# Geochemical characterization of the Oman Crust-Mantle transition zone, OmanDP Holes CM1A and CM2B

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#### Abstract

The transition from the gabbroic oceanic crust to the residual mantle harzburgites of the Oman ophiolite has been drilled at Holes CM1A and CM2B (Wadi Tayin massif) during Phase 2 of the International Continental Scientific Drilling Program (ICDP) Oman Drilling Project (OmanDP) (Nov. 2017-Jan. 2018). In order to unravel the formation processes of ultramafic rocks in the Wadi Tayin massif (CM) crust-mantle transition zone and deeper in the mantle sections beneath oceanic spreading centers, our study focuses on the whole rock major and trace element compositions (together with CO2 and H2O concentrations) of these ultramafic rocks (56 dunites and 49 harzburgites). Despite extensive serpentinization and some carbonation, most of the trace element contents (REE, HFSE, Ti, Th, U) record high temperature, magnatic process-related signatures. Two major trends are observed, with good correlations between (1) Th and U, Nb and LREE on one hand, and between (2) HREE, Ti and Hf on the other hand. We interpret the first trend as the signature of late melt/peridotite interactions as LREE are known to be mobilized by such processes ('lithospheric process'), and the second trend as the signature of the initial mantle partial melting ('asthenospheric process'), with little or no overprint from melt/rock reaction events.

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- 25 Keywords: Partial melting vs. melt-rock reaction; Dunites and harzburgites serpentinization and
- 26 carbonation; Oman ophiolite Crust-Mantle transition; Holes CM1A and CM2B; ICDP Oman
  - 27 Drilling Project;
  - 28 Key Points:
  - The transition from the oceanic crust to the mantle of Oman has been drilled in the CM
     Holes during Phase 2 of the ICDP Oman Drilling Project

- There is large petrological and chemical variability in the dunites and harzburgites from
   Holes CM1A and CM2B
- Partial melting *vs*. melt-rock reaction, and the effects of serpentinization and carbonation
   of dunites and harzburgites are investigated

# 35 Abstract

The transition from the gabbroic oceanic crust to the residual mantle harzburgites of the Oman 36 ophiolite has been drilled at Holes CM1A and CM2B (Wadi Tayin massif) during Phase 2 of the 37 International Continental Scientific Drilling Program (ICDP) Oman Drilling Project (OmanDP) 38 (Nov. 2017-Jan. 2018). In order to unravel the formation processes of ultramafic rocks in the Wadi 39 40 Tayin massif (CM) crust-mantle transition zone and deeper in the mantle sections beneath oceanic spreading centers, our study focuses on the whole rock major and trace element compositions 41 (together with CO<sub>2</sub> and H<sub>2</sub>O concentrations) of these ultramafic rocks (56 dunites and 49 42 harzburgites). Despite extensive serpentinization and some carbonation, most of the trace element 43 contents (REE, HFSE, Ti, Th, U) record high temperature, magmatic process-related signatures. 44 Two major trends are observed, with good correlations between (1) Th and U, Nb and LREE on 45 one hand, and between (2) HREE, Ti and Hf on the other hand. We interpret the first trend as the 46 signature of late melt/peridotite interactions as LREE are known to be mobilized by such processes 47 ('lithospheric process'), and the second trend as the signature of the initial mantle partial melting 48 ('asthenospheric process'), with little or no overprint from melt/rock reaction events. 49

# 50 Plain Language Summary

We focus on the transition from the oceanic crust to the Earth's mantle by studying Holes CM1A 51 and CM2B, drilled in the Oman ophiolite during Phase 2 of the International Continental Scientific 52 53 Drilling Program (ICDP) Oman Drilling Project (OmanDP). Despite extensive serpentinization and some carbonation, the dunites and harzburgites from the transition zone and the mantle section 54 55 show a large variability in their petrological and chemical compositions. Results indicate that most of the trace element contents (REE, HFSE, Ti, Th, U) record high temperature, magmatic process-56 related signatures. Two major trends are observed, with good correlations between (1) Th and U, 57 Nb and LREE on one hand, and between (2) HREE, Ti and Hf on the other hand. We interpret the 58 first trend as the signature of late interactions between a percolating melt and the harzburgites 59 and/or dunites, and the second trend as the signature of the initial mantle partial melting, with little 60 or no overprint from melt/rock reaction events. 61

# 62 **1 Introduction**

Melts play a fundamental role in the lithospheric mantle chemical and mineralogical 63 heterogeneities, and have a large effect on mantle rheology, viscosity and seismic anisotropy 64 (Batanova & Savelieva, 2009; Kelemen et al., 1997; Tommasi & Vauchez, 2015). Numerous 65 studies have been dedicated to melt-rock interaction characterization in both the continental and 66 the oceanic lithospheric mantle (e.g. Bodinier et al., 1990; Dalton et al., 2017; Dygert et al., 2016; 67 Godard et al., 2008; Kelemen et al., 1998, 1990; Kelemen & Ghiorso, 1986; Parkinson & Pearce, 68 1998; Takazawa et al., 1992; Vauchez et al., 2005; Morgan et al., 2008; Navon & Stolper, 1987; 69 70 Niu, 1997; Warren et al., 2009; Warren and Shimizu, 2010). Several studies demonstrated that trace element variations coupled with microstructural, mineralogical and petrological 71 72 observations, and trace element numerical modeling, are a pertinent way to evaluate melt transport and constrain melt-peridotite processes (e.g. Navon & Stolper, 1987; Batanova et al., 1998; Godard 73 et al., 1995; Kelemen et al., 1995; Kelemen & Ghiorso, 1986; Kourim et al., 2014; Oliveira et al., 74 2020). Despite all these studies, the nature of melt and/or fluids involved in the reactional processes 75 76 in the oceanic upper mantle below spreading centers remains debated. One of the biggest challenges to understanding these processes is the collection of representative natural sample 77

suites. Finding locations where samples that have not been affected by either incomplete melt

extraction or interaction with melt, coexisted at the same site with samples that were affected by

80 either partial melt and/or interaction with melts migrating through the mantle is difficult. Abyssal

peridotites (e.g. Godard et al., 2008; Johnson et al., 1990; Niu, 1997; Parkinson & Pearce, 1998)
and mantle xenoliths (e.g. Bedini & Bodinier, 1999; Dalton et al., 2017; Fitzpayne et al., 2018;

Grégoire et al., 2001) are good candidates to study mantle processes in present-day oceanic and

continental settings, but their sampling is exceptional, lacking second-order geologic context and

is limited to the uppermost oceanic and continental mantle.

Oman ophiolite has been instrumental in elucidating the accretion and evolution of oceanic 86 lithosphere in present-day oceans and exhibits the largest ophiolitic exposures of oceanic 87 lithosphere worldwide. The mantle section of the Oman ophiolite is mainly composed of depleted 88 harzburgites and of some dunites, and has been the subject of many petrological, geochemical and 89 structural studies (e.g. Boudier & Coleman, 1981; Ceuleneer et al., 1988; Dygert et al., 2017; 90 91 Kelemen et al., 1995; Godard et al., 2000; Le Mée et al., 2004; Monnier et al., 2006; Nicolas et al., 2000; Takazawa et al., 2003). The general consensus stands that the dunites, as channels in the 92 mantle section or massive at the crust-mantle transition, are residues of reaction between a melt 93 undersaturated in silica at low pressure and mantle harzburgites; this reaction leads to the complete 94 consumption of orthopyroxene and to the concomitant precipitation of olivine (e.g. Abily & 95 Ceuleneer, 2013; Boudier & Nicolas, 1995; Braun et al., 2002; Godard et al., 2000; Kelemen et 96 al., 1995, 1997; Koga et al., 2001; Ouick, 1981b; Rabinowicz et al., 1987; Rospabé et al., 2017, 97 98 2018a, 2019a). However, the relationship between the harzburgites and the dunites, the nature of the reactant melt, and the chemical budgets related to the 'dunitization' process itself, are still 99 debated. Oman ophiolite exposes large portions of the mantle and crust-mantle transition zone, 100 suitable to understanding local to large scale studies of mantle heterogeneities and melt/peridotite 101 reaction processes. The Oman Drilling Project (OmanDP) enabled sampling of a continuous 102 section of the crust-mantle transition at Holes CM1A and CM2B (Wadi Tayin Massif, during 103 104 Phase 2 of the ICDP OmanDP, Nov. 2017-Jan. 2018), starting from the base of the layered gabbroic crust and going through the uppermost harzburgitic mantle (Kelemen et al., 2020a, 105 2020b; Proceedings available at https://www.omandrilling.ac.uk/). In this paper, we characterize 106 the major and trace element contents of the dunites and harzburgites from the Hole CM1A and 107 108 CM2B drill cores to better constrain dunitization processes by first, gaining insights into Wadi Tayin mantle and crust-mantle geochemical characteristics, then, comparing these characteristics 109 110 to previously studied Maqsad diapir harzburgites and mantle-crust transition zone, taking advantage of the continuous and regular, high resolution sampling performed in the Oman Drilling 111 Project. 112

# 113 2 Geological setting and context of CM drill cores

# 114 **2.1 Geology of the Samail ophiolite**

115 The Samail ophiolite, located in the Sultanate of Oman and the United Arab Emirates (Fig. 1a),

exposes a relatively continuous section of oceanic lithosphere, with, from top-to-depth, a 5-7 km-

thick crust made of pillow basalts, a sheeted dike complex and gabbros, overlying the crust-mantle

transition at the top of the upper mantle peridotites (e.g. Coleman & Hopson, 1981; Glennie et al.,

119 1974; Searle & Malpas, 1980, Lippard et al., 1986; Nicolas et al., 1988; Nicolas, 2012 and

120 references therein). According to the ages of pelagic sediments interbedded with basalts and of

zircons in evolved gabbroic and plagiogranitic rocks, the accretion event that led to the formation 121 122 of the Oman ophiolite has been estimated at around 94-97 Ma ago (Rioux et., 2012, 2013, 2016; Tilton et al., 1981; Tippit et al., 1981; Warren et al., 2005). The tectonic setting in which the Samail 123 124 ophiolite evolved is still debated. The spatial distribution along the ophiolite of the nature of the (1) mafic dikes cutting-across the mantle section, and (2) lower crustal cumulates, as well as their 125 geochemical signature, attest that both MORB like and depleted calc-alkaline series coexisted 126 during the igneous evolution of the ophiolite (e.g. Benoit et al., 1996; 1999; Ceuleneer et al., 1996; 127 Clénet et al., 2010; Python & Ceuleneer, 2003; Python et al., 2008). The dikes belonging to the 128 MORB-like volcanic units have been mapped mainly in the SE of the ophiolite and in other more 129 restricted spots, whereas the depleted calc-alkaline series were observed at a more widespread 130 131 scale (Python & Ceuleneer, 2003).

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Along the Samail ophiolite, spatially constrained vertical flow structures frozen within the 133 mantle section were interpreted as former asthenospheric diapirs distributed along the oceanic 134 ridge (Ceuleneer, 1991; Ceuleneer et al., 1988; Jousselin et al., 1998; Nicolas et al., 1988, 2000). 135 This mantle section is mainly composed of harzburgites (85 to 95%), relatively depleted with a 136 typical orthopyroxene content of 15-25% and locally grading into lherzolites, and to a lesser extent 137 of dunites (5-15%) (Boudier & Coleman, 1981; Lippard et al., 1986). The crust-mantle transition 138 is mainly made of dunites and wehrlites and its thickness varies from a few meters to a few hundred 139 140 meters (e.g. Abily & Ceuleneer, 2013; Boudier & Nicolas, 1995; Ceuleneer & Nicolas, 1985; Jousselin & Nicolas, 2000; Koga et al., 2001; Rospabé et al., 2017, 2018a). 141

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Geochemical studies have demonstrated the overprint of partial melting and of 143 melt/peridotite reaction processes in the mantle harzburgites' signatures (Gerbert-Gaillard, 2002; 144 Girardeau et al., 2002; Godard et al., 2000; Hanghøj et al., 2010; Kanke & Takazawa, 2014; Khedr 145 et al., 2014; Le Mée et al., 2004; Monnier et al., 2006; Takazawa et al., 2003). In this context, 146 mantle dunites and dunites from the crust-mantle transition zone (CMTZ) have mostly been 147 interpreted as replacive in origin, products of melt-harzburgite reaction leading to the complete 148 consumption of orthopyroxene and concomitant precipitation of olivine (e.g. Abily & Ceuleneer, 149 2013; Boudier & Nicolas, 1995; Gerbert-Gaillard, 2002; Godard et al., 2000; Kelemen et al., 1995, 150 1997; Koga et al., 2001; Rabinowicz et al., 1987; Rospabé et al., 2018a). This dunitization process 151 may have been enhanced by the involvement of a hydrous component in the reaction (Rospabé et 152 al., 2017, 2018a, 2019a). However, a reaction origin and a cumulate origin are not mutually 153 exclusive as it has been shown that the uppermost part ( $\sim 20\%$ ) of the crust-mantle transition may 154 have a composition consistent with cumulates while the main lower part (~80%) has a composition 155 supporting the replacive origin (Abily & Ceuleneer, 2013). Furthermore, as olivine-saturated melt 156 begins to cool conductively, hybrid processes, termed 'relative crystallization' (Collier & Kelemen 157 2010) produce reactive characteristics (e.g. Benn et al., 1988; Boudier and Nicolas, 1995; Koga et 158 159 al., 2001; Abily and Ceuleneer, 2013; Rospabé et al., 2018a).

# 160 **2.2 The crust-mantle transition at Sites CM1 and CM2**

Samples studied in this paper were drilled in the Wadi Tayin massif in the SE of the ophiolite during Phase 2 of the ICDP OmanDP (Nov. 2017-Jan. 2018). According to structural and petrological maps (Gerbert-Gaillard, 2002; Python & Ceuleneer, 2003; Nicolas et al., 2000), this site is located near the border - or in an intermediate position between the border and the axis of the frozen paleo-spreading center centered on the Maqsad (Sumail massif) paleo-mantle diapir - of the MORB segment in this part of the ophiolite. The crust-mantle transition zone (CMTZ) is relatively well exposed in this area, with a clear transition from harzburgite to the north to dunites then gabbros to the south (Fig. 1b). The two sites CM1 and CM2, separated by about 400 m, have been drilled twice: with one Hole for core recovery and a second wider Hole for geophysical logging. At these sites 400 m of core was recovered from CM1A and about 300 m of core was recovered from CM2B.

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The geological map produced by the OmanDP group during the preparation of Phases 1 and 2 173 shows a general tilt of the units by about 30° to the south (Fig. 1c). Considering these petrological 174 and structural configurations, Hole CM1A borehole was cored 400 m with an inclination of 60° 175 trending to the north, in order to cut perpendicularly across the mantle-crust transition (Fig. 1d). It 176 crosses from the gabbroic lower crust (~150 m; the Layered Gabbro "Crustal Sequence", CS), 177 through the dunite-rich crust-mantle transition zone (~150 m CMTZ) that includes the Dunite (DS) 178 and Dunite with Gabbro Sequences (DGS), to the residual upper mantle harzburgites (~100 m, 179 Mantle Sequence, MS). At Site CM2, the fully cored borehole CM2B is vertical, parallel to the 180 wider rotary borehole for geophysical logging (Hole CM2A). Hole CM2B starts within the crust-181 182 mantle transition zone (~110 m) and extends deeper in the underlying residual mantle peridotites than Hole CM1A (~180 m) (Fig. 1d). The main rock types sampled in Holes CM1A and CM2B 183 are olivine-gabbro, gabbro, dunite, harzburgite and wehrlite, associated with minor gabbronorite, 184 185 troctolite, websterite, anorthosite, and chromitite layers (Fig. 1d). The crust-mantle transition zone sampled in Hole CM1A has been divided into two parts according to the rock types present: the 186 upper half is mainly made of dunites containing rare melt migration features (DS for Dunite 187 Sequence); in the lower half, the dunites alternate with thin bands containing a higher proportion 188 of interstitial plagioclase (+/- clinopyroxene), which has been called the Dunite with Gabbro 189 Sequence (DGS). In the present article we focus on the geochemical compositions of dunites from 190 the crust sequence (CS), crust-mantle transition zone (CMTZ) and mantle sequence (MS) and of 191 mantle harzburgites. 192

# 193 **3 Results**

The sample selection strategy and the analytical methods are detailed in Supporting Information 194 (see also Kelemen et al., 2020a, 2020b, 2020c). In summary, one sample was taken every 10 m 195 along Holes CM1A and CM2B to cover the entire crust-mantle transition and mantle sections. 196 Additional samples were collected to better characterize some specific levels (e.g. to document 197 local, minor lithologies). The samples were analyzed for their major (as well as volatile) and trace 198 element compositions. Sample lithology, macroscopic and microscopic observations and mineral 199 modes calculated from major elements are reported in supplementary data table 1. Whole rock 200 major and volatile element compositions are reported in supplementary data table 2. Whole rock 201 trace element compositions are reported in supplementary data table 3. 202

# **3.1. Sample description**

The studied samples represent the ultramafic lithologies (harzburgites and dunites) of the mantle section (46 harzburgites and 12 dunites), the crust-mantle transition zone (45 dunites) and the crustal Layered Gabbro Sequence (2 dunites) drilled at Holes CM1A (46 samples) and CM2B (59 samples). Lithological classifications were made based on macroscopic and microscopic observations, and mineral modes calculated from XRF measurements (supplementary data table 1). Four rock groups have been defined, harzburgites (36), carbonate-bearing harzburgites (8), pure
dunites (31) and impregnated dunites (29) (Fig. 1e, Fig. SD-1, supplementary data table 1).

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212 The harzburgites and carbonate-bearing harzburgites contain olivine (ol), orthopyroxene (opx) and minor spinel (sp) as primary minerals (Fig. 2 e, f, g, h, i, j, k and l, Fig. SD-1, supplementary data 213 table 1), displaying a porphyroclastic texture. Olivine abundance (primary mode) ranges from 72% 214 to 93% in the harzburgites and from 64% to 75% in carbonate-bearing harzburgites. The olivine 215 typically show subhedral shape and equant habits. Orthopyroxene abundance ranges from 10 to 28 216 % with mostly subhedral shapes; some orthopyroxene show sigmoidal crystal-plastic deformation 217 features (Fig. 2i and j). Grains of spinel are present in all samples (up to 1-3%). Compared to the 218 dunites, the harzburgites are less altered, the alteration becoming complete only in highly veined 219 zones and at the bottom of both Holes where carbonate (carb)/serpentine (serp) associations occur 220 (Fig. 2k and 1). The most abundant minerals in the harzburgite background alteration are serpentine 221 and magnetite (mag). The carbonate alteration occurs at the bottom of both Holes starting at 222 388.3m depth in CM1A and at 279.5m depth in CM2B. Harzburgite sample CM2B 129Z1 5-10 223 cm (depth 299.7 m) is the most carbonate vein-rich harzburgite. Secondary amphibole (amph), 224 225 chlorite (chl) and hydrogrossular are also present as trace alteration minerals (abundance < 3%, supplementary data table 1). No patch or deformation features specifically related to the alteration 226 were observed. 227

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The pure dunites are primarily composed of ol >97% and generally contain <1 % sp (Fig. 2a, b, c 229 and d, supplementary data table1). Mineral modes calculated using bulk rock data (Fig. SD-1, 230 supplementary data table1) indicate that many dunites have >10% normative pyroxene. The 231 macroscopic and microscopic descriptions indicate that some dunites contain plagioclase (pl) 232 and/or pyroxenes. We refer to this group of dunites as 'impregnated dunites' in contrast to the pure 233 dunites containing only ol and sp. The ol in almost all samples has been completely replaced by 234 serp (Fig. 2a, b, c and d). In many dunites and in impregnated dunites, no relics of porphyroclastic 235 opx are present, indicating no textural relics of a porphyroclastic texture. The pure and impregnated 236 dunites' primary texture was a fine- to medium-grained granular microstructure, characterized by 237 euhedral ol forming a mosaic of equigranular grain size distribution as preserved by the equant 238 contacts between serp mesh cores (Fig. 2a, b, c and d). Alteration minerals mainly consist of serp 239 and mag (Fig. 2a, b, c and d). In addition, brucite (brc) after ol was detected by X-Ray Diffraction 240 (XRD, performed during the ChikyuOman 2018 Leg 3) in some serpentinized dunites from the 241 crust-mantle transition zone and in dunites from the mantle sequence (absent in harzburgites). 242 Where relics of ol are present, in rare cases, they are surrounded by serp and mag. If ol is 243 completely serpentinized, the mesh cores are mainly composed of serp with minor mag and 244 accessory grains of sulfides. Serpentine and more abundant magnetite at mesh rims trace former 245 ol grain and sub-grain boundaries (Fig. 2a, b, c and d). 246

# **3.2. Loss on ignition, CO<sub>2</sub> and H<sub>2</sub>O contents**

Samples from Holes CM1A and CM2B display high loss on ignition (LOI) values. The LOI varies from 8.29 to 14.92 wt.% in the harzburgites, from 9.02 to 23 wt.% in the carbonate-bearing

harzburgites, from 10.14 to 15.31 in the pure dunites and from 6.14 to 15.53 wt.% in the

impregnated dunites (Fig. 3 and supplementary table 2). The averaged H<sub>2</sub>O concentration is 12.0

 $\pm 1.7$  wt.% in harzburgites,  $11.0 \pm 5.5$  wt.% in carbonate-bearing harzburgites,  $14.3 \pm 1.2$  wt.% in

dunites and  $13.7 \pm 1.6$  wt.% in impregnated dunites. The LOI values correlate with measured water 253 254 concentration (H<sub>2</sub>O), which slightly decreases downhole (Fig. 3). The concentrations of CO<sub>2</sub> measured in harzburgites, carbonate-bearing harzburgites excluding the carbonate vein-rich 255 harzburgite CM2B 129Z1 5-10 cm (depth 299.7 m, CO<sub>2</sub> = 19.54 wt.%), pure dunites and 256 impregnated dunites vary from 0.12 to 0.31 wt.%, from 0.36 to 1.30 wt.%, from 0.13 to 0.34 wt.% 257 and from 0.03 to 0.41 wt.% respectively. The averaged CaCO3 concentration excluding the 258 carbonate-rich harzburgite mentioned previously (22.77 wt.%), is 0.40 wt.% in harzburgites, 1.57 259 wt.% in carbonate-bearing harzburgites, 0.51 wt.% in pure dunites and 0.46 wt.% in impregnated 260 dunites. The downhole profile of CaCO<sub>3</sub> shows an increase in concentration in the deepest part of 261 Holes CM1A and CM2B, with recovered harzburgites having higher CO<sub>2</sub> concentrations, LOI and 262 H<sub>2</sub>O contents, consistent with the particularly high alteration degree and high carbonate content in 263 these samples (serp, brc, carb, see Fig. 3 XRD). 264

# **3.3. Whole rock major element compositions**

Major element analyses were performed on 36 harzburgites, 8 carbonate-bearing harzburgites, 31 pure dunites, and 29 impregnated dunites sampled along Holes CM1A and CM2B.

268 The mantle section sequence (MS)

Harzburgite is the most abundant lithology in the MS (36 harzburgites analysed) followed by 269 impregnated dunite (9 impregnated dunites analysed) and pure dunite (4 pure dunites analysed). 270 Mg# (Mg# =  $100 \times \text{molar Mg/(Mg+Fe}_{total})$ ) in mantle harzburgites is between 90.7 and 92.4. CaO, 271 Al<sub>2</sub>O<sub>3</sub>, and TiO<sub>2</sub> contents range from 0.19 to 2.07 wt.%, 0.53 to 0.91 wt.%, and 0.02 to 0.04 wt.%, 272 respectively (Fig. 4). Harzburgite with dunite patches contains lower CaO concentrations than the 273 average value (0.30 to 0.81 wt.%). Carbonate-bearing harzburgites are characterized by similar 274 Mg#, Al<sub>2</sub>O<sub>3</sub>, and TiO<sub>2</sub> contents (90.6, 0.70 wt.% and 0.03 wt.% on average respectively), and very 275 276 high CaO concentrations ranging from 0.68 to 5.42 wt.% compared to carbonate-free harzburgites. Dunites from the MS have high Mg#, on average 90.2 for pure dunites and 91.1 for impregnated 277 dunites (Fig. 4). The pure dunites have CaO, Al<sub>2</sub>O<sub>3</sub> and TiO<sub>2</sub> contents from 0.30 to 0.94 wt.%, 278 0.32 to 0.72 wt.%, and 0.02 to 0.05 wt.% respectively, where the impregnated dunites have CaO 279 280 contents ranging from 0.14 to 1.59 wt.%, Al<sub>2</sub>O<sub>3</sub> from 0.21 to 0.95 wt.% and TiO<sub>2</sub> from 0.02 to 0.04 wt.% (Fig. 4). Along the MS, TiO<sub>2</sub> and Al<sub>2</sub>O<sub>3</sub> contents do not show any systematic variation 281 with depth. On the other hand, the CaO content shows some variations downhole. The vertical 282 evolution of the CaO content in harzburgites from Hole CM1A is different from other elements; 283 284 successive trends of increasing and decreasing CaO with depth form a well-defined zigzag pattern In detail, it increases from 0.97 to 2.1 wt. % between 311 and 340 m and from 0.54 to 1.8 wt.% 285 between 360 and 388 m, and decreases from 2.1 to 0.54 wt. % between 340 and 360 m then from 286 1.8 to 0.8 wt.% at most from 388 to around 400 m. CaO contents in dunites and nearby harzburgites 287 are correlated. Downhole intervals with the highest CaO contents are characterized by high 288 carbonate vein concentrations (Fig. 3 XRD, and Fig. 4). 289

290 The crust-mantle transition zone sequence (CMTZ)

The CMTZ is composed mainly of pure and impregnated dunites (Fig. 1d). 27 pure dunites and 18

impregnated dunites from Holes CM1A and CM2B CMTZ were analyzed. Both pure and impregnated dunites from the CMTZ display slightly lower Mg# (89.9 on average, Fig. 4) and

CaO concentrations (pure dunites = 0.16 wt.%, and impregnated dunites = 0.28 wt.% on average, 294 295 Fig. 4) compared to the dunites from the MS. Some dunites from CM1A DGS (between 252.96 and 271.42 m) display lower Mg# (Mg# = 86.8-88.0) that seems mostly controlled by the increase 296 297 in the FeO content. The TiO<sub>2</sub> contents in pure and impregnated dunites from the CMTZ (0.04 wt.% on average) are slightly higher than the TiO<sub>2</sub> contents in the MS dunites (0.03 wt.% on average). 298 Al<sub>2</sub>O<sub>3</sub> concentrations are lower in CMTZ pure dunites (0.42 wt.% on average) and higher in 299 impregnated dunites (1.12 wt.% on average) compared to pure (0.53 wt.% on average) and 300 impregnated (0.51 wt.% on average) dunites respectively from the MS. A similar zigzag pattern to 301 the one observed in CaO along the MS is irregularly observed in CMTZ dunites (i.e. in CM1A, 302 decreasing from 0.17 to 0.12 wt. % between 170 and 245 m and, increasing from 0.08 to 0.52 wt. 303 % between 253 and 310 m, Fig. 4). This CaO zigzag variation in CM1A CMTZ dunites is 304 associated with Mg# zigzag variation (decreasing from 91 to 89 wt. % between 170 and 245 m 305 and, increasing from 87 to 91 wt. % between 253 and 310 m, Fig. 4). The dunites' CaO varies 306 over ~ 20 m at the base of the CMTZ following the mantle harzburgites zigzag variation, whereas 307 the Mg# varies over  $\sim 60$  m, along with the FeO variations. 308

# 309 The crustal sequence (CS)

The two analyzed dunites from the Layered Gabbro Sequence (LGS) (C5707A-51Z-1 W, 31.0-39.0 cm, 125.60 m in depth, and C5707A-58Z-2 W, 1.0-6.0 cm, 143.93 m in depth) are impregnated (Fig. 4), they have relatively low Mg# (85.6 and 85.5 respectively), compared to dunites in the other sequences. This tendency is similar to the dunites from CM1A DGS. CaO (0.04 wt.% and 0.19 wt.% respectively), Al<sub>2</sub>O<sub>3</sub> (0.15 wt.% and 0.93 wt.% respectively) and TiO<sub>2</sub> (0.03 wt.%) contents are similar to the impregnated dunites values from the MS and the CMTZ (Fig. 3).

CM Holes harzburgites, carbonate-bearing harzburgites, pure dunites and impregnated dunites 317 (except CM1A LGS impregnated dunites) show similar major element compositions to previously 318 reported harzburgites, dunites and impregnated dunites from the Oman ophiolite mantle and crust-319 mantle transition zones (Fig. 5). Most of the pure and impregnated dunites from MS, and CMTZ 320 plot above the terrestrial array (Earth differentiation trend, Jagoutz et al., 1979) in the MgO/SiO<sub>2</sub> 321 vs. Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub> diagram, whereas the harzburgites, carbonate-bearing harzburgites and some pure 322 and impregnated dunites plot below. CM harzburgites have similar MgO and FeO to harzburgites 323 324 from other massifs of the Oman ophiolite (with a slightly higher Mg# in some CM samples). The MgO-FeO variations in most of the CM dunites mimic the stoichiometric variation of the ol Mg-325 Fe composition, similar to other pure/slightly impregnated dunites elsewhere, while only a few 326 CM samples fall in the domain of the more highly impregnated dunites. Lower CaO and wider 327 range of Al<sub>2</sub>O<sub>3</sub> values characterize all the dunites compared to the harzburgites (Fig. 5). 328

# **329 3.4.Whole rock trace element contents**

Chondrite-normalized Rare Earth Element (REE) and primitive mantle-normalized trace element variations and patterns are shown in Figures 6 and 7 respectively. Similar to other refractory peridotites from the Oman ophiolite mantle section and crust-mantle transition, the studied harzburgites, carbonate-bearing harzburgites, pure dunite and impregnated dunite whole rock concentrations are lower than chondritic (CN) and primitive mantle (PM) values (Figs. 6 and 7). Their REEs are characterized by three types of chondrite-normalized patterns: linear or flat linear,

U- or V-shaped, and concave-upward patterns. CM Holes show vertical trends, continuous over 336

337 tens of meters, in their REE content. These are characterized by ~ 50 m-thick alternations between

increasing and decreasing concentrations ('zigzag' patterns) (e.g. La, Yb, LREE patterns; Fig. 6). 338

339 These trends along CM Holes are observed in both dunites and harzburgites, independent of the

lithology, and the changes from one trend to another are commonly correlated with the presence 340 of faults described by the structural team during the ChikyuOman2018 Leg 3 (Kelemen et al.,

- 341
  - 2020a, 2020b). 342

#### 343 *The mantle section sequence*

344 The mantle harzburgites show two types of chondrite-normalized REE patterns: (1) 19 harzburgites display linear REE patterns characterized by a progressive depletion from heavy REE 345 346 (HREE) (Gd<sub>CN</sub>/Yb<sub>CN</sub>=  $0.25 \pm 0.21$ , Yb<sub>CN</sub> =  $0.22 \pm 0.11$ ) to middle REE (MREE: Sm, Eu and Gd)  $(Sm_{CN} = 0.06 \pm 0.05 \text{ and } Gd_{CN} = 0.06 \pm 0.05)$  and light (LREE)  $(La_{CN}/Sm_{CN} = 0.98 \pm 0.74, La_{CN})$ 347 =  $0.06 \pm 0.05$ ). 6 samples show a positive Eu anomaly ((Eu/Eu\*)<sub>CN</sub> =  $2.36 \pm 0.66$ , with (Eu/Eu\*)<sub>CN</sub> 348 = Eucn/ $\sqrt{(Sm_{CN} \times Gd_{CN})}$ ). In detail, 12 harzburgites display linear LREE-depleted patterns (3 from 349 350 CM1A and 9 from CM2B) characterized by a progressive depletion from HREE to LREE, and 7 harzburgites (5 from CM1A and 2 from CM2B) display flat linear REE patterns characterized by 351 slightly lower LREE concentrations compared to HREE concentrations. (2) 15 harzburgites (4 352 from CM1A and 11 from CM2B) display U- or V-shaped REE patterns reflecting significant 353 MREE depletion relative to LREE (Lacn/Smcn =  $3.21 \pm 1.18$ ) and HREE (Gdcn/Ybcn =  $0.07 \pm$ 354

0.06). 7 samples show a positive Eu anomaly ((Eu/Eu\*)<sub>CN</sub> =  $3.36 \pm 1.69$ ). 355

The mantle pure dunites and impregnated dunites show two types of chondrite-normalized REE 356 patterns: linear REE patterns (3 pure dunites from CM2B and 6 impregnated dunites (3 from 357 CM1A and 3 from CM2B)) and U- or V-shaped REE patterns (5 pure dunites and 1 impregnated 358 dunites from CM2B). In detail the linear REE pattern are subdivided to: (1) flat linear REE patterns 359 displayed by the three CM2B impregnated dunites, they are characterized by roughly similar 360 LREE (La<sub>CN</sub> =  $0.14 \pm 0.07$ ), MREE (Sm<sub>CN</sub> =  $0.10 \pm 0.05$ ), and HREE (Yb<sub>CN</sub> =  $0.18 \pm 0.05$ ) 361 concentrations, together with a positive Eu anomaly ((Eu/Eu\*)<sub>CN</sub> =  $2.06 \pm 0.65$ ). (2) LREE-362 depleted linear REE patterns displayed by 3 pure dunites and 3 impregnated dunites from CM2B 363 show a progressive depletion from HREE (Gd<sub>CN</sub>/Yb<sub>CN</sub>=  $0.25 \pm 0.16$  and  $0.18 \pm 0.04$ , Yb<sub>CN</sub> = 0.22364  $\pm 0.05$  and  $0.28 \pm 0.07$  respectively) to MREE (Sm<sub>CN</sub> =  $0.04 \pm 0.03$  and  $0.04 \pm 0.02$ , Gd<sub>CN</sub> = 0.06365  $\pm$  0.04 and 0.05  $\pm$  0.02 respectively) to LREE (La<sub>CN</sub>/Sm<sub>CN</sub> = 0.71  $\pm$  0.28 and 0.37  $\pm$  0.23, La<sub>CN</sub> = 366  $0.03 \pm 0.02$  and  $0.02 \pm 0.01$  respectively), 1 sample shows a positive Eu anomaly (Eu/Eu\*)<sub>CN</sub> = 367 1.60). The U- or V-shaped REE patterns (5 pure dunites and 1 impregnated dunites from CM2B) 368 are characterized by significant MREE depletion relative to LREE (Lacn/Smcn =  $2.66 \pm 1.76$ ) and 369 HREE (Gd<sub>CN</sub>/Yb<sub>CN</sub> =  $0.08 \pm 0.12$ ), 2 pure dunites display a positive Eu anomaly ((Eu/Eu\*)<sub>CN</sub> = 370 2.91 and 2.54). 371

#### *The crustal-mantle transition zone sequence (CMTZ)* 372

The pure dunites from the CMTZ show two types of chondrite-normalized REE patterns: (1) U-373

374 shaped REE pattern displayed by 11 pure dunites (6 from CM1A and 5 from CM2B), with MREE

depletion relative to LREE (Lac<sub>N</sub>/Sm<sub>CN</sub> < 3.34) and HREE (0.07 < Gd<sub>CN</sub>/Yb<sub>CN</sub> < 0.27). Some 375

samples also have positive Eu anomalies (4 samples from CM1A and 3 samples from CM2B, 376

 $(Eu/Eu^*)_{CN} = 3.62 \pm 2.49$ ). (2) Linear LREE-depleted or slightly concave-upward REE patterns 377

# displayed by 10 pure dunites from Hole CM1A (La<sub>CN</sub>/Yb<sub>CN</sub> = $0.07 \pm 0.05$ ) with similar HREE

concentrations (Yb<sub>CN</sub> =  $0.13 \pm 0.02$ ) and lower LREE and MREE variation compared to HREE (La<sub>CN</sub> =  $0.01 \pm 0.01$ ; Gd<sub>CN</sub> =  $0.02 \pm 0.01$ ). Most samples display a positive Eu anomaly (6 samples,

381  $(Eu/Eu^*)_{CN} = 4.43 \pm 2.44$ ).

The CMTZ impregnated dunites display three types of chondrite-normalized REE patterns: (1) 382 The linear REE patterns displayed by 5 impregnated dunites (2 from CM1A and 3 from CM2B) 383 are characterized by a progressive depletion from HREE (Gd<sub>CN</sub>/Yb<sub>CN</sub>=  $0.38 \pm 0.20$ , Yb<sub>CN</sub> = 0.23384  $\pm$  0.09) to MREE (Sm<sub>CN</sub> = 0.07  $\pm$  0.06 and Gd<sub>CN</sub> = 0.10  $\pm$  0.09) to LREE (La<sub>CN</sub>/Sm<sub>CN</sub> = 0.77  $\pm$ 385 0.19, Lacn =  $0.05 \pm 0.06$ ), 2 samples show a positive and negative Eu anomaly ((Eu/Eu\*)<sub>CN</sub> = 2.23) 386 and 0.45 respectively). (2) 3 samples from CM2B display U-shaped REE pattern, with MREE 387 depletion relative to LREE (La<sub>CN</sub>/Sm<sub>CN</sub> < 3.93) and HREE ( $0.09 < Gd_{CN}/Yb_{CN} < 0.25$ ). (3) One 388 sample from CM1A and three samples from CM2B display concave-upward patterns characterized 389 by a nearly flat slope of the HREE segment (Gd<sub>CN</sub>/Yb<sub>CN</sub> =  $0.86 \pm 0.39$ ) followed by a progressive 390 depletion from MREE to LREE (Lac<sub>N</sub>/Sm<sub>CN</sub> =  $0.13 \pm 0.07$ ). One CM1A impregnated dunite shows 391 a negative Eu anomaly ( $(Eu/Eu^*)_{CN} = 0.42$ ). 392

393 The crustal sequence (CS)

The two impregnated dunites from the Hole CM1A CS (C5707A-51Z-1 W, 31-39.0 cm, 125.60 m

depth, and C5707A-58Z-2 W, 1.0-6.0 cm, 143.94 m depth) display relatively linear (REE) patterns characterized by a steady decrease of REE abundances HREE to LREE, as well as by a clear

397 positive Eu anomaly ( $(Eu/Eu^*)_{CN} = 3.8-5.0$ ).

The PM-normalized multi-element patterns of most harzburgites and carbonate-bearing 398 harzburgites exhibit strong to moderate enrichments in LILE, Th, U, Nb and Ta relative to LREE 399 (e.g. averaged Rbpmn/Lapmn = 12.95 and 18.24; Upmn/Lapmn = 4.64 and 4.45; Nbpmn/Lapmn = 2.02 400 401 and 4.30 respectively). The carbonate-bearing harzburgites display stronger Pb, Sr and Ti positive anomalies (averaged Pbpmn/Cepmn = 145.23; Srpmn/Ndpmn = 373.50; Tipmn/Gdpmn = 11.09) 402 403 compared to the harzburgite (averaged PbpMN/CepMN = 25.83; SrpMN/NdpMN = 30.21, TipMN/GdpMN = 3.52). Most pure dunites show similar enrichments exhibited by the harzburgites and carbonate-404 405 bearing harzburgites in LILE, Th, U, Nb and Ta relative to LREE (e.g. averaged Rb<sub>PMN</sub>/La<sub>PMN</sub> = 20.0; UPMN/LaPMN = 5.12; NbPMN/LaPMN = 4.10), but with smaller Pb and Sr positive anomalies 406 (averaged Pbpmn/Cepmn = 19.68; Srpmn/Ndpmn = 29.05). The pure dunites display stronger Ti 407 positive anomalies compared to the harzburgites and smaller compared to the carbonate-bearing 408 409 harzburgites (Ti<sub>PMN</sub>/Gd<sub>PMN</sub> = 8.98). The impregnated dunites exhibit moderate LILE, Th, U, Nb and Ta enrichments relative to LREE (e.g. averaged Rb<sub>PMN</sub>/La<sub>PMN</sub> = 3.94; U<sub>PMN</sub>/La<sub>PMN</sub> = 3.12; 410 Nb<sub>PMN</sub>/La<sub>PMN</sub> = 1.05), and the smallest Pb, Sr and Ti positive anomalies compared to all groups of 411 rocks (averaged Pbpmn/Cepmn = 9.99; Srpmn/Ndpmn = 14.03; Tipmn/Gdpmn = 1.62). 412

# 413 **4 Discussion**

# 414 4.1 Effects of serpentinization and carbonation on the composition of dunites and 415 416 417

In some cases, alteration, especially serpentinization of ultramafic rocks (sometimes associated with mineralization related to hydrothermal activity), may significantly modify bulk-rock

chemical composition (e.g. Beinlich et al., 2020; de Obeso & Kelemen, 2018; Gruau et al., 1998; 418 419 Hodel et al., 2018; Malvoisin, 2015; Paulick et al., 2006; Snow & Dick, 1995). As described above, the dunites and harzburgites from the crust-mantle transition and mantle section of Holes CM1A 420 and CM2B are extensively serpentinized (Fig. 2), up to 100% in many samples. In addition, strong 421 carbonate-veining affected parts of the cores, especially at the base of Hole CM2B where it is 422 intensely faulted (carbonate-bearing harzburgites, Fig. 3, CO2 and CaCO3 logs). To decipher the 423 effects of such significant fluid/rock interactions on the composition of OmanDP CM samples is 424 critical (especially on the trace elements). It particularly concerns the dunites, because of their 425 more altered character (perhaps related to higher ol mode), and because their trace element budget 426 was more significantly controlled by primary ol and trace phases such as pyroxenes before 427 alteration (unlike the harzburgites that contain abundant residual pyroxenes). 428

The LOI content is generally higher in carbonate-bearing harzburgite (averaged LOI =  $13.04 \pm$ 429 9.57 wt.%), pure dunites (averaged LOI =  $13.92 \pm 1.21$  wt.%) and impregnated dunites (averaged 430 LOI =  $13.33 \pm 2.02$  wt.%) than in harzburgites (averaged LOI =  $11.91 \pm 1.53$  wt.%). 431 Serpentinization of many of the studied samples appears to have led to enrichment in  $SiO_2$  as 432 already observed in abyssal peridotites affected by a Si addition, and/or magnesium loss (e.g. de 433 Obeso & Kelemen, 2018; Paulick et al., 2006; Snow & Dick, 1995). This open system behavior is 434 confirmed by the plot of some dunites (3 from CM2B) and impregnated dunites (4 from CM1A 435 and 4 from CM2B) below the mantle fractionation array at the same field as the harzburgites and 436 the carbonate-bearing harzburgites (Fig. 5a), suggesting MgO loss and/or SiO<sub>2</sub> enrichment as 437 438 reported in pervasively serpentinized abyssal peridotites or talc-bearing serpentinites (de Obeso & Kelemen, 2018; Snow & Dick, 1995; Paulick et al., 2006). This may be the reason why there are 439 elevated normative pyroxene modes in samples that were classified as dunites based on 440 macroscopic (hand specimen) and microscopic (thin sections) observations (Fig. SD-1, 441 supplementary data table 1). Some of these dunites have no pyroxenes or pyroxene pseudomorphs 442 in thin section. This is supported by XRD analyses performed during ChikyuOman 2018 Phase 2 443 444 Leg 3 which revealed the widespread occurrence of brucite associated with other alteration minerals (Kelemen et al., 2020a, 2020b). It also appears clear that the higher CaCO<sub>3</sub> (averaged 445  $CaCO_3 = 3.93 \pm 10.85$  wt.% compared to  $0.40 \pm 0.54$  wt.% in harzburgites) together with higher 446 CO<sub>2</sub> contents (averaged CO<sub>2</sub> =  $2.82 \pm 9.501$  wt.% compared to  $0.20 \pm 0.25$  wt.% in harzburgites) 447 448 in the carbonate-bearing harzburgites at the base of CM1A and CM2B (4 from CM1A and 4 from CM2B) are related to carbonate-veins. However, the covariation of Ni and Co contents with the 449 450 XMg suggests that the possible precipitation of sulfides related to these strong water/rock interactions did not erase the primary compositions. 451

The plots of the concentration of several trace elements as a function of the LOI (Fig. 8a-c) and of 452 the CO<sub>2</sub> and CaCO<sub>3</sub> contents show no clear correlation. On the contrary, some reasonably good 453 covariations are observed between Th and U, Nb and especially La on one hand (Fig. 8 h-j), and 454 Yb, Ti and Hf on the other hand (Fig. 8 n-o); the Zr content is partially correlated with both Th 455 and Yb (Fig. 8 f, k). Since Th and Ti are generally considered immobile during alteration processes 456 (e.g. Kogiso et al., 1997; Niu, 2004; Paulick et al., 2006), these trends probably reflect one or more 457 overprinted geochemical signatures acquired during high temperature, magmatic processes rather 458 than during a later serpentinization event. The large ion lithophile elements as well as Li and Pb 459 do not correlate with the LOI nor with other elements (Fig. 8 b, c, g, l, m), and their compositions 460 may result from the overprint of several processes having operated over a large range of 461 temperatures and conditions, from igneous to alteration events. Accordingly, only the 462

463 concentrations in REE, HFSE and Th-U, will be used to discuss the igneous processes that led to
464 the formation of the dunites from the crust-mantle transition zone and mantle sequence sampled
465 by the CM Holes.

# 466 **4.2. Partial melting** *vs.* **melt-rock reaction in the Oman ophiolite mantle section**

The mantle section of CM Holes (Wadi Tayin massif) is composed of refractory harzburgites with 467 relatively homogeneous modal and major element compositions (excluding some major elements 468 e.g. CaO and Na<sub>2</sub>O, Figs. 4 and 5) and more variable trace element contents (Figs. 6, 7a and 7b). 469 Similar to other previously studied Wadi Tayin massif and Magsad diapir harzburgites (Godard et 470 al., 2000; Hanghøj et al., 2010; Monnier et al., 2006) and most refractory abyssal peridotites 471 (Godard et al., 2008; Niu, 1997; Warren et al., 2009), CM harzburgites plot near the most depleted 472 473 end of the mantle fractionation array (Fig. 5a). They display similar low  $Al_2O_3/SiO_2$  ratios in comparison to other harzburgites from Wadi Tayin and Magsad harzburgites (0.01-0.02, 0.01-0.04 474 and 0.01-0.08 respectively), and a high MgO/SiO<sub>2</sub> ratios typical of refractory peridotites (0.98-475 1.10, 0.95-1.10 and 0.96-1.15 respectively). Oman harzburgites are characterized by narrow FeO 476 477 and MgO contents compared to the pure and impregnated dunites (Fig. 5b), with slightly higher Mg# in CM harzburgites (90.7-92.4) compared to Wadi Tayin and Maqsad harzburgites (89.6-478 479 91.5 and 88.4-91.1 respectively). Generally, Al<sub>2</sub>O<sub>3</sub> and CaO show broad positive correlation in harzburgites from Wadi Tayin massif and Maqsad diapir (Fig. 5c), Godard et al, (2000) 480 demonstrating that the observed Al<sub>2</sub>O<sub>3</sub>/CaO ratio variability displayed by Wadi Tayin and Maqsad 481 harzburgites decreases with increasing cpx content in the main harzburgites sequence to lower 482 483 values in the cpx-harzburgites at the base of the mantle section. Al<sub>2</sub>O<sub>3</sub>/CaO ratio variability is also observed along cores CM1A and CM2B, however CaO variability at the bottom of CM Holes is 484 related to CO<sub>2</sub>-bearing fluids interactions with CM harzburgites (see section 4.1.). 485

Most pure and impregnated dunites from CM mantle and crust-mantle transition zone plot above 486 487 the mantle fractionation array, similar to previously studied dunites and impregnated dunites from the mantle section of Wadi Tayin massif (Godard et al., 2000; Hanghøj et al., 2010) and the CMTZ 488 at the top of Maqsad diapir (Rospabé et al., 2018a, 2019a) (Fig. 5a). CM pure and impregnated 489 dunites show similar FeO, MgO, Al<sub>2</sub>O<sub>3</sub> and CaO contents to other pure and impregnated dunites 490 from Wadi Tayin and Magsad harzburgites (Fig. 5b, 5c). The CM mantle harzburgites, pure and 491 impregnated dunites show variable trace element compositions and contrasting shapes in their REE 492 493 and extended trace element patterns (Fig. 7a-b) (their major element compositions are much more homogeneous, Figs. 4 and 5). Significant geochemical variability of the mantle peridotites has also 494 been observed all along the Oman ophiolite. In some cases, previous geochemical studies inferred 495 the overprint of partial melting and melt/peridotite reaction processes in the mantle harzburgite 496 signatures (e.g. Gerbert-Gaillard, 2002; Girardeau et al., 2002; Godard et al., 2000; Hanghøj et al., 497 2010; Kanke & Takazawa, 2014; Khedr et al., 2014; Le Mée et al., 2004; Monnier et al., 2006; 498 Takazawa et al., 2003). 499

Two different geochemical trends are combined in the trace element contents of all samples studied here. On one hand, Yb shows a good correlation with Ti and some HFSE (Fig.8o). Heavy REE have been demonstrated to be less impacted than MREE and especially LREE during meltperidotite reactions such as melt/rock re-equilibration during melt migration or in response to conversion of harzburgite to dunite (e.g. Godard et al., 1995; Kelemen et al., 1990; Navon & Stolper, 1987; Prinzhofer & Allègre, 1985; Spiegelman & Kelemen, 2003; Vernières et al., 1997;

Rospabé et al., 2018a). Therefore, these correlations may have been formed during partial melting, 506 507 with little or no overprint by subsequent melt-peridotite interactions at shallow depth. On the other hand, we observe a good correlation between Th, U and La, and partially with the HFSE (e.g. Nb, 508 Fig. 8h-j), that does not correlate with Yb or Ti concentrations. As the U-/V-shape of the REE 509 patterns, resulting from the selective enrichment in LREE relative to MREE and HREE in 510 peridotites, may be attributed to chromatographic fractionation associated with interstitial melt 511 percolation and/or to the transformation of harzburgite into dunite (e.g. Godard et al., 1995; Navon 512 & Stolper, 1987; Prinzhofer & Allègre, 1985; Vernières et al., 1997; Rospabé et al., 2018a), we 513 interpret this second geochemical signature as the overprint of melt-peridotite reaction processes. 514

## 515 4.2.1. Geochemical logs

516 Vertical chemical trends are observed along CM Holes, especially in REE (e.g. La, Yb; Fig. 6), and in some major elements (e.g. FeO, MgO, CaO, Na<sub>2</sub>O; Fig. 4). These chemical trends are 517 continuous over tens of meters, and alternate between increasing and decreasing ('zigzag' patterns) 518 with a characteristic thickness of ~ 50 m. Abrupt changes in these trends, particularly in Yb, LREE 519 520 and HREE concentrations, are mostly associated with the presence of faults (Fig. 6). The trends observed in CM1A and CM2B dunites and harzburgites do not show any significant correlation as 521 a function of LOI, CO<sub>2</sub>, or CaCO<sub>3</sub> contents (Figs. 3 and 8). This, together with the continuity of 522 each individual trend, suggests that the trends were imprinted at high temperature and are not 523 524 related to post-magmatic, low temperature events. Furthermore, the trends are observed in both dunites and harzburgites in the mantle section, independent of dunite and harzburgite alternations. 525 526 Similar trend changes attributed to the focus of the percolation/migration along faults have been observed for the Magsad CMTZ dunites, with the same characteristic thickness of about 50 m 527 (Rospabé et al. 2019a, 2020). Our results seem to confirm the significant impact of deep-seated 528 syn-magmatic faults on the development of the crust-mantle transition at the expense of the 529 shallower mantle and the recorded whole rock chemical signatures, in addition to their impact on 530 the formation of the lower oceanic crust (Abily et al., 2011; see also Sauter et al., 2021). Such 531 structural characters must have developed early and are not just restricted to the Maqsad area (e.g. 532 Rospabé (2018) for other Oman areas and Sauter et al. (2021) for present-day oceans). 533

## 534 4.2.2. The origin of REE pattern shapes in Oman ophiolite harzburgites

CM mantle harzburgites display two types of chondrite-normalized REE patterns: (1) linear (3 535 harzburgites from CM1A, and 9 from CM2B) to flat linear REE patterns characterized by a 536 537 progressive depletion from HREE to MREE to LREE, some samples show similar LREE and HREE concentrations and slightly lower LREE and MREE variation compared to HREE (5 538 harzburgites from CM1A and 2 from CM2B); (2) U- or V-shaped REE patterns reflecting 539 significant MREE depletion relative to LREE and HREE (4 harzburgites from CM1A and 11 from 540 CM2B, and all the 8 carbonate-bearing harzburgites from CM1A and CM2B). Most of CM 541 harzburgites are enriched in LREE relative to MREE (7 harzburgites display flat linear REE 542 543 patterns and 15 display U- or V-shaped REE patterns) with 12 harzburgites (3 from CM1A and 9 from CM2B) that are depleted in LREE relative to MREE and HREE, this is not expected for 544 mantle residues after near-fractional partial melting (Godard et al., 1995, 2000, 2008; Gruau et al., 545 1998; Johnson & Dick, 1992; Johnson et al., 1990; Kelemen et al., 1997; Navon & Stolper, 1987; 546 Prinzhofer & Allègre, 1985; Vernières et al., 1997). The linear REE patterns observed in CM 547 harzburgites are similar to the main harzburgites mantle section REE patterns from Wadi Tayin 548

described by Godard et al. (2000); these harzburgites have relatively homogeneous modal and 549 major element compositions. The Maqsad diapir area harzburgites and Samail massif cpx-550 harzburgites display concave-upward and 'spoon-shaped' REE patterns respectively, these 551 patterns are not observed in CM Wadi Tayin harzburgites. The Magsad diapir has been interpreted 552 as frozen-upwelling mantle that fed a former spreading centre (Ceuleneer et al., 1988). The 553 concave-upward REE patterns and the higher Al<sub>2</sub>O<sub>3</sub>/CaO ratios and TiO<sub>2</sub> contents in the diapir 554 harzburgites result from the feedback between deformation and melt percolation, and the *cpx*-555 harzburgites REE patterns and chemical characteristics are interpreted as a result of a cpx forming 556 melt-rock reaction at decreasing melt mass (at near-solidus conditions) along the lithosphere-557 asthenosphere boundary (Godard et al., 2000). 558

Figure 9 shows the modal composition of CM1A, CM2B (harzburgites and carbonate-bearing 559 harzburgites) and Nakhl-Samail-Wadi Tayin massif samples (harzburgites, Godard et al., 2000) 560 plotted with two published melting models: model 1 represents the Niu (1997) polybaric melting 561 model (1a) with and (1b) without excess ol; model 2 represents the Walter et al. (1995) isobaric 562 melting at 11 (2a), 16 (2b) and 17 kbar (2c). The figure indicates that CM harzburgites could result 563 from high degrees of partial melting and melt extraction in the range of 15-30 % (e.g. Asimow et 564 al., 2001; Kelemen et al., 1990, 1992, 1995; Niu, 1997; Walter et al., 1995), as suggested by 565 Godard et al. (2000) for the main harzburgite section of the MORB-like, NW-SE paleo spreading 566 segment (Nakhl-Samail-Wadi Tayin massifs). This range is higher than melting degrees producing 567 MORB in present-day oceans (5-10 %; Langmuir et al., 1992) and high-Ti magmas such as those 568 569 forming the dyke complex and the MORB-like lava sequence in Oman (Godard et al., 2006; Lippard et al., 1986). However, Niu (1997) and Dick & Natland (1996) have reconciled this 570 inconsistency by considering that abyssal peridotites represent only the shallowest part of the 571 mantle column affected by partial melting and therefore record the highest melting degrees. In 572 contrast, MORB are thought to represent integrated, mixed melt fractions from polybaric 573 decompression melting over 60-100 km at ascent melting column and therefore record average 574 575 melting degrees.

The linear REE patterns of some studied harzburgites do not show features of strong LREE 576 577 depletion, but most of the other CM harzburgites show U-shaped REE patterns, characterized by LREE enrichment. This, together with the high (primary) proportion of ol (75-90 %) observed in 578 some samples, point to the fact that near fractional partial melting alone fails to explain the 579 harzburgite geochemical signatures along the CM cores. Vernières et al. (1997) noted that the 580 relatively unfractionated REE distribution may simply result from melt transport through the 581 melting peridotites, as an "open-system melting process". This process would result in a negative 582 correlation between LREE/HREE ratios and peridotite fertility, as commonly observed in other 583 ophiolites (e.g. Prinzhofer & Allègre, 1985). It results from the competition between the partial 584 melting continuously depleting the mantle residue on one hand and chromatographic effects 585 related to the melt extraction that enrich the residue in the most incompatible elements on the other 586 hand (e.g. Johnson et al., 1990; Navon & Stolper, 1987; Niu, 2004; Takazawa et al., 1992; 587 Spiegelman & Kelemen, 2003). 588

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590 CM harzburgites show good correlations among Yb, Ti and Hf concentrations, indicating that the

- variability in HREE was more likely controlled by the melting process rather than by the overprint
- of melt-rock reaction processes (Fig. 8 n-o). We compared the linear REE patterns in CM
- harzburgites to Vernières et al. (1997) plate model calculations performed by Godard et al. (2000),

who modeled trace elements in peridotites form the Nakhl-Samail-Wadi Tavin massifs to explore 594 595 whether the REE variations observed in the studied harzburgites resulted from reactive porous flow at increasing melt mass, or from partial melting coupled with melt transport (Fig. 10). The 596 597 authors first simulated a standard incremental melting model (experiment a, Fig. 10a, Johnson et al., 1990). Then, they simulated reactive porous flow at increasing melt mass in a second model 598 (experiment b, Fig. 10b). Experiment (a) produces strongly LREE-depleted peridotite residues, 599 quite different in shape from CM and other Oman harzburgites. This experiment does not provide 600 a better fit to the data when the presence of the trapped melt in the residue is included (Fig. 10c). 601 Experiment (b) produces peridotite residues moderately depleted in LREE and with a small amount 602 of trapped melt (0.5-1 %) (Fig. 10d). The linear REE patterns of harzburgites from Holes CM1A 603 and CM2B are similar to Nakhl-Samail-Wadi Tayin harzburgites of Godard et al. (2000) and well 604 reproduced by experiment (b). The model suggests that part of the studied harzburgites were 605 pervasively percolated by diffuse melt flow which affected their geochemical signature. However, 606 the presence of many dunite intervals at the top of the mantle section requires an orthopyroxene-607 consuming reaction between the residual peridotites and infiltrated melts. 608

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# 610 4.2.3. Cryptic and modal mantle refertilization in Oman ophiolite

The CMTZ pure dunites from CM holes (Wadi Tavin massif) display two types of chondrite-611 normalized REE patterns: (1) U-shaped REE patterns (6 samples from CM1A and 5 from CM2B); 612 (2) Linear LREE-depleted or slightly concave-upward REE patterns (10 samples from Hole 613 CM1A) with similar HREE, MREE and LREE concentrations, with slightly and lower LREE and 614 MREE concentrations compared to HREE in some samples. The CMTZ impregnated dunites 615 display three types of chondrite-normalized REE patterns: (1) linear REE patterns (2 samples from 616 CM1A and 3 from CM2B); (2) U-shaped REE patterns (3 samples from CM2B); (3) concave-617 upward patterns characterized by a nearly flat slope of the HREE segment followed by a 618 progressive depletion from MREE to LREE. The Magsad mantle-crust transition zone pure dunites 619 trace elements are characterized by U-shaped to concave-upward REE patterns (Godard et al., 620 2000; Rospabé et al., 2018a, 2019a), similar to Maqsad diapir harzburgites (Godard et al., 2000) 621 and to CM mantle dunites and harzburgites trace element patterns but with larger range of LREE 622 623 variations in Maqsad MTZ dunites compared to Wadi Tayin CMTZ (CM cores). Maqsad impregnated dunites described by Rospabé et al. (2018a, 2019a) are characterized by similar trace 624 element patterns to CM impregnated dunites, varying between linear LREE-depleted to variably 625 626 concave-upward trace elements patterns. Maqsad MTZ pure and impregnated dunites have been interpreted as end-members that recorded different stage of an initially shared same igneous 627 processes (Rospabé et al., 2018a, 2019a). Boudier & Nicolas (1995) and Godard et al (2000) attest 628 that Magsad MTZ dunites are diapir harzburgites that were strongly modified by ol-forming melt-629 rock reactions at high melt/rock ratios. Furthermore, Rospabé et al. (2018a) argue that the pure 630 dunites are residues left after extraction of a percolating melt, whereas, the impregnated dunites 631 correspond to a frozen stage before complete melt extraction. 632

Relatively good covariations are observed in CM dunites and harzburgites between Th and U, Nb
 and especially La (Fig. 8 h, j), whereas their concentrations are not correlated with the HREE.
 Following many previous works, we interpret these correlations as the result of melt/peridotite
 reaction contemporaneously with, and/or subsequent to, the partial melting event discussed above.

The crust-mantle transition zone pure and impregnated dunites from CM Holes display similar 637 REE patterns to the dunites and harzburgites in the mantle section, U-/V-shaped REE patterns 638 (displayed by 11 pure dunites and 3 impregnated dunites from the CM, CMTZ) cannot be 639 explained by a pure cumulate origin (Fig. 7a-b). Several studies argued that the LREE enrichment 640 relative to MREE cannot be explained by REE partition coefficients between ol and melt (Frey et 641 al., 1978; Hauri & Hart, 1995; Kelemen et al., 1993; Lee et al., 2007; McKenzie & O'Nions, 1991; 642 Sun & Liang, 2014) and may better be explained by peridotite metasomatism as a result of melt-643 peridotite reactions (e.g. Agranier & Lee, 2007; Godard et al., 1995; Navon & Stolper, 1987; 644 Vernières et al., 1997). Most CM harzburgites display U-shaped REE patterns (15 harzburgites 645 and 8 carbonate-bearing harzburgites) characterized by a LREE-enrichment compared to the linear 646 REE patterns (12 harzburgites), with strong LREE enrichments indicating extensive interaction 647 with a pervasive melt (Fig. 6a-b; e.g. Gerbert-Gaillard, 2002; Godard et al., 2000; Monnier et al., 648 2006). CM samples, in particular in Hole CM2B, show downhole variations that indicate a 649 decreasing degree of melting with increasing depth (see Section 4.3.3. below geochemical logs). 650 Two intervals from 170 m to 260 m, and 230 m to 300 m depth, are particularly good examples of 651 this (Fig. 6, CM2B e.g. LREE, La, U). The correlations observed between LREE enrichment and 652 an increasing fraction of trapped melt - calculated from experiment (b and d) in the most residual 653 peridotites from 'Plate model' of Vernières et al. (1997) applied by Godard et al. (2000) 654 (experiment b and d, see Fig.10) - in CM samples (e,g, La, Fig. 6), suggest that CM samples have 655 experienced extensive interaction with a pervasive melt or fluid (for fluid-interaction at Maqsad 656 CMTZ see also Rospabé et al. 2017, 2018a, 2019a). In the mantle section sampled by the CM 657 cores, crosscutting dunites are widespread and represent end products of the opx-consuming 658 reaction. The harzburgites experienced a more extensive melt flow at a shallow level that 659 contributed to their ol enrichment (Fig. 1e, logs). The dunites were probably individualized by 660 channeled percolation at the top of the melting column. 661

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# 4.3. General geological and magmatic context of the OmanDP CM sites

664 Most studies of the ultramafic rocks from the Oman ophiolite consider the dunites to be channels in the mantle or massive at the crust-mantle transition, and replacive in origin. In this model, the 665 dunites represent residues of reaction between a melt undersaturated in silica at low pressure and 666 the host mantle harzburgites during a melt percolation event that led to the complete consumption 667 of orthopyroxene and to the concomitant precipitation of olivine (e.g. Abily & Ceuleneer, 2013; 668 Boudier & Nicolas, 1995; Braun et al., 2002; Godard et al., 2000; Kelemen et al., 1995, 1997; 669 Koga et al., 2001; Rabinowicz et al., 1987; Rospabé et al., 2017, 2018a, 2019a). The hypothesis 670 of the replacive origin of the dunites, also proposed for other ophiolitic sections and for dunites 671 associated with abyssal peridotites (e.g. Dick & Natland, 1996; Kelemen et al., 1990; Godard et 672 al., 2008; Quick, 1981a, 1981b), contrasts with an older cumulative model, in which ol 673 crystallization and accumulation created the lowermost part of the oceanic crust (e.g. Elthon, 1979; 674 O'Hara, 1965; Smewing, 1981). Alternatively, it has been proposed that the crust-mantle transition 675 of the Oman ophiolite may be of double origin, with the lower 80% as replacive and the upper 676 20% as cumulates (Abily & Ceuleneer, 2013). In the case of the OmanDP CM cores, the 677 alternations between dunites and mantle harzburgites at the top of the mantle sequence is consistent 678 679 with a melt/rock reaction origin, a feature observed at crust-mantle transitions of many massifs

along the Oman ophiolite (e.g. Boudier & Nicolas, 1995) and that cannot be accounted for by a
 simple fractional crystallization process.

The CM ultramafic rocks studied in this paper (Wadi Tayin) compared to previously studied 682 Maqsad MTZ dunites and diapir harzburgites show that: (1) the dunites have a replacive origin, 683 they are the products of melt-harzburgite reaction leading to the complete consumption of 684 orthopyroxene and concomitant precipitation of ol (e.g. Abily & Ceuleneer, 2013; Boudier & 685 Nicolas, 1995; Gerbert-Gaillard, 2002; Godard et al., 2000; Kelemen et al., 1995, 1997; Koga et 686 al., 2001; Rabinowicz et al., 1987; Rospabé et al., 2018a, and references therein). (2) The 687 alternations between dunites and mantle harzburgites observed at the top of the mantle sequence 688 of CM Holes were also observed at the base of the crust-mantle transition in many other massifs 689 along the Oman ophiolite (e.g. Boudier & Nicolas, 1995), recording a snapshot of melt/harzburgite 690 reaction frozen at the time of the uppermost mantle dunitization. (3) The vertical chemical trend 691 changes related to the focus of the percolation/migration along faults observed at CM Holes 692 samples have been previously observed at the Maqsad CMTZ dunites, with the same characteristic 693 thickness of about 50 m (Rospabé et al. 2019a, 2020), confirming the control of synmagmatic 694 faulting on melt/peridotite reactions and the petrological and geochemical structuration of the 695 CMTZ. 696

Structural and petrological mappings of the Oman ophiolite have revealed contrasting domains 697 along the ophiolite. Especially, the spatially varying nature and composition of the dikes cross-698 cutting the mantle section reflect formation involving a MORB-like melt mainly in the south-699 700 eastern Nakhl, Samail and Wadi Tayin massifs (troctolite and ol-gabbro dikes), contrasting with a more widespread depleted, calc-alkaline magma composition elsewhere (mostly gabbronorite and 701 pyroxenite dikes) (Python & Ceuleneer, 2003; Python et al., 2008). The MOR-like area 702 characterizes a NW-SE oriented paleo-spreading segment that seems to have developed within 703 older, already accreted lithosphere of depleted calc-alkaline affinity (Ceuleneer et al., 1988; 1996; 704 Gerbert-Gaillard, 2002; Godard et al., 2000; Nicolas et al., 2000; Python & Ceuleneer, 2003). This 705 MORB segment is hypothesized to have been centered on, and fed with melts by, the fossil mantle 706 diapir of the Maqsad area in the Samail massif (Rabinowicz et al., 1987; Ceuleneer et al., 1988; 707 Jousselin et al., 1998). According to the published structural and petrological maps, the drilling 708 site is located within the area where melts were MORB-like, near the NE limit of the paleo 709 spreading segment. This delimitation has been defined as the Makhibiyah shear zone in the Wadi 710 Tayin massif, making the contact with the older lithosphere (Nicolas & Boudier, 2008). 711 Accordingly, it is reasonable to consider that a MORB-like melt played an important role in the 712 formation of the CM dunites and harzburgites. 713

The two main differences with the crust-mantle transition and mantle section of the neighboring 714 Maqsad area, that have been extensively studied due to their diapir-related features, are that (1) if 715 the two sites (Maqsad and OmanDP CM) are both more or less related to the rise of the mantle 716 diapir, the CM site may have had less pronounced magmatic activity due to its distance to the axis 717 of the diapir, and (2) the CM site may have been contaminated by possible remelting of the base 718 of the old lithosphere during the development of the MORB segment, as evidenced in the Samail 719 massif at the borders of the area influenced by the diapir (Amri et al., 1996; Benoit et al., 1999; 720 Clénet et al., 2010). In cores from Holes CM1A and CM2B, core description revealed the 721 widespread presence of magmatic impregnations in the dunites (i.e. minerals that crystallized 722 723 interstitially between ol grain during melt migration; e.g. Benn et al., 1988; Dick, 1989), and of magmatic segregation or more intruding dikes (Kelemen et al., 2020a, 2020b). This evidence for melt migration was described within both the crust-mantle transition and the mantle sequence. The impregnations in dunites are composed of plagioclase and clinopyroxene, and the magmatic segregations and dikes are mostly troctolitic and gabbroic, consistent with a MORB-like parent melt produced by decompression melting in the mantle. On the other hand, more exotic websterite and anorthositic gabbros to anorthosites were also observed in the mantle sequence and the dunite/harzburgite alternations.

CM ultramafic rocks studied in this paper (Wadi Tayin) show thus multiple differences compared 731 to the previously studied Magsad CMTZ dunites and diapir harzburgites: (1) the CM samples have 732 been more intensively affected by low temperature alteration features (serpentinization up to 100% 733 in many samples). (2) The crust-mantle transition zone is thinner along CM drilled Holes (~150 734 m, Kelemen et al., 2020a, 2020b) compared to the Magsad CMTZ (300-400 m, Abily and 735 Ceuleneer, 2013, Boudier and Nicolas, 1995; Jousselin and Nicolas, 2000; Rabinowicz et al., 1987; 736 Rospabé, 2018; Rospabé et al., 2019a). (3) CM pure dunites display lower LREE concentrations 737 compared to the Magsad ones. (4) Two major trends are observed in the trace element signatures 738 of CM dunites and harzburgites, with good correlations between the Th and U, Nb and LREE on 739 one hand, and between the HREE, Ti and Hf on the other hand, that are not observed at Magsad. 740 (5) The CM site CMTZ is mostly composed of pure dunites at the top and impregnated dunites at 741 742 the bottom, whereas the typical structuration of the Maqsad CMTZ is generally composed by impregnated dunites at the top and pure dunites at the bottom (Rospabé, 2018; Rospabé et al. 743 2019a). (6) The dunites' impregnation characteristic appears to be different in Magsad (e.g., mostly 744 plagioclase and clinopyroxene (Boudier and Nicolas, 1995; Koga et al., 2001; Abily and 745 Ceuleneer, 2013) but also widespread opx and amphibole impregnations in the higher level of the 746 transition zone (Rospabé et al., 2017, 2018a, 2019a) compared to CM Hole dunites (plagioclase 747 and clinopyroxene impregnations only). Similarly, exotic silicate inclusions (opx, amph, mica for 748 the more abundant) enclosed in disseminated chromite grains in the dunites from the Maqsad area 749 as well as in associated chromitite ore bodies (Lorand and Ceuleneer, 1989; Leblanc and 750 751 Ceuleneer, 1991; Schiano et al., 1997; Borisova et al., 2012; Rollinson et al., 2018; Zagrtdenov et al., 2018; Rospabé et al., 2019b; 2020, 2021), suggest the involvement of a fluid or fluid-rich melt 752 in the melt/rock reactions, which were not investigated in existing work on cores CM1A and 753 CM2B - a few chromite schlierens have been sampled along the cores but not studied in details 754 yet. Further investigation of CM ultramafic rock mineral chemistry is need to evaluate these 755 discrepancies between the two CM and Maqsad sites. 756

All the above similarities and differences between CM Wadi Tayin and Magsad ultramafic rocks 757 758 point to a lighter imprint of melt-rock reaction at CM compared to Maqsad (e.g., thinner transition zone, lighter LREE enrichment in dunites, two distinct geochemical trends in trace elements, that 759 perhaps are totally overprinted by stronger melt/rock reaction at Maqsad). This could be the 760 consequence of one or the several factors: It could be related to the structural position of Wadi 761 Tayin (at the periphery of the diapir) and Maqsad (centered on the diapir) leading to different 762 melt/rock ratios (i.e., the geological context), to timing of the occurrence of the melt-rock 763 764 interaction, and/or to the nature of the percolating magma/fluid involved in the melt-rock reactions.

Further detailed mineral chemistry, trace element geochemical modelling, structural and microstructural studies will be of use for addressing the above questions.

# 767 Conclusions

Continuous sampling of the Oman crust-mantle transition zone at Holes CM1A and CM2B, 768 recovered by the Oman Drilling Project allows the study of the large range of petrological and 769 geochemical variations in Oman ultramafic rocks with an unprecedented high resolution. Volatile 770 (H<sub>2</sub>O and CO<sub>2</sub>), and major and trace elements of 56 dunites and 49 harzburgites from Holes CM1A 771 and CM2B have been analyzed. CM1A and CM2B volatile element contents reflect extensive 772 773 serpentinization (+/- carbonation) linked to the late-stage interaction with H2O- and/or CO2bearing fluids. The refractory samples are characterized by relatively homogeneous modal and 774 775 major element compositions, whereas other samples show primary cryptic and modal refertilization. Bulk rock Mg, Si and Al systematics and normative mineral modes suggest that 776 open system behavior during alteration (Mg loss and/or Si gain) affected many samples. However, 777 the trace element concentrations are interpreted as reflecting magmatic processes and exhibit 778 779 significant variations: the refractory harzburgites characterized by linear REE patterns are interpreted as mantle residues after  $\geq 15$  % melt extraction. The REE signatures in these samples 780 can be explained by melt transport associated with partial melting. Other harzburgites displaying 781 U-/V-shaped REE patterns are interpreted as the result of interstitial melt percolation. The pure 782 and impregnated dunites from the mantle and the crust-mantle transition zones are characterized 783 by similar trace element patterns as the mantle harzburgites, they are interpreted as reflecting 784 785 different stages of ol-forming melt/rock reactions at high melt/rock ratio. The plagioclase and pyroxene minerals present in impregnated dunites indicate that the impregnated dunites represent 786 the stage before complete extraction of the melt from the dunites, whereas the pure dunites 787 represent the melt/rock reaction end-product after complete melt extraction (i.e. compaction of the 788 ol matrix and system closure before/without interstitial plagioclase and clinopyroxene 789 crystallization). 790

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# 810 Data availability

Data are available as Supporting Information Tables 1-3. Sample lithology, macroscopic and microscopic observations and mineral modes calculated from major elements are reported in supplementary data table 1. Whole rock major and volatile element compositions are reported in supplementary data table 2. Whole rock trace element compositions are reported in supplementary data table 3. All data will become available online on PANGAEA (www.pangaea.de). These include analysis results of major, as well as volatiles and trace element compositions and the calculated mineral modes from XRF data of all analyzed samples.

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# 1245 Figure captions







Figure 2 microphotographs of petrographic details of a selection of CMTZ dunites (a, b, c and d) 1257 and harzburgites (e, f, g, h, i, j, k and l) from Holes CM1A and CM2B. The dunites are 1258 characterized by fine- to medium grained granular texture and the harzburgites by porphyroclastic 1259 texture. The dunites and harzburgites have generally preserved their primary high temperature 1260 texture (granular or porphyroclastic textures) after complete (a, and b), or partial (c, d, e, f, g, h. i, 1261 **j**, **k** and **l**) replacement of olivine crystals by serpentine. **i** and **j** illustrate microphotographs of a 1262 plastically deformed orthopyroxene surrounded by neoblasts. k and l show harzburgite CM2B-1263 104Z2 crosscut by low-temperature veins (Carbonate). 1264

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Figure 3 Downhole plots of (from left to right) the lithology, LOI (wt.%), H<sub>2</sub>O (wt.%), CO<sub>2</sub> (wt.%) and calculated CaCO<sub>3</sub> (wt.%) contents in pure dunites, impregnated dunites, harzburgites and carbonate-bearing harzburgites recovered samples at Holes CM1A and CM2B. The thick solid red lines indicate the faults. XRD: X-Ray Diffraction, serp: serpentine, brc: brucite, mag: magnetite.

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1274 **Figure 4** Downhole plots of (from left to right) the lithology, Mg# (cationic 100 x Mg/(Mg+Fe<sub>total</sub>);

1275 calculated assuming all Fe as FeO) (mol%),  $SiO_2$  (wt.%),  $TiO_2$  (wt.%),  $Al_2O_3$  (wt.%), FeO (wt.%),

1276 MgO (wt.%), MnO (wt.%), CaO (wt.%), Na<sub>2</sub>O (wt.%), K<sub>2</sub>O (wt.%), and P<sub>2</sub>O<sub>5</sub> (wt.%) in whole

1277 rock samples recovered at Holes CM1A and CM2B. The thick solid red lines indicate the faults.



**Figure 5** Whole rock major compositions of samples recovered at Holes CM1A and CM2B compared to other crust-mantle transition dunites and mantle harzburgites/lherzolites from the

Oman ophiolite (Gerbert-Gaillard, 2002; Godard et al., 2000; Hanghøj et al., 2010; Khedr et al.,
2014; Monnier et al., 2006; Nicolle et al., 2016; Rospabé et al., 2018a, 2019a; Takazawa et al.,

2003). (a) MgO/SiO<sub>2</sub> versus Al<sub>2</sub>O<sub>3</sub>/SiO<sub>2</sub>, (b) total iron as FeO versus MgO, and (c) Al<sub>2</sub>O<sub>3</sub> versus

1285 CaO. Compositions are recalculated on a volatile-free basis. Red bar in panel (a) represents the





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1289 Figure 6 Downhole plots of (from left to right) (LREE: (La+Ce+Pr+Nd)<sub>CN</sub>), (HREE: (Ho+Er+Tm+Yb+Lu)cn), (La)cn, (Yb)cn, (La/Sm)cn, (Gd/Yb)cn, (U)<sub>PMN</sub> and (Th)<sub>PMN</sub> in whole 1290 rock samples recovered at Hole CM1A and CM2B (CN: chondrite-normalized; PMN: primitive 1291 mantle-normalized). The thick solid red lines indicate the faults, and the thicker solid black lines 1292 indicate the (La)<sub>CN</sub> and (La/Sm)<sub>CN</sub> concentrations in the most residual peridotites from 'Plate 1293 model' of Vernières et al. (1997) applied by Godard et al. (2000) (experiment b and d, see Fig.9), 1294 1295 the numbers on the black lines indicate the proportions of trapped melt (in percentage) issued from the models. Normalizing chondrite and Primitive Mantle values are from Barrat et al. (2012) and 1296 Sun and McDonough (1989) respectively. 1297

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Figure 7a Chondrite-normalized REE and Primitive Mantle-normalized multi-element patterns of 1301 pure dunites, impregnated dunites, harzburgites and carbonate-bearing harzburgites from the crust, 1302 crust-mantle transition zone and mantle sections recovered at Hole CM1A. Other pure dunites 1303 (dark gray field formed) and impregnated dunites (field formed by black line) from the crust-1304 1305 mantle transition, and harzburgites (light gray field) and lherzolites (field formed by dashed line) patterns from the mantle section of the whole Oman ophiolite are reported for comparison 1306 (Gerbert-Gaillard, 2002; Girardeau et al., 2002; Godard et al., 2000; Hanghøj et al., 2010; Khedr 1307 et al., 2014; Lippard et al., 1986; Monnier et al., 2006; Nicolle et al., 2016, Rospabé et al., 2018a, 1308 2019a; Takazawa et al., 2003). Normalizing chondrite and Primitive Mantle values are from Barrat 1309 et al. (2012) and Sun and McDonough (1989) respectively. 1310



Figure 7b Chondrite-normalized REE and Primitive Mantle-normalized multi-element patterns of 1312 pure dunites, impregnated dunites, harzburgites and carbonate-bearing harzburgites from the crust-1313 mantle transition zone and mantle sections recovered at Hole CM2B. Other pure dunites (dark gray 1314 field formed) and impregnated dunites (field formed by black line) from the crust-mantle 1315 transition, and harzburgites (light gray field) and lherzolites (field formed by dashed line) patterns 1316 from the mantle section of the whole Oman ophiolite are reported for comparison (Gerbert-1317 Gaillard, 2002; Girardeau et al., 2002; Godard et al., 2000; Hanghøj et al., 2010; Khedr et al., 2014; 1318 Lippard et al., 1986; Monnier et al., 2006; Nicolle et al., 2016, Rospabé et al., 2018a, 2019a; 1319 Takazawa et al., 2003). Normalizing chondrite and Primitive Mantle values are from Barrat et al. 1320 1321 (2012) and Sun and McDonough (1989) respectively.



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**Figure 8** Plots of LOI (wt.%) vs. Zr (ppm), Cs (ppm), Li (ppm), MgO (wt.%), and SiO<sub>2</sub> (wt.%); Th (ppm) vs. Zr (ppm), Cs (ppm), Nb (ppm) , La (ppm), and U (ppm); Yb (ppm) vs. Zr (ppm), Cs (ppm), Li (ppm), Hf (ppm), and Ti (ppm) in dunite and harzburgites recovered at Hole CM1A and CM2B. Other pure dunites, impregnated dunites, harzburgites and lherzolites compositions from the crust-mantle and the mantle section of the whole Oman ophiolite are reported for comparison (Gerbert-Gaillard, 2002; Girardeau et al., 2002; Godard et al., 2000; Hanghøj et al., 2010;

1330 Khedr et al., 2014; Lippard et al., 1986; Monnier et al., 2006; Nicolle et al., 2016, Rospabé et al., 2018a, 2019a;

1331 Takazawa et al., 2003).





**Figure 9** Modal compositions of the analyzed CM harzburgites plotted on a cpx/opx vs. olivine diagram. The field defined by the mantle harzburgites studied in Godard et al., 2000 is reported for comparison. Published melting models are also shown for comparison: model 1 represents the polybaric melting model after Niu (1997), (1a) with and (1b) without excess olivine; model 2 represents the isobaric melting after Walter et al. (1995) at 11 (2a), 16 (2b) and 17 kbar (2c). The initial modal composition is given by Niu (1997) for polybaric melting and was fixed for isobaric melting as: 55% ol, 28% opx, 15% cpx and 2% sp. Numbers refer to percent melting degrees.





Figure 10 CM1A and CM2B REE linear flat shaped REE patterns compared to the 'Plate model' 1344 1345 of Vernières et al. (1997) applied by Godard et al. (2000) to simulate REE variations in a peridotite affected by partial melting with (a) or without (b) melt infiltration. The chondrite-normalized REE 1346 patterns of the Oman harzburgites from Godard et al. (2000) (main harzburgite section) are also 1347 1348 shown for comparison. The authors simulate standard incremental melting in model (a) and the percolation of fixed N-MORB composition melt through molten peridotites in model (b). The 1349 initial modal composition was (spinel neglected): 57% ol, 28% opx and 15% cpx. The melting 1350 reaction was taken from Walter et al. (1995). Mineral/melt partition coefficients are the same as 1351 those selected by Bedini & Bodinier (1999). Numbers on the chondrite-normalized REE patterns 1352 indicate olivine proportion (in percentage) in residual peridotites. Thicker lines indicate the REE 1353 patterns of the less residual peridotites. In model (a), the most residual peridotite (76% olivine) is 1354 produced after 21.1% melt extraction. In model (b), the ratio of infiltrated melt to peridotite varies 1355 from 0.02 to 0.19. (Bottom) Modifications of the REE patterns of residual peridotites due to the 1356 presence of equilibrium, trapped melt. Models (c) and (d) show the effect of trapped melt on the 1357 most residual peridotites of models (a) and (b), respectively (thicker solid lines). Numbers on the 1358 REE patterns indicate the proportions of trapped melt (in percentage). Normalizing chondrite 1359 values are from Barrat et al. (2012). 1360 1361