and down-dip parts of the megathrust, emissions up-dip of the main asperities are more  
33 frequent than suggested by earlier results.

## 34 **Plain Language Summary**

35 Back-projection is an earthquake imaging method based on seismic waveforms recorded re-  
36 motely at a group of seismometers (seismic array). We present a new approach that combines  
37 waveforms of multiple seismic arrays to improve spatial resolution. We provide a catalog of  
38 large earthquake rupture histories from 2010 to 2021, producing a map view of the high-  
39 frequency radiation emitted along the fault plane. The method estimates semi-automatic  
40 earthquake rupture length, speed, directivity, and aspect ratio. Based on these estimates,  
41 we determined scaling relations between earthquake magnitude and rupture length that  
42 agree with classical relationships. We identified several strike-slip earthquakes propagating  
43 at supershear, i.e., faster than the shear wave speed, the usual limit for self-sustaining rup-  
44 ture propagation. We observed complex rupture behaviors, e.g., multiple faults activated,  
45 bilateral ruptures, and triggering of the main phase of a rupture by a primary (P) wave  
46 from the earliest part of the rupture. For subduction earthquakes, high-frequency emission  
47 points were often observed forming a ring around the fault interface patches (asperities)  
48 where the main slip occurs. There is a preference for high-frequency emissions to emanate

49 from central and deeper parts of the subduction plate interface, but shallower emissions are  
50 more frequent than expected from previous literature.

## 1 Introduction

Since the 26 December 2004 Sumatra-Andaman ( $M_W$  9.1) earthquake, back-projection rupture imaging has become a complementary method to finite-fault source inversions for determining the kinematics of very large ruptures (e.g., Krüger & Ohrnberger, 2005; Ishii et al., 2005; Walker et al., 2005). Back-projection exploits the coherence of (usually) teleseismic P waveforms with limited prior assumptions on the fault geometry. Applications have targeted, for example, megathrust subduction earthquakes (e.g., Lay et al., 2010; Palo et al., 2014; Meng et al., 2011; Koper et al., 2011), aftershock detection (e.g., Kiser & Ishii, 2013; Feng et al., 2020; Tilmann et al., 2016), complex ruptures along multiple faults (e.g., Meng et al., 2012a; Lay et al., 2018; Ruppert et al., 2018), intermediate-depth earthquakes (e.g., Kiser et al., 2011; Ye et al., 2020), moderate size earthquakes (e.g., D’Amico et al., 2010; Taymaz et al., 2021), and near-field tsunami prediction (e.g., An & Meng, 2016; Xie & Meng, 2020), among others.

Back-projection is generally applied to large earthquakes. It uses short-period filtered waveforms, which can resolve details of the earthquake rupture better than long-period waveforms (e.g., Kiser & Ishii, 2017). The frequency band must be considered in the interpretation of the inferred rupture history. Short-period radiation is related to variations of slip and rupture velocity (e.g., Madariaga, 1983, 1977; Marty et al., 2019). This implies

69 that the back-projected rupture history can be related to fault heterogeneities. For example,  
70 subduction zone megathrust earthquakes were found to radiate short-period energy predom-  
71 inantly in deeper parts of the megathrust, down-dip of the maximum slip regions derived  
72 from finite-fault inversions, which are sensitive to the low-frequency behavior of the rupture  
73 (e.g., Lay et al., 2012; Yao et al., 2011, 2013). The apparently dominant down-dip short-  
74 period radiation is proposed to occur at the transition between brittle and ductile regions  
75 (e.g., Simons et al., 2011; Lay et al., 2012). Recently, Wang et al. (2020) correlated the 27  
76 February 2010 Maule ( $M_W$  8.8) earthquake short-period rupture with down-dip segmenta-  
77 tion at the base of the overriding mantle wedge. In contrast to this view of the along-dip  
78 segmentation of seismic emission during large earthquakes, Meng et al. (2018) provided  
79 the first evidence for an encircling rupture, with high-frequency emissions both up-dip and  
80 down-dip of the main slip patch for the 16 September 2015 Illapel ( $M_W$  8.3) earthquake.  
81 This observation is in contrast to earlier back-projection studies for this earthquake, which  
82 had shown the expected, predominantly down-dip rupture pattern (e.g., Melgar et al., 2016;  
83 Tilmann et al., 2016; Yin et al., 2016). It raises the question if this type of complex behavior  
84 is actually more common and just requires higher resolution rupture images.

85 Fracture mechanics predicts that instabilities such as earthquakes either propagate at  
86 sub-Rayleigh speed or exceed the shear wave speed (e.g., Burridge, 1973; Andrews, 1976;  
87 Das & Aki, 1977), where the latter is only expected for mode II cracks. In mode II cracks the

88 rupture propagates in the direction of displacement, which is also the direction of the initial  
89 shear stress resolved onto the fault. Such supershear earthquakes are associated with simple  
90 fault geometry and homogeneous stress-strength conditions (e.g., Bouchon et al., 2010), but  
91 also to damage zones under relatively low stress leading to unstable supershear ruptures,  
92 that is, between the shear wave speed and  $\sqrt{2}$  times the shear wave speed (e.g., Burridge  
93 et al., 1979). Many strike-slip earthquakes with supershear speeds have been reported in  
94 the literature, see Robinson et al. (2010) for a review, e.g., the 1979 Imperial Valley (e.g.,  
95 Archuleta, 1984; Spudich & Cranswick, 1984), 1999 Izmit (e.g., Bouchon et al., 2001), 2002  
96 Denali (e.g., Dunham & Archuleta, 2004; Walker & Shearer, 2009), 2001 Kunlunshan (e.g.,  
97 Bouchon & Vallée, 2003), 2013 Craig (e.g., Aderhold & Abercrombie, 2015; Yue et al., 2013)  
98 and the 2018 Palu earthquake (e.g., Bao et al., 2019; Socquet et al., 2019).

99 A disadvantage of the back-projection method is that the array configuration can cause  
100 notable artifacts in the recovered rupture (e.g., Meng et al., 2012b). Back-projection often  
101 leads to a persistent time-space trade-off of the earthquake rupture towards the seismic array,  
102 often called ‘swimming’ artifact. The swimming artifact arises from the low curvature of  
103 the time-distance travel time curve. Sources closer to the receivers but activated later will  
104 have the same arrival time as earlier sources farther away, resulting in a point source in  
105 space and time to appear as an extended source drifting towards the array. In animations  
106 showing the evolution of the back-projected energy with time, this slightly irregular drifting

107 looks like a swimming motion, giving the artifact its name. Because of the dependence on  
108 the array azimuth, it is easy to understand why the swimming artifact can be reduced by  
109 combining multiple array images.

110 Depth phases can cause additional artifacts in the form of ‘ghost’ emitters correspond-  
111 ing approximately to the bounce points of the surface-reflected phase, but they also can  
112 carry additional information. For large intermediate-depth earthquakes, the time-delay be-  
113 tween P and depth phases (e.g., pP and sP) allowed to improve the resolution in depth  
114 by combining both P and depth phase backprojections (e.g., Kiser et al., 2011). For more  
115 shallow earthquakes (40–100 km), depth phases can contribute significantly to uncertainties  
116 (e.g., Zeng et al., 2019). To our knowledge, however, a systematic imaging method that in-  
117 tegrates depth phases and multiple arrays for the shallow depth range has not been reported  
118 yet.

119 This study presents such a method and its application to derive a catalog of rupture  
120 histories of recent large earthquakes (2010 to October 2021) in the 0.5–2.0 Hz frequency  
121 range, which is complete for  $M_W \geq 7.5$  and depths less than 200 km. Specifically, we ex-  
122 tended the multi-array approach of Rössler et al. (2010) and included depth phases (for  
123 earthquakes deeper than 40 km) and weighted seismic array images. We provide an algo-  
124 rithm for automatically estimating rupture length, directivity, speed, and aspect ratio from

125 back-projection results. Short-period rupture lengths were used to calculate scaling relations  
126 and compare them to established relationships. The analysis focuses on complex ruptures  
127 and depth-varying short-period radiation for large subduction earthquakes and the detec-  
128 tion of supershear ruptures for strike-slip events. We also show that short-period ruptures  
129 encircling asperities, as observed for the 2015 Illapel earthquake in Chile, are frequent in  
130 subduction megathrust earthquakes. The results suggest that short-period rupture complex-  
131 ities, e.g., encircling rupture around slip patches, are related to asperity stress conditions  
132 (and seismogenic barriers) rather than the overall along-dip megathrust segmentation.

## 133 **2 Methods**

### 134 **2.1 Multi-Array Multi-Phase Back-Projection**

135 The back-projection method is similar to beamforming in maximizing the coherency  
136 of time-shifted waveforms at an array. Unlike in beamforming, there is no assumption of  
137 a planar wavefield. Instead, the time shifts are calculated from the predicted travel times  
138 for a grid spaced around the hypocenter. In practice, the grid is usually two-dimensional,  
139 chosen to be either a horizontal plane at the hypocentral depth, or a plane aligned with  
140 one of the nodal planes of the focal mechanism, or an *a priori* known fault surface, e.g.,  
141 the slab interface in subduction zones. The waveforms are back-projected onto this grid.  
142 The theoretical arrival of a target seismic wave (e.g., P wave) based on a reference velocity

143 model controls the beamforming delays. For each grid point, the resulting array beam is:

$$b_i(t) = \frac{1}{N_k} \sum_{k=1}^{N_k} u_k(t + t_{ik} + \Delta t_k), \quad (1)$$

144 where  $b_i(t)$  is the beam for the  $i$ th grid point,  $u_k$  the vertical component waveform recorded  
 145 at station  $k$ ,  $t_{ik}$  the travel time between the grid-point  $i$  and station  $k$  in a reference velocity  
 146 model, and  $\Delta t_k$  the station-specific static correction term accounting for differences between  
 147 the reference and true velocity model.

148 The station correction terms can, to a large extent, absorb the effect of 3D Earth  
 149 heterogeneities on arrival times. They are usually determined by cross-correlating the first  
 150 few seconds of the rupture recorded by each receiver. The resulting time-shifted arrivals are  
 151 then compared to those predicted for the catalog hypocenter, and the differences correspond  
 152 to the necessary correction terms. Thus, the back-projection image retrieves the rupture  
 153 nucleation at the catalog hypocenter by definition. For very large earthquakes, aftershocks  
 154 can alternatively be used to correct source-receiver paths away from the hypocenter (e.g.,  
 155 Ishii et al., 2007; Palo et al., 2014; Meng et al., 2016). The advantage over the hypocenter-  
 156 based calibration is that the location errors of several events are averaged, such that a  
 157 possible bias from mislocation of the mainshock hypocenter is reduced and furthermore  
 158 spatially varying station terms due to 3D structure effects can be accommodated. The  
 159 calibration with aftershocks is particularly important when the rupture pattern is compared

160 with aftershocks. In contrast, the mainshock calibration offers advantages in near-real-time  
 161 applications with automatic routines since only the hypocenter is required. For simplicity, in  
 162 this work, we adopt the hypocenter calibration, but it would be easy to adopt the method to  
 163 aftershock calibration. Alternatively, travel calculations could be performed in a 3D Earth  
 164 model (e.g., Liu et al., 2017).

165 In order to carry out the actual rupture tracking, the maxima of beamformed energy  
 166 ( $\mathbf{E}_i$ ) and semblance ( $\mathbf{S}_i$ ), defined in equations (2) and (3) below, are used to locate the most  
 167 intense emission at each time step. The energy represents the amount of radiation emitted,  
 168 and semblance provides a measure of the coherence of waveforms which is not affected by  
 169 the amplitudes of individual traces and is, therefore, more effective for tracking the location  
 170 of earthquake rupture (e.g., Neidell & Taner, 1971; Rössler et al., 2010; Palo et al., 2014).  
 171 For both measures, a time window of length  $W$  needs to be defined, which should contain  
 172 at least two periods of the longest period analyzed.

$$\mathbf{E}_i(t) = \int_t^{t+W} |\mathbf{b}_i(\tau)|^2 d\tau \quad (2)$$

$$\mathbf{S}_i(t) = \frac{1}{N} \frac{\mathbf{E}_i(t)}{\int_t^{t+W} [\sum_{k=1}^{N_k} \mathbf{u}_k^2(\tau + t_{ik} + \Delta t_k)] d\tau} \quad (3)$$

173 Both measures provide an image of the earthquake rupture. Given a grid of sources,  
 174 the tracking of the local semblance maxima provides a way to map the rupture propagation.  
 175 The energy peak in each time window provides a relative measure of the source time function

176 of short-period seismic energy, which is related to but not necessarily proportional to the  
 177 moment rate, e.g., as derived from finite fault solutions. We also note that the absolute  
 178 values of the energy function depend on the array configurations, and a comparison of the  
 179 absolute energy amplitudes is not physically meaningful unless the same arrays are utilized,  
 180 but that the time history for any given earthquake is related to the physical rupture process.  
 181 The time-integrated energy maps thus provide a summary view of the high-frequency energy  
 182 radiation.

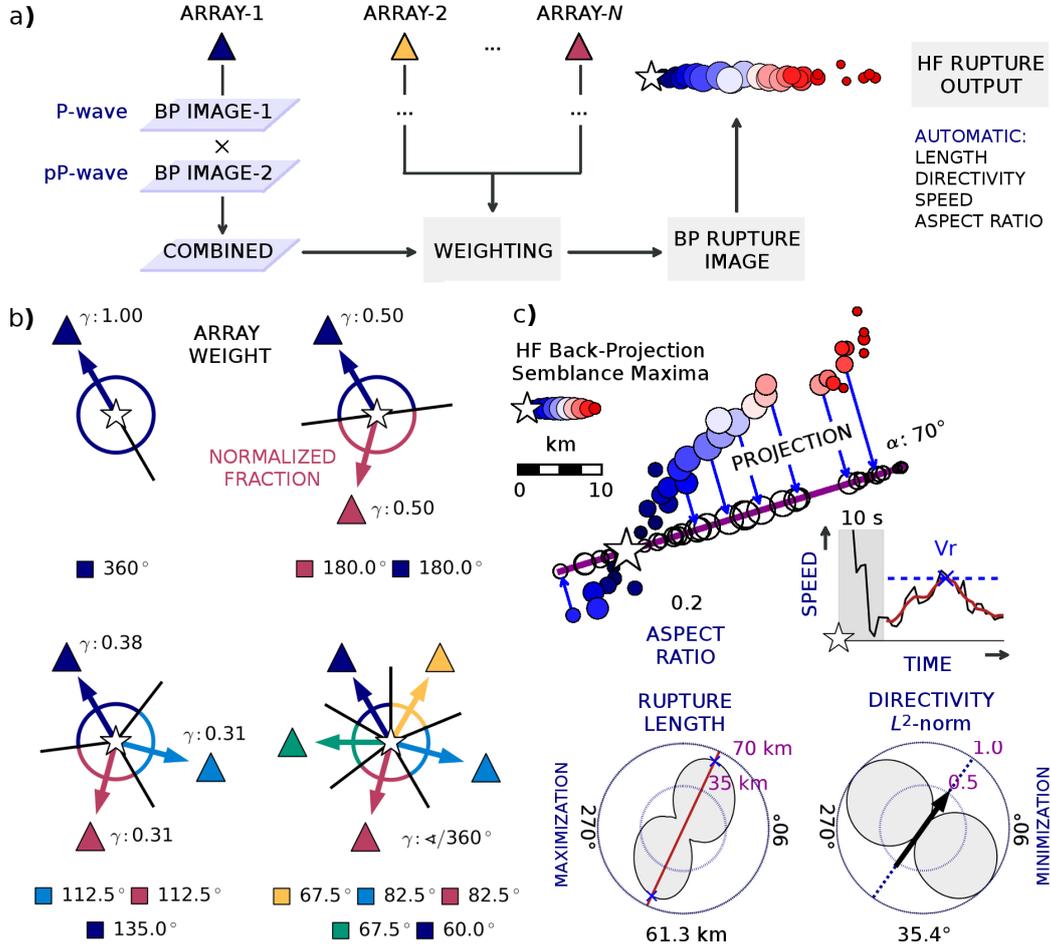
183 For the processing of data from multiple arrays, we initially follow the approach from  
 184 Rössler et al. (2010) where semblance and energy maps from  $N_a$  arrays are multiplied to  
 185 provide the rupture and energy radiated, respectively:

$$\hat{\mathbf{E}}_i(t) = \prod_{a=1}^{N_a} \mathbf{E}_i(t) \quad (4)$$

$$\hat{\mathbf{S}}_i(t) = \prod_{a=1}^{N_a} \mathbf{S}_i(t) \quad (5)$$

186 Here, we introduce two modifications: (i) we use exponents (equivalent to weights in log-  
 187 space) to balance the contributions of individual arrays based on their azimuthal distribution  
 188 in order to avoid artifacts due to clustering of arrays in certain azimuthal ranges, and (ii) for  
 189 earthquakes deeper than 40 km, we combine P and pP backprojections to reduce artifacts  
 190 related to the depth phase (see Fig. 1a). The modified expressions are:

$$\hat{\mathbf{E}}_i(t) = \prod_{a=1}^{N_a} \left[ \prod_{p=1}^{N_p} f(t - t_0) \cdot \mathbf{E}_i(t) \right]^{\gamma_a} \quad (6)$$



**Figure 1.** Multi-array multi-phase back-projection. a) Workflow for rupture imaging using P and pP waveform arrivals and multiple seismic arrays. pP arrivals are only included for earthquakes with depth  $\geq 40$  km. The last step contains the extraction of a few rupture parameters based on the timing and locations of rupture maxima. b) Array weighting. The sum of azimuthal half-angles between the target and its two neighboring arrays is proportional to the weights  $\gamma$ . c) Automatic estimation of the rupture aspect ratio, speed, length, and directivity; see text for details.

$$\hat{\mathbf{S}}_i(t) = \prod_{a=1}^{N_a} \left[ \prod_{p=1}^{N_p} f(t-t_0) \cdot \mathbf{S}_i(t) \right]^{\gamma_a} \quad (7)$$

191 We assign  $p = 1$  to P and  $p = 2$  to pP waveforms, and  $N_P$  is the number of phases used.

192 Here  $N_p = 1$  is used for shallow earthquakes and 2 for deep earthquakes, but the method

193 is open to experimenting with alternate seismic wave arrivals. The weighting exponent  $\gamma$

194 is set proportionally to the sum of the two half-angles between the azimuths of target and

195 neighboring arrays (see Fig. 1b), where the median of the array receiver coordinates is

196 assumed to be the reference location for the weighting estimation, and the normalization

197 is chosen such that  $\sum_i^{N_a} \gamma_i = 1$ . Therefore, for a single array  $\gamma = 1$ , and for two, we

198 always have  $\gamma = 0.5$ , but for irregularly distributed arrays, the weighting depends on the

199 distribution. The term  $f(t-t_0)$ , with  $t_0$  the origin time of the earthquake, is a taper function

200 adapted from Kiser et al. (2011) to mute the waveform before the first arrival when multiple

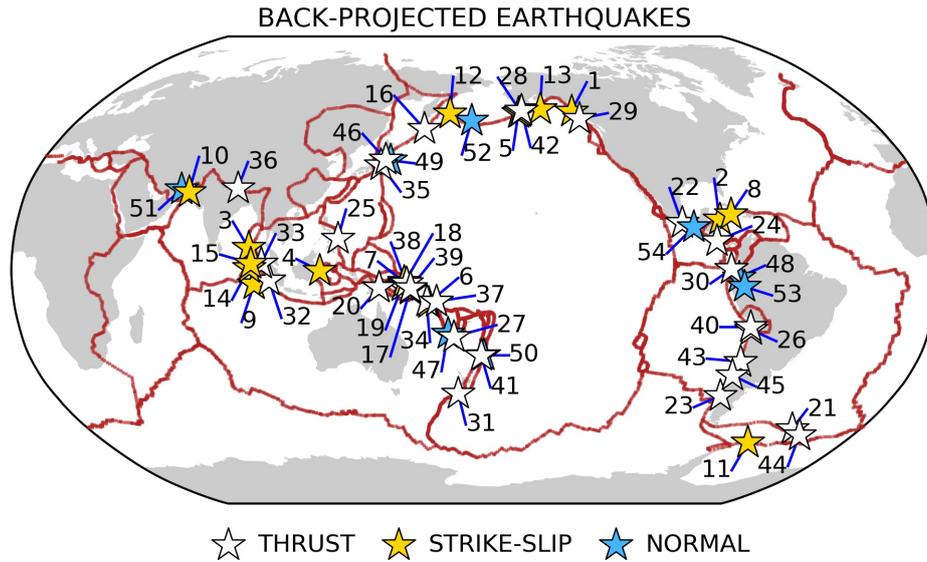
201 seismic phases are included:

$$f(\tau) = \begin{cases} 0 & \text{for } \tau \leq -T/2 \\ \frac{1}{2} [\cos(\frac{2\tau\pi}{T}) + 1] & \text{for } -T/2 < \tau < 0 \\ 1 & \text{for } \tau \geq 0 \end{cases} \quad (8)$$

202 If  $\tau = 0$ , the rupture has initiated, and the results are fully incorporated to image the source;

203 otherwise, they are suppressed.  $T$  is the period of the cosine taper function controlling the

204 transition between admitted and suppressed intervals.



**Figure 2.** Global distribution of back-projected earthquakes colored by focal mechanism. Labels designate the earthquake ID, with details given in Fig. 5 and Tables S1–S3. Red segments show major plate boundaries (Bird, 2003).

## 2.2 Processing details

Fifty-four large earthquakes in the time range 01/2010–10/2021 (all events with  $M_W \geq 7.5$ , depth  $\leq 200$  km according to Global CMT catalog, <https://www.globalcmt.org>) were back-projected using the multi-array multi-phase approach. Fig. 2 shows the distribution of these earthquakes color-coded by faulting type. Thrust earthquakes dominated (57%) the catalog, followed by strike-slip (28%), and normal (15%) faulting mechanisms. Tables S1, S2 and S3 in the supporting information present the earthquake source parameters for strike-slip, thrust, and normal faulting earthquakes, respectively.

213 The number of (ad-hoc) arrays weighted and combined in the analysis depended on their  
214 availability at epicentral distances between 30–100°, where the P waveforms are not subject  
215 to complications like triplications as observed at smaller distances. The array selection  
216 and weighing prioritize even azimuthal distribution, while as many arrays as possible are  
217 incorporated.

218 Dense arrays are formed depending on configuration and geometry of permanent broad-  
219 band networks, e.g., North America (e.g., US: United States National Seismic Network;  
220 AK: Alaska Regional Network), Japan (NIED Hi-net Network), Europe (many national  
221 and global networks distributed by ORFEUS-EIDA; Strollo et al. (2021)), and Africa (e.g.,  
222 AF: AfricaArray). Additionally, local and regional networks (often temporary deployments)  
223 are evaluated (e.g., data availability, waveform coherence) to form smaller aperture arrays  
224 and maximize the azimuthal coverage. The extent to which a good azimuthal distribution  
225 can be achieved therefore differs between the different regions, but also with the time and  
226 time gap after the event (as temporary network data are often only openly available a few  
227 years after the experiment). Earthquakes in Indonesia are favorably located to be imaged  
228 with several arrays, e.g., networks in Asia, Europe, Africa, Australia, and Antarctica. In  
229 contrast, earthquakes in northern and central Chile suffer from limited coverage, although  
230 still a somewhat reasonable azimuthal distribution can be achieved by combining networks

231 in North America, Africa, and Antarctica. For events in southernmost Chile, Australian  
232 networks supplement the coverage.

233 After downloading, the instrument response is removed from the waveforms. P ar-  
234 rival times are predicted based on the IASP91 velocity model (Kennett & Engdahl, 1991).  
235 Static corrections are determined by measuring the relative time shifts of first arrivals on  
236 bandpass-filtered (0.4–3.0 Hz) vertical velocity waveforms with the adaptive stacking method  
237 of Rawlinson and Kennett (2004) based on the first 15 s after the P-wave onset. This fre-  
238 quency band is a little wider than the band used for the back-projection (0.5–2.0 Hz) and  
239 optimizes coherence while enough high frequencies are retained for a precise alignment. We  
240 also removed anomalous traces that could impact the waveform stack during the adaptive  
241 stacking as part of the input quality control.

242 For non-subduction megathrust earthquakes, the target grid was placed on a plane at  
243 the hypocentral depth, with grid points every 5 km. For subduction megathrust earthquakes,  
244 the grid followed the depth variations of the SLAB1.0 model (Hayes et al., 2012). For  
245 bilateral ruptures, we additionally mapped semblance maxima over a pre-defined sub-region  
246 of the grid to probe secondary rupture patterns. After visual inspection, we examined  
247 bilateral ruptures for the 27 February 2010 Maule ( $M_W$  8.8), 11 March 2011 Tohoku-Oki  
248 ( $M_W$  9.1), and 17 July 2017 Komandorsky Islands ( $M_W$  7.8) earthquakes. Similarly, for the

249 16 September 2015 Illapel ( $M_W$  8.3) earthquake, we tracked simultaneous up-dip and down-  
250 dip emissions (relative to the main slip area) by introducing a separate grid for the up-dip  
251 area. Although the 12 August 2021 South of Sandwich Islands earthquake was reported as  
252 a doublet by several agencies, we simply processed it as a single event, using the GEOFON  
253 hypocenter of the first event for calibration.

254 The back-projection considered a time window  $W$  of 6 s, chosen to be three times the  
255 dominant period, moved forward in 1 s increments. Longer values for  $W$  would have over-  
256 smoothed the rupture image. The back-projection frequency band (0.5–2.0 Hz) corresponded  
257 to the highest range for which consistently sufficient waveform coherency for rupture imaging  
258 has been obtained in prior studies (e.g., Palo et al., 2014; Meng et al., 2015, 2016, 2018).  
259 Finally, the end of the rupture was determined manually; frequent reactivation of earlier  
260 peaks or scattered semblance maxima are indicators for the end of the rupture. The apparent  
261 source time function of radiated energy was also considered, as small values compared to its  
262 peak also indicate the end of the rupture.

### 263 **2.3 Estimation of basic source parameters**

264 We have fully automated the earthquake rupture length, directivity, and speed based  
265 on the obtained rupture image from semblance peaks. For estimating rupture length (see  
266 Fig. 1c), firstly, candidate lengths  $L$  are estimated by the projection of the semblance

267 maxima (blue-red circles in Fig. 1c) on lines passing through the epicenter (white star)  
268 for all azimuths 0–180° (magenta line and open black circles show the realization for a 70°  
269 azimuth). The maximum value of  $L$  over all azimuths is chosen as the rupture length (red  
270 line in the bottom left subplot in Fig. 1c). Directivity is measured similarly, with lines of  
271 all azimuths 0–180° pivoting through the epicenter. However, the quantity minimized here  
272 is the sum of the squares of the perpendicular distances of all semblance maxima to the line  
273 (i.e., average squared lengths of blue arrows in Fig. 1c, which are shown symbolically for a  
274 few semblance peaks only). For simple ruptures, the azimuths returned by the length and  
275 directivity measurements will be very similar, but the length estimate is controlled by the  
276 end points, whereas the directivity estimate is controlled by all points simultaneously. The  
277 ambiguity in the actual directivity is resolved by considering an imaginary line perpendicular  
278 to the rupture direction and passing through the epicenter; we then choose the directivity  
279 based on which side of this line more semblance maxima are found. Additionally, the aspect  
280 ratio of the rupture is defined by the quotient between the minimum and maximum length  
281 estimates.

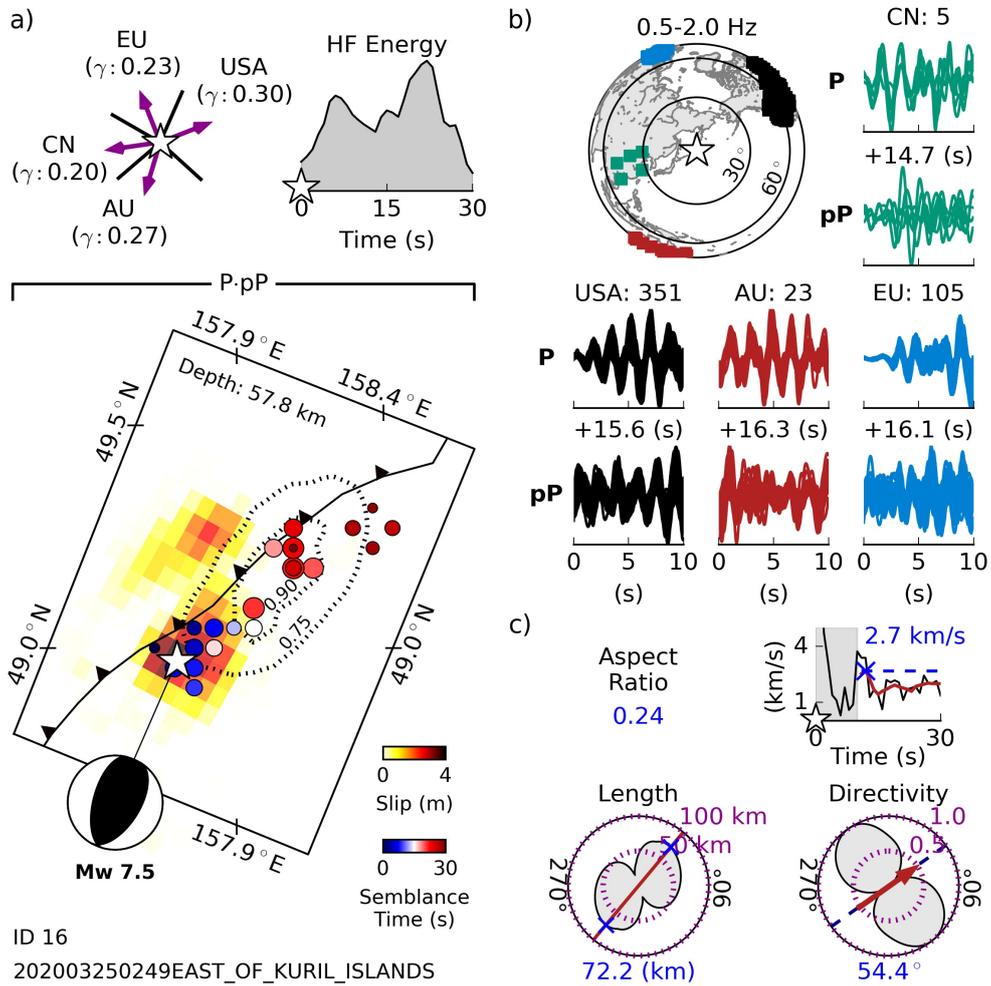
282 To determine the rupture speed, we first calculate estimates of instantaneous rupture  
283 velocities by dividing the distance between the epicenter and a subsequent semblance peak  
284 by the time elapsed since the origin time, which really represents the average rupture velocity  
285 from nucleation to the current time step. The resulting time series is strongly affected by

286 location uncertainties for small distances and times but stabilizes quickly (see Fig. 1c). The  
287 time series of rupture velocities is smoothed within a 4 s wide smoothing window, and its  
288 peak value is then taken as the event rupture speed; maxima during the first unstable 10 s are  
289 ignored. Although strictly speaking, this estimate represents the maximum average velocity  
290 (where the average is taken over all preceding times), it represents a reasonable estimate  
291 for most ruptures. Based on theoretical considerations, for supershear ruptures on simple  
292 faults, supershear velocities are thought to be attained quickly and maintained until the end  
293 of the rupture (Das & Aki, 1977), such that they can be detected straightforwardly by this  
294 estimation procedure. Nevertheless, the estimated rupture velocities can be significantly  
295 lower than the peak rupture speeds obtained for complex ruptures with directional changes,  
296 particularly those with a slow start. For the 17 December 2016 Solomon Islands earthquake,  
297 instead of referring back to the epicenter for the whole rupture duration, we reset the  
298 reference point to an emission point later in the rupture.

### 299 **3 Results**

#### 300 **3.1 Example: 2020 Kuril Islands Earthquake**

301 We introduce the presentation of our results using the 25 March 2020 East of Kuril  
302 Islands ( $M_W$  7.5) earthquake as an example, see Fig. 3. The results for all earthquakes  
303 are presented in the supplementary material, Figures S2–S60, with a summary of derived



**Figure 3.** The 25 March 2020 East of Kuril Islands earthquake back-projection (0.5–2.0 Hz). a) Earthquake rupture image. Blue-red dots show semblance maxima tracking the earthquake rupture color-coded by time and scaled by energy radiated. Black dotted contours outline the short-period energy radiated (normalized to 1). The yellow-red polygons present the USGS-NEIC finite fault slip solution for comparison. Trench line derived from SLAB1.0 model (Hayes et al., 2012). Focal mechanism from Global CMT catalog. Inset: Array weights and energy radiated source time function. b) Multi-array configuration and time-shifted P and pP waveforms. c) Automatic rupture parameter determination; see Fig. 1c for further information on the format.

304 parameters in Fig. 5 and Table S4. We selected the Kuril Islands event because of the  
305 relatively simple rupture and because at the time of writing, to our knowledge, no other  
306 back-projection analysis had been published yet for this event. This event is an intraplate  
307 thrust earthquake, which probably was triggered by compressional bending stresses in the  
308 deep interior of the subducting Pacific plate (Ye et al., 2021).

309 The left inset of Fig. 3a shows the distribution of arrays, where the arrays in North  
310 America and Australia are weighted more strongly as they cover a larger backazimuthal  
311 range. In Fig. 3b the waveform coherency near the rupture initiation is visualized. It is  
312 generally good except for the pP phase at the China array. The main plot compares the  
313 back-projected rupture with the USGS-NEIC finite fault slip solution ([https://earthquake](https://earthquake.usgs.gov/earthquakes/eventpage/us70008fi4/finite-fault)  
314 [.usgs.gov/earthquakes/eventpage/us70008fi4/finite-fault](https://earthquake.usgs.gov/earthquakes/eventpage/us70008fi4/finite-fault)); for other events, fre-  
315 quently finite slip models from the literature are shown instead. Because this event is  
316 not a subduction megathrust earthquake, we back-projected onto a horizontal plane at 58  
317 km depth, the hypocentral depth of the event. While in theory, it might be desirable to  
318 use the focal mechanism to define an inclined plane for back-projection, in practice, it is  
319 not easy to determine which nodal plane is the fault plane, except for megathrust events.  
320 For the East of Kuril Islands earthquake, the USGS-NEIC finite fault solution preferred  
321 the south-east dipping plane, whereas the finite fault solution of Ye et al. (2021) favored  
322 the northwest-dipping fault plane. A further advantage of a horizontal plane is that this

323 assumption will not fail in case of more than one planar fault being activated (e.g., 2018  
324 Gulf of Alaska earthquake; Fig. S45). For the Kuril event, the main slip patch appeared  
325 close to the epicenter, and at the beginning of the rupture, both finite slip and short-period  
326 rupture spatially agreed. After 15 s, the high-frequency rupture propagated to the east of  
327 the secondary slip patch. This time period revealed the most energetic short-period emis-  
328 sion. Close to this area, the most intense aftershock activity was observed (see Fig. 1 of Ye  
329 et al., 2021). The final high-frequency emissions occurred even further to the east-northeast,  
330 apparently far from the finite-slip area, but also here aftershocks occurred nearby.

331       The rupture propagated unilaterally in a linear manner, making the length and direc-  
332 tivity estimation unambiguous (Fig. 3c). The implied strike direction of the rupture track  
333 is clockwise rotated by some 10–20° with respect to the strike direction indicated by the  
334 moment tensor. At face value, this would imply that the short-period emission points did  
335 not strictly propagate along strike but moved additionally up-dip (or down-dip, depending  
336 on which fault plane is the correct one).

337       We can compare the time history of emitted short-period energy for the Kuril Islands  
338 earthquake with the moment rate obtained from the USGS finite slip solution (highlighted  
339 in Fig. 4). Both show two peaks, and the timing of the first peak and subsequent trough  
340 agree. However, at short periods, the second peak of energy exceeded the first one, and the

341 rupture appeared to continue for longer. Interestingly, in the source time function estimated  
342 by Ye et al. (2021) (their Fig. 1), the moment rate of the second peak also exceeded the  
343 first one.

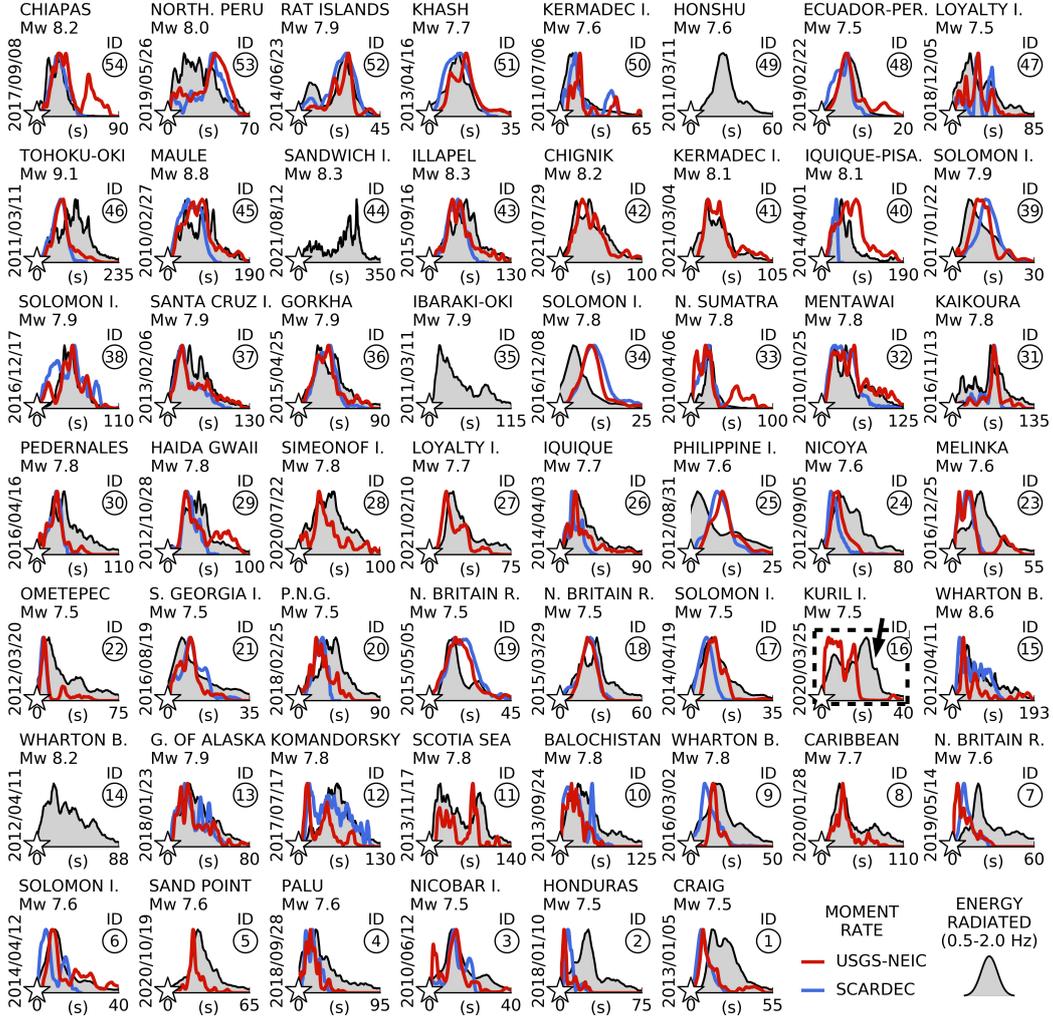
### 344 **3.2 Short-period source time functions**

345 The short-period energy time functions for all events are compared with USGS and  
346 SCARDEC (Vallée & Doeut, 2016) moment rates in Fig. 4. The time history of radiated  
347 high-frequency seismic energy frequently agree (e.g., 2021 Chignik, ID 42; 2020 Caribbean,  
348 ID 8; 2016 Kaikoura, ID 31; 2013 Scotia Sea, ID 11; 2012 Wharton Basin, ID 15; 2010  
349 Mentawai, ID 32) with the seismic moment rate functions, but it is sometimes apparently  
350 time-shifted, following (e.g., 2016 Melinka, ID 23; 2013 Craig, ID 1; 2011 Tohoku-Oki, ID  
351 46) or, less often, preceding (e.g., 2016 Solomon, ID 34; 2012 Philippine Islands, ID 25)  
352 the seismic moment rate. A few earthquakes presented a quite different shape (e.g., 2014  
353 Iquique-Pisagua, ID 40; 2019 Northern Peru, ID 53). However, it has to be noted that the  
354 USGS and SCARDEC moment rates also do not always agree. One example is the  $M_W$   
355 8.1 Iquique-Pisagua earthquake (ID 40; Fig. 4), where the SCARDEC estimate suggested  
356 a much shorter source time function. For the  $M_W$  8.2 Chiapas earthquake (ID 54; Fig. 4),  
357 SCARDEC (and the short-period energy function) completely lack a secondary peak seen  
358 in the USGS estimate. Because of this variety, it is difficult to judge whether the differences

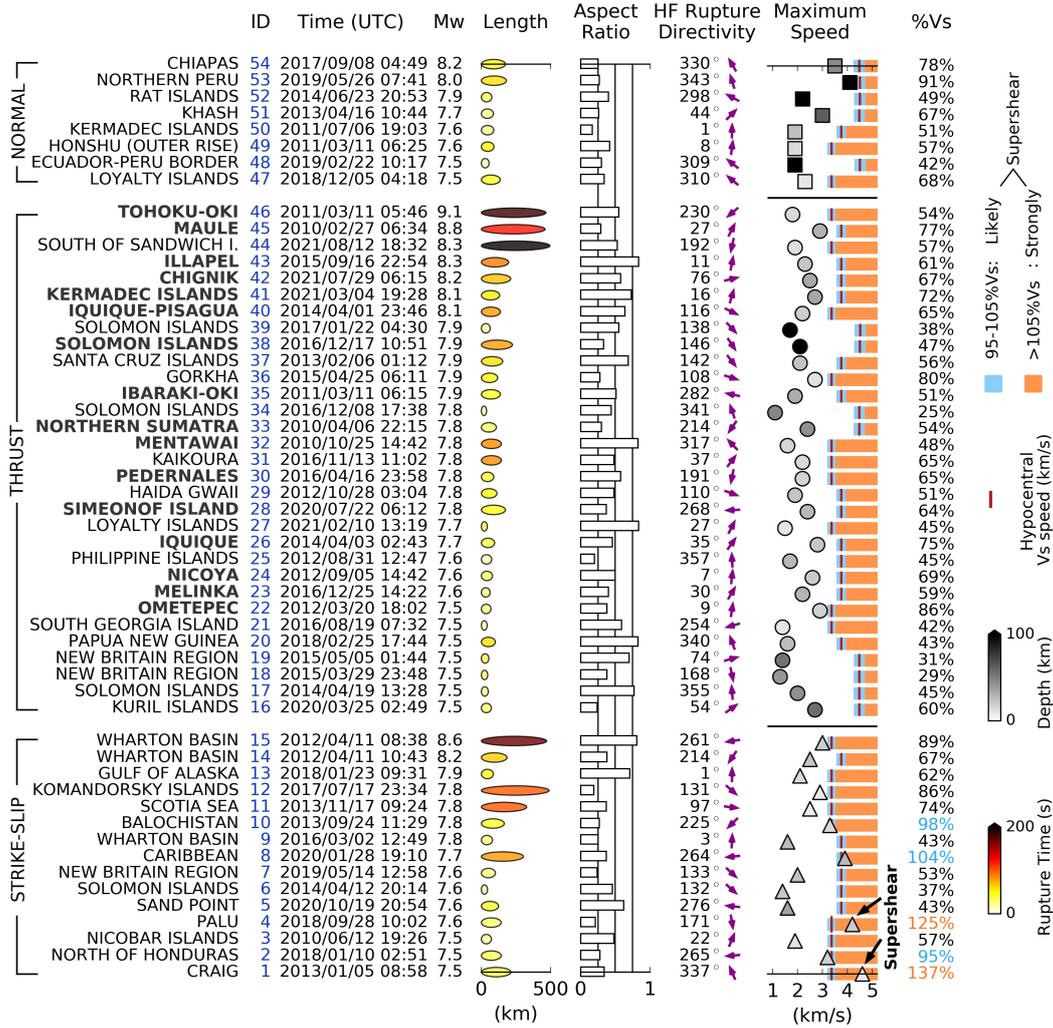
359 between moment rate and short-period energy function are physical, i.e., relate to the time-  
360 varying ratio of slip to high energy emissions, methodological artifacts, or both. A possible  
361 physical explanation is that the ratio of short-period to long-period spectral energy is known  
362 to vary significantly between earthquakes (or equivalently, the seismic energy to moment  
363 relation, which is related to the variability in stress drop). It seems likely that this variability  
364 can also extend to the case of different asperities within a given rupture, which would then  
365 cause different time histories of moment rate function and short-period energy function.

### 366 **3.3 Subduction zone megathrust ruptures**

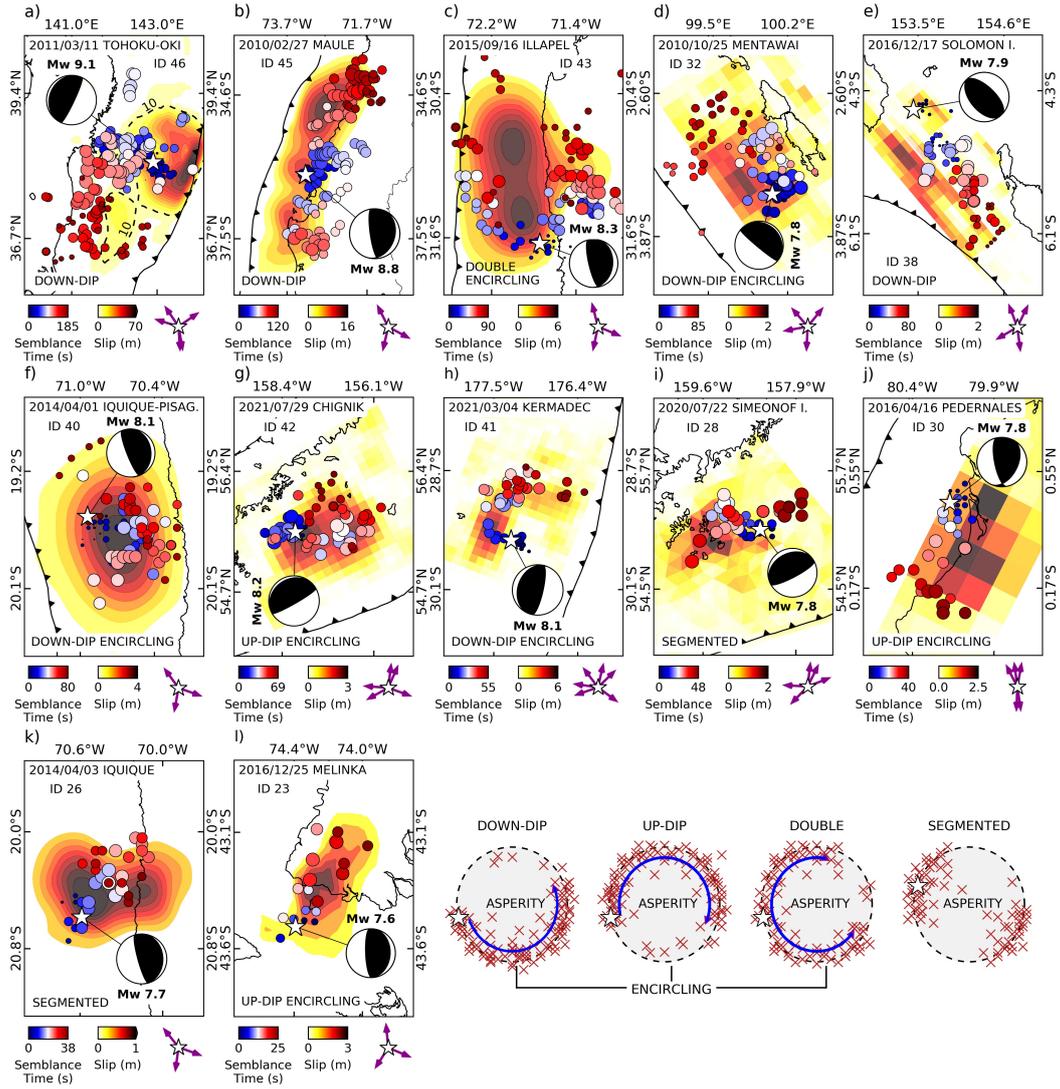
367 We now consider the rupture patterns of subduction megathrust earthquakes, specifi-  
368 cally the spatial relationship between finite-slip models of rupture displacement and short  
369 period emissions. As mentioned, previous observations of short-period ruptures indicated a  
370 strong preference for emissions down-dip of the main slip asperity (e.g., Lay et al., 2012).  
371 Fig. 6 summarizes the rupture patterns for the largest subduction zone megathrust events  
372 in our catalog. We can distinguish a variety of typical patterns referred to as down-dip,  
373 (down-dip, up-dip, or double-)encircling, and segmented, defined by the relation of short-  
374 period radiation and the main asperity (see Fig. 6 inset on the bottom right, for a graphical  
375 definition; and supporting Movie S1).



**Figure 4.** Short-period (0.5–2.0 Hz) energy radiated source time functions derived from back-projection. The energy radiated maxima evaluated over the rupture time provides the source time function (in gray). Additionally, USGS-NEIC (in red) and SCARDEC (in blue) moment rate functions are shown for comparison. The event highlighted with a dashed box is used as an example and discussed in detail in the text.



**Figure 5.** Earthquake rupture parameters from back-projection (0.5–2.0 Hz). Event ID (for cross-referencing to other figures, tables, and supporting information), moment magnitude, rupture length (colored by rupture time), aspect ratio, directivity, and rupture speed (colored by hypocentral depth). The events are categorized into normal (upper), thrust (middle), and strike-slip (bottom) earthquakes, with the order in each category determined by magnitude. Events in bold text are assumed to be subduction megathrust earthquakes. The vertical red bars in the rupture speed panel show the expected shear wave speed ( $V_S$ ) at the hypocentral depth in the IASP91 velocity model (Kennett & Engdahl, 1991). Ruptures in the range 95–105% $V_S$  and >105% $V_S$  define ‘likely’ and ‘strongly’ supershear earthquakes, respectively.



**Figure 6.** Back-projected earthquake rupture patterns (0.5–2.0 Hz) for major subduction zone megathrust earthquakes. Blue-red dots show semblance maxima tracking the earthquake rupture color-coded by time and scaled by energy radiated. The subplots are sorted by rupture duration. The dark-magenta arrows (lower-right corner) indicate the multi-array distribution relative to the epicenter (white star). Trench location from SLAB1.0 model (Hayes et al., 2012); focal mechanisms from Global CMT catalog. The yellow-red background shows the slip distribution of a) Inuma et al. (2012), b) Moreno et al. (2012) c) Tilmann et al. (2016), d), e), g) and h) USGS-NEIC finite fault solution, f) and k) Schurr et al. (2014), i) Crowell and Melgar (2020), j) Heidarzadeh et al. (2017), and l) Moreno et al. (2018). Bottom inset: Canonical high-frequency rupture patterns. The dominant type is noted in the bottom of each plot.

376 The 2010 Maule earthquake showed a classical down-dip rupture with a bilateral prop-  
377 agation ( $\sim 29\%$  southward; Fig. 6b and S2). The 2011 Tohoku-Oki earthquake also ruptured  
378 bilaterally and mainly down-dip (Fig. 6a and S7). However, we draw attention to the rup-  
379 ture near the trench where large shallow slip occurred (up to 50 m). In the last phase of  
380 the rupture (135–180 s), a subsidiary pattern with up-dip and down-dip emissions encircling  
381 the second larger slip area off the coast of Fukushima Prefecture can be discerned (see the  
382 southern 10 m contour in Fig. 6a).

383 Rupture images were not always limited to only the deeper portion of the megathrust.  
384 The observations included several short-period ruptures showing encircling patterns (see  
385 inset in Fig. 6), as previously only clearly reported for the 2016 Illapel earthquake. The  
386 2021 Kermadec Islands earthquake first propagated to the northeast and down-dip of the  
387 main asperity, followed by a circular shape with the rupture moving up-dip in the final  
388 phase (Fig. 6h and S58). Similarly, the 2010 Mentawai earthquake first propagated along  
389 the down-dip edge of the main asperity and then ruptured up-dip to the near trench region  
390 (Fig. 6d and S6). The 2014 Iquique-Pisagua earthquake ruptured with a half-ellipse pattern  
391 surrounding the larger slip area (Fig. 6f and S22). For the 2016 Pedernales earthquake, the  
392 reverse of the conventional pattern occurred, with short periods being radiated up-dip of  
393 the main asperity, followed by a down-dip propagation near the southern tip of the rupture  
394 (Fig. 6j and S33); a similar sequence defined the 2016 Melinka earthquake rupture (Fig. 6l

395 and S39). The 2021 Chignik earthquake (Fig. 6g and S59) showed an eastward propagation  
396 that fully encircled the slip area, but up-dip emissions dominated the rupture. More or less  
397 parallel and contemporaneous rupture ‘tracks’ up-dip and down-dip of the main asperity  
398 were observed for the 2015 Illapel earthquake in central Chile (Fig. 6c and S31), as also  
399 pointed out by the dedicated back-projection study of Meng et al. (2018) for this event.  
400 The main rupture propagated northward along the down-dip. A secondary front in the  
401 shallow part of the megathrust (less than  $\sim 15$  km depth) followed the region parallel to the  
402 trench for over  $\sim 150$  km, completing a double encircling pattern.

403 Other observations exhibited segmented ruptures, with short-period emissions concen-  
404 trated on asperity edges. The 2014 Iquique event showed a unilateral northeast short-period  
405 rupture partitioned around the asperity but outlining the maximum slip region (Fig. 6k and  
406 Fig. S23). Offshore of the Alaska Peninsula, the 2020 Simeonof Island earthquake first  
407 propagated down-dip to the northwest and west (Fig. 6i and S55), but the last emissions  
408 originated east of the epicenter.

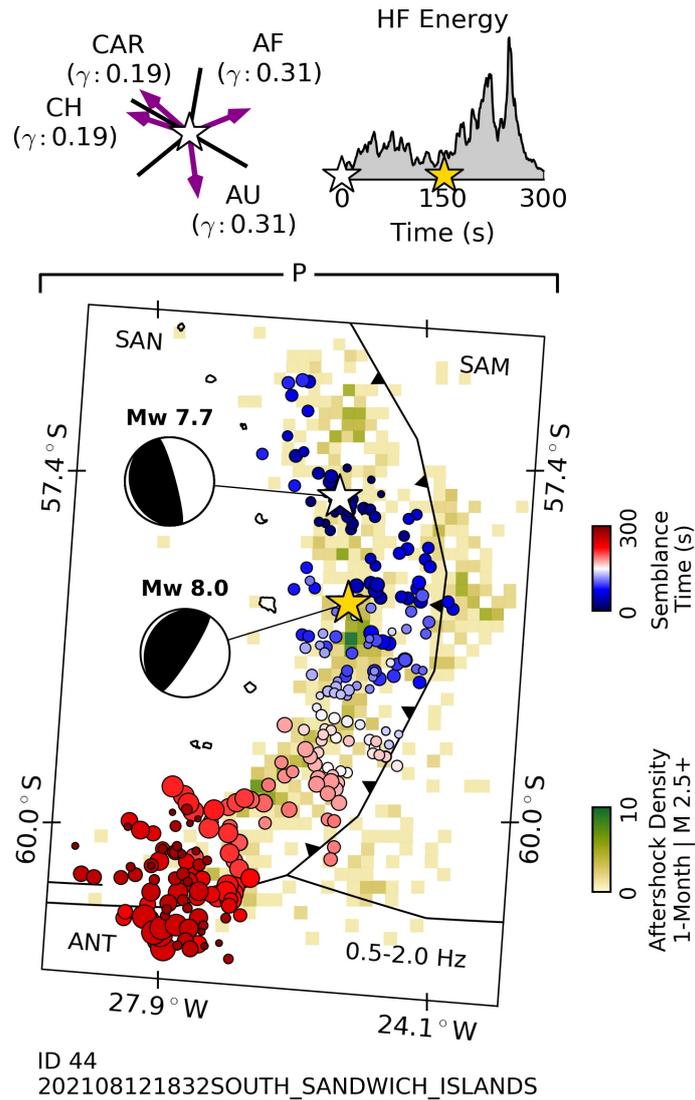
409 The 2021 South of Sandwich Islands earthquake doublet, the most recent event in our  
410 catalog, is another complex rupture (Fig. 7 and S60). Although major agencies all identified  
411 the earthquake as a doublet, estimates of the partitioning of moment between the sub-events  
412 and the onset of second sub-event differ considerably, including large discrepancies in the

413 estimated total moment (GCMT:  $M_W$  8.3 and 7.9 for the first and second sub-event, origin  
414 times separated by 178 s; GEOFON:  $M_W$  7.7 and 8.0, +153 s; USGS:  $M_W$  7.5 and 8.1, +148  
415 s). The event showed a predominantly SW rupture propagation with an initial short phase  
416 of NW-directed propagation. For about 200 s, the rupture followed the megathrust bend  
417 formed by the subducting South American Plate and the overriding Sandwich Plate. At  
418 the southern edge of the rupture, final emissions (up to 300 s) occurred near the transform  
419 fault margin where the Sandwich Plate borders the Antarctic Plate. However, uncertainties  
420 in short-period emission points are too large to unambiguously apportion slip to either the  
421 transform section or the subduction megathrust.

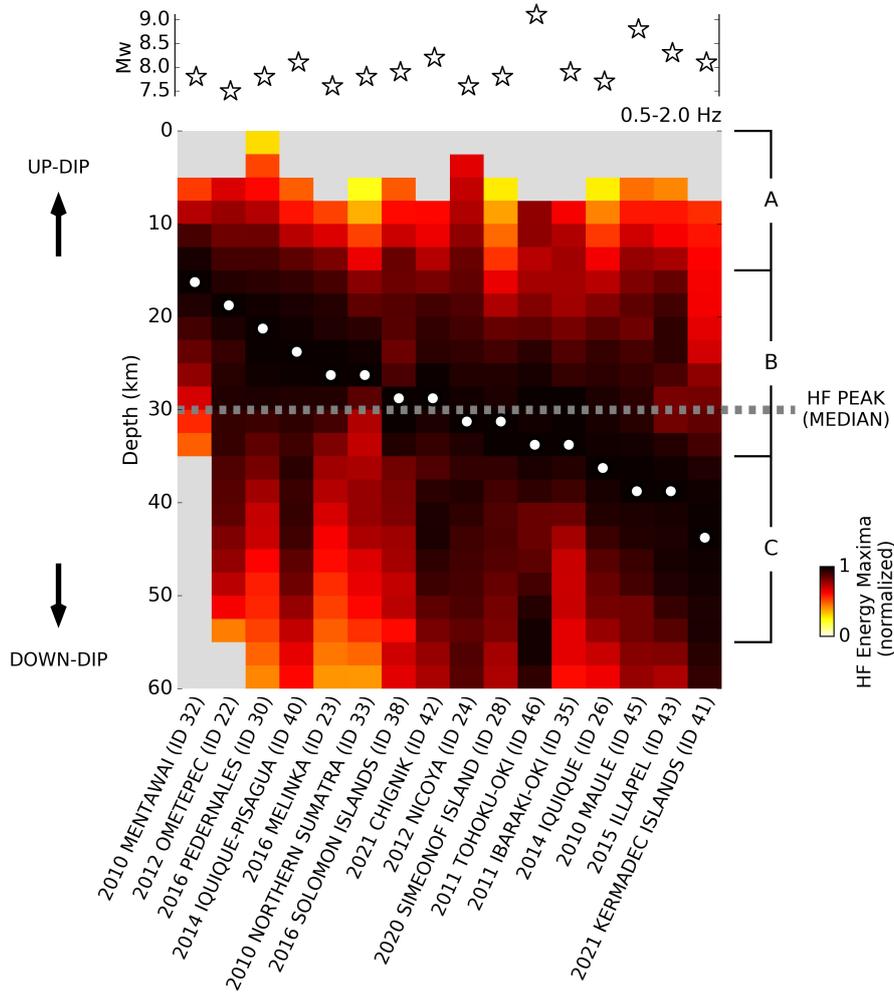
422       The resulting inferred rupture length of 496 km covered nearly three-quarters of the  
423 South Sandwich subduction zone, and is at least twice as long as expected for earthquakes  
424 of this magnitude based on scaling relations by Wells and Coppersmith (1994) and Blaser  
425 et al. (2010), see section 3.5 for quantitative analysis of short-period rupture lengths. The  
426 radiated short-period energy function is clearly separated into two phases (top inset in  
427 Fig. 7), presumably corresponding to the two sub-events; the origin time of the second  
428 sub-event in the GEOFON and USGS split coincides approximately with the minimum of  
429 the short-period energy function between the two peaks, and the second peak is clearly  
430 larger, in qualitative agreement with the relative moment sizes of the two sub-events in  
431 these catalogs. Additionally, Fig. 7 compares the back-projected rupture with a month of

432 cumulative seismicity after the first event. The short-period rupture followed mainly the  
433 down-dip band of seismicity, while a burst of emissions east of the epicenter outlined the  
434 area close to the trench, adjacent to a patch of outer-rise aftershocks.

435 In order to evaluate the depth distribution of seismic radiation more systematically,  
436 Fig. 8 presents the depth distribution of radiated short-period energy (more specifically,  
437 the maximum energy per depth level, normalized for each event). We remind the reader  
438 that depth is not resolved explicitly, but the horizontal position is translated to depth,  
439 assuming slip occurs on the slab interface defined by the SLAB1.0 model. The maximum  
440 energy distribution occurred between 15 and 45 km depth, with all depths in between  
441 represented fairly evenly but a higher concentration at depths larger than 25 km, as visible  
442 in the slope break in the figure. The resulting median depth is  $\sim 30$  km, which is close  
443 to the typical depth of the transition between domains B and C of Lay et al. (2012), i.e.,  
444 down-dip of the main asperity, but the large variation also demonstrated that the pattern  
445 of dominant down-dip high-frequency radiation has many exceptions. Also, events with  
446 maximum radiation at similar depths do not necessarily share the same characteristics. For  
447 example, the 2010 Mentawai, 2012 Ometepec, and 2016 Pedernales earthquakes all showed  
448 dominant radiation in the shallow part of the megathrust ( $< 25$  km depth). But the 2010  
449 Mentawai event was a tsunamigenic earthquake with an unusually shallow main asperity,  
450 therefore, this shallow depth still corresponds to dominant radiation down-dip or down-dip



**Figure 7.** The 2021 South of Sandwich Islands earthquake back-projection (0.5–2.0 Hz). Blue-red dots show semblance maxima tracking the earthquake rupture color-coded by time and scaled by energy radiated. The earthquake is a composite event, with the white and yellow star representing the (GFZ) epicenters and origin times of the two sub-events. Focal mechanisms are from the GEOFON (GFZ) moment tensor catalog. The yellow-green polygons present a month of cumulative seismicity (M 2.5+; USGS-NEIC) (patches of 10 × 10 km; M 2.5+ from the USGS-NEIC earthquake catalog; <https://earthquake.usgs.gov/earthquakes/search/>). Trench distribution from Styron and Pagani (2020). Tectonic setting: South American (SAM), Sandwich (SAN), and Antarctic (ANT) plates. Inset: Array weights and energy radiated source time function.



**Figure 8.** Depth distribution of radiated short-period energy (0.5–2.0 Hz) for all megathrust earthquakes. The maximum back-projected energy is extracted and displayed for each depth, normalized to 1 separately for each earthquake (yellow-red shading). White dots mark the depths for the peak values, showing the depth at which most short-period energy was emitted. The gray dashed line presents the median depth for short-period emissions. Gray polygons show areas with no data, i.e., depths, which were not included in the search grid for the respective earthquake. Inset: Earthquake magnitude (upper) and indicative depths of A-B-C megathrust faulting domains (right) following the definition of Lay et al. (2012); actual depths will differ somewhat between subduction zones.

451 encircling pattern. In contrast, the main rupture asperities for the 2016 Pedernales and  
452 2012 Ometepepec earthquakes were in domain B or even C, such that the strong shallow  
453 short-period emissions indicated an up-dip or up-dip encircling patterns. Most of the great  
454 subduction earthquakes with  $M_W > 8.0$ , e.g., 2010 Maule, 2011 Tohoku-Oki, 2015 Illapel,  
455 2021 Kermadec Islands, and 2021 Chignik earthquakes, radiated predominantly in deeper  
456 regions at 30–45 km depth, consistent with the classic down-dip segmentation described by  
457 Lay et al. (2012).

### 458 **3.4 Earthquake rupture speeds**

459 Fig. 5 shows the speed estimations, additionally expressed as a percentage of the shear  
460 velocity at the hypocentral depth in the IASP91 model. The underlying rupture speed  
461 time series are summarized in Fig. S1. Because the rupture speed estimates are based on  
462 the horizontal projection, rupture velocity might be underestimated if rupturing up-dip or  
463 down-dip dipping faults, e.g., for a dip angle of  $30^\circ$  this effect would amount to  $\sim 15\%$ .  
464 However, this effect will only affect a small subset of intra-plate thrust or normal faults.

465 Thrust earthquakes averaged a rupture propagation speed of 2.1 km/s and 56% of the  
466 shear wave speed, with a range between 25% and 86% of the shear wave speed, which places  
467 propagation speeds firmly in the expected sub-Rayleigh regime. The 2010 Mentawai, 2011  
468 Tohoku-Oki, and 2021 South of Sandwich Islands earthquakes were slower than the average.

469 The Mentawai and Tohoku-Oki events are characterized by large displacement at very shal-  
470 low depths, implying a low normal stress environment and large tsunamis. The outer-rise  
471 seismicity and notable near trench short-period rupture emissions for the Sandwich Islands  
472 event could also indicate a shallower rupture for at least part of the rupture propagation.  
473 However, finite fault modeling results are required for an extended interpretation.

474 A special case is presented by the 17 December 2016 Solomon Islands earthquake  
475 (Fig. 6e and S37). A megathrust rupture released the bulk of the moment, but its hypocenter  
476 is at 105 km depth, and the first motion mechanism indicated down-dip intraslab faulting,  
477 which is thought to have triggered the main megathrust event (e.g., Lay et al., 2017; Lee et  
478 al., 2018). In the rupture image, this sequence was visible as a gap of  $\sim 55$  km between the  
479 first emissions and the dominant megathrust pattern, which also outlined a southeast prop-  
480 agation of the rupture along the strike of the megathrust. The short time interval needed  
481 to cross the gap implied an extremely fast rupture speed of 5.7 km/s equivalent to  $\sim 71\%$   
482 of the P wave speed at the hypocentral depth, indicating that most likely P waves from the  
483 intraslab sub-event triggered the megathrust rupture (note that the relative P wave speed is  
484 calculated in the reference model; P wave speeds in the down-going oceanic crust would be  
485 slower, thus higher percentages are possible). To explore rupture propagation in the main  
486 phase, we therefore placed the reference point at the megathrust rupture initiation; then, a  
487 more typical rupture propagation speed of 2.1 km/s was obtained (Fig. S38).

488 For normal earthquakes, the average was 2.6 km/s and 63% of the shear wave speed,  
489 with a range between 42% and 91% of the shear wave speed. The 2019 Northern Peru  
490 (106 km depth; Fig. S51) earthquake showed the fastest rupture (4.1 km/s and 91% $V_s$ ),  
491 which is at the upper limit of the sub-Rayleigh speed. Although this intermediate-depth  
492 event was fast, the depth does not seem to have a systematic effect on rupture velocity. For  
493 example, the 2014 Rat Islands event at nearly the same depth (109 km; Fig. S27) and 2019  
494 Ecuador-Peru Border Region earthquake (at 146 km depth; Fig. S49) ruptured comparably  
495 slowly at 2.2 and 1.9 km/s. For the Rat Island event, previous back projections and finite  
496 fault modeling showed a similar rupture velocity in the range of  $\sim 1.5$ – $2.5$  km/s (e.g., Ye  
497 et al., 2014; Twardzik & Ji, 2015). However, the lack of solid evidence for a shallow or  
498 steeply-dipping fault plane preference makes the projection effects in the rupture velocity  
499 difficult to evaluate.

500 Strike-slip earthquakes showed the greatest variability. The average rupture speed was  
501 2.7 km/s and 78% of the shear wave speed, but the range spans from 37% to 137% of the  
502 shear wave speed. Two events were clearly within the unstable supershear range (between  
503  $V_s$  and  $\sqrt{2}V_s$ ): the 2013 Craig (4.6 km/s and 137% $V_s$ ; Fig. S17) and 2018 Palu (4.2 km/s  
504 and 125% $V_s$ ; Fig. S47) earthquakes. The Palu event was supershear from early on as re-  
505 ported also by Bao et al. (2019) using the high-resolution MUSIC back-projection method,  
506 while the oceanic interplate Craig initiated at subshear speed and transitioned to supershear

507 after  $\sim 20$  s (see time series and supershear reference in Fig. S1). The average rupture speed  
508 for the Craig earthquake inferred by us agrees with the finite-fault modeling results (4–5  
509 km/s) of Aderhold and Abercrombie (2015). Previously, Yue et al. (2013) provided faster  
510 peak velocity estimates from Sg and Sn arrivals ( $> 4.5$  km/s) and finite-fault inversion  
511 (5.5–6 km/s), but as we do not consider time-variable propagation rates, these observa-  
512 tions are not in conflict. The 2013 Balochistan event propagated at a speed of 3.3 km/s  
513 (Fig. S20), representing 98% of the shear wave speed, that is, around the shear velocity but  
514 (very likely) faster than Rayleigh waves. Additionally, the 2018 North of Honduras (95% $V_s$ ;  
515 Fig. S44) and 2020 Caribbean (104% $V_s$ ; Fig. S53) earthquakes propagated at ‘likely’ super-  
516 shear (95–105% $V_s$ ; by assuming expected rupture velocity errors in the estimates). For the  
517 Caribbean earthquake, Tadapansawut et al. (2021) inferred supershear rupture fronts with  
518 peak velocities larger than 5 km/s from finite-fault modeling, i.e., significantly faster than  
519 our average-based estimates.

520 Errors in emission point location arise from the array configuration and frequency band  
521 and can affect high-frequency rupture speed estimates, while finite-fault modeling operates  
522 at relatively low frequencies with imposed conditions, i.e., imposed speed boundaries and  
523 fault geometry. Marty et al. (2019) showed that rupture fronts derived from short-period  
524 backprojections propagate close to the rupture speed in laboratory experiments. It is an  
525 open question whether differences in rupture speed between back-projection and finite-

526 fault modeling seen here might be related to methodological concerns or reflect the more  
527 heterogeneous nature of natural fault systems.

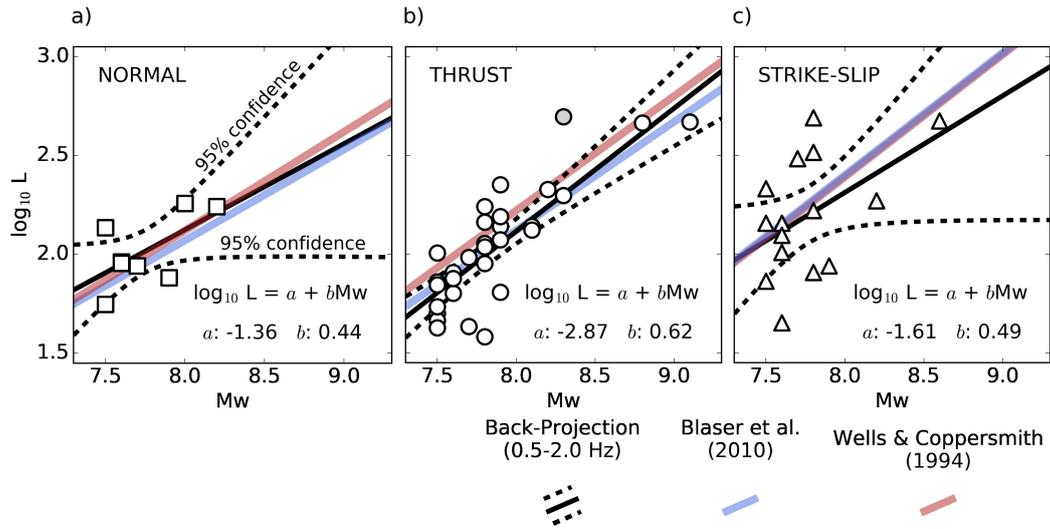
### 528 **3.5 Rupture length and aspect ratio estimates**

529 The earthquake rupture length estimates are used to determine magnitude-length scal-  
530 ing relations for each faulting slip type (normal, thrust, and strike-slip). The logarithm of  
531 rupture length and moment magnitude controlled the regression coefficients for the back-  
532 projected scaling relations (Fig. 9 and supporting Table S5). Due to the composite nature  
533 of the earthquake and variability in moment estimates, the Sandwich Islands event was not  
534 included in the regression for thrust magnitude-length scaling relation.

535 Thrust and normal earthquakes showed similar magnitude-length dependencies, while  
536 strike-slip earthquakes had a large data dispersion on average. Fig. 5 shows a comparative  
537 view of the rupture lengths colored by rupture time and sorted by moment magnitude for  
538 normal (upper), thrust (middle), and strike-slip (bottom) earthquakes.

539 We further explored the relationship between aspect ratio and complexity. Simple  
540 elongated ruptures generally are characterized by small aspect ratios, e.g., the 2017 Koman-  
541 dorsky Islands (ratio  $\sim 0.19$ ; Fig. S41), 2016 Wharton Basin (ratio  $\sim 0.25$ ; Fig. S32), and  
542 2011 Kermadec Islands (ratio  $\sim 0.17$ ; Fig. S10) earthquakes, all of them bilateral ruptures.  
543 In contrast, complex events with no straightforward interpretation of ‘rupture length’ tend

544 to have large values. For strike-slip events, events with aspect ratio close to 1 often indicate  
545 the rupture of multiple faults. Prominent examples are the 2012 Wharton Basin (ratio  $\sim$   
546 0.81; Fig. S12) and 2018 Gulf of Alaska (ratio  $\sim$  0.71; Fig. S45) earthquakes that ruptured a  
547 conjugate fault system. For the 2020 Sand Point strike-slip event, Santellanes et al. (2021)  
548 proposed that considerable slip occurred on the megathrust; the back-projection showed  
549 northward and eastward complex rupture emissions, but mostly eastward nearly along the  
550 megathrust strike (ratio  $\sim$  0.62; Fig. S56), which is relatively westward and up-dip of the  
551 2020 Simeonof megathrust rupture (Fig. S55). For thrust earthquakes, the aspect ratio also  
552 quantified complexity, but the complexity usually does not arise from a rupture on different  
553 faults, but from the distribution of high-frequency on the megathrust, e.g., the 2015 Illapel  
554 earthquake and its double encircling pattern (ratio  $\sim$  0.84; Fig. S31). However, regardless  
555 of the faulting slip type, a compact rupture can also result in an aspect ratio close to 1  
556 due to the greater importance of scatter for smaller ruptures, e.g., the 2021 Loyalty Islands  
557 (ratio  $\sim$  0.84; Fig. S57) and 2014 Solomon Islands (ratio  $\sim$  0.77; Fig. S25) earthquakes.  
558 Therefore, rupture complexity is linked to aspect ratios  $>\sim$  0.5, but such a larger value  
559 does not necessarily imply complexity but can also indicate compactness, particularly for  
560 earthquakes with  $M_W <\sim$  8. Finally, we note that aspect ratios are always evaluated in  
561 the horizontal plane, implying that results for strongly dipping faults might not be directly  
562 comparable.



**Figure 9.** Rupture length scaling relations for short-period back-projection based estimates. a) Normal, b) thrust and c) strike-slip earthquakes. The black line shows the back-projection relationships (solid line) and the 95% of confidence interval predicted (dashed line). The squares, circles, and triangles present the data distribution for each faulting type. The gray circle shows the 2021 South of Sandwich Islands earthquake not included in the regression. Blue and red lines show the well-known scaling relationships of Blaser et al. (2010) and Wells and Coppersmith (1994) for comparison.

## 4 Discussion

### 4.1 Multi-array multi-phase imaging method

A multi-array configuration provides wide coverage in azimuth with benefits in resolvability, but a heterogeneous distribution can promote biases, e.g., ‘swimming’ artifacts. As an example, we consider the 19 April 2014 Solomon Islands earthquake ( $M_W$  7.5; 43 km depth): the multi-array back-projection with the inclusion of the depth phase approach showed a short-period rupture outlining the finite fault slip model (see Fig. S25). However, the rupture differed significantly when the arrays were restricted in number, azimuthal distribution, or depth phases were not considered (Fig. S26). For example, P wave backprojections from one or two arrays showed a significant time-space trade-off of the earthquake rupture in the form of swimming artifacts, e.g., a drift towards the arrays in Alaska and Japan (Fig. S26a–b). These artifacts are progressively reduced with more and well-distributed arrays (Fig. S26c–d). If depth phases are incorporated, the same array coverage offered a moderate reduction of artifacts (Fig. S26e–h). This is also noticeable for the source time function, which is strongly affected by depth phases when the arrays are poorly distributed, and only P waveforms are included. The depth phase appears as a spurious secondary peak at  $\sim 15$  s (see inset in Fig. S26a–c).

580 Further detailed analyses of the effects of artificially restricted coverage are shown in  
581 the supplementary material for selected events and exhibit similar effects: the 2010 Northern  
582 Sumatra (Fig. S4), 2017 Chiapas (Fig. S43), 2019 Northern Peru (Fig. S52), and 2020 East  
583 of Kuril Islands (Fig. S54) earthquakes.

## 584 **4.2 Complex megathrust ruptures and depth-varying short-period radia-** 585 **tion**

586 Seismicity prior to a great earthquake is often demonstrated to be active on the rupture  
587 zone edge while a seismic quiescence dominates the region later experiencing the maximum  
588 slip according to the ‘Mogi doughnut’ hypothesis (e.g., Mogi, 1969; Kelleher & Savino, 1975).  
589 According to this hypothesis, the spots at the asperity edge become seismically active as  
590 the stress on the asperity increases, outlining the potential rupture area (e.g., Kanamori,  
591 1981). For example, Schurr et al. (2020) observed a Mogi doughnut in the years prior to the  
592 2014 Iquique-Pisagua earthquake. Sippl et al. (2020) showed microseismicity forming three  
593 half-ellipses along the central Chile megathrust, with one of them encircling the 2015 Illapel  
594 earthquake area. In back-projection, short-period ruptures are related to stress conditions  
595 and rupture velocity variations (e.g., Marty et al., 2019). Therefore, back-projected rupture  
596 patterns might likewise be an indication of the high-stress gradient around the asperity,  
597 equivalently to the Mogi doughnut pattern of seismicity.

598 Many megathrust earthquakes exhibited encircling patterns, where the back-projection  
599 highlights propagation to shallower regions. This pattern could be related to the presence of  
600 seismogenic barriers under the assumption of a strong rupture deceleration or stopping phase  
601 (e.g., Madariaga, 1983). The 2015 Illapel earthquake showed a double encircling pattern; it  
602 was first observed by Meng et al. (2018) using the high-resolution MUSIC back-projection  
603 method and interpreted as a splitting of rupture fronts surrounding a large asperity or  
604 barrier. The 2021 Chignik earthquake is another example of a rupture forming a ring  
605 around the slip patch with both up-dip and down-dip limits included. The 2010 Mentawai  
606 and 2011 Tohoku-Oki earthquakes were tsunamigenic with a rupture that propagated up to  
607 the trench (e.g., Lay et al., 2011; Hill et al., 2012; Iinuma et al., 2012). This is compatible  
608 with the back-projected rupture, where semblance peaks are (unusually) partly located at  
609 shallow regions below accretionary structures. The transition from the seismogenic to a  
610 shallow aseismic sliding region might explain a strong rupture velocity variation and the  
611 shallow source of short-period radiation. For the 2016 Pedernales event, Agurto-Detzel et  
612 al. (2019) proposed, in the up-dip limit of the earthquake rupture, a barrier mechanically  
613 controlled by the subduction of a rough oceanic relief. Similarly, the back-projection revealed  
614 a rupture up-dip of the slip distribution (at 15–20 km depth; Fig. S33) that agrees with high  
615 residual ( $> \sim 1$  km) bathymetry data (see Fig. 9 of Agurto-Detzel et al., 2019), indicating a

616 first-order relationship between the along-dip barrier structure and short-period earthquake  
617 radiation.

618 In summary, short-period energy maxima emitted from the shallow part of the megathrust  
619 (< 35 km depth; A and B regions) resulted in being more frequent than suggested  
620 by earlier results. The 2010 Mentawai, 2012 Ometepec, and 2016 Pedernales earthquakes  
621 emitted their maximum short-period radiation from very shallow regions (< 25 km depth),  
622 unlike in the classic Lay et al. (2012) dip-segmentation. The median depth for the energy  
623 radiated peak falls at the end of domain B (15–35 km; Fig. 8) of Lay et al. (2012), that  
624 is, before the transition from regions with modest to high amounts of coherent short-period  
625 energy, but, as we have seen, large deviations are not uncommon between large subduction  
626 earthquakes.

### 627 **4.3 Rupture length scaling relations at short periods**

628 We compare the newly derived short-period scaling relation to established scaling rela-  
629 tions (e.g., Wells & Coppersmith, 1994; Blaser et al., 2010) which were derived from large  
630 data sets using rupture lengths estimated from aftershocks sequences, geodetic modeling,  
631 and empirical relationships (Fig. 9, Table S5). For normal faulting earthquakes, our scaling  
632 relation is very similar to those derived from both Wells and Coppersmith (1994) and Blaser  
633 et al. (2010) (Fig. 9a). For thrust earthquakes, our scaling relation predicts shorter thrust

634 rupture lengths (Fig. 9b) than Wells and Coppersmith (1994) but is close to the predictions  
635 of Blaser et al. (2010) (which is the newer reference based on a larger database). Lengths for  
636 strike-slip earthquakes presented a large dispersion (Fig. 9c), providing estimates generally  
637 shorter than Blaser et al. (2010) and Wells and Coppersmith (1994) but still within the  
638 expected, albeit large, uncertainty. A first limitation here is the sparsity of strike-slip events  
639 with  $M_W > \sim 8$ . Second, several of the large strike-slip earthquakes were complex events  
640 rupturing multiple conjugate faults, leading to large ambiguity in the interpretation of the  
641 rupture length determined by our algorithm, as the cumulative fault length ruptured will be  
642 much larger than this estimate. This pattern is often found for intraplate earthquakes in the  
643 oceanic lithosphere, for example, the great 2012 Wharton Basin (e.g., Meng et al., 2012a;  
644 Duputel et al., 2012; Satriano et al., 2012; Hill et al., 2015) and 2018 Gulf of Alaska (e.g.,  
645 Lay et al., 2018; Krabbenhoft et al., 2018; Ruppert et al., 2018) earthquakes, where the  
646 back-projection results are broadly compatible with the more comprehensive assessments  
647 carried out in these references, which also included aftershocks, seafloor topography and/or  
648 finite fault modeling in the analysis.

## 649 **5 Conclusions**

650 We presented a catalog of short-period rupture histories for 54 large earthquakes (0.5–  
651 2.0 Hz;  $M_W \geq 7.5$ ; 01/2010-10/2021) based on a new implementation of the teleseismic

652 back-projection method, which takes into account multiple arrays and combined P and pP  
653 waveforms (for earthquakes deeper than 40 km). Based on the back-projection results,  
654 rupture length, directivity, rupture speed, and aspect ratio are estimated algorithmically.

655 The main findings are as follows:

- 656 1. We find distinct differences in rupture patterns between finite fault slip models and  
657 short-period emissions for subduction megathrust earthquakes. We confirm a prefer-  
658 ence for short-period seismic energy to be emitted down-dip of the main slip asperity  
659 (as previously reported in the literature), but additionally identify many examples of  
660 segmented and encircling configurations with down-dip, up-dip, or double patterns  
661 outlining coseismic slip patches.
- 662 2. Short-period rupture patterns for megathrust earthquakes, e.g., encircling patterns  
663 around slip patches, might result from the stress gradient around asperities (including  
664 along-strike seismogenic barriers).
- 665 3. Earthquake rupture speeds were consistently in the sub-Rayleigh regime for thrust  
666 and normal faulting earthquakes, with a median of 56% and 63% of the shear wave  
667 speed, respectively. Strike-slip earthquakes showed the greatest variability with a me-  
668 dian of 78% and a range between 37% and 137% of the shear wave speed. The 2013  
669 Craig (137% $V_s$ ) and 2018 Palu (125% $V_s$ ) events propagated in the unstable supers-

670 hear range, while the 2013 Balochistan ( $98\%V_s$ ), 2018 North of Honduras ( $95\%V_s$ ),  
671 and 2020 Caribbean ( $104\%V_s$ ) earthquakes were ‘likely’ supershear ( $95\text{--}105\%V_s$ ; by  
672 assuming expected rupture velocity errors in the estimates) and in any case faster  
673 than Rayleigh waves.

674 4. Finally, we presented new scaling relations from short-period backprojections com-  
675 parable to finite difference methods. Thrust and normal earthquakes showed a sim-  
676 ilar magnitude-length relationship compared to established ‘long-period’ relations.  
677 Strike-slip events presented a large dispersion but still within the expected uncer-  
678 tainty. Limitations for strike-slip earthquakes were the lack of events with a large  
679 moment magnitude ( $M_W > \sim 8$ ) and the underestimation of rupture lengths due to  
680 complexities in the rupture, e.g., complex fault systems as observed for the great 2012  
681 Wharton Basin ( $M_W$  8.6) and 2018 Gulf of Alaska ( $M_W$  7.9) earthquakes.

682 Whereas overall the method has been optimized for near-real-time processing, the array  
683 selection and picking of the end of the rupture still need to be done manually, as is the  
684 definition of secondary search grids. This paper has presented and analyzed the fault rupture  
685 histories for 54 earthquakes, but we intend to continue to analyze future earthquakes with  
686  $M_W \geq 7.5$  and depth less than 200 km. These analyses will be made available under the  
687 link listed below.

## 688 **6 Datasets**

689       The catalog of ruptures is included in the supporting information as map views and in  
690 machine-readable format (supporting Data Set S1). The automatic earthquake parameters  
691 derived from back-projection are also available (supporting Data Set S2). Both datasets are  
692 distributed through GFZ data services (<https://dx.doi.org/...> )

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