Crust and Upper Mantle Structure Beneath the Eastern United States

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Abstract

The Eastern United States has a complex geological history and hosts several seismic active regions. We investigate the subsurface structure beneath the broader eastern United States. To produce reliable images of the subsurface, we simultaneously invert smoothed P-wave receiver functions, Rayleigh-wave phase and group velocity measurements, and Bouguer gravity observations for the 3D shear wave speed. Using surface-wave observations (3-250 s) and spatially smoothed receiver functions, our velocity models are robust, reliable, and rich in detail. The shear-wave velocity models fit all three types of observations well. The resulting velocity model for the eastern U.S. shows thinner crust beneath New England, the east coast, and the Mississippi Embayment. A relatively thicker crust was found beneath the stable North America craton. A relatively slower upper mantle was imaged beneath New England, the east coast, and western Mississippi Embayment. A comparison of crust thickness derived from our model against four recent published models shows first-order consistency. A relatively small upper mantle low-speed region correlates with a published P-waves analysis that has associated the anomaly with a 75 Ma kimberlite volcanic site in Kentucky. We also explored the relationship between the subsurface structure and seismicity in the eastern U.S. We found earthquakes often locate near regions with seismic velocity variations, but not universally. Not all regions of significant subsurface wave speed changes are loci of seismicity. A weak correlation between upper mantle shear velocity and earthquake focal mechanism has been observed.

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12	Key Points:
13 14	• Inverting smoothed receiver functions, surface-wave dispersion and gravity for a 3D shear wave velocity model for the eastern US
15	• Our velocity model is broadly consistent with published results for the region
16 17	• Earthquakes often but not universally locate near areas with seismic speed variation, but not all velocity changes are loci of seismicity
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anomaly with a 75 Ma kimberlite volcanic site in Kentucky. We also explored the relationship

between the subsurface structure and seismicity in the eastern U.S. We found earthquakes often

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39 Plain Language Summary

40 The Eastern United States experienced a complex series of geological activities. Earthquakes in

the Eastern United States have been recorded at several localized regions. A detailed subsurface

42 structure can help us recover the geological history and studying earthquakes. We use multiple

43 types of geophysical observations to reliably image the subsurface. Our images of the subsurface

44 confirmed many findings from previous studies. The crust is thinner beneath New England, the

45 east coast, and the south-central United States. The interior of North America has a thicker crust.

46 The upper mantle seismic speed is shower beneath New England, the east coast, and the western

47 portion of the south-central United States. A smaller region of slower upper mantle speed in

Kentucky agrees with a published study, which linked the slower speed with a 75 Ma volcanic site. We compared images of subsurface against earthquake locations. Earthquakes often locate

near regions with lateral subsurface structure changes. Lateral subsurface structure changes do

not always collocate with earthquakes. The type (faulting) of earthquakes weakly correlates with

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52 the upper mantle seismic speed.

53 **1 Introduction**

54 The eastern United States (EUS) has a long and complex geological history that includes

at least two episodes of continental collision and breakup. The eastern part of North America

56 grew from the continent's Archean core through arc and continental collisions in the Early-

57 Middle Proterozoic (e.g., Whitmeyer, 2007). The resulting supercontinent Rodina broke up

⁵⁸ during the opening of the Iapetus Ocean in the Late Proterozoic (Hynes & Rivers, 2010).

59 Renewed plate convergence and episodic terrane accretion formed the supercontinent Pangea 60 during the Paleozoic and the early Mesozoic (e.g., Hatcher, 2010; Whitmeyer, 2007). The

during the Paleozoic and the early Mesozoic (e.g., Hatcher, 2010; Whitmeyer, 2007). The
 opening of the Atlantic Ocean started in the Triassic-Jurassic (Hames et al., 2000). The east coast

became a passive margin around 180 Ma (Faill, 1998). Several stages of rifting modified the

passive margin (e.g., Whitmeyer, 2007). Except for first-order observations of a relatively fast

upper mantle, the relationship between the geological history and subsurface structure is still notwell known.

Though the EUS is naturally less seismically active than the western U.S., several 66 intraplate seismic zones (see Figure 1) including the New Madrid seismic zone (NMSZ), the 67 Wabash Valley seismic zone (WVSZ), the South Carolina seismic zone, the Central Virginia 68 69 seismic zone, and the West Quebec seismic zone (WQSZ) show localized seismic activity in the east coast. Three great-to-major earthquakes (Magnitude 7-to-8) struck the NMSZ in 1811-1812 70 and caused severe damage (Hough et al., 2000; Johnston, 1996). Paleoseismic studies suggest 71 that at least two large earthquakes occurred thousands of years prior to 1800 (Tuttle, 2002). Past 72 destructive earthquakes and slow deformation (Newman, 1999) provoke debate over the seismic 73 hazard assessment (Stein, 2007). The mechanism of stress concentration that is proposed to have 74 75 caused these large events is unsolved. Some geodynamic models predicted a stress concentration in the seismogenic zone (e.g., Levandowski et al., 2016; Pollitz, 2001). However, the assumed 76 structural models differ significantly due to poor images of the subsurface, especially in the 77 crust. Detailed structure images beneath the EUS are required to answer outstanding questions 78 79 such as what factors control the occurrence of intraplate earthquakes.



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Figure 1. Seismicity (black and blue circles, magnitude 3 and larger prior May 2021 from the 81 USGS catalog) and cross-section locations (red lines). Black cirlces represent earthquakes that 82 are shown in cross-sections. Blue circles are earthquakes that are located 100 km away from any 83 cross-sections. The size of circles is proportion to magnitude. Abbreviations: CVSZ - Central 84 Virginia Seismic Zone; ETSZ – Eastern Tennessee Seismic Zone; NMSZ – New Madrid Seismic 85 Zone; OES - Oklahoma Earthquake Swarm; SCSZ - South Carolina Seismic Zone; WQSZ -86 West Quebec Seismic Zone; WVSZ – Wabash Valley Seismic Zone. (For interpretation of the 87 references to color in this figure, the reader is referred to the web version of this article.) 88 89 The deployment of EarthScope USArray Transportable Array (TA) provided

90 unprecedented station coverage in the EUS for more than 10 years. The TA has crept

continuously across the U.S. and collected seismic observations at more than 2000 unique

locations. With a nearly uniform station spacing (~70 km) and high-quality sensors, the array

93 was designed to improve our understanding of subsurface structure and have led to many seismic

velocity models for the lithosphere structure of the EUS and surround regions (e.g., Biryol et al.,

95 2016; Bollmann et al., 2019; Dong & Menke, 2017; Golos et al., 2018; Menke et al., 2018; B.

Savage et al., 2017; Brian Savage, 2021; Wagner et al., 2018; B. B. Yang et al., 2014). To better
utilize the improved data coverage and compare the subsurface structure across a broad region,

we investigate the seismic velocity variations beneath the EUS with a model parameterization

99 that is suitable for TA stations.

100 Even with dense seismic networks, tightly constraining subsurface 3D geologic variations is a challenge. But the seismological community has worked for decades with less data to extract 101 information valuable enough to steadily advance our understanding of the lithosphere. P-wave 102 receiver functions processed from teleseismic P-waves provide us a way to repeatedly sample 103 subsurface structure beneath seismic stations as source-side effects are removed by 104 deconvolution (e.g., Langston, 1979). To extract detailed subsurface structural parameters from 105 receiver functions, an inversion is often used to estimate model parameters from receiver 106 function observations. The inversion of receiver functions is non-unique (e.g., Ammon et al., 107 1990), but incorporating complementary observations has shown to ease the non-unique problem 108 and improve the sensitivity (e.g., Chai et al., 2015; Chong et al., 2016; Julià et al., 2000, 2003; 109 Özalaybey et al., 1997; Sun et al., 2014). The construction of 3D models adds additional 110 complexity. Different data sets may average 3D heterogeneity differently, making one-111 dimensional models that fit all the data difficult to construct. For example, 3D scattering in 112 receiver functions can introduce hard-to-identify high-wavenumber artifacts during an inversion. 113 An often used and generally successful approach to reduce these effects is to target smooth earth 114 models that represent averages of the true structure. We used the receiver function 115 smoothing/interpolation technique (Chai et al., 2015) to reduce the scattering noise in the data 116 prior to the inversion to avoid mapping noise into artifacts. Simultaneous inversion using 117 smoothed receiver functions has produced reliable images of subsurface seismic velocity 118

119 variations (Chai et al., 2015).

120 We simultaneously inverted smoothed/interpolated P-wave receiver functions, Rayleighwave phase and group velocity measurements, and Bouguer gravity observations for subsurface 121 structure beneath the EUS using linearized uniform-cell-based 3D inversions. Compared with 122 123 recent studies (e.g., Porter et al., 2016; Schmandt et al., 2015; Shen & Ritzwoller, 2016), we used a wider period range for surface-wave dispersion observations, included longer receiver function 124 signals, and reduced scattering noise in receiver functions through spatially smoothing. We 125 126 describe the procedures used to process each observation type in the next section. Details of the receiver function smoothing/interpolation technique are discussed in the following section and 127 are followed with a description of the theory and processes for the simultaneous inversion. The 128 129 fits to each type of observations are illustrated in a separate section to show an important component of the quality control applied to our results. We end with a discussion on the resulting 130 3D shear-wave velocity models that compares our results with recently published models and an 131 investigation of the relationship between seismicity and subsurface material variations. 132

133 **2 Data**

134 We used three types of geophysical observations to image the subsurface structure including P-

135 wave receiver functions, Rayleigh-wave dispersion measurements, and Bouguer gravity

observations. The receiver functions were computed using observations from about 1700 seismic

137 stations operated within the broad EUS (Figure S1).

138 2.1 Receiver function observations

P-wave receiver functions were processed using teleseismic waveforms recorded at TA, 139 and many other seismic networks (a complete list of networks can be found in Table S1). 140 Seismic events (around 3,700) with body-wave magnitude mb larger than 5.7 during station 141 operation times (prior June 2015) and at epicentral distances from 30 to 100° (see Figure S2) 142 were selected to avoid P-waveform interference with upper mantle triplications and the core 143 shadow zone. We downloaded three-component P-waveforms along with metadata for these 144 events from the Data Management Center of Incorporated Research Institutions for Seismology 145 (IRIS DMC). The seismograms were rotated to the radial-transverse-vertical system from the 146 147 original north-east-vertical recording coordinates. P-wave receiver functions (Langston, 1979) were obtained by deconvolving the vertical component from the radial and transverse 148 components using the time domain iterative deconvolution algorithm (Ligorría & Ammon, 149

150 1999). Gaussian filters (Ammon, 1991) were used in the deconvolution process to limit the

bandwidth of the receiver functions and preserve absolute amplitudes of receiver functions.

152 Specifically, we computed receiver functions with a Gaussian filter of 1.0 and 2.5 that

153 correspond roughly to pulse widths of 1.67 and 0.67 s at half the maximum, respectively.

Similar to receiver functions obtained for the western U.S region (e.g., Chai et al., 2015), 154 155 many receiver functions processed from even high-quality stations are quite noisy. In order to exclude noisy outliers, we used three programmable waveform selection criteria. Other 156 waveform selection criteria have been used by previous studies (e.g., Yang et al., 2016), but 157 visual examinations of our receiver function waveforms confirm that these three criteria are 158 sufficient to remove problematic receiver functions from the data. First, receiver functions with 159 signal-to-noise ratios (measured using a 170 s time window that ends 10 s earlier than the 160 expected P arrival time as noise and a 130 s time window immediately after as signal) less than 161 10 were discarded. Second, since the deconvolution can be unstable, we reject receiver functions 162 with convolution fit less than 85%; the convolution fit is computed as a signal power ratio 163 between the radial component and a convolved signal of receiver function and the vertical 164 component. Even with these relatively strict criteria, occasionally, some problematic receiver 165 functions (e.g. unusual large amplitude, wrong polarity, low-frequency noise) could pass through 166 and influence our observations. To identify these signals, we computed the signal difference of 167 each receiver function with respect to the single-station averaged waveform to further clean up 168 our receiver function data set and excluded receiver functions with signal difference larger than 169 300% of the stacked waveform signal power. The distribution of receiver function signal 170 differences follows the extreme value distribution for the EUS data set (Figure S3), which 171 suggests a small portion of receiver functions is significantly different from the single-station 172 averaged receiver functions. We obtained around 228,000 (38% of the raw receiver functions) 173 174 acceptable receiver functions in total. The accepted receiver functions were then binned into three ray parameter bins (smaller than 0.05 s/km, larger than 0.07 s/km and in between) for both 175

Gaussian 1.0 and 2.5 receiver functions. Therefore, we have six receiver function waveforms at a station when data are abundant.

178 2.2 Surface-wave dispersion observations

We used both Rayleigh wave group and phase velocity observations in the simultaneous 179 inversion. We avoid Love waves to minimize complexity that may arise from anisotropy. Our 180 results are thus an approximation of the potentially anisotropic Earth. Surface-wave dispersion 181 observations from two studies were blended to combine the short period dispersion values from 182 Herrmann et al. (2021) and longer period dispersion observations from Ekström (2011). The 183 short period dispersion observations were measured using both earthquake signals and ambient 184 noise cross-correlations. The measured dispersion observations were localized to grid points 185 through a surface wave tomography (e.g., Ekström, 2011; Herrmann et al., 2021). The short 186 period dispersion model used a 1°-by-1° grid with a node at the center of each grid cell. Only the 187 188 dispersion data corresponding to cells with good ray path coverage (at least 50 km of ray path) were used. We adopted the following formula from Maceira & Ammon (2009) for a smooth 189 blending of the dispersion curves. 190

(1)

191
$$s(T) = \cos^2 \varphi \, s_2(T) + \, \sin^2 \varphi \, s_1(T)$$

in which

193

$$\varphi = \frac{\pi}{2} \frac{1 + \tanh[\varepsilon(T - T_c)]}{2}$$

In the expressions above, s(T) is the blended dispersion value (group or phase velocity). T 194 is the period, $s_1(T)$ is a value from the long-period dispersion tomography (Ekström, 2011), and 195 $s_2(T)$ is a value from the short-period dispersion tomography (Herrmann et al., 2021) at the same 196 period. φ is a control parameter that is a function of T_c and ε . ε is set equal to 0.5. T_c is the period 197 around which we want to switch from short- and long-period surface wave observations. Since 198 the available period range varies from location to location (due to ray path coverage differences), 199 we used a grid search to determine the best transitional period (T_c) . We tried a range of T_c and 200 computed the gradient of the resulting dispersion curves. In order to minimize artificial 201 anomalies caused by the blending, the optimal T_c is the one with a minimum gradient (the 202 smoothest dispersion curve). All blended dispersion curves were visually examined and minor 203 adjustments were made as needed. Generally, a dispersion-curve transition period between 30-40 204 205 s was chosen. The blended dispersion data set spans from 3 s to 250 s in the best case. As an example, the blended dispersion curves are compared with recent dispersion models (Bensen et 206 al., 2007; Ekström, 2011, 2014; Herrmann et al., 2021; Jin & Gaherty, 2015) in Figure S4. The 207 blending of dispersion curves greatly extended the period range and are consistent with 208 alternative recent dispersion models at most places. 209

210 2.3 Gravity observations

Gravity observations were extracted from a global Bouguer gravity model WGM2012 (Balmino et al., 2012). This Bouguer gravity data set was computed by spherical harmonic analysis using ETOPO1 topography-bathymetry data and gravity observations from the EGM2008 global gravity model (Pavlis et al., 2012). The lateral resolution of the gravity data is 5 arc minutes (~9 km), which is higher than what we can resolve with a 1°-by-1° grid in our inversion. The Bouguer gravity observations are averaged within a 1°-by-1° grid for the EUS to reduce gravity anomalies that are due to smaller-scale structure perturbations. Long wavenumber 218 gravity signals are primarily caused by deep density changes. However, interpretation of deep

219 gravity signals is nonunique and can be associated with density variations, plate flexure, and

220 upper mantle dynamic effects. We wavenumber filtered the gravity observations with a box-car

filter so that the remaining gravity signals are mainly sensitive to the shallow density structure (upper \sim 15 km). Initial gravity observations from WGM2012 and wavenumber filtered gravity

(upper ~15 km). Initial gravity observations from WGM2012 and wavenumber filtered gravity
 values are compared in Figure S5. The filtered gravity image shows fewer small scale and large

223 values are compared in Figure 55. The intered gravity image shows rewer small scale 224 wave-number variations than the raw image as expected.

wave-number variations than the raw image as expected

3 Receiver-function smoothing/interpolation

P-wave receiver functions and surface-wave dispersion observations have different 226 spatial (lateral) sensitivity. Receiver functions are sensitive to sharp changes in vertical and 227 lateral structure. In particular, large, complicated (and hopelessly aliased) variations near the 228 surface often strongly influence P-wave receiver functions. Surface-wave dispersion 229 230 observations are more sensitive to longer wavenumber variations in the structure. We smoothed the receiver-function wavefield spatially to reduce near-surface scattering effects and to better 231 complement the surface-wave dispersion measurements. We have developed receiver function 232 233 smoothing/interpolation to attack this issue by simplifying the receiver function wavefield (Chai et al., 2015). The smoothing/interpolation reduces the scattering noise on receiver function 234 observations and isolates the spatially coherent components of the signals (see Figure 2 and 235 Figure S6). Compared with the traditional single-station-averaged receiver functions, 236

237 interpolated/smoothed receiver functions are more consistent laterally for different lag times.

238 A spatially smoothed receiver function is computed by averaging receiver functions recorded at adjacent stations with distance-dependent weights. Smoothing/interpolation also 239 equalizes the lateral sensitivity of receiver functions and surface-wave dispersion measurements, 240 which can reduce potential inconsistencies between surface-wave observations and body-wave 241 receiver functions. Interpolation also provides us a way to approximate receiver functions at grid 242 points where surface-wave observations are commonly estimated by tomography. This 243 functionality is quite useful especially for 3D inversions using multiple types of observations. 244 We can express the receiver function wavefield as R(x,y,t), where x and y are spatial coordinates 245 (e.g. latitude and longitude), and t is lag time after the direct P wave. Then, the interpolated 246 receiver function wavefield is computed using the following formula repeatedly for all locations 247 of interest (e.g. grid points) 248

249
$$R_i(x_i, y_i, t) = \frac{\sum_{j=1}^n \omega_j R_j(x_j, y_j, t)}{\sum_{j=1}^n \omega_j}$$

250 with

251
$$\omega_j = \begin{cases} 1 & \text{if } r_{ij} < d_1 \\ \frac{r_{ij} - d_1}{d_1 - d_2} + 1 & \text{if } d_1 < r_{ij} < d_2 \\ 0 & \text{if } r_{ij} > d_2 \end{cases}$$

where r_{ij} is the distance between the point of interest (x_i, y_i) and station location (x_j, y_j) , and *n* is total number of stations. d_1 and d_2 ($d_1 < d_2$) are two distance parameters that control the distance range for the stations included in the averaging and the weight (w_j) of the different stations. Receiver functions recorded at stations at distances less than or equal to d_1 are weighted

- as 1. Stations located at distances between d_1 and d_2 are weighted between 0 and 1 with the
- weight linearly decreased from 1 at distance d_1 to 0 at distance d_2 . Chai et al. (2015) compared
- smoothed receiver functions computed using several pairs of distance parameters for different
- 259 seismic network configurations in the western U.S. In our analysis, we choose a d_1 of 110 km 260 and a d_2 of 160 km for the EUS to match the 1° grid used for surface-wave dispersion
- and a d_2 of 160 km for the EUS to match the 1° grid used for surface-wave dispersion tomography. By spatially smoothing the receiver function wavefield, we sacrifice spatial
- resolution for simplicity and better average properties, but we may blur some geological
- transitions and boundaries. The resulting smoothed receiver functions are more complementary
- to the surface-wave dispersion data and comprise a data set more consistent with the other
- 265 observations used in the joint inversion and with less near-surface scattering effects that can
- complicate the inverse models.



Figure 2. Time slice at time -1.5 s (a, b), 3.4 s (c, d) and 18.1 s (e, f) from receiver function 268 wavefield in the EUS and adjacent Canada. Each circle represents a seismic station at which we 269 have computed a stacked receiver function. The color indicates the amplitude of (a, c, e) the 270 receiver function averaged at each station or (b, d, f) the smoothed/interpolated value of the 271 receiver function wavefield. Inset shows the stack of all RFs (clipped to show details), and the 272 red line shows the corresponding lag time of the time slice. See Movie S1 for all the time slices. 273 (For interpretation of the references to color in this figure, the reader is referred to the web 274 version of this article.) 275

4 Simultaneous inversion 277

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To estimate the subsurface shear wave speeds beneath the EUS, we simultaneously 278 inverted the smoothed/interpolated receiver function wavefield, Rayleigh wave group and phase 279 velocities in the period range from 3 to 250 s, and wavenumber-filtered Bouguer gravity 280 observations. The study region was divided into 950 1°-by-1° size cells for the EUS so the best 281 282 lateral resolution of our model is 110 km by 110 km. The resulting 3D grid was used for our simultaneous inversions in a hybrid 1D-3D manner. Receiver-function and surface-wave-283 dispersion calculations were performed with a 1D formalism for each cell across the region 284 where observations are available. Gravity calculations were computed in 3D using rectangular 285 prisms. The lateral dimension of the prisms is the same as the grid size whereas the vertical 286 dimension varies from 1 km near the surface to 50 km in the mantle. A linearized discrete 287 geophysical inversion developed from (Chai et al., 2015; Julià et al., 2000; Maceira & Ammon, 288 2009) was used with smoothness-based stabilization. A jumping strategy is used for 289

regularization to allow constraints directly on model parameters, not changes in model 290

parameters (e.g., Constable et al., 1987). The linearized inversion can be expressed as 291

$$\begin{pmatrix} \sqrt{\frac{p(T)}{\omega_s^2}} D_s \\ \sqrt{\frac{1}{\omega_r^2}} D_r \\ \sqrt{\frac{1}{\omega_r^2}} D_g \\ \frac{1}{\sqrt{\omega_g^2}} D_g \\ \frac{\eta \Delta}{W} \\ S \end{pmatrix} m_1 = \begin{pmatrix} \sqrt{\frac{p(T)}{\omega_s^2}} R_s \\ \sqrt{\frac{1}{\omega_r^2}} R_r \\ \sqrt{\frac{1}{\omega_r^2}} R_g \\ 0 \\ 0 \\ 0 \end{pmatrix} + \begin{pmatrix} \sqrt{\frac{1}{\omega_r^2}} D_r \\ \sqrt{\frac{1}{\omega_g^2}} D_g \\ 0 \\ 0 \\ 0 \end{pmatrix} m_0 + \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \\ W \\ 0 \end{pmatrix} m_a$$

where D_s , D_r and D_g are matrices containing the partial derivatives of the seismic shear velocity 293 model corresponding to the dispersion, receiver function, and gravity estimates, respectively. m_{θ} 294 is the 3D velocity model from previous iteration. m_a is the a priori velocity model. m_1 is the 295 updated velocity model for the current iteration. R_s , R_r and R_g are the corresponding data residual 296 vectors, and ω_s^2 , ω_r^2 and ω_g^2 are global weights assigned to the three data sets. Those weights are 297 defined in the same way as Julià et al. (2000) as $N\sigma_k^2$ where N is the number of data points for 298 the specific data set and σ_k^2 is the kth data point variance. p(T) was introduced by Maceira & 299 Ammon (2009) to control the trade-off between fitting both gravity and surface-wave dispersion 300 measurements. The matrix Δ applies vertical smoothing with a weight η to make the 1D velocity 301 302 profiles vary smoothly - necessary when data constraints are not sufficient. W is a diagonal matrix to constrain the velocities from varying too far from the a priori values in m_a with 303 associated weights. Lateral smoothing is added through the matrix S using the first differences 304 305 between shear velocity values in adjacent grid cells. The lateral smoothing does not smooth across the ocean-continent boundary to allow a sharp change in material properties across this 306 well-defined (for our cell size) feature. The inverse equation is solved using a Conjugate-307 Gradient LSQR solver for sparse linear equations (Paige & Saunders, 1982). The weights and 308 smoothing parameters are determined by performing suites of inversions and comparing data 309 misfits and model properties such as roughness. Interactive visualization tools were used to help 310 the comparison of inversion results with different weighting parameters (Chai et al. 2018). The 311 vertical smoothness constraints increase with depth to reflect a loss in data resolution and the 312

increase in surface wavelengths sampling the greater depth in the model. Partial derivatives for

- 314 surface wave dispersion were computed using finite-difference approximations. The simulation
- of surface-wave dispersion is based on algorithms from Saito (1988). The gravity derivatives
- were computed using the equations from Plouff (1976) and the chain rule. The inversion is performed to estimate shear-wave speed, which is related to P-wave speed by the V_p/V_s ratio,
- performed to estimate shear-wave speed, which is related to P-wave speed by the V_p/V_s ratio, and to the density using formulas described in Maceira & Ammon (2009) and Chai et al. (2015).
- The V_p/V_s ratio was fixed throughout the inversion (inherited from the initial model).

The inversion for the EUS started with an initial model that was based on Crust Model 320 1.0 (Laske et al., 2013). Velocities below the crust were initialized with the AK135 velocity 321 model (Kennett et al., 1995). Since flat-Earth codes were used to compute dispersion, the initial 322 model was flattened using the formulas from Biswas (1972) with a slightly modified density 323 transformation exponent (see the supplements of Chai et al., 2015 for details). Velocities and 324 densities were unflattened prior to the gravity partial calculation and flattened after the gravity 325 computation since the density-shear velocity relationship is based on laboratory measurements 326 (corresponding to the spherical model). 327

Based on the results from hundreds of 3D inversions (multiple cells) and thousands of 1D 328 inversions (single cell), it became clear that tuning the weights for each data set and constraints 329 is necessary to produce good fits to the observations. We experimented with weight selection 330 using numerous 1D inversions with observations selected from several samples in the overall 331 data set. We searched for optimal ranges of weight parameters by randomly testing 1D inversion 332 parameters, and examining the range of velocity models that resulted from the simulation. 333 Within this suite of velocity models, some fit the data poorly. For many choices of the weights, 334 the results were quite similar - those that fit the observations well often differed from the average 335 model by less than 0.1 km/s. The tests suggest we can use data fits as a guide to select both the 336 data and smoothness weights for the larger 3D inversions. To be sure that the resulting model fits 337 the observations reasonably well, we used interactive visualization tools (Chai et al., 2018) to 338 339 visually check data fits for all three types of observations.

The EUS results were obtained after eight total iterations of the 3D inversion (Figure S7). We began the inversion without gravity observations and modeled the seismic data for a total of four iterations. We then added the gravity observations for the final four iterations. The final EUS model has been unflattened for interpretation and easy comparison with published spherical Earth models.

345 **5 Data fits**

Fitting data is the most important quality control for our inversion. We discuss not only 346 the overall data fit of the three types of observations we use but also the spatial distribution of 347 data misfits. Fit to surface-wave dispersion and receiver functions is significantly improved in 348 the first four iterations and are stable for the last four iterations (Figure S7). Gravity misfit was 349 reduced greatly after one iteration. In general, our inverted model fits all three types of 350 observations much better than the starting model. Figure 3 shows the receiver function and 351 surface-wave dispersion misfit distributions of all cells for the initial and final models. In all 352 cases, an improvement in fit is clear. We use a 95% threshold to quantify the overall misfit. The 353 95% misfit threshold represents the fractional misfit value equal to or larger than that found for 354 95% of the cells. The metric is used to eliminate the influence of the worst fitting 5% of the cells 355 356 (which are often in the ocean, where data coverage is substantially more limited). As shown in

Figure 3, the 95% misfit threshold was improved from 15.5% (initial model) to 6.5% (inverted model) for group-velocity measurements, from 8% (initial model) to 2.5% (inverted model) for phase-velocity measurements, from 102% (initial model) to 33% (inverted model) for Gaussian 1.0 receiver functions, and from 132% (initial model) to 54% (inverted model) for Gaussian 2.5 receiver functions. The modest reduction in dispersion misfits reflects the relatively small range of speeds within the data. Any model with a roughly correct average dispersion value performs reasonably well with this metric. As expected, phase-velocity measurements are fit better than

- 364 group-velocity measurements. Narrowband receiver functions were fit better than broadband
- 365 receiver functions.



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Figure 3. Comparison of misfit using the initial model (lighter color) and the inverted model (darker color) for (a) phase velocity, (b) Gaussian 1.0 receiver functions, (c) group velocity, and (d) Gaussian 2.5 receiver functions. Dashed lines show the misfit value that is larger than misfits for 95% of measurements. A root-mean-squared (RMS) misfit is normalized with the RMS of the corresponding observation.

373 5.1 Receiver function misfits

In Figure S8, we show the spatial distribution (at available grid points) of receiver 374 function misfits for both the initial model and the inverted (final) model. Receiver function 375 misfits computed from the initial model are quite large at most grid points. The misfit shown in 376 the maps (Figure S8a and b) is the average of those corresponding to all available (up to six) 377 378 receiver functions. For example, at the location corresponding to Figure S8c, the misfit (represented as the circle size in Figure S8a and b) is computed from six receiver functions (three 379 ray parameter bins and two Gaussian widths). Comparing the observed and predicted receiver 380 functions in Figure S8c, we found the converted phase and multiples arrive later in initial-model 381 receiver functions. Receiver functions computed from the inverted model agree well with 382 observations. At many grid points, the receiver functions computed with the initial model differ 383 significantly from the observations, while the inverted model fit the observations nicely (see 384 Figure S8d for an example). Receiver function misfits corresponding to the inverted model are 385 uniformly small with slightly larger values in Gulf of Mexico coastal regions and the Williston 386 basin where thick sediments complicate receiver function waveforms. The relatively large misfits 387 in the eastern Tennessee and western North Carolina may due to anisotropy or less-optimal 388 V_p/V_s ratios. 389

390 5.2 Surface-wave dispersion misfits

The spatial distribution of surface-wave dispersion misfits (an average of group and 391 phase speed misfits) is shown in Figure S9. Group and phase velocities computed from the initial 392 393 model differ significantly from surface-wave dispersion observations for most model grid points as indicated by cells with large circles in Figure S9a. For example, we compared observed and 394 simulated dispersion curves at a grid point located in northeast Mississippi in Figure S9c. 395 Predicted dispersion curves based on the inverted model agree well with the observations while 396 simulated group and phase velocities from the initial model are too small at most periods. Even 397 for regions that the initial-model-derived surface-wave dispersion curves are similar to 398 observations (see Figure S9d for an example), the inverted model predicts the surface-wave 399 measurements better, especially at short periods (less than 20 s). As shown in Figure S9c, 400 surface-wave dispersion misfits corresponding to the inverted model are much smaller than those 401 from the initial model except for a few off-coast locations. Large misfits at these off-coast grid 402 points (at the gulf coast of Florida and near northern NewJersey) are likely due to limited 403 observations and a poor starting model. 404

405 5.3 Gravity misfits

Gravity observations are much fewer in number compared to receiver functions and 406 surface-wave dispersion. For this reason, we use a smaller weight for gravity observations and 407 include gravity data at a later stage of the inversion. The predicted gravity values agree with 408 Bouguer gravity observations well (Figure S10). Since the gravity calculations are performed in 409 3D, we used a buffer zone (gray color filled regions in Figure S10) to avoid edge effects. In 410 general, our goal with the gravity data is to fit the relatively high-wavenumber features that are 411 likely associated with crustal density variations. The addition of gravity to the inversion does not 412 introduce significant changes to the model. However, these changes are necessary to fit the 413 gravity observations. Numerical tests showed that small changes of roughly 0.1 km/s to the upper 414

415 15 km account for an improved fit to the gravity. We only fit the first-order features in gravity416 observations to avoid overfitting.

417 6 Results

Representative shear-velocity depth slices are shown in Figure 4. Maps in Figure 4a-b are 418 a plot of the average speed within the specified depth range. At shallow depth (0-5 km, Figure 419 4a), low velocities correlate well with major sedimentary basins and the coastal plains. Low 420 velocity layers near the surface are thick along the coastal plains and within the basins of the 421 Great Plains, compared with the Michigan, Illinois, and Appalachian basins. Average lower 422 upper-crustal shear velocities are lower beneath the Mid-continent rift, which interrupts the 423 relatively fast shallow crusts of Wisconsin and northern Minnesota and the Canadian Shield 424 regions to the north and continues southward through southern Minnesota and Iowa. There is a 425 hint (solid ellipse in Figure 4b) of a sharper transition along the northern (western) Mississippi 426 427 Embayment (ME) boundary than along the east. Seismicity extends from the lower shear wave speeds into a region of more normal speeds. On average, the Appalachian region and the 428 Midwest appear slightly faster than the region to the west that includes Wisconsin, Iowa, 429 430 Nebraska, Kansas, and Oklahoma. The fast region in the northwest lies beneath the Williston Basin and is not well resolved, but the model, driven by dispersion measurements at these 431 depths, suggests perhaps an unusually fast middle crust beneath the basin (Figure 4b). Two east-432 coast regions, the Carolinas and southern Virginia, and Maine appear slower than regions 433

immediately west in the middle crust (Figure 4b).

435 Figure 4c is a map of the shear-velocity model at 37 km depth. This depth was chosen to provide first-order information on crustal thickness - in areas of thin crust, we see mantle-like 436 speeds (dark blue colors, V_s larger than 4.4 km/s), while in many other areas, we simply see the 437 lower crust. The map identifies the southern ME and coastal regions as regions of unusually thin 438 crust. Regions of relatively fast deep crust are suggested in the eastern Dakotas, beneath 439 Michigan, and parts of Canada and New England (see Figure 5 for crust thickness variations). 440 We discuss specific estimates of crustal thickness later. Figure 4d is a map of the shear-wave 441 model at a depth of 63 km, which provides a sample of the shear-wave speeds in the uppermost 442 mantle. The range of velocity variation is about 6%. The image indicates a relatively slower 443 upper mantle along the east coast from New England to South Carolina. A low velocity feature 444 extends from the northern ME south to Louisiana and east Texas. The fastest speeds are to the 445 north, into the Canadian Shield region. 446

Figure 5 includes a crustal thickness map along with example velocity profiles that were extracted from the model. Example velocity profiles at eight different locations (Figure 5a-d, f-i) show detailed velocity changes as a function of depth. To define crustal thickness, we measured the depth corresponding to P-wave velocity larger than 7.8 km/s (based on visual inspections of velocity profiles). Only a few cells failed to reach this speed before a depth of 53 km. The automatically measured crustal thicknesses are shown as dashed gray lines in Figure 5a-d,f-i, which match the crust-mantle transition well visually at all presented locations.

Our smoothing approach to receiver functions and reliance on surface-wave dispersion tomography to constrain the deeper features in the model increase the consequences of the assumption of sharp features such as the crust-mantle boundary in the initial model (Crust 1.0) on the location of the crust-mantle transition. However, as shown earlier, even when the model is good, slight adjustments are made to improve the alignment of the converted phases originating from the crust-mantle boundary region. Although we did not perform a rigorous analysis, our

examination of the model suggests that when the constraints used to build Crust 1.0 were based

on good nearby data, little adjustment in crustal thickness was needed by our inversion. In

regions where we believe that Crust 1.0 relied heavily on interpolation, the inversion changes from our initial model were larger. Although adjustments are generally not large, typically 3 km,

464 our model is almost systematically thinner than Crust 1.0 by a few kilometers (both models share

465 the same V_p/V_s ratio).



466

Figure 4. Shear-velocity maps showing depth range (a) 0-5 km, (b) 15-31 km, (c) 37 km, and (d) 467 63 km from the 3D model. Thick black lines in (a) indicate major sedimentary units. Thick black 468 lines in (b), (c) and (d) show the physiographic boundaries. Thin black lines show state 469 boundaries. Warm colors show relatively slower regions, cool colors indicate relatively faster 470 regions. Although the colors are constant, the velocity range in each figure varies substantially. 471 The anomalies at 0-5 km depth are primarily corresponding to sedimentary basins. The velocity 472 changes at 15-31 km depth are related to lateral variations in mid-lower crust structure. 473 Abbreviations: AH – Appalachian Highlands; AM – Appalachian Mountains; AN – Anadarko 474 Basin; AP – Appalachian Basin; AT – Atlantic and Gulf Coastal Plains; BR – Basin and Range; 475 CS - Canadian Shield; DE - Denver Basin; GM - Gulf of Mexico; IH - Interior Highlands; IL -476 Illinois Basin; IN – Interior Plains; ME – Mississippi Embayment; MI – Michigan Basin; MR – 477 Midcontinent Rift; NAC - North America craton; OZ - Ozark Uplift; PA - Palo Duro Basin; PE 478

479 – Permian Basin; WI – Williston Basin. (For interpretation of the references to color in this

480 figure, the reader is referred to the web version of this article.)

481 Crustal thickness patterns for eastern North America are familiar. The crust is thinner
 482 near the coast and thicker landward. The average crustal thickness in the region is 43 km with a
 483 standard deviation of 5 km. Relatively thinner crust (< 43 km) is also imaged beneath eastern

- 484 North Dakota. The crust beneath the ME and into east Texas appears significantly thinner (< 35
- km km) than the continental interior. Thicker crust (> 43 km) is found beneath the Appalachian
- 486 Mountains and the Great Plains. The depth of the crustal-mantle transition is less well
- 487 constrained beneath thick sedimentary basins due to the dominance of basin reverberations in the 488 receiver functions. One potential application of our model is to form the basis of a wavefield
- receiver functions. One potential application of our model is to form the basis of a wavefield
 downward continuation and decomposition (Chai et al., 2017; Langston, 2011) to extract
- 489 receiver function signals generated from the crust-mantle transition from teleseismic P-wave
- 491 seismograms assuming the shallow structure is known. We leave that analysis for future efforts.



Figure 5. Crustal thickness map of (e) the EUS and (a-d, f-i) example velocity profiles at eight locations. Warm colors show relatively thinner crust, cool cools indicate relatively thicker crust. The average crustal thickness is 43 km with a standard deviation of 5 km. Green boxes on the map show the location of the corresponding velocity profile. In the profile plots (a-d, f-i), black lines represent shear-wave velocity while gray lines correspond to P-wave velocity. The dashed gray lines indicate the measured crustal thickness. (For interpretation of the references to color in this figure, the reader is referred to the web version of this article.)

We compared our crustal thickness measurements with several published models 501 502 (Crotwell & Owens, 2005; Porter et al., 2016; Schmandt et al., 2015; Shen & Ritzwoller, 2016). The EARS (Crotwell & Owens, 2005) project measured crustal thickness from automatically 503 computed P-wave receiver functions using H-k stacking method (Zhu & Kanamori, 2000) by 504 assuming a constant velocity in the crust. The original results may suffer from scattering noise in 505 receiver functions (e.g., Chai et al., 2015). We spatially smoothed the measured crustal thickness 506 from EARS to show first-order features (see Figure S11). Comparing our crustal thickness map 507 (Figure 5a) with four recently published models (Figure S12), we note that a thinner crust was 508 imaged along the east coast and beneath the ME for all models. However, notable differences 509 exist among these models. The average crustal thicknesses and the standard deviations from the 510 Shen-2016 model (Shen & Ritzwoller, 2016), the Schmandt-2014 model (Schmandt et al., 2015), 511 and our model are the same. The EARS model (Crotwell & Owens, 2005) may underestimate 512 crustal thicknesses due to oversimplified model parameterization (constant-velocity crustal 513 layer). The Porter-2015 model (Porter et al., 2016) used the EARS results to constrain crustal 514 thickness and shared the same issue. The common conversion point (CCP) stacking procedure 515 (Zhu, 2000) used for the Schmandt-2014 model reduced some scattering but had difficulty in 516 517 challenging regions (sedimentary basins for example) where stacking energy is not focused. To a large degree, the Schmandt-2014 model, the Shen-2016 and our model are consistent. Small 518 differences exsit between the Shen-2016 and our model (see Figure S13). The EARS model is 519 certainly suitable for a starting model. We believe that our model is less contaminated by 520 scattering in receiver functions and includes simultaneously-modeled constraints from Rayleigh 521 waves (similar to the Shen-2016 model) and near-surface gravity. 522

Figure 6a and b are maps of the average crustal and upper-mantle shear-wave speed, 523 respectively. The mean of the average crustal V_s velocity across the model continental grid 524 points is 3.7 km/s (standard deviation 0.1 km/s). The corresponding mean crustal V_p velocity is 525 6.5 km/s (standard deviation 0.2 km/s, see Figure S14). These mean values agree reasonably well 526 with the global compilation of Christensen & Mooney (1995). Relatively fast average crustal 527 528 velocity is found beneath the stable North America craton. Regions of very slow average crustal shear-wave speeds are associated with thick sediments in the southern ME. The mean uppermost 529 mantle shear-wave speed is the average of the upper 60 km of the mantle and it is 4.50 km/s 530 (standard deviation 0.05 km/s). The average uppermost mantle P-wave speed in the study area is 531 8.1 km/s (standard deviation 0.1 km/s), which is also consistent with the global compilation 532 (Christensen & Mooney, 1995) and regional surveys (Li et al., 2007). Note these P-wave speeds 533 were derived from shear-wave speed using V_p/V_s ratios inherited from Crust 1.0. The uppermost 534 mantle shear-wave velocity is slowest beneath New Mexico and the associated southern Basin 535 and Range province. We imaged relatively slow average shear-wave speeds in the uppermost 536 mantle beneath New England and the southeastern US, regions that may have interacted with 537 Great Meteor (e.g., Eaton & Frederiksen, 2007) and Bermuda (e.g., Cox & Van Arsdale, 2002) 538 hotspots respectively. Slow speeds along the east coast seem to follow the trend of eastern North 539 America rift basins but extend further into northern New England than the surface expression of 540 541 rifting. A slightly slower uppermost mantle is imaged beneath the western ME. Localized low average mantle speeds (dashed ellipses in Figure 6b) near the Kansas-Nebraska border and the 542 Kentucky-Ohio border are situated near kimberlite volcanic site and may represent modified 543 lithosphere associated with these volcanic structures. The relatively slow upper mantle weakly 544 correlates with the region of thrust faulting along the eastern margin of the U.S. and within 545

southeast Canada. The thrust faulting environment also correlates with the thin crust for much ofthe region, but not northern New York or southeastern Canada (Figure S15).

548

Figure 6. A summary of average crustal and mantle shear-wave velocity. (a) Average crustal 549 shear-wave velocity map, (b) uppermost mantle shear-wave velocity map, and histograms of 550 crustal V_s velocity (c) and uppermost mantle V_s velocity (d). The circles and rectangular boxes in 551 (c) show locations of dated kimberlite intrusions and lithosphere anomalies from Chu et al. 552 (2013), respectively. The mean (Avg.) of crustal V_s is 3.7 km/s with a standard derivation (Std.) 553 of 0.1 km/s. The mean of uppermost mantle V_s is 4.51 km/s with a standard derivation of 0.03 554 km/s. The dashed ellipses show the locations of two slow velocity anomalies described in the 555 text. (For interpretation of the references to color in this figure, the reader is referred to the web 556 557 version of this article.)

558 We do not see any evidence for slower mantle resulting from lithospheric delamination 559 and upwelling that is evident in geochemical observations (Mazza et al., 2014). Thus if 560 lithospheric delamination and mantle upwelling occurred as the rocks suggest, the feature must be quite localized. The West Virginia-Virginia border overlies a relatively abrupt transition in

- crustal thickness. The crust is relatively flat towards the Appalachian interior from west to east
- but begins to thin near the area containing the kimberlite intrusions (see Figure 6b for the
- 564 locations). More evidence for a significant change in lithospheric structure across the region is 565 the difference in mid-crustal speed, which decreases by about 0.2-0.3 km/s from the
- the difference in mid-crustal speed, which decreases by about 0.2-0.3 km/s from the
 Appalachians eastward (see Figure 4b). Although the transition is smoother, the upper mantle
- 567 speed also decreases towards the east in the same region. The differences in structure suggest
- 568 interesting geologic differences throughout the lithosphere from the Appalachians to central and
- 569 eastern Virginia.

In Figure S16, we merged the EUS model with the published model in the western U.S. (Chai et al., 2015). The smooth transition from west to east indicates that our EUS model is compatible with the published western U.S. model. The anomalous upper mantle beneath the western ME, the east coast, and New England is modest compared with that beneath the Basin and Range province in the western U.S.

An automated cluster analysis (Chai et al., 2015) was used to group the 1-D shear 575 velocity profiles into clusters of similar Earth structure. The spatial distribution of the clusters 576 using shear velocities between the depth range of 5 and 200 km is shown in Figure S17. The 577 clustered velocity models are shown in Figure S18. The crust is thinner beneath western ME, 578 southern Basin and Range region, southern Interior Plains, and Atlantic Plain. A thicker crust 579 was imaged beneath western Interior Plains, Appalachian Highlands, Interior Highlands, central 580 Interior Plains. The southeastern Canadian Shield shows a slighter thinner crust. The upper 581 mantle is faster beneath the stable North America craton. The seismicity in the region does not 582 show a simple correlation to the subsurface seismic speed variation on the broad image in Figure 583 S18. The clusters obtained with shear velocities between the depth range of 0 to 20 km is shown 584 in Figure S19 and S20. To a certain extent, the distribution of clusters (Figure S19) can be used 585 as a guide to the choice of a local 1D velocity model for initial earthquake location or moment 586 587 tensor inversion. For instance, the same 1D velocity model may be used for the region specified by the same cluster. We use more detailed cross-sections to explore the relationship between the 588 spatial distribution of earthquakes and subsurface structural variations. 589

590 7 Discussion

Large-scale regional lithospheric models contain many features that are often described 591 in a list of short commentary related to geology. Many of the structural features that we observed 592 have been pointed out in earlier studies (e.g., Chen Chen et al., 2016; Porter et al., 2016; 593 Schmandt et al., 2015; Shen & Ritzwoller, 2016); we noted several of these in the Results 594 section. One aspect not covered in other papers is the fact that a detailed crustal model provides 595 an opportunity to explore the relationship(s) between seismicity and subsurface structure (if any 596 exist). We use the cross-sections defined in Figure 1 to review the 3D model in details and to 597 explore potential relationships between structure and seismicity. The cross-sections were chosen 598 to cover most of the seismically active regions. Despite intensive studies, the cause of localized 599 seismicity in the region remains poorly understood. Many ideas have been proposed to explain 600 the intraplate seismicity of the EUS, such as stress concentration in regions of crustal weakness 601 or changes in lithospheric structure (Grana & Richardson, 1996; Grollimund & D. Zoback, 2001; 602 Kenner, 2000; Levandowski et al., 2016; Pollitz, 2001). Numerical models have shown that 603 crustal deformation can be induced at shallow depth when a dense anomaly resides in the crust 604

605 (Levandowski et al., 2016; Pollitz, 2001; Zhan et al., 2016) above a relatively weak upper

mantle. The depth and cause of a weak zone are still under debate (Chuanxu Chen et al., 2014;

607 Pollitz & Mooney, 2014). Since seismicity in Oklahoma may be related to human activity, we

skip that region (see Chai et al., 2021 for a detailed study of subsurface structure in that region).
 But note that ultimately, the Oklahoma activity is a result of the geologic structure associated

- 609 But note that ultimately, the Oklahoma activity is a result of the geologic structure associated 610 with the hydrocarbon-rich basins in the area. The EUS model contains no significant features
- deep beneath the Oklahoma seismic activity. Two regions of particular interest are the northern
- 612 ME and the Appalachian Mountains. We discussed some of the seismicity pattern and structure
- ⁶¹³ relationships for New Madrid in the Results section.

The cross-section A1-A2 (Figure 7) passes through seismically active regions in 614 Oklahoma, the NMSZ, and the Eastern Tennessee Seismic Zone (ETSZ). The NMSZ is 615 underlain by a relatively fast lower crust, which has been interpreted as a mafic intrusion 616 (Catchings, 1999; Mooney et al., 1983). Reconstruction of the feature details is difficult with the 617 coarse 100-km sampling, but the broader scale model shows the slower mantle in comparison 618 with regions to the east and west. Numerical tests and Rayleigh-wave sensitivity kernels suggest 619 that our resolution begins to degrade at about 150 km (Figure S21), so we do not image the entire 620 feature. Nyamwandha et al. (2016) suggested that a Cretaceous thermal event producing these 621 anomalies is associated with upwelling fluids from the Farallon Slab, along with an already 622 weakened and thinned lithosphere as a result of interaction with the Bermuda Hot Spot roughly 623 80-100 Ma (Cox & Van Arsdale, 2002). The crustal thickness is a maximum beneath the Valley 624 and Ridge and decreases by about 10 km at the coastal region to the east. Although not 625 associated with seismicity, the cross-section indicates relatively higher mid-crust speeds beneath 626 the Piedmont region of North Carolina. 627

Seismicity in the ETSZ occurs on basement faults that have no surface expression 628 (Steltenpohl et al., 2010). Seismicity is located in the Valley and Ridge Province west of the 629 highest Appalachian elevations. A recent local earthquake tomography study imaged the 630 631 existence of the buried fault (Powell et al., 2014) as a low-velocity anomaly. The lateral extent of the low anomaly is beyond the resolution of our velocity model. Our model contains a velocity 632 change in the lower crust beneath the ETSZ (ellipse in Figure 7), which may be associated with 633 the transition from the Granite-Rhyolite basement (west) to the Grenville basement (east) (Fisher 634 et al., 2010). Upper mantle shear velocity beneath the region is 2-3% faster than the mantle 635 beneath the NMSZ at depths of about 100 km. 636

The northwest-southeast striking cross-section B1-B2 (Figure 8) shows apparently 637 thinned crust beneath western North and South Dakota (solid ellipse) which is consistent with 638 639 Thurner et al. (2015) who argued that the apparently thin crust is actually a very fast underplated crust in the vicinity of the 2 Ga Trans-Hudson Orogen. The model at what could be considered 640 uppermost mantle depths is relatively slow, so it is plausible to interpret the material as fast crust. 641 A region of apparently thin crust in southeastern North Dakota is more perplexing. Thurner et al. 642 (2015) estimated a crustal thickness of 30-35 km in this region. Our apparent crustal thickness is 643 slightly larger, about 38-40 km. To be fair, reverberations in the thin surface sedimentary cover 644 interfere with the Ps arrival from the apparent crust-mantle boundary, which may affect an 645 accurate estimation of the crust-mantle boundary. The multiples in receiver functions from the 646 apparent crust-mantle boundary are free of near-surface interference, which provide some 647 contraints on the boundary. The P-wave speeds in the apparent lower crust are in the 6.6-7.0 648 km/s range, not unusual, and the values increase to 7.8 sharply and reach values of 8.0-8.1 km/s 649

by a depth of 43 km. The region has relatively low heat flow (50-60 mW/m², Blackwell et al.,

- 2011) though the data coverage is sparse. The apparently thinned crust may be a result of
- eclogite facies mafic material in the lowermost crust rendering the petrologic crust-mantle
- boundary seismically transparent (Furlong & Fountain, 1986; Griffin & O'Reilly, 1987). In
 cratonic areas, eclogite facies mafic material in the lowermost crust can have a P-wave speed
- larger than 8 km/s, which can lead to bias in the estimation of the crust-mantle boundary. Near
- the South Dakota Iowa border, the mid-crustal shear wave speed decreases relatively abruptly.
- 657 The relatively slow mid-crust (see also in Visualization S1) from Iowa through Missouri may be
- related to a batholith inferred from the Missouri gravity low (Hildenbrand et al., 1996). The structure from Iowa through Missouri is relatively uniform until interrupted by the relatively fast
- lower crust beneath the NMSZ. The upper mantle beneath Iowa and Missouri is relatively fast
- and uniform across this region of relatively slow middle crust. To the southeast of the NMSZ, the
- middle crust again is relatively slow and the crust begins to thin (dashed ellipse in Figure 8) and
- 663 the upper crust speed decreases near the Alabama-Georgia Border. The upper mantle below 100 664 km, from the NMSZ to the southeast is one of the slowest profiles in the model, and includes the
- region believed to have been crossed by the Bermuda Hot Spot (Chu et al., 2013; Cox & Van

667

Figure 7. Shear velocity cross-sections along A1-A2. The panel (a) used a color palette for

suitable crustal speeds. The panel (b) shows shear-wave speed changes in the upper mantle.

670 Circles are earthquakes located within 100 km of the cross-section. Black circles are events with

depth uncertainties less than 5 km. Gray circles represent earthquakes with larger depth

- uncertainties or without uncertainties. Note the image is vertically exaggerated.
- 673

Additional cross-sections can be found in Figure S22-S25 and discussed in Text S1. The inverted 3D seismic velocity model for the eastern United States is available in Data Set S1. The 3D seismic velocity model for the western United States from Chai et al. (2015) is provided in Data Set S2. We also provide an interactive tool (Chai et al., 2018) to easily view the 3D model with depth slides and depth profiles side by side for both the eastern United States (Visualization S1) and the western United States (Visualization S2).

680

Figure 8. Same as Figure 7 but for cross-section B1-B2. The ellipses indicate anomalies described in the text.

683 7 Conclusions

Using spatially smoothed P-wave receiver functions, surface-wave dispersion, and Bouguer gravity observations, we construct 3D shear-wave velocity models in the EUS. The average crustal thickness of the EUS model is 43 km; the average crustal shear speed is 3.7 km/s; the average uppermost mantle shear-wave velocity in the model is 4.5 km/s. We imaged thinner crust beneath New England, the east coast, and the ME. The relative slow average shear-wave speeds in the mantle beneath New England and the southeastern US may be linked to hotspots. 690 Comparing to a compilation of basement age (Lund et al., 2015), regions with thin crust were 691 formed after 670 Ma, thicker crust in the cratonic region formed before 1000 Ma.

A comparison of seismicity and subsurface shear velocity suggests that often, but not 692 universally, earthquakes locate near regions with seismic velocity variation. However, not all 693 regions of significant subsurface seismic speed changes are loci of seismicity. The eastern 694 695 seaboard mantle appears slow, consistent with coarser, but deeper sampling models that have been used as a basis for estimating dynamic topographic changes along the eastern seaboard 696 (Rowley et al., 2013). A weak correlation between upper mantle shear velocity and earthquake 697 focal mechanism has been observed. A relatively small upper mantle low-speed region in eastern 698 Kentucky and southwestern Ohio correlates with the area of perturbed upper mantle P-waves 699 analyzed by (Chu et al., 2013) which they associated with the circa 75 Ma kimberlite volcanic 700 site near Elliot Kentucky. The northern ME, and in particular the region of the large earthquakes 701 in 1811-12 appears to be underlain by a relatively fast lower crust and a relatively slow 702 uppermost mantle. Levandowski et al. (2016) suggested that such a structure can focus stress in 703 the upper crust, and our model is consistent with the idea. 704

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@AGUPUBLICATIONS

1	
2	Geochemistry, Geophysics, Geosystems
3	Supporting Information for
4	Crust and Upper Mantle Structure Beneath the Eastern United States
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24	Introduction

This supplement includes details on the formulas to convert seismic velocities to densities (Text S1); paragraphs describing cross-sections B1-B2, D1-D2, E1-E2, F1-F2, G1-G2, and H1-H2 (Text S2); a map showing the seismic stations used for receiver functions (Figure S1); a map showing the seismic events used for receiver functions (Figure S2); a figure showing the relative data variance of receiver functions (Figure S3); a figure

30 showing example surface-wave dispersions (Figure S4); a figure showing Bouguer

31 gravity maps before and after wavenumber filtering (Figure S5); a comparison of single-32 station receiver functions and smoothed receiver functions (Figure S6); a figure showing 33 the convergence of the simultaneous inversion (Figure S7); a figure showing receiver 34 function misfits (Figure S8); a figure showing surface-wave dispersion misfits (Figure 35 S9); a comparison of the observed gravity maps and the predicted gravity maps (Figure 36 S10); a comparison of EARS crust thickness before and after smoothing (Figure S11); a 37 comparison of crustal thickness maps from recent seismic velocity models (Figure S12); 38 a comparison of 1D velocity profiles (Figure S13); a figure showing the distributions of 39 crustal and uppermost mantle P-wave velocities (Figure S14); a comparison of the 40 uppermost mantle V_s and earthquake focal mechanisms (Figure S15); a comparison of the 41 upper mantle shear speed at 63 km depth in western and easter U.S. (Figure S16); maps 42 showing cluster locations (Figure S17 and S19); velocity profiles for each cluster (Figure 43 S18 and S20); Rayleigh wave sensitivity kernels (Figure S21); V_s cross-section C1-C2 44 (Figure S22); V_s cross-section D1-D2 (Figure S23); V_s cross-section E1-E2 and F1-F2 45 (Figure S24); V_s cross-section G1-G2 and H1-H2 (Figure S25); the final seismic velocity 46 model for the eastern United States in a text file (Data Set S1); the seismic velocity model 47 for the western United States in a text file (Data Set S2); a list of seismic networks used 48 in an excel file (Table S1); an animation compares the single-station-averaged receiver 49 functions and the smoothed version (Movie S1); and interactive tools to view the seismic 50 velocity model for the eastern United States (Visualization S1) and for the western

51 United States (Visualization S2).

52 **Text S1.**

We used the following formulas from Maceira & Ammon (2009) to convert seismic
 velocities to densities.

$$\rho(\alpha) = \cos^2 \phi \, \rho_2(\alpha) + \sin^2 \phi \, \rho_1(\alpha)$$
$$\alpha + 2.40$$

56

57

55

$$\rho_1(\alpha) = \frac{\alpha + 2.16}{3.125}$$

$$\rho_2(\alpha) = \frac{7.55 - \sqrt{57.00 - 10.56(6.86 - \alpha)}}{5.20}$$

 $\phi = \frac{\pi}{2} \frac{1 + \tanh[\xi(\alpha - \alpha_c)]}{2}$

59 In the formulas, ξ is 0.5 and α_c is 6.2 km/s. α is the seismic P wave velocity in km/s and ρ is density in g/cm³.

61 **Text S2.**

62 The cross-section C1-C2 (Figure S22) is parallel to the longer arm of the New Madrid 63 seismic zone (NMSZ) and passes through the Wabash Valley seismic zone (WVSZ) of 64 Illinois and Indiana. Compared to the region to the north (left in Figure), the crust hosting 65 modern WVSZ seismicity is relatively faster, with a smaller velocity gradient in the mid-66 to-lower crust. The WVSZ is underlain by a slightly thicker crust than the NMSZ. 67 However, the upper mantle beneath the WVSZ is faster than that beneath the NMSZ. A 68 broad higher velocity anomaly is imaged beneath the WVSZ about 70-150 km depth,

69 which agrees with a recent local study (Chen Chen et al., 2016). The continuation of a

- 70 relatively fast lower crust beneath the NMSZ northward to the WVSZ may suggest these
- 71 two seismic active regions are connected (Chen Chen et al., 2016). To the south-
- 72 southwest of the NMSZ, deeper into the Mississippi Embayment, the sedimentary cover
- 73 thickens to at least 10 km and the crust thins to roughly 30 km.

74 The cross-section D1-D2, E1-E2, F1-F2, G1-G2, and H1-H2 pass through and across the 75 Appalachian Mountains. Cross-section D1-D2 (Figure S23) clearly shows a slower upper 76 mantle beneath New England, which is consistent with other studies (Pollitz & Mooney, 77 2016; Schmandt et al., 2015; Shen & Ritzwoller, 2016). The decrease seismic velocity 78 has been interpreted as due to interaction with the Great Meteor hotspot roughly 50 Ma 79 (Eaton & Frederiksen, 2007). Along the profile, which samples the Valley and Ridge 80 Province of central and eastern Pennsylvania, the crustal thickness increases by roughly 81 10 km from eastern Pennsylvania into the West-Virginia border. Crustal thicknesses in 82 western Pennsylvania are comparable to those to the south (see Figure 5). Although the 83 mid crust varies in speed along the profile, the change in crustal thickness appears to arise 84 from an increase in thickness of lower crustal material. Depths of earthquakes along the 85 profile are generally above 25 kilometers and show no systematic variation with the

86 structure.

87 Cross-section E1-E2 (Figure S24) samples from southeastern Canada into New England 88 and crosses the West Ouebec seismic zone (WOSZ). Magnitude 3 and larger earthquakes 89 extend to 20 km depth in the WQSZ region and appear to shallow slightly along the 90 profile in New York and New England. Along this direction, the WQSZ locates near the 91 edge of the Canadian Shield as is evident in the mantle speed cross-section. Cross-section 92 F1-F2 (Figure S24) shows a crustal thickness change beneath central Pennsylvania. 93 Crossing from the Appalachian Plateau to the Valley and Ridge Province, the upper crust 94 slows and the lower crust thins. At roughly the same position the mantle speeds decrease. 95 Seismicity in the region shows transitions from reverse faulting in the thinner southeast 96 part of the state to strike-slip faulting in the northwest. Whether the stress change is 97 associated with the structure within the crust and/or upper mantle is difficult to tell. The 98 pattern of reverse faulting continues down the eastern seaboard along the area of 99 relatively thin crust. But reverse faulting in northern New York and southeastern Canada 100 occurs with crust with more typical interior thicknesses. The pattern is slightly better 101 matched with reverse faulting occurring above the regions of the relatively slow 102 uppermost mantle (Figure S24), so perhaps the change (from South Carolina to Ottawa) 103 is a result of an overall variation in lithospheric strength.

104 Cross-section G1-G2 (Figure S25) crosses eastern Ohio and through central Virginia and 105 into northwestern North Carolina. Crustal thickness in eastern Ohio is comparable to that 106 under the Appalachians, or perhaps slightly thinner. As discussed earlier in the Results 107 section, the crustal thickness changes quickly as you exit the Appalachians to the east. 108 Mantle speeds decrease modestly, but steadily from Ohio to the Appalachians. The 2011 109 M5.7 Virginia earthquake was located near an edge of a faster lower crust anomaly (the 110 solid ellipse in Figure S25) and a change in crustal and upper mantle structure. Cross-111 section H1-H2 (Figure S25) crosses from the northeast WVSZ to the Charleston region 112 and the South Carolina Seismic Zone. The crustal thickness increases slightly from the 113 midwest into the Appalachians, and seismicity appears to extend slightly deeper in the

114 ETSZ near the profile. Near the southeastern margin of the Appalachians, into the coastal 115 plain, the depth range of slower crustal material increases. The material that could be 116 called lower-mid crust in the midwest and Appalachians disappears as the material with 117 typical lower-crustal speeds shallows with a thinning of the crust. At the same position, 118 the mantle speeds decrease along the east coast. Mantle speeds decrease modestly, but 119 steadily from Indiana to the Appalachians, crossing the region that (Chu et al., 2013) 120 suggested a hidden hot-spot track. Along the profile, we see no evidence for a slow upper 121 mantle. However, our model includes a slight reduction in average upper mantle speed in 122 northeast Kentucky and southwest Ohio (see Figure 6), directly above the turning points 123 of the rays that showed delayed travel times and frequency-dependent amplitudes 124 analyzed by (Chu et al., 2013) (the signals were generated by the Virginia earthquake and 125 recorded on midwest Transportable Array stations described in that work). The slow 126 upper mantle anomaly in our model is shallower than where they placed the anomaly but 127 may reflect the same feature. However, we do not see it extend to the west, towards the northern Mississippi Embayment, as they suggested. 128

129

Figure S1. Map of the study region and seismic stations (dots) used in the receiver function
 wavefield smoothing/interpolation. Red dots show the stations used in Figure S6.

Figure S2. Seismic events (dots) used for the receiver function calculations.

Figure S3. Relative data variance distribution of receiver functions recorded in the Eastern U.S.
 region. The insets show a detailed view for relative variance ranges between 165% and 400%
 (dashed box). Black lines indicate extreme value distributions. The relative data variance is
 computed as the signal variance between an individual receiver function and the single-station averaged receiver function.

144 **Figure S4.** Example dispersion measurements and blended curves in the eastern U.S. (latitude

44.5°N, longitude 73.5°E) for (a) Rayleigh-Wave group velocity and (b) Rayleigh-Wave phase

velocity. The blended curves were computed using values from Ekström (2011) and Herrmann et

al. (2021). Dispersion models from other sources (Bensen et al., 2007; Ekström, 2014; Jin &

- 148 Gaherty, 2015) are only shown for reference.
- 149

151 **Figure S5.** Bouguer gravity maps before (a) and after (b) wavenumber-filtering to emphasize

152 gravity anomalies related to shallow structure and features have spatial dimension larger than

153 1°. Note the color scale changes.

Figure S6. A comparison of (a) single station averaged and (b) smoothed/interpolated receiver

- 157 functions in a cross-section view. The locations of the stations used are shown in Figure S1.

160 **Figure S7.** Convergence of the simultaneous inversion. The gravity observations were included in

161 the last three iterations. Misfits for each type of observations are normalized with the

162 maximum. For each iteration, misfits were averaged over the entire grid. At each grid location,

163 dispersion misfit was averaged between group and phase velocity measurements while receiver

- 164 function misfit was averaged from all available receiver functions. RF stands for receiver
- 165 function.

167

Figure S8. Maps showing the spatial distribution of receiver-function misfits for (a) the initial model and (b) the final model. The size of dots represents the normalized misfit at each grid point. Receiver functions at two grid points, (34.5°, -88.5°) and (35.5°, -84.5°), are showing in (c) and (d) with the locations indicated by the dashed black lines. Receiver functions with different ray parameters are displayed for both the narrow-band (Gaussian 1.0, top frame) and the broad-band (Gaussian 2.5, bottom frame). Since the broad-band receiver functions are noisier, we used a shorter time window.

Figure S9. Maps showing the spatial distribution of the surface-wave dispersion misfits for (a) the initial model and (b) the inverted model. The size of dots represents normalized misfit at each grid points. Rayleigh-wave dispersion curves at two grid points, (34.5°, -88.5°) and (35.5°, -84.5°), are shown in (c) and (d) with the location indicated by the dashed black lines. The upper panel of (c) and (d) shows phase velocities while the lower panel shows group velocities.

Figure S10. A comparison of (a) the observed gravity against (b) the predicted gravity for the

185 eastern U.S.

187 Figure S11. A comparison of EARS crustal thickness results (Crotwell & Owens, 2005) from (a)

188 the raw data and (b) the spatially smoothed version.

191 Figure S12. A comparison of crustal thickness maps from four recent models, (a) Shen-2016

192 model (Shen et al., 2016), (b) Schmandt-2014 model (Schmandt et al., 2015), (c) Porter-2015

193 model (Porter et al., 2016), and (d) the smoothed EARS model (Crotwell & Owens, 2005). The 194 results of this paper are shown in Figure 5.

197 Figure S13. A comparison of 1D shear-wave velocity profiles at two grid points. The Shen-2016

198 model is extracted from Shen et al. (2016). The initial model is obtained from Crust 1.0.

Corresponding receiver functions and surface-wave dispersions can be found in Figure S8 andS9, respectively.

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- 202

Figure S14. Maps showing (a) the averaged crustal V_p velocities, (b) uppermost mantle Vp velocities and histograms of (c) crustal and (d) uppermost mantle V_p velocity.

Figure S15. A comparison of the uppermost mantle V_s and focal mechanisms from the Saint Louis University (SLU) catalog (prior 2021/05/01).

Figure S16. A comparison of the upper mantle shear speed at 63 km in western and eastern U.S.

The dashed line indicates the transition from the western U.S. model (Chai et al., 2015) to the eastern U.S. model (this study).

219 Figure S17. Spatial distributions of the automated model clusters generated using a simple 220 hierarchical clustering algorithm. Similar 1D velocity profiles were grouped together as one 221 cluster. Velocity profiles with large differences were assigned to different clusters. Each 1D 222 velocity profile consists of shear velocities between 6 km and 200 km in depth. Note the oceanic 223 profiles were assigned based on geographic location. The rough correspondence of the clusters 224 to geologic regions are (0) Oceanic; (1) Southern Basin and Range Region; (2) Western 225 Mississippi Embayment; (3) Central Basin and Range Region; (4) Atlantic Plain and Northern 226 Appalachian Highlands; (5) and (6) Interior Plains, Central and Southern Appalachian Highlands, 227 and Southern Canadian Shield. The velocity profiles within each cluster are summarized in Figure 228 S18. Cirlces represent seismic events with a magnitude larger than 3 from the USGS NEIC catalog 229 before May 2021.

Figure S18. Shear velocity profiles of Earth model clusters corresponding to Figure S17. Each 1D velocity profile consists of shear velocities between 6 km and 200 km in depth. The label above each panel corresponds to a cluster in Figure S17. In each panel, individual 1D shear velocity profiles belong to the cluster are shown. Velocity profiles within the cluster are sorted from north to south by row (like in a book). Dots shows seismicity (magnitude 3 and larger) from the USGS NEIC catalog before May 2021.

Figure S19. Same as Figure S17 but used the shear velocities of the upper 20 km.

Figure S20. Same as Figure S18 but the clustering used the shear velocities of the upper 20 km.

- 243 The label above each panel correspond to that in Figure S19.

247 Figure S21. Rayleigh wave sensitivity to shear wave velocity.

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Figure S22. Shear velocity cross-sections along C1-C2. Top panel used a color palette that is suitable for crustal speeds. The lower panel shows shear-wave velocity using a color palette more suitable for mantle speeds. Circles are earthquakes located within 100 km of the crosssection. Black circles are events with depth uncertainties less than 5 km. Gray circles represent earthquakes with larger depth uncertainties or without uncertainties. Note the image is vertically exaggerated.

Figure S23. Same as Figure S22 but for cross-section D1-D2.

261 Figure S24. Same as Figure S22 but for cross-sections E1-E2 and F1-F2.

Figure S25. Same as Figure S22 but for cross-section G1-G2 and H1-H2. The solid ellipse

265 incidates a faster lower curst anomaly.

267 **Table S1.** A list of seismic networks used for the receiver function calculation.

- Data Set S1. The final seismic velocity model for the eastern United States derived from the
 inversion. The first column is latitude. The second column is longitude. The third column is depth
 (top of the cell) in kilometers. The fourth, fifth and sixth column are P-wave velocity (km/s), S-
- 272 wave velocity (km/s), and density (g/cm³), respectively.
- 273
- Data Set S2. The seismic velocity model for the western United States from Chai et al. (2015).
 The first column is latitude. The second column is longitude. The third column is depth (top of
 the cell) in kilometers. The fourth, fifth and sixth column are P-wave velocity (km/s), S-wave
 velocity (km/s), and density (g/cm³), respectively.
- 278
- 279 **Movie S1.** An animation compares the single-station-averaged receiver functions against the spatially smoothed/interpolated receiver functions.
- 281
- Visualization S1. An interactive tool to view S-wave velocities of the 3D model for the eastern United States as depth slides and depth profiles side by side. The visualization was created with
- a Python script developed by Chai et al. (2018).
- 285
- 286 Visualization S2. An interactive tool to view S-wave velocities of the 3D model for the western
- 287 United States (from Chai et al., 2015) as depth slides and depth profiles side by side. The
- visualization was created with a Python script developed by Chai et al. (2018).
- 289