A Model for Positive Corona Inception from Charged Ellipsoidal Thundercloud Hydrometeors

Silke Alexandra Peeters¹, Shahriar Mirpour¹, Christoph Köhn², and Sander Nijdam¹

¹Eindhoven University of Technology ²Technical University of Denmark

November 22, 2022

Abstract

Lightning is observed to incept in thundercloud electric fields below the threshold value E_k for discharge initiation. To explain this, the local enhancement of the electric field by hydrometeors is considered. The conditions for the onset of positive corona discharges are studied in air for ellipsoidal geometries. A hydrometeor is simulated as an individual charged conductor in zero ambient field; there is only a field generated by the charge on the hydrometeor surface. By doing so, the feasibility of corona inception from ellipsoidal hydrometeors can be formulated based on the self-sustaining condition of electron avalanches. For representative hydrometeor volumes and typical thundercloud pressure, values between \$1.2\,E_k\$ and \$37\,E_k\$ were found for the onset electric field at the tip of the ellipsoid. From simulations the required ambient electric field for corona onset from an uncharged hydrometeor can then be derived. This results in values between \$0.07\,E_k\$ and \$0.8\,E_k\$ for semi-axes aspect ratios between 0.01 and 1. The charge required on the hydrometeor volume. For the simulated hydrometeors, the values of this onset charge for typical pressures are between 1500\,pC and 3200\,pC. Including a size-correction for comparison to in situ measurement shows agreement with measured precipitation charges. From the results it is concluded that corona onset from ellipsoidal hydrometeors of a realistic volume can be achieved in thundercloud conditions for certain aspect ratios.

A Model for Positive Corona Inception from Charged Ellipsoidal Thundercloud Hydrometeors

S. A. Peeters¹, S. Mirpour¹, C. Köhn², S. Nijdam¹

 $^1\rm Eindhoven$ University of Technology, Department of Applied Physics, The Netherlands $^2\rm DTU$ Space, National Space Institute, Technical University of Denmark, Denmark

Key Points:

1

2

3

4 5

6

7	•	Corona onset through hydrometeors is modelled using the self-sustaining condi-
8		tion of electron avalanches.
9	•	An optimal ellipsoidal aspect ratio of 0.1 for corona inception for representative
10		conditions is found.
11	•	Lightning inception via ellipsoidal hydrometeors is found to be achievable in thun-
12		dercloud conditions.

Corresponding author: S. A. Peeters, s.a.peeters@student.tue.nl

13 Abstract

Lightning is observed to incept in thundercloud electric fields below the threshold value 14 E_k for discharge initiation. To explain this, the local enhancement of the electric field 15 by hydrometeors is considered. The conditions for the onset of positive corona discharges 16 are studied in air for ellipsoidal geometries. A hydrometeor is simulated as an individ-17 ual charged conductor in zero ambient field; there is only a field generated by the charge 18 on the hydrometeor surface. By doing so, the feasibility of corona inception from ellip-19 soidal hydrometeors can be formulated based on the self-sustaining condition of electron 20 avalanches. For representative hydrometeor volumes and typical thundercloud pressure, 21 values between $1.2 E_k$ and $37 E_k$ were found for the onset electric field at the tip of the 22 ellipsoid. From simulations the required ambient electric field for corona onset from an 23 uncharged hydrometeor can then be derived. This results in values between $0.07 E_k$ and 24 $0.8 E_k$ for semi-axes aspect ratios between 0.01 and 1. The charge required on the hy-25 drometeor surface for corona onset is minimum for semi-axes aspect ratios between 0.04 26 and 0.07 depending on the considered hydrometeor volume. For the simulated hydrom-27 eteors, the values of this onset charge for typical pressures are between 1500 pC and 3200 pC. 28 Including a size-correction for comparison to in situ measurement shows agreement with 29 measured precipitation charges. From the results it is concluded that corona onset from 30 ellipsoidal hydrometeors of a realistic volume can be achieved in thundercloud conditions 31 32 for certain aspect ratios.

33 1 Introduction

One of the greatest unanswered questions in lightning physics is how lightning is 34 initiated in a thunderstorm (Mazur, 2016; Petersen et al., 2008; Dwyer & Uman, 2014). 35 From in situ measurements it is found that lightning initiates in thundercloud electric 36 fields which are considerably lower than the breakdown electric field required for the in-37 ception of electric discharges (Stolzenburg & Marshall, 2009; Marshall et al., 1995). One 38 of the most popular and widely corroborated theories explaining how this is possible is 39 the hydrometeor theory (Mazur, 2016; Petersen et al., 2008). This theory states that hy-40 drometeors - ice and water particles in thunderclouds - locally enhance the electric field, 41 such that the breakdown field is exceeded, and lightning inception is enabled. In recent 42 observations of narrow bipolar events in thunderstorms, which generally coincide with 43 lightning initiation, clear evidence supporting the involvement of hydrometeors was ob-44 tained (Rison et al., 2016). The role of hydrometeors in lighting inception has been in-45 vestigated in laboratory experiments (Petersen et al., 2015; Coquillat et al., 1995; R. F. Grif-46 fiths & Latham, 1974) with a main focus on corona onset, which is the initial stage of 47 the formation of a lightning leader. 48

A corona discharge is the result of electrical breakdown, which occurs at the volt-49 age where the insulating gas surrounding the electrode becomes electrically conductive. 50 An electrical discharge is thus only possible when a critical voltage, the onset voltage V_0 , 51 on the electrode is reached. Equivalently, the electric field E in the discharge region should 52 exceed the breakdown threshold field E_k and thus the field on the surface of the elec-53 trode should exceed the onset field E_0 . The breakdown field scales linearly with pres-54 sure, and at a typical thundercloud altitude pressure of 0.4 atm this field has a value of 55 about 10 kV/cm (Raizer, 1991a). The main mechanism of electrical breakdown is the elec-56 tron avalanche. For electric fields above E_k , electrons can multiply by means of impact 57 ionization of air molecules, thereby forming avalanches. In order to have a self-sustaining 58 discharge, a constant source of seed electrons is required, which can be supplied by photo-59 ionization (Raizer, 1991b). The first group of electrons that collides with the gas molecules 60 and leads to photoionization is known as the primary electron avalanche, and the sub-61 sequently formed second group of electrons that can give further photoionization is known 62 as the secondary electron avalanche (see Figure 1). In air, the gas molecules that are dom-63 inant in emitting photons after collisions with free electrons are nitrogen molecules, and 64 the gas molecules that are predominantly photoionized by these photons are oxygen molecules. 65

It should be noted that avalanche formation is a stochastic process, such that electron
 multiplication can also take place in fields (slightly) below the breakdown field. These
 contributions are briefly investigated but otherwise neglected in this work.

The onset of a corona discharge is typically defined by the discharge becoming selfsustaining. The self-sustaining criterion that is often applied is the amount of photons produced by the secondary avalanche being at least equal to those produced by the primary avalanche (Liu et al., 2012; Naidis, 2005). This condition is also adapted by the current work, which closely follows the structure of the work by Liu et al. (2012).

Depending on electrode polarity, corona discharges can be positive or negative. A
popular hypothesis for lightning initiation is that the development of a positive streamer
system, developed from a seed positive streamer from the corona on a hydrometeor, precedes and leads to negative breakdown (L. P. Babich et al., 2016; Petersen et al., 2008;
C. T. Phelps & Griffiths, 1976; C. Phelps, 1974; Loeb, 1966). Therefore, positive corona
discharges are of great interest when investigating the initial stage of lightning initiation.

Laboratory experiments have revealed that the onset of a corona discharge strongly 80 depends on the size and shape of the hydrometeor. In their study on corona initiation 81 from small ice crystals, Petersen et al. (2015) reported that the onset field E_0 decreases 82 with hydrometeor length and that ice crystals with sharper tips promote glow coronae 83 while inhibiting positive streamer formation. Moreover, they noted that the onset field 84 increases linearly with the relative gas density $\delta = N/N_0$ (where N and N₀ are the ac-85 tual and standard gas densities), meaning $E_0 \sim p/T$, with p the pressure and T the tem-86 perature. The decrease of the onset field with size is also found in many point-to-plane 87 and rod-to-plane experiments using metal electrodes (D'Alessandro & Berger, 1999; Wa-88 ters & Stark, 1975; Nasser & Heiszler, 1974; Schumann, 1923; Kip, 1938), which are ob-89 served to give corona onset voltages very similar to ice electrodes (Bandel, 1951). 90

In simulations, similar conclusions were reached. Dubinova et al. (2015) investigated 91 discharge inception conditions for dielectric ellipsoidal hydrometeors and concluded that 92 an increase in hydrometeor length yields stronger field enhancement, as does a decrease 93 in hydrometeor tip radius. Hence, a longer, sharper hydrometeor generally requires a lower 94 background electric field for the initiation of a discharge. Likewise, in simulations of streamer 95 initiation from charged water drops L. P. Babich et al. (2016) found a lower threshold 96 ambient field for larger drop sizes. Dubinova et al. (2015) also observed an optimal semi-97 axes aspect ratio for inception; though longer hydrometeors produce a higher electric field, 98 the probability of discharge initiation decreases when they become too sharp, because qq the field enhancement becomes too localized at the tip. As this ratio fixes the ellipsoidal 100 hydrometeor's shape, an optimal shape can be determined. Simulations (Riousset et al., 101 2020) also show the experimentally observed linear pressure dependence of discharge ini-102 tiation. This is expected, as it follows from the pressure dependence of the breakdown 103 field 104

In addition to size, shape and air density, the onset of a corona discharge has been 105 found to depend on the orientation, surface features and initial charge of the hydrom-106 eteor. R. F. Griffiths and Latham (1974) concluded from experimental studies on ice par-107 ticles that onset fields in thundercloud regions are probably in the range of $400-500 \, \text{kV/cm}$, 108 which was later corrected by R. Griffiths (1975) to $350-450 \,\mathrm{kV/cm}$ when taking into ac-109 count the effect of charge on ice particles. Furthermore, R. F. Griffiths and Latham (1974) 110 suggested that continuous corona discharges could be generated from thundercloud ice 111 crystals at temperatures above -18 °C only. Of course, the gas density increases with de-112 creasing temperature, explaining the subsequent increase of the onset field. Moreover, 113 the surface conductivity decreases with decreasing temperature such that corona onset 114 becomes less likely (R. F. Griffiths & Latham, 1974; Petersen et al., 2006), and gener-115 ally smaller ice crystals are formed at lower temperatures (Petersen et al., 2006). In 2006, 116 Petersen et al. (2006) demonstrated that corona discharges can initiate in temperatures 117 down to -38 °C, showing that corona and streamer discharges can initiate from hydrom-118 eteors at thundercloud altitudes relevant for lightning initiation. Moreover, from numer-119 ical simulations L. Babich et al. (2017) observed that the required charge on hydrom-120

eteors at these representative temperatures and altitudes agrees with measured thundercloud precipitation charges, which are generally between 10 and 200 pC and for a small fraction of hydrometeors between 200 and 400 pC (Marshall & Winn, 1982).

To conclude, these studies reveal that the onset of a corona discharge from a hydrometeor depends on its size, shape and surface charge, and on environmental conditions such as pressure, temperature and the ambient electric field. Experimental results and in situ measurements indicate the essential role of hydrometeors in lightning initiation. These findings are supported by simulations of lightning inception from ice and water particles.

The comparison of experimental work on corona onset from ice point electrodes to measurements on metal point electrodes has shown the corona onset voltage to be very comparable (Bandel, 1951). To simulate the onset of a positive corona discharge from a metal electrode in air, Naidis (2005) introduced a model giving a corona inception criterion taking into account the ambient pressure and the size and shape of the electrode. This model was applied to spherical and cylindrical electrodes, and later revisited by Liu et al. (2012) for the spherical case.

The main goal of this paper is to extend this model to include another represen-137 tative shape, the prolate spheroid, as ice and water particles in a thundercloud can have 138 a wide variety of shapes depending on thundercloud conditions. Their sizes range from 139 a few micrometers to several centimeters (MacGorman et al., 1998). The size distribu-140 tion of hydrometeors is little investigated within thunderclouds due to difficulties of in 141 situ measurements (Mazur, 2016), but it is expected that the extreme cases of several 142 centimeters are rare, and that a millimeter range is more representative (Weinheimer et 143 al., 1991; Gardiner et al., 1985). When these hydrometeors fall downwards due to grav-144 ity, they are extended along the vertical direction. The shape of the hydrometeor in the 145 direction perpendicular to the thundercloud electric field has a negligible contribution 146 to the field enhancement. More precisely, the enhancement at the tips is mainly deter-147 mined by the length of the hydrometeor and the radius of curvature of the tip of the hy-148 drometeor (Köhn & Ebert, 2015; Dubinova et al., 2015). Taking this into account, it can 149 prove fruitful to investigate ellipsoidal hydrometeors. More specifically, assuming cylin-150 drically symmetric thundercloud conditions, a prolate ellipsoid of revolution, or prolate 151 spheroid, is considered. 152

Thus, the purpose of this study is to simulate positive corona discharges originat-153 ing from a positively charged spheroidal hydrometeor tip. In doing so, the feasibility of 154 lightning initiation from a spheroidal hydrometeor is studied. The simulation of this con-155 figuration is done using the model for the onset of positive corona discharges introduced 156 by Naidis (2005) and further elaborated by Liu et al. (2012). The investigated hydrom-157 eteor is isolated and without ambient electric field. Thus, there is only an electric field 158 generated by the charge on the hydrometeor, which differs from realistic lightning oc-159 currences, where there is also an external field present due to the large-scale charge dis-160 tribution. However, the effects of the field induced by a charged particle can already re-161 veal a lot about the role of particle shape and size in discharge inception. Hence, for the 162 charged hydrometeor the dependence of corona onset on its semi-axes aspect ratio and 163 volume is reported for various ambient pressures by varying its major and minor axes. 164

¹⁶⁵ 2 Model Description

As elaborated, a corona discharge is the result of electrical breakdown via direct 166 impact ionization within avalanches. The resulting avalanches are seeded by electrons 167 supplied through photoionization. Taking loss by attachment processes into account, $\alpha =$ 168 η defines electrical breakdown, where α is the number of ionizing collisions per unit length 169 and η the number of electron attachments per unit length. Formulating the net ioniza-170 tion coefficient $\alpha_{eff} = \alpha - \eta$, breakdown is defined by $\alpha_{eff} = 0$. Of course these co-171 efficients depend on the electric field E, meaning $\alpha_{eff} = 0$ determines the breakdown 172 field E_k . 173

The number of photons produced by a primary avalanche is denoted by N_1 , and 174 those produced by a secondary avalanche by N_2 . N_2 depends on N_1 through $N_2 = \gamma N_1$, 175 where γ is the mean number of photons from the secondary avalanche produced by one 176 of the photons from the primary avalanche (see Figure 1). In short, γ is the multiplica-177 tion factor. Naidis (2005) formulates the criterion for corona inception as the secondary 178 avalanche producing at least as many photons as the primary avalanche, so $N_2 = N_1$, 179 or equivalently $\gamma = 1$. Then, the discharge is self-sustaining; it can proceed without ex-180 ternal ionization sources. This criterion does not take into account the stochastic nature 181 of discharge inception. The region around the hydrometeor where the breakdown field 182 is exceeded is sufficiently small such that individual electron avalanches, which have an 183 intrinsically random nature, should be considered. Here, the randomness, and therefore 184 contributions from outside this region, is neglected, as only the total amount of electrons 185 in the avalanche is investigated. The inclusion of stochastic effects would soften the cri-186 terion, as then electrons can 'tunnel' to higher energies (Rutjes, 2018). 187

For point and wire electrodes, most of the electrons and photons are produced near 188 the surface of the electrodes. It is therefore a reasonable assumption that all photons that 189 lead to photoionization are produced at the electrode surface. This assumption overes-190 timates the effect of photoionization, as the effect of photons on electron production is 191 now maximized. As will be further substantiated, this paper studies the minimum con-192 ditions for the onset of a corona discharge, such that this assumption is acceptable. Be-193 sides inducing photoionization and thereby triggering secondary electron avalanches, a 194 photon can also fall back to the electrode surface or leave the ionization region and con-195 sequently not contribute to the secondary avalanche. Different factors, such as the pho-196 ton absorption probability, affect this balance and thus play a role in satisfying the $\gamma =$ 197 1 criterion for positive corona onset. 198

To formulate the $\gamma = 1$ criterion, a spherical coordinate system (r, θ, ϕ) is intro-199 duced with its origin at the surface of the electrode. This is illustrated in Figure 2 for 200 the ellipsoidal electrode, with major axis a and minor axis b, considered in this paper. 201 In the region near the electrode tip the electric field E reaches its maximum. Consequently, 202 the number of electrons in the primary avalanche and the probability of photon emis-203 sion are also maximum near the tip. For simplicity, it is assumed that the primary pho-204 ton is emitted at the origin of the spherical coordinate system. Taking into account all 205 possible directions in which this photon can move, the photon absorption region can be 206 defined as the part of the ionization region $(E \ge E_k)$ where $\theta \le \pi/2$. In other words, 207 the photon absorption region is the region that can be reached by the photon and where 208 the field is sufficiently high such that an electron avalanche can be created. This region, 209 highlighted in deep yellow in Figure 2, is thus the region of interest for the initiation of 210 211 a corona discharge.

The corona inception criterion $\gamma = 1$, derived by Naidis (2005) using the above self-sustaining criterion, is then formulated as

$$\gamma \approx \xi \beta(\rho_0) \int_0^{2\pi} d\phi \int_0^{\pi/2} \sin\theta d\theta \int_0^{r_{max}(E)} r^2 P(r) \cdot \left[\exp\left(\int_{\rho_0(\theta,\phi)}^{\rho_{ab}(r,\theta,\phi)} \alpha_{eff}(\rho, E) d\rho \right) - 1 \right] dr = 1.$$
(1)

The coordinates ρ, r and θ are defined in Figure 2. Because of the cylindrical symmetry of the prolate spheroid, there is no ϕ -dependence. Besides the spherical coordinate system (r, θ, ϕ) with the origin at the tip of the ellipsoid, the coordinate ρ , which is given by the direction of the electric field and starts from the z-axis, is introduced as well, as is the radial coordinate ρ' from the center of the ellipsoid.

The term ξ is the ionization probability of an oxygen molecule at photon absorption. The distance $\rho_{ab}(r, \theta, \phi)$ is the distance between the point of photoionization (equivalent to the position of photon absorption) and the symmetry axis of the ellipsoid along the direction of the electric field in the point of photoionization. It is thus the length of



Figure 1. An illustrative image (not to scale) of the inception process in which the primary avalanche releases energetic photons, leading to the production of a photo-ionized electron. The secondary avalanche is formed by the multiplication of the photo-ionized electron via direct impact ionization. Inception occurs when the number of electrons in the secondary avalanche and the primary avalanche are equal. All processes occur in the photon absorption area, where the electric field is higher than the breakdown field (E_k) .

the line along the ρ coordinate that ends at the point of photon absorption (see Figure 223 2). Similarly, the distance from the symmetry axis of the ellipsoid to its surface along 224 the surface electric field direction is given by ρ_0 (for a sphere this would be its radius). 225 The position where the electric field has decreased to the breakdown field E_k is given 226 by r_{max} in the spherical coordinate system (r, θ, ϕ) . Naidis (2005) uses the expression for 227 the photon absorption probability P(r) in air where photoionization of oxygen molecules 228 takes place at absorption of radiation of wavelengths 98 - 102.5 nm, emitted by nitro-229 gen molecules (Zhelezniak et al., 1982) 230

$$P(r) = \frac{\exp\left(-\kappa_1 r \delta\right) - \exp\left(-\kappa_2 r \delta\right)}{4\pi r^3 \log\left(\kappa_2/\kappa_1\right)},\tag{2}$$

where $\kappa_1 = 5.6 \text{ cm}^{-1}$ and $\kappa_2 = 320 \text{ cm}^{-1}$. The term $\xi \beta(\rho_0)$ can be found from

231

$$\xi\beta = \left(0.03 + \frac{3.78}{E}\right)\frac{\delta_q}{\delta + \delta_q},\tag{3}$$

where $\delta_q = 0.04$ and E is the electric field (Zhelezniak et al., 1982). Here β is the coefficient of production of ionizing photons scaled to the net ionization coefficient α_{eff} . Because of its weak dependence on the electric field and the high fields at the electrode surface, β is approximated by its value $\beta(\rho_0)$ at the surface. To apply the corona inception criterion $\gamma = 1$ to a prolate spheroid, analytical expressions should be derived for the distances r_{max} , ρ_0 and ρ_{ab} , and the electric field E, on which the ionization probability and the net ionization coefficient depend.



Figure 2. A schematic of the photon absorption area around a positive ellipsoidal electrode.

To determine ρ_{ab} , the direction of the electric field is needed. This direction is given by the bisector of the two straight lines from the focal points of the prolate spheroid to the observation point (Curtright et al., 2020). Using various trigonometric relations, which are given in the supporting information, it can be derived that

$$\rho_{ab} = \frac{\sqrt{2}\rho_1^2 \sqrt{\frac{\left(4\sqrt{a^2 - b^2}(a + r\cos(\theta)) + \rho_1^2\right)(2ar\cos(\theta) + b^2 + \rho_1\rho_2 + r^2)}{\rho_1\rho_2}}}{\rho_1^2 + \rho_1\rho_2},\tag{4}$$

with ρ_1 and ρ_2 the straight lines from the two focal points of the ellipsoid to the observation point (see also the supporting information) given by

$$\rho_{1,2} = \sqrt{r^2 + (a \mp \sqrt{a^2 - b^2})^2 + 2(a \mp \sqrt{a^2 - b^2})r\cos\theta}$$
(5)

Using the derived expression for ρ_{ab} , the distance ρ_0 can be formulated. This is done by formulating the equation of the ellipsoid with the origin at its tip, using the coordinate system (r,θ,ϕ) . By solving the ellipsoid equation $(\frac{x^2}{b^2} + \frac{y^2}{b^2} + \frac{z^2}{a^2} = 1$ rewritten in the considered coordinates) for r and substituting r in the expression for ρ_{ab} , ρ_{ab} is constrained to the surface of the ellipse and thus ρ_0 is obtained. Because the surface of an ellipsoid and \mathbb{R}^3 do not form a diffeomorphic pair, two expressions for r are obtained and therefore two expressions for ρ_0 . These expressions are valid separately for $\theta \leq \pi/2$ and $\theta > \pi/2$ and are given in Appendix A.

The electric field of a conducting ellipsoid has been derived analytically by Köhn and Ebert (2015) for the prolate spheroid case and by Curtright et al. (2020) for arbitrary dimensions. The derivation of the electric field strength E yields

$$E(x, y, z) = \frac{Q}{4\pi\epsilon_0} \left(\prod_{k=1}^3 \frac{1}{\sqrt{a_k^2 + \Theta(\vec{r})}} \right) / \sqrt{\left(\sum_{m=1}^3 \frac{x_m^2}{(a_m^2 + \Theta(\vec{r}))^2} \right)},$$
(6)

where Q is the total charge on the ellipsoid surface, ϵ_0 is the vacuum permittivity, $\Theta(\vec{r})$ the equipotential surfaces and $a_1 = a_x = b$, $a_2 = a_y = b$ and $a_3 = a_z = a$ are the semi-axes of the considered spheroid of Figure 2. Moreover, the Θ -equipotentials follow from

$$\sum_{k=1}^{3} \frac{x_k^2}{a_k^2 + \Theta(\vec{r})} = 1, \text{ for } \Theta(\vec{r}) > 0.$$
(7)

The above electric field expression can be rewritten in the considered coordinates 260 (ρ', r, θ, ϕ) as defined in Figure 2. Here r can be converted to ρ' using the trigonomet-261 ric relation $\rho' = \sqrt{a^2 + r^2 + 2ar\cos(\theta)}$. The field is also reformulated to contain the 262 electric field at the ellipsoid tip $(z = a \text{ in equation } (6)) E_0 = \frac{Q}{4\pi b^2 \epsilon_0}$. To obtain the 263 final expression for the electric field, the ρ' coordinate is converted to the ρ coordinate 264 along the electric field direction as required for the $\gamma = 1$ criterion. This is done us-265 ing the derived ρ_{ab} expression. In order to have analytically solvable equations in this 266 derivation, r is not converted to ρ' in the conversion from ρ' to ρ . This means some am-267 biguity remains in the expression of the electric field $E = E(\rho, r, \theta)$. As eventually the 268 equation $\gamma = 1$ is solved numerically, this ambiguity is not a problem as long as the re-269 sulting α_{eff} (which is calculated using the electric field) behaves correctly. The used for-270 mulation is 271

$$E(\rho, r, \theta) = \frac{2b^2 E_0 \rho'}{\sqrt{a^2 - b^2 + q + {\rho'}^2} \left(-a^2 + b^2 + q + {\rho'}^2\right) \sqrt{\frac{(b^2 - a^2)(a^2 + 2ar\cos(\theta) + r^2\cos(2\theta))}{a^2 + 2ar\cos(\theta) + r^2}} + q + {\rho'}^2},$$
(8)

with the shorthand $q = \sqrt{\frac{2\rho'^2(b^2 - a^2)(a^2 + 2ar\cos(\theta) + r^2\cos(2\theta))}{a^2 + 2ar\cos(\theta) + r^2}} + (a^2 - b^2)^2 + {\rho'}^4$ and with

$$\rho' = \sqrt{\frac{2ar\cos(\theta)\left(3a^2 - 3b^2 + \rho^2\right) + 2a^2\rho^2 + a^2\rho_1\rho_2 - b^2\rho^2 - b^2\rho_1\rho_2 + \rho^2r^2 + \rho^2\rho_1\rho_2 + p}{2a^2 + 2ar\cos(\theta) - b^2 + r^2 + \rho_1\rho_2}},$$
(9)

with $p = 2a^4 - a^2b^2 + 2a^2r^2\cos(2\theta) + a^2r^2 - b^4 - 2b^2r^2\cos(2\theta) - b^2r^2$. The derivations of these expressions are presented in the supporting information.

Finally, the distance r_{max} from the tip of the prolate spheroid to the position where 276 $E = E_k$ can be determined. Because there is no explicit solution for r_{max} in the con-277 sidered geometry, this is done by approximating the surface $E = E_k$ as forming an el-278 lipsoid surface near the tip, as is validated in simulations in the supporting information. 279 Then, finding r_{max} specifically for $\theta = 0$ and $\theta = \pi/2$ is sufficient to obtain r_{max} for 280 arbitrary θ . These expressions are found by reformulating the electric field in terms of 281 r and θ only and solving $E(r, \theta = 0) = E_k$ and $E(r, \theta = \pi/2) = E_k$, the latter leading 282 to a case known as 'Casus irreducibilis' (Wantzel, 1843). This yields an analytical ex-283 pression for r_{max} which can be validated using the aforementioned simulations: 284

$$r_{max}(\theta) = \frac{r_{max}(\theta = \pi/2) \left(\sqrt{r_{max}(\theta = 0)^2 (2a + r_{max}(\theta = 0))^2 \sin^2(\theta) + r_{max}(\theta = \pi/2)^2 (a + r_{max}(\theta = 0))^2 \cos^2(\theta)}{r_{max}(\theta = 0) (2a + r_{max}(\theta = 0)) \sin^2(\theta) + r_{max}(\theta = \pi/2)^2 \cos^2(\theta)}$$
(10)

with the expressions for $r_{max}(\theta = 0)$ and $r_{max}(\theta = \pi/2)$ derived and given in the supporting information.

Using the now known required expressions, the surface electric field E_0 at the tip of the ellipsoidal hydrometeor required for the onset of a positive corona discharge can be calculated from equation (1) at the known electric field distribution $E(\rho, r, \theta)$ for different values of the relative gas density δ and major and minor axes a and b. As noted by Liu et al. (2012), it is more convenient to, instead of using $\gamma = 1$, define a new quantity:

$$Y \equiv \gamma - 1 = 0. \tag{11}$$

The onset surface electric field E_0 at the tip can now be computed by finding the 293 zero of Y. This cannot be done analytically due to the complexity of the integrals. Moreover, since the integration limits in equation (1) also depend on the unknown E_0 and the 295 integration variables, numerical integration by itself is also not sufficient. However, this 296 numerical integration can be combined with a numerical function that finds the root of 297 an expression, such as the MATLAB function 'fzero', as used by Liu et al. (2012) and 298 this work, or the Mathematica function 'FindRoot'. Substituting the numerical integra-299 tion of equation (11) into the find root function means the numerical integration can be 300 solved even though the integration limits are not numbers. Thus, in combination with 301 this method the model determines the corona onset field E_0 at the hydrometeor tip through 302 equation (11), equivalent to equation (1). 303

After applying the find root function, the found onset field E_0 can be used to evaluate the ionization integral K, given by

$$K = \int_{\rho_0}^{\rho_c} \alpha_{eff}(\rho, E) d\rho, \qquad (12)$$

where $\rho_c = \rho_{ab}(r_{max}, \theta, \phi)$ gives the position of the breakdown field E_k . Exponentiation of K yields the number of electrons produced by an avalanche from the edge of the ionization region to the surface of the hydrometeor. Equation (12) is thus a criterion for the onset of a positive corona discharge with K a threshold value that needs to be reached to enable initiation. It is important to note that the above integration is taken along the field line from the surface of the electrode to the edge of the ionization region, because the avalanche follows the direction of the electric field. Per the convention used by Naidis (2005), the model is set up to output the onset field E_0 , from which the onset voltage V_0 , onset charge Q, and ionization integral Kcan be derived. This order is thus kept in the following results section.

As stated, the used model gives the minimum condition for the onset of a corona 316 discharge. Besides assuming all photons are emitted at the surface, it neglects the pres-317 ence of space charge created in the discharge. Furthermore, the onset criterion is only 318 imposed on the secondary avalanche; further avalanches are assumed to take place when 319 this criterion is satisfied. While these factors generally increase the threshold for corona 320 inception, including its stochastic nature would lower this threshold. The validity of the 321 model depends on the relevant dimensions. For the model to be reliable, the largest pho-322 ton absorption length (r_{max}) should be smaller than the length of the ellipsoid. Other-323 wise, the equilibrium between the ionization coefficients with the local electric field can-324 not be guaranteed. 325

326 **3** Results and discussions

327 328

3.1 The effects of varying aspect ratio and volume of spheroidal hydrometeors on the corona inception criterion

To calculate the required effective ionization coefficient α_{eff} in equation (1), we 329 need to use the air plasma-chemical reactions which are listed in Table 1. All electron 330 impact ionization, excitation, elastic and attachment reactions (except three-body at-331 tachment) that are included in the list were taken from Itikawa database (Itikawa database, 332 www.lxcat.net, retrieved on Sep 15, 2020., n.d.; Itikawa, 2005, 2008). The three-body at-333 tachment with O_2 as the third body was taken from Phelps database (*Phelps database*, 334 www.lxcat.net, retrieved on Sep 15, 2020., n.d.) and scaled to the different δ . Next, the 335 reactions were used as input for BOLSIG+ (Hagelaar & Pitchford, 2005; BOLSIG+ solver 336 ver. Windows 12/2019, n.d.) to calculate the ionization and attachment coefficients. The 337 results are depicted in Figure 3 and the effective ionization coefficient is defined as the 338 subtraction of the attachment coefficient from the ionization coefficient. 339

 Table 1.
 List of plasma-chemical reactions used for calculation the ionization and attachment coefficients.

	Reaction
Elastic	$ \begin{array}{c} e^- + N_2 \rightarrow e^- + N_2 \\ e^- + O_2 \rightarrow e^- + O_2 \end{array} $
Ionization	$ \begin{vmatrix} e^- + N_2 \to 2e^- + N_2^+ \\ e^- + N_2 \to 2e^- + N^+ + N \\ e^- + N_2 \to 3e^- + N^{2+} + N \\ e^- + O_2 \to 2e^- + O_2^+ \\ e^- + O_2 \to 2e^- + O^+ + O \\ e^- + O_2 \to 3e^- + O^{2+} + O \end{vmatrix} $
Attachment	$\begin{vmatrix} e^{-} + O_{2} + O_{2} \to O_{2}^{-} + O_{2} \\ e^{-} + O_{2} \to O^{-} + O \end{vmatrix}$
Excitation	$\begin{vmatrix} e^- + O_2 \to e^- + O_2^* \\ e^- + N_2 \to e^- + N_2^* \end{vmatrix}$

Using the corona inception criterion of equation (1), equation (11) is solved numerically in MATLAB for varying hydrometeor volume $\frac{4}{3}\pi C$, with $C = ab^2$ the volume parameter, and varying aspect ratios b/a. The aspect ratio b/a is considered instead of, for



Figure 3. Reduced attachment (η/N) and ionization (α/N) coefficients as a function of reduced electric field in an N₂ : O₂ = 80:20 mixture at δ =1. The breakdown field is determined where $\alpha - \eta = 0$.

example, the major axis *a*, such that the effects of varying volume and shape can be investigated separately.

The studied hydrometeor geometries have volume parameters of C = 0.01, 0.05,345 and $0.1 \,\mathrm{cm}^3$ and aspect ratios from b/a = 0.01 to 1, where b/a = 1 represents a sphere 346 (a = b). First, positive corona inception is investigated at atmospheric pressure ($\delta =$ 347 1). The onset field E_0 at the tip of the ellipsoid, found directly from solving equation 348 (11), is presented in Figure 4a. It can be seen that E_0 decreases with volume for a fixed 349 aspect ratio. For the smallest hydrometeor, $C = 0.01 \,\mathrm{cm}^3$, the onset field at b/a = 0.045350 is $366 \,\mathrm{kV/cm}$, while for the largest hydrometeor, $C = 0.1 \,\mathrm{cm}^3$, this is $248 \,\mathrm{kV/cm}$. The 351 decrease of the onset field with increasing volume is expected as a larger hydrometeor, 352 simulated as an electrode, provides more surface for photon emission to the photon ab-353 sorption region. Here, it should be noted that in the model it was assumed that all pho-354 tons that lead to photoionization are emitted at the surface. Thus, for a smaller hydrom-355 eteor less photons are emitted and therefore less electrons are produced by photoioniza-356 tion, such that to satisfy the corona onset criterion a larger onset field E_0 is required. 357

From Figure 4a it can also be concluded that for a fixed volume, a sharper ellip-358 soid has a larger onset field E_0 at its tip. Because a sharper ellipsoid has less surface near 359 the photon absorption region, a larger E_0 is needed to meet the inception criterion. In 360 the spherical limit, b/a = 1, the onset field for $C = 0.01 \text{ cm}^3$ is about 17% larger than 361 that for $C = 0.1 \,\mathrm{cm}^3$, and for the much sharper tip at $b/a \approx 0.015$ this difference has 362 increased to about 70%. The onset field thus increases much stronger with sharpness for 363 a smaller hydrometeor, which is expected as a smaller object has more surface area com-364 pared to its volume. It should be noted, however, that Figure 4a does not give the whole 365 story. This onset field is only at the tip of the ellipsoid. Moreover, as charges on a con-366 ductor tend to move away from each other as much as possible on its surface, the elec-367 tric field is enhanced more strongly near a sharper tip. Hence, even though a sharper 368 ellipsoid has a larger E_0 , this does not necessarily mean corona inception from sharper 369

hydrometeors in thunderstorms is less likely. On the contrary, Petersen et al. (2015) ob-370 served sharper hydrometeors promote glow coronae. R. F. Griffiths and Latham (1974) 371 suggested in their paper on coronae from ice hydrometeors that the onset ambient field 372 decreases with increasing combined length of the liquid filament, which was confirmed 373 by Crabb and Latham (1974), who also found that the elongated filament resulting from 374 raindrop collision promotes corona onset. This seems to contradict the decrease of E_0 375 with elongation in Figure 4a, but taking into account the mentioned effect of only con-376 sidering the tip this discrepancy is explained. To draw clearer conclusions, other quan-377 tities such as potential, surface charge and the ionization integral should be considered 378 as well when studying corona inception from an ellipsoid. 379



Figure 4. The a) onset field, and b) onset voltage for positive coronae at the tip of the ellipsoidal hydrometeor for C = 0.01, 0.05, and 0.1 cm^3 for varying aspect ratio b/a at atmospheric pressure. For clarity, two ellipsoid shapes are given at different b/a. Results are compared in the spherical limit (b = a) with Liu et al. (2012).

From the onset field E_0 the onset voltage, or inception voltage, along the major 380 axis (from the tip to infinity) can be calculated by the integration of the electric field. 381 This onset voltage V_0 is shown in Figure 4b. The inception voltage increases with hy-382 drometeor volume. This is also found by Liu et al. (2012) for a spherical electrode. Note 383 that the onset field decreases with volume while the onset voltage increases, which can 384 be quickly understood by looking at the simpler configuration of a sphere, where $V_0 =$ 385 $E_0\rho_0$, with ρ_0 its radius. Figure 4b also shows that the onset voltage is lower for a sharper 386 ellipsoid. For a very sharp tip this difference is less noticeable, and the onset voltage is 387 about 4 kV for the three hydrometeors. In the spherical limit, the largest hydrometeor 388 $(C = 0.1 \,\mathrm{cm}^3)$ requires 30 kV for corona onset, while the smallest hydrometeor $(C = 0.1 \,\mathrm{cm}^3)$ 389 $0.01 \,\mathrm{cm}^3$) requires 16 kV. 390

Besides the onset voltage V_0 , the onset charge Q can also be derived from the on-391 set field E_0 through $E_0 = \frac{Q}{4\pi b^2 \epsilon_0}$. Of course, the onset field is a result of the onset charge, making this the more fundamental parameter. The onset charge, which is the total charge 392 393 on the electrode surface, is depicted in Figure 5. A size-dependent optimum aspect ra-394 tio b/a is observed at which the onset charge is lowest. While a sharper ellipsoid has a 395 higher onset field and thus requires more charge at the tip to reach this E_0 , a larger frac-396 tion of the total charge is collected at its tip because of the optimization of charge sep-397 aration. In simulations of corona inception from hydrometeors modelled as dielectrics 398 in an external electric field, Dubinova et al. (2015) also found a size-dependent aspect 399

ratio for which the onset background field is minimum. From Figure 5 the range of onset charge for hydrometeors with volumes between 0.042 cm^3 and 0.42 cm^3 is found to

 $_{402}$ be 2367 pC to 15,467 pC at atmospheric pressure.



Figure 5. The onset charge for positive coronae at the tip of the ellipsoidal hydrometeor for $C = 0.01, 0.05, \text{ and } 0.1 \text{ cm}^3$ for varying aspect ratio b/a at atmospheric pressure.

Finally, the ionization integral K along the major axis can be calculated from the 403 onset field as well, through equation (12). The result is presented in Figure 6. At a fixed 404 volume, the ionization integral decreases with b/a, meaning that less electrons are re-405 quired in an avalanche from the edge of the photon absorption region to the electrode 406 surface. To interpret these results the dependence of the photon absorption area and length 407 on the electrode dimensions are studied in COMSOL for some data points, of which the 408 results are given in Table 2. From this data it can be concluded that for a fixed aspect 409 ratio, a smaller electrode has a smaller photon absorption area and length, as does a sharper 410 electrode for a fixed volume. However, for a very sharp electrode the photon absorption 411 area and length are approximately equal, as can be seen for b/a = 0.014 in Table 2. 412

As an ellipsoid with smaller b/a has a smaller photon absorption region, photons 413 are absorbed closer to the electrode compared to its size, such that stronger avalanches 414 are required to satisfy the inception criterion. A similar argument was made by Naidis 415 (2005) to explain the ionization integral dependence on radius for a spherical and cylin-416 drical electrode. Comparing the data points for different volumes, two regions can be dis-417 cerned in Figure 6, separated by a cross-over point around b/a = 0.55. At large b/a, 418 where K drops below 14, the largest ellipsoid has the largest value for the ionization in-419 tegral, again because photons are absorbed closer to the electrode with respect to its size. 420 When K increases above 14 for decreasing b/a it is observed that the smallest ellipsoid 421 has the largest K value. An explanation for this could be that when b/a becomes small 422 enough, the photon absorption region becomes so small that its absolute size instead of 423 its relative size determines the value of the ionization integral. Stronger avalanches are 424 then required for a smaller electrode. For very small b/a the data points for different vol-425

Volume parameter $C \ (\mathrm{cm}^3)$	Aspect ratio b/a	Semi axis a (cm)	$\begin{vmatrix} \text{Semi axis} \\ b \ (\text{cm}) \end{vmatrix}$	$\begin{vmatrix} Area \\ (mm^2) \end{vmatrix}$	$\begin{array}{c c} \text{Length} \\ r_{max} \ (\text{mm}) \end{array}$
0.1	0.014	7.93	0.11	0.21	0.12
0.1	0.045	3.68	0.16	0.33	0.22
0.1	0.141	1.70	0.24	1.02	0.47
0.1	0.447	0.79	0.35	6.20	1.02
0.05	0.014	6.29	0.08	0.25	0.11
0.05	0.045	2.92	0.13	0.28	0.20
0.05	0.141	1.35	0.19	0.84	0.41
0.05	0.447	0.62	0.28	4.91	0.88
0.01	0.014	3.68	0.05	0.25	0.09
0.01	0.045	1.70	0.07	0.28	0.16
0.01	0.141	0.79	0.11	0.58	0.31
0.01	0.447	0.36	0.16	2.71	0.61

 Table 2.
 Photon absorption area and length for various ellipsoidal electrode aspect ratios and volumes.

umes appear to converge again. A likely explanation is that when the ellipsoid becomes very sharp, a photon is absorbed so close to the tip such that the total volume of the electrode has no effect; only the sharpness of the tip determines the value of the ionization integral. This is supported by the photon absorption area being approximately equal for the different volumes at b/a = 0.014 in Table 2.

431

3.2 Variation of the corona inception criterion with pressure

Next, the dependence of corona onset from an ellipsoidal hydrometeor on the am-432 bient pressure is investigated by varying the relative gas density δ . More specifically, the 433 values $\delta = 10, 1, \text{ and } 0.1, \text{ analogous to the works by Naidis (2005) and Liu et al. (2012),}$ 434 and $\delta = 0.5$, representative for thundercloud altitudes, are considered. The volume pa-435 rameter is fixed at $C = 0.01 \text{ cm}^3$ and the aspect ratio b/a varies again from b/a = 0.01436 to 1. The results for the onset field at the hydrometeor tip are shown in Figure 7a. As 437 expected, a higher pressure leads to a higher onset field E_0 . As explained by Liu et al. 438 (2012), at a higher pressure more of the excited nitrogen molecules responsible for emit-439 ting the ionizing photons are quenched, leading to a lower photon production such that 440 a higher field is required. To briefly examine how the results are affected by the afore-441 mentioned photoionization outside of the ionization region, the computations are redone 442 with an integration upper limit of $10r_{max}$ instead of r_{max} . It follows that the difference 443 in outcome is generally well below 1%, only rising above 5% for $\delta = 0.5$ for the small-444 est volume parameter $C = 0.01 \,\mathrm{cm}^3$, and only for very blunt tips, nearing $b/a \approx 1$. 445 Neglecting this stochastic effect thus seems justified. 446

Similarly, the onset charge increases with pressure, as depicted in Figure 7b. For 447 $C = 0.01 \text{ cm}^3$ the onset charge is between 547 pC and 2400 pC for $\delta = 0.1$ and between 448 1500 pC and 3100 pC for $\delta = 0.5$. For $\delta = 0.5$ the onset charge is minimum at an as-449 pect ratio of approximately 0.1. A pressure above atmospheric pressure, at $\delta = 10$, is 450 not representative for thunderstorms, but is included for completeness. Again, the ion-451 ization integral K can be calculated from the onset field and is plotted in Figure 7c. The 452 pressure dependence can be explained as before; due to increased quenching of excited 453 nitrogen molecules at higher pressures the photon production is lowered. Therefore, stronger 454 avalanches are required to satisfy the inception criterion. 455



Figure 6. The ionization integral along the major axis for positive corona onset at the tip of the ellipsoidal hydrometeor for C = 0.01, 0.05, and 0.1 cm^3 for varying aspect ratio b/a at atmospheric pressure.

456 457

3.3 Dependence of the derived ambient electric field on the aspect ratio for thundercloud pressure

The ambient field E_{bq} required for corona onset can be derived from the onset field 458 E_0 at the hydrometeor tip. This is done by simulating the hydrometeor as a conductor 459 without surface charge in an ambient electric field in COMSOL, and increasing this field 460 until the determined E_0 is obtained at the tip. The relative gas density of $\delta = 0.5$ and 461 the most representative size of $C = 0.01 \,\mathrm{cm}^3$ (as hydrometeors are generally found in 462 the millimeter range (Weinheimer et al., 1991; Gardiner et al., 1985)) are chosen. The 463 results are presented in Figure 8. It is seen that the required background field E_{bq} is be-464 low the breakdown field E_k , between 0.07 E_k and 0.8 E_k , and is lowest for the sharpest 465 hydrometeor tips. 466

467 4 Summary, Conclusions and Outlook

The corona inception criterion set up by Naidis (2005) is applied through numer-468 ical simulations to spheroidal electrodes of various dimensions at different pressures. By 469 doing so, the theoretical onset of a positive corona from an ellipsoidal hydrometeor is stud-470 ied. It is found that the onset electric field at the hydrometeor tip decreases with hy-471 drometeor volume and tip bluntness, as the hydrometeor surface near the photon absorp-472 tion region increases with these factors. Moreover, the onset field increases with pres-473 sure due to the quenching of excited nitrogen molecules. For a hydrometeor of $0.042 \,\mathrm{cm}^3$ 474 volume ($C = ab^2 = 0.01 \,\mathrm{cm}^3$) and thundercloud pressure ($\delta = 0.5$), the onset field at 475 the tip varies approximately from $2.4E_k$ (limiting case sphere) to $70E_k$ (sharpest case 476 considered), where E_k is the breakdown field. However, the onset field at the tip is not 477



Figure 7. The a) onset field, b) onset charge, c) ionization integral for positive coronae at the tip of the ellipsoidal hydrometeor for $\delta = 10, 1, 0.5$, and 0.1 for varying aspect ratio b/a at a fixed volume parameter $C = 0.01 \text{ cm}^3$. Results are compared in the spherical limit (b = a) with Liu et al. (2012).

deemed representative for the likeliness of corona onset as it does not provide information on the entire surface. These values were also obtained without the inclusion of an ambient electric field. Instead, the onset potential difference, V_0 , can provide a more realistic picture for corona onset, since it can be compared with experimental results. As we can observe, sharper hydrometeors need a lower voltage to initiate a discharge.

Another way to better predict the feasibility of corona onset in thundercloud elec-483 tric fields is by the derivation of the required ambient electric field E_{bg} from the com-484 puted onset field E_0 . This yields values between $0.07 E_k$ and $0.8 E_k$ for semi-axes aspect 485 ratios between 0.01 and 1. Hence, the found ambient electric field is well below the break-486 down field. It should be noted that this derived ambient field is neither an upper limit 487 nor lower limit on the field required for onset. While the model gives a minimum con-488 dition for corona onset, the found E_0 would be lower if an ambient electric field was in-489 cluded in the model in the first place. Thus, the used assumptions and simplifications 490 should be kept in mind when interpreting these results. However, the ambient field be-491 ing significantly lower than the breakdown field for representative shapes is very promis-492 ing. 493



Figure 8. The required background electric field, E_{bg} , to have an enhanced electric field of E_0 at the tip of the hydrometeor ($C = 0.01 \,\mathrm{cm}^3$). The values are calculated at $\delta = 0.5$, where $E_k = 17.9 \,\mathrm{kV/cm}$.

Whereas the onset field only provides information on the hydrometeor tip and the 494 onset voltage only on the major axis, the onset charge is the total charge on the hydrom-495 eteor surface. This onset charge reveals, depending on hydrometeor volume, an optimal 496 semi-axes aspect ratio of the ellipsoidal hydrometeor for which the least amount of charge 497 is required for positive corona onset. The minimum in the onset charge curve is caused 498 by the interplay between required onset field and geometry; a sharper hydrometeor has 499 a larger onset field at the tip and thus requires more charge at the tip, but a larger part 500 of its total charge is located at the tip. As this optimum was not found for the onset field 501 or onset voltage, this suggests that considering only the major axis, which is often done 502 in models for simplification, may not be sufficient when investigating corona onset con-503 ditions. Interestingly, in their study on lightning inception from hydrometeors, simulated 504 as dielectrics in an ambient electric field, Dubinova et al. (2015) obtain a length-dependent 505 optimum aspect ratio of the hydrometeor that requires the lowest ambient field for dis-506 charge inception. In addition, the obtained results can be compared to measured pre-507 cipitation charges. Generally the hydrometeor charge is measured below 400 pC (Marshall 508 & Winn, 1982). For the volume closest to the measured precipitation, $C = 0.01 \,\mathrm{cm}^3$, 509 and a relative gas density of $\delta = 0.5$ the onset charge is found to be between and 1500 pC 510 and 3100 pC. However, these charges were measured for estimated hydrometeor diam-511 eters between 1 and $3 \,\mathrm{mm}$, whereas in the spherical limit the simulated hydrometeors 512 have diameters between 4 mm and 9 mm. In their simulations on spherical hydromete-513 ors using the same corona inception criterion as this paper, Liu et al. have shown that 514 the onset charge varies over several orders of magnitude in the estimated size range of 515 hydrometers. For spherical hydrometeors of 9 mm diameter, the simulated onset charge 516 was near ten times larger than for a 3 mm diameter. With this size correction (roughly 517 a factor 10) onset charge values are close to the hydrometeors charges obtained from in 518 situ measurements. Moreover, the considered configuration is an isolated hydrometeor 519 with zero ambient field. Interaction between hydrometeors (see for example (Rutjes et 520 al., 2019)) and a non-zero ambient field would lower the amount of charge required for 521

⁵²² corona inception, which explains why the found onset charge is higher than expected from ⁵²³ in-situ measurements.

Besides the onset charge, the ionization integral K also displays different behaviour 524 in different b/a regions. For hydrometeors with very blunt tips, close to a spherical shape, 525 a larger hydrometeor has a larger K value for onset, as photons are absorbed closer to 526 the hydrometeor with respect to its size. However, for hydrometeors with sufficiently sharp 527 tips, the absolute size of the photon absorption region seems to be more important than 528 its relative size, such that a smaller hydrometeor has a larger value of the ionization in-529 530 tegral. For any ellipsoidal shape, the value of the ionization integral is larger at higher pressures, because of the quenching of excited nitrogen molecules, which leads to less-531 ened photon emission and therefore a need for stronger avalanches for corona onset. 532

To investigate the validity of the results, the approximations and assumptions of 533 the model should be evaluated. Firstly, the distance r_{max} from the tip to the edge of the 534 photon absorption region should be smaller than the hydrometeor length. Using the ex-535 pression derived in the supporting information, it is found that for all data points the 536 maximum ratio of this distance to length is $r_{max}/L = 0.2$, meaning this condition for 537 the model to hold is satisfied. Furthermore, the presence of space charges is ignored in 538 the model, leading to an overestimation of the electric field magnitude. When the ion-539 ization integral K, or equivalently number of electrons in the avalanche, is large enough, 540 the perturbation of the electric field by the space charge becomes comparable to the mag-541 nitude of the electric field itself, such that space charge cannot be neglected. This is ac-542 companied by the transformation of the avalanche into a streamer. In literature, it is of-543 ten taken that K should be below 14-22 (Naidis, 2005; Raizer, 1991c) for the perturba-544 tion of the electric field by space charge to be neglected. In the results, the value of K545 is below this threshold for sufficiently blunt hydrometeors. Near b/a = 0.555 in Fig-546 ure 6, which is also the cross-over point of the three curves, this value rises above 14. Hence, 547 for sharper ellipsoids possibly more physics should be added to the model to obtain more 548 accurate results. 549

In the model of the current work, it is assumed that there are sufficient free elec-550 trons present for the primary electron avalanche. To be able to draw conclusions on whether 551 corona onset is possible in thunderclouds, it should be considered how these free elec-552 trons are supplied, and if this supply is large enough. The source of free electrons for light-553 ning initiation is a widely researched subject, see for example (Dubinova, 2016; Rutjes 554 et al., 2019). At least one primary electron is required for discharge initiation, but more 555 electrons lower the inception threshold. When more electrons are available, the require-556 ments on the other factors, such as the aspect ratio and volume of the hydrometeor or 557 amplitude of the ambient field, will be softened. 558

From the above considerations, it can be concluded that lightning initiation from 559 a spheroidal hydrometeor is feasible. While the onset field at the tip of the charged hy-560 drometeor without ambient field was not found to be below the breakdown field in the 561 considered configuration, the derived onset ambient electric field for the uncharged hy-562 drometeor is lower than this threshold. Further enhancement could be provided by the 563 interaction between hydrometeors. For representative dimensions and pressures, the amount 564 of charge required for corona onset provided by the model is comparable to measured 565 hydrometeor charges. Whether sharper hydrometeors promote lightning onset is a delicate discussion, which depends on which parameters are considered. From our results, 567 it appears that only considering the major axis is not sufficient to reach conclusions on 568 this matter. To further investigate the corona onset from hydrometeors using this model, 569 570 more physics could be included. Most importantly, the thundercloud ambient electric field could be added to the model. Furthermore, the method can be applied to a hydrom-571 eteor cluster. The role of humidity, which was studied by Liu et al. (2012) for spherical 572 hydrometeors, and the low-temperature environment can also be investigated. Finally, 573 the model could be adjusted to account for space charge effects. 574

575 Appendix A Derivation of the distance ρ_0

To find the distance ρ_0 from the major axis to the surface of the ellipsoid along the surface electric field direction, the equation defining the ellipsoid (with the origin at the tip of the ellipsoid)

$$\frac{x^2}{b^2} + \frac{y^2}{b^2} + \frac{(z+a)^2}{a^2} = 1$$
(A1)

is reformulated in spherical coordinates, which yields

$$\frac{r^2 \sin^2(\theta)}{b^2} + \frac{(a + r \cos(\theta))^2}{a^2} = 1.$$
 (A2)

Solving equation (A2) for r gives two solutions, valid separately for $\theta \le \pi/2$ and $\theta > \pi/2$, namely

$$r = \begin{cases} 0 & \theta \le \pi/2\\ \frac{2ab^2 \cos(\theta)}{(a^2 - b^2) \cos^2(\theta) - a^2} & \theta > \pi/2, \end{cases}$$
(A3)

as the range $\theta \leq \pi/2$ is the ionization region, which only encompasses the tip, r =0, of the ellipsoidal surface (see also Figure 2). Substituting these solutions into the expression for ρ_{ab} (equation (4) and supporting information), thus constraining ρ_{ab} to the surface of the ellipsoid, gives

$$\rho_{0} = \begin{cases} \frac{(a - \sqrt{a^{2} - b^{2}})\sqrt{2a(\sqrt{a^{2} - b^{2}} + a) - b^{2}}}{a} & \theta \leq \pi/2 \\ \frac{4(2a(\sqrt{a^{2} - b^{2}} - a) + b^{2} + p_{2})\sqrt{\frac{\left(\frac{4ab^{2}(2\sqrt{a^{2} - b^{2}})\cos^{2}(\theta)}{(a^{2} - b^{2})\cos^{2}(\theta) - a^{2}} + 2a\left(\sqrt{a^{2} - b^{2}} - b^{2} - p_{2}\right)\left(\frac{4ab^{2}\sqrt{a^{2} - b^{2}}\cos^{2}(\theta)}{(a^{2} - b^{2})\cos^{2}(\theta) - a^{2}} + b^{2} + \frac{p_{1}}{8} - p_{2}\right)}{p_{1}}}{2a(\sqrt{a^{2} - b^{2}} - a) + b^{2} - \frac{p_{1}}{8} + p_{2}} & \theta > \pi/2, \end{cases}$$

586 with

579

$$p_1 = 8\rho_1\rho_2 \left(r = \frac{2ab^2\cos(\theta)}{(a^2 - b^2)\cos^2(\theta) - a^2}\right)$$

587

$$p_2 = -\rho_1^2 \left(r = \frac{2ab^2 \cos(\theta)}{(a^2 - b^2)\cos^2(\theta) - a^2} \right) - 2a \left(\sqrt{a^2 - b^2} - a \right) - b^2$$

588 Acknowledgments

and

We would like to thank Ute Ebert from Centrum Wiskunde & Informatica (CWI), who provided insight and expertise that greatly assisted the research. This project has received funding from the European Union's Horizon 2020 research and innovation program under the Marie Sklodowska-Curie grant agreement 722337.

The data that support the findings of this study are available from the corresponding author, S. A. Peeters, upon reasonable request. The MATLAB scripts used to generate the data for this paper and the full derivations of the indicated expressions are included in the supporting information.

597	References
	10010101000

- Babich, L., Bochkov, E., & Neubert, T. (2017). The role of charged ice hydrom eteors in lightning initiation. Journal of Atmospheric and Solar-Terrestrial
 Physics, 154, 43 46. Retrieved from http://www.sciencedirect.com/
 science/article/pii/S1364682616304564 doi: https://doi.org/10.1016/
 j.jastp.2016.12.010
- Babich, L. P., Bochkov, E. I., Kutsyk, I. M., Neubert, T., & Chanrion, O. (2016).
 Positive streamer initiation from raindrops in thundercloud fields. Journal of Geophysical Research: Atmospheres, 121(11), 6393-6403. Retrieved from https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1002/
 2016JD024901 doi: https://doi.org/10.1002/2016JD024901
- Bandel, H. W. (1951, Oct). Point-to-plane corona in dry air. *Phys. Rev.*, 84, 92–99.
 Retrieved from https://link.aps.org/doi/10.1103/PhysRev.84.92 doi: 10
 .1103/PhysRev.84.92
- Bolsig+ solver ver. windows 12/2019. (n.d.). Retrieved from http://www.bolsig 12 .laplace.univ-tlse.fr/
- ⁶¹³ Coquillat, S., Chauzy, S., & Médale, J.-C. (1995). Microdischarges between ice
 ⁶¹⁴ particles. Journal of Geophysical Research: Atmospheres, 100(D7), 14327 ⁶¹⁵ 14334. Retrieved from https://agupubs.onlinelibrary.wiley.com/doi/
 ⁶¹⁶ abs/10.1029/95JD00986 doi: https://doi.org/10.1029/95JD00986
- 617Crabb, J. A., & Latham, J.(1974).Corona from colliding drops as a possi-618ble mechanism for the triggering of lightning.Quarterly Journal of the619Royal Meteorological Society, 100(424), 191-202.Retrieved from https://620rmets.onlinelibrary.wiley.com/doi/abs/10.1002/qj.49710042406doi:621https://doi.org/10.1002/qj.49710042406
- Curtright, T. L., Cao, Z., Huang, S., Sarmiento, J. S., Subedi, S., Tarrence, D. A., &
 Thapaliya, T. R. (2020, apr). Charge densities for conducting ellipsoids. *European Journal of Physics*, 41(3), 035204. Retrieved from https://doi.org/
 10.1088%2F1361-6404%2Fab806a doi: 10.1088/1361-6404/ab806a
- 626
 D'Alessandro, F., & Berger, G. (1999, oct).
 Laboratory studies of corona emis

 627
 sions from air terminals.
 Journal of Physics D: Applied Physics, 32(21), 2785

 628
 2790.
 Retrieved from https://doi.org/10.1088%2F0022-3727%2F32%2F21%

 629
 2F311
 doi: 10.1088/0022-3727/32/21/311
- Dubinova, A. (2016). *Modeling of streamer discharges near dielectrics* (Unpublished doctoral dissertation). Department of Applied Physics. (Proefschrift)
- Dubinova, A., Rutjes, C., Ebert, U., Buitink, S., Scholten, O., & Trinh, G. T. N.
 (2015, Jun). Prediction of lightning inception by large ice particles and
 extensive air showers. *Phys. Rev. Lett.*, 115, 015002. Retrieved from
 https://link.aps.org/doi/10.1103/PhysRevLett.115.015002 doi:
 10.1103/PhysRevLett.115.015002
- ⁶³⁷ Dwyer, J. R., & Uman, M. A. (2014). The physics of lightning. *Physics Reports*, ⁶³⁸ 534(4), 147–241.
- Gardiner, B., Lamb, D., Pitter, R. L., Hallett, J., & Saunders, C. P. R. (1985). Measurements of initial potential gradient and particle charges in a montana summer thunderstorm. Journal of Geophysical Research: Atmospheres, 90(D4), 642 6079-6086. Retrieved from https://agupubs.onlinelibrary.wiley.com/ 643 doi/abs/10.1029/JD090iD04p06079 doi: https://doi.org/10.1029/
- Griffiths, R. (1975). The initiation of corona discharges from charged ice particles in
 a strong electric field. Journal of Electrostatics, 1(1), 3 13. Retrieved from
 http://www.sciencedirect.com/science/article/pii/0304388675900030
 doi: https://doi.org/10.1016/0304-3886(75)90003-0
- Griffiths, R. F., & Latham, J. (1974). Electrical corona from ice hydrometeors.
 Quarterly Journal of the Royal Meteorological Society, 100(424), 163-180. Re trieved from https://rmets.onlinelibrary.wiley.com/doi/abs/10.1002/

	ai 10710012101 doi: https://doi.org/10.1002/ci.10710042404
652	$q_{j}.49710042404$ doi: https://doi.org/10.1002/ $q_{j}.49710042404$
653	mann equation to obtain electron transport coefficients and rate coeffi-
655	cients for fluid models Plasma Sources Sci Technol 1/(4) 722–733 doi:
656	10 1088/0963-0252/14/4/011
657	Itikawa V (2005 December) Cross Sections for Electron Collisions with Nitro-
658	gen Molecules Journal of Physical and Chemical Reference Data 35(1) 31–
659	53. (Publisher: American Institute of Physics) doi: 10.1063/1.1937426
660	Itikawa V (2008 December) Cross Sections for Electron Collisions with Oxy-
661	gen Molecules. Journal of Physical and Chemical Reference Data, 38(1), 1–20.
662	(Publisher: American Institute of Physics) doi: 10.1063/1.3025886
663	Itikawa database, www.lxcat.net, retrieved on sep 15, 2020. (n.d.).
664	Kip, A. F. (1938, Jul). Positive-point-to-plane discharge in air at atmospheric pres-
665	sure. Phys. Rev., 54, 139-146. Retrieved from https://link.aps.org/doi/
666	10.1103/PhysRev.54.139 doi: 10.1103/PhysRev.54.139
667	Köhn, C., & Ebert, U. (2015). Calculation of beams of positrons, neutrons, and
668	protons associated with terrestrial gamma ray flashes. Journal of Geophysical
669	Research: Atmospheres, 120(4), 1620-1635. Retrieved from https://agupubs
670	.onlinelibrary.wiley.com/doi/abs/10.1002/2014JD022229 doi: https://
671	doi.org/10.1002/2014JD022229
672	Liu, N., Dwyer, J. R., & Rassoul, H. K. (2012). Effects of pressure and humidity
673	on positive corona inception from thundercloud hydrometeors. Journal of
674	Atmospheric and Solar-Terrestrial Physics, 80, 179 - 186. Retrieved from
675	http://www.sciencedirect.com/science/article/pii/S1364682612000260
676	doi: https://doi.org/10.1016/j.jastp.2012.01.012
677	Loeb, L. B. (1966). The mechanisms of stepped and dart leaders in cloud-to-ground
678	lightning strokes. Journal of Geophysical Research (1896-1977), 71 (20), 4711-
679	4721. Retrieved from https://agupubs.onlinelibrary.wiley.com/doi/abs/
680	10.1029/JZ071i020p04711 doi: https://doi.org/10.1029/JZ071i020p04711
681	MacGorman, D., MacGorman, R., Rust, W., & Rust, W. (1998). The electrical
682	nature of storms. Oxford University Press. Retrieved from https://books
683	.google.nl/books:id=_NbHNj/KJecU
684	Marshall, T. C., McCarthy, M. P., & Rust, W. D. (1995). Electric field mag-
685	nitudes and lightning initiation in thunderstorms. Journal of Geophysical $P_{\text{constrained}}$ the second from https:///
686	Research: Atmospheres, 100(D4), 7097-7103. Retrieved from https://
687	agupubs.online11b1a1y.w11ey.com/do1/abs/10.1029/95JD00020 doi.https://doi.org/10.1020/05JD00020
688	Marshall T C & Winn W P (1082) Measurements of charged precipitation in
689	a new mexico thunderstorm: lower positive charge centers
601	<i>a new mexico uninderstorm. Tower positive enarge centers:</i> $50a/mar of Geophysical Research: Oceans 87(C9) 7141-7157 Betrieved from https://$
692	agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/JC087iC09p07141
693	doi: https://doi.org/10.1029/JC087iC09p07141
694	Mazur, V. (2016). <i>Principles of lightning physics</i> . IOP Publishing. Retrieved from
695	http://dx.doi.org/10.1088/978-0-7503-1152-6 doi: 10.1088/978-0-7503
696	-1152-6
697	Naidis, G. V. (2005, jun). Conditions for inception of positive corona discharges
698	in air. Journal of Physics D: Applied Physics, 38(13), 2211–2214. Retrieved
699	from https://doi.org/10.1088\%2F0022-3727\%2F38\%2F13\%2F020 doi:
700	10.1088/0022- $3727/38/13/020$
701	Nasser, E., & Heiszler, M. (1974, August). Mathematical-physical model of the
702	streamer in nonuniform fields. Journal of Applied Physics, 45(8), 3396-3401.
703	doi: $10.1063/1.1663791$
704	Petersen, D., Bailey, M., Beasley, W. H., & Hallett, J. (2008). A brief review
705	of the problem of lightning initiation and a hypothesis of initial lightning
706	leader formation. Journal of Geophysical Research: Atmospheres, 113(D17).

707	Retrieved from https://agupubs.onlinelibrary.wiley.com/doi/abs/
708	10.1029/2007JD009036 doi: https://doi.org/10.1029/2007JD009036
709	Petersen, D., Bailey, M., Hallett, J., & Beasley, W. (2015). Laboratory investigation
710	of corona initiation by ice crystals and its importance to lightning. <i>Quarterly</i>
711	Journal of the Royal Meteorological Society, 141(689), 1283-1293. Retrieved
712	from https://rmets.onlinelibrary.wiley.com/doi/abs/10.1002/qj.2436
713	doi: https://doi.org/10.1002/qj.2436
714	Petersen, D., Bailey, M., Hallett, J., & Beasley, W. H. (2006). Laboratory in-
715	vestigation of positive streamer discharges from simulated ice hydrometeors.
716	Quarterly Journal of the Royal Meteorological Society, 132(615), 263-273. Re-
717	trieved from https://rmets.onlinelibrary.wiley.com/doi/abs/10.1256/
718	qj.05.32 doi: https://doi.org/10.1256/qj.05.32
719	Phelps, C. (1974). Positive streamer system intensification and its possible role
720	in lightning initiation. Journal of Atmospheric and Terrestrial Physics,
721	36(1), 103 - 111. Retrieved from http://www.sciencedirect.com/
722	science/article/pii/0021916974900701 doi: https://doi.org/10.1016/
723	0021-9169(74)90070-1
724	Phelps, C. T., & Griffiths, R. F. (1976). Dependence of positive corona streamer
725	propagation on air pressure and water vapor content. Journal of Applied
726	<i>Physics</i> , 47(7), 2929-2934. Retrieved from https://doi.org/10.1063/
727	1.323084 doi: 10.1063/1.323084
728	Phelps database, www.lxcat.net, retrieved on sep 15, 2020. (n.d.).
729	Raizer, Y. P. (1991a). Gas discharge physics. Fizika gazovogo razryada. Berlin:
730	Springer. Retrieved from https://cds.cern.ch/record/264408 (Translated
731	from Russian by Vitaly I Kisin)
732	Raizer, Y. P. (1991b). Gas discharge physics. Fizika gazovogo razryada. Berlin:
733	Springer. Retrieved from https://cds.cern.ch/record/264408 (Translated
734	from Russian by Vitaly I Kisin)
735	Raizer, Y. P. (1991c). Gas discharge physics. Fizika gazovogo razryada. Berlin:
736	Springer. Retrieved from https://cds.cern.ch/record/264408 (Translated
737	from Russian by Vitaly I Kisin)
738	Riousset, J. A., Nag, A., & Palotai, C. (2020). Scaling of conventional breakdown
739	threshold: Impact for predictions of lightning and tles on earth, venus, and
740	mars. <i>Icarus</i> , 338, 113506. Retrieved from http://www.sciencedirect.com/
741	science/article/pii/S0019103519305688 doi: https://doi.org/10.1016/
742	J.icarus.2019.113506
743	Rison, W., Krehbiel, P. R., Stock, M. G., Edens, H. E., Shao, XM., Thomas,
744	R. J., Zhang, Y. (2016). Observations of narrow bipolar events reveal
745	how lightning is initiated in thunderstorms. Nature communications, 7,
746	10/21. Retrieved from https://europepmc.org/articles/PMC4/56383
747	$\frac{\text{doi: 10.1038/ncomms10/21}}{\text{Detting } C_{1}(2018)} = M_{1} L_{1} $
748	(Unpublished doctored discontation) Department of Applied Direction (Uppublished doctored discontation)
749	(Unpublished doctoral dissertation). Department of Applied Physics. (Proef-
750	Schrift) Dertier C. Ehert H. Deritinh, C. Schelter, O. & Trinh, T. N. C. (2010). Con-
751	Rutjes, C., Ebert, U., Bultink, S., Schölten, O., & Irinn, I. N. G. (2019). Gen-
752	contion problem including parrow bipolar events
(53	cal Research: Atmospheres 19/(13) 7955 7960 Rotriouod from https://
755	agunubs onlinelibrary wiley $com/doi/abs/10.1020/2018 ID020040 doi:$
756	401/205/201000029040 (01)
100	Schumann W (1023 June) Über das minimum der durchbruchfeldetärke
758	bei kugelektroden Archiv f Elektrotechnik 593-608 Retrieved from
759	https://doi.org/10.1007/BF01656766
760	Stolzenburg M & Marshall T C (2009) Electric field and charge structure
761	in lightning-producing clouds. In H D Betz U Schumann & P Laroche

762	(Eds.), Lightning: Principles, instruments and applications: Review of mod-
763	ern lightning research (pp. 57–82). Dordrecht: Springer Netherlands. Re-
764	trieved from https://doi.org/10.1007/978-1-4020-9079-0_3 doi:
765	10.1007/978-1-4020-9079-0_3
766	Wantzel, L. (1843). Classification des nombres incommensurables d'origine
767	algébrique. Nouvelles annales de mathématiques : journal des candidats
768	aux écoles polytechnique et normale, 1e série, 2, 117-127. Retrieved from
769	http://www.numdam.org/item/NAM_1843_1_2117_1
770	Waters, R. T., & Stark, W. B. (1975, mar). Characteristics of the stabilized glow
771	discharge in air. Journal of Physics D: Applied Physics, 8(4), 416–426. Re-
772	trieved from https://doi.org/10.1088%2F0022-3727%2F8%2F4%2F014 doi:
773	10.1088/0022- $3727/8/4/014$
774	Weinheimer, A. J., Dye, J. E., Breed, D. W., Spowart, M. P., Parrish, J. L., Hoglin,
775	T. L., & Marshall, T. C. (1991). Simultaneous measurements of the charge,
776	size, and shape of hydrometeors in an electrified cloud. Journal of Geo-
777	physical Research: Atmospheres, $96(D11)$, 20809-20829. Retrieved from
778	https://agupubs.onlinelibrary.wiley.com/doi/abs/10.1029/91JD02262
779	doi: $https://doi.org/10.1029/91JD02262$
780	Zhelezniak, M. B., Mnatsakanian, A. K., & Sizykh, S. V. (1982, November). Pho-
781	toionization of nitrogen and oxygen mixtures by radiation from a gas dis-
782	charge. High Temperature Science, $20(3)$, 357-362.

JOURNAL OF GEOPHYSICAL RESEARCH

@AGUPUBLICATIONS

Supporting Information for

A Model for Positive Corona Inception from Charged Ellipsoidal Thundercloud Hydrometeors

S. A. Peeters¹, S. Mirpour¹, C. Köhn², S. Nijdam¹

¹Eindhoven University of Technology, Department of Applied Physics, The Netherlands

 $^2\mathrm{DTU}$ Space, National Space Institute, Technical University of Denmark, Denmark

Corresponding author: S. A. Peeters, (s.a.peeters@student.tue.nl)

Contents of this file

- 1. Text S1 to S3 $\,$
- 2. Figures S1 to S4

Additional Supporting Information (Files uploaded separately)

- 1. effective_ ionization_ coefficient.m
- 2. onset_ field.m
- 3. ionization_ integral.m
- 4. onset_ voltage.m
- 5. Towsend_coeff_delta_Extended.xlsx

Introduction The supporting information included in this document pertains to the calculations used for solving the corona onset criterion for the onset surface field of the charged hydrometeor. While these full derivations are not necessary in the paper and can in principle be done by the reader, providing them supports reproduction of the study. Specifically, derivations for the distance ρ_{ab} to the point of photon absorption (Text S1), the electric field $E(\rho, r, \theta)$ of a conducting ellipsoid (Text S2), and the distance r_{max} between the tip of the ellipsoid and the position of the breakdown field E_k (Text S3) are presented. The calculations are supported by Figures S1 to S4. The MATLAB code used to implement the model is given as additional supporting information and uploaded separately. Comments are provided in the MATLAB scripts for clarity and an Excel file is also included to be able to run these scripts.

Text S1. Derivation of the distance ρ_{ab}

The distance ρ_{ab} to the point of photon absorption, which will now be referred to as the 'observation point', is in the direction of the electric field. This direction is given by the bisector of the two straight lines from the focal points of the spheroid to the observation point (Curtright et al., 2020). The bisector ρ_{ab} is depicted in Figure S1.

Since the origin of the spherical coordinate system in Figure S1 is placed at the tip of the ellipsoid on the positive z-axis, the lines from the focal points to the observation point given by $\rho^2 = x^2 + y^2 + z^2$ become in spherical coordinates:

$$\rho_1^2 = (r\sin\theta)^2 + (a - d + r\cos\theta)^2 = r^2 + (a - d)^2 + 2(a - d)r\cos\theta,$$
(1)

$$\rho_2^2 = (r\sin\theta)^2 + (a + d + r\cos\theta)^2 = r^2 + (a + d)^2 + 2(a + d)r\cos\theta,$$
with coordinates as given in Figure S1 and the linear eccentricity $d^2 = a^2 - b^2.$

Using the law of cosines, the angle 2β between these two lines ρ_1 and ρ_2 is given by:

$$\cos(2\beta) = \frac{1}{2\rho_1\rho_2} \left(\rho_1^2 + \rho_2^2 - 4d^2\right).$$
(2)

Using the trigonometric identity $\cos(2\beta) = 2\cos^2\beta - 1$ the cosine of the angle between ρ_{ab} and ρ_1 is found:

:

$$\cos\left(\beta\right) = \sqrt{\frac{1}{2} + \frac{1}{4\rho_1\rho_2}\left(\rho_1^2 + \rho_2^2 - 4d^2\right)}.$$
(3)

The acute angle ϕ between ρ_1 and the z-axis follows from:

$$\sin\phi = \frac{r\sin\theta}{\rho_1}.\tag{4}$$

Since the acute angle between ρ_{ab} and the z-axis is given by $\phi - \beta$, it can be concluded that

$$\rho_{ab} = \frac{r \sin \theta}{\sin \left(\phi - \beta\right)},\tag{5}$$

where β and ϕ are given by Equations 3 and 4, respectively. The final expression for ρ_{ab} can be formulated more compactly by applying the trigonometric identity $\sin(\phi - \beta) =$ $\sin(\phi)\cos(\beta) - \cos(\phi)\sin(\beta)$. Figure S1 shows that $\cos\phi = (r\cos(\theta) + a - d)/\rho_1$. Moreover, the law of sines applied to the triangle with the observation point and the two focal points as vertices in Figure S1, in combination with the trigonometric identity $\sin(2\beta) = 2\sin(\beta)\cos(\beta)$, gives $\sin(\beta) = \sin(\phi)d/\cos(\beta)\rho_2$. Finally, the following expression for ρ_{ab} is obtained:

$$\rho_{ab} = \frac{\sqrt{2}\rho_1^2 \sqrt{\frac{\left(4\sqrt{a^2 - b^2}(a + r\cos(\theta)) + \rho_1^2\right)(2ar\cos(\theta) + b^2 + \rho_1\rho_2 + r^2)}{\rho_1\rho_2}}}{\rho_1^2 + \rho_1\rho_2},\tag{6}$$

X - 4 :

with ρ_1 and ρ_2 the straight lines from the two focal points of the ellipsoid to the observation point (see also the supplementary materials) given by

$$\rho_1 = \sqrt{r^2 + (a - \sqrt{a^2 - b^2})^2 + 2(a - \sqrt{a^2 - b^2})r\cos\theta},\tag{7}$$

$$\rho_2 = \sqrt{r^2 + (a + \sqrt{a^2 - b^2})^2 + 2(a + \sqrt{a^2 - b^2})r\cos\theta},$$
(8)

This expression reduces to $\rho_{ab} = \sqrt{(r\sin\theta)^2 + (\rho_0 + r\cos\theta)^2}$ for a sphere of radius $a = b = \rho_0$.

Text S2. Derivation of the electric field $E(\rho, r, \theta)$

The electric field of a conducting ellipsoid has been derived analytically by Köhn and Ebert (2015) for the prolate spheroid case and by Curtright et al. (2020) for arbitrary dimensions. The derivation of the electric field strength E yields

$$E(x,y,z) = \frac{Q}{4\pi\epsilon_0} \left(\prod_{k=1}^3 \frac{1}{\sqrt{a_k^2 + \Theta}} \right) / \sqrt{\left(\sum_{m=1}^3 \frac{x_m^2}{(a_m^2 + \Theta)^2} \right)},\tag{9}$$

where Q is the charge on the ellipsoid surface, ϵ_0 is the vacuum permittivity, and $a_1 = a_x = b, a_2 = a_y = b$ and $a_3 = a_z = a$ are the semi-axes of the ellipsoid. Moreover, the Θ -equipotentials follow from

$$\sum_{k=1}^{3} \frac{x_k^2}{a_k^2 + \Theta} = 1, \text{ for } \Theta > 0.$$
(10)

Solving equation (10) for Θ gives

$$\Theta = \frac{1}{2} \left(\sqrt{\left(-a^2 - b^2 + x^2 + y^2 + z^2 \right)^2 + 4 \left(-a^2 b^2 + a^2 x^2 + a^2 y^2 + b^2 z^2 \right)} - a^2 - b^2 + x^2 + y^2 + z^2 \right)$$
(11)

Substituting this in equation (9) and converting to spherical coordinates (ρ', θ', ϕ') with the origin at the center of the ellipsoid yields an expression for the electric field in terms of ρ' , the radial coordinate from the center of the ellipsoid, and θ' , the azimuthal angle for the origin at the center of the ellipsoid. Using the trigonometric relations

$$\theta' = \tan^{-1} \left(\frac{r \sin(\theta)}{a + r \cos(\theta)} \right) \tag{12}$$

and

$$\cos\left(2\tan^{-1}u\right) = (1-u^2)/(1+u^2),\tag{13}$$

with $u = \frac{r \sin(\theta)}{a + r \cos(\theta)}$, the obtained electric field expression can be rewritten in the desired θ coordinate. Note that the conversion from θ' to θ can be done in multiple (equivalent) ways using the tangent, sine or cosine. Now, the radial coordinate ρ' needs to be converted to the ρ coordinate along the electric field direction. This is done using the expression for ρ_{ab} , which is the distance to point of photoionization along the ρ coordinate and is given by equation (6). By simple trigonometry the following relation can be derived

$$\rho' = \sqrt{2a\Delta - \Delta^2 + \rho^2 + 2\Delta r \cos(\theta)},\tag{14}$$

with $\Delta = \Delta(r, \theta) = a - \sqrt{\rho_{ab}^2 + \frac{1}{2}r^2\cos(2\theta) - \frac{r^2}{2}} + r\cos(\theta)$ the distance between the center of the ellipsoid and the intercept of ρ_{ab} with the major axis and with ρ_{ab} given by equation (6). Writing out equation (14) gives

$$\rho' = \sqrt{\frac{2ar\cos(\theta)\left(3a^2 - 3b^2 + \rho^2\right) + 2a^2\rho^2 + a^2\rho_1\rho_2 - b^2\rho^2 - b^2\rho_1\rho_2 + \rho^2r^2 + \rho^2\rho_1\rho_2 + p}{2a^2 + 2ar\cos(\theta) - b^2 + r^2 + \rho_1\rho_2}},$$
(15)

with $p = 2a^4 - a^2b^2 + 2a^2r^2\cos(2\theta) + a^2r^2 - b^4 - 2b^2r^2\cos(2\theta) - b^2r^2$. Thus, equation (14) allows us to convert ρ' to ρ . However, as r is also a function of ρ' through $\rho' = \sqrt{a^2 + r^2 + 2ar\cos(\theta)}$ (or equivalently $r = \sqrt{\frac{1}{2}a^2\cos(2\theta) - \frac{a^2}{2} + {\rho'}^2} - a\cos(\theta)$), this derivation is not complete. Substituting the expression for r in terms of ρ' in equation (14) gives an equation that is not analytically solvable. Ideally, an electric field would be obtained that is only a function of ρ and θ . As an analytical expression is required for the model, the expression for the electric field in terms of ρ , r and θ is now accepted, given by

$$E(\rho, r, \theta) = \frac{2b^2 E_0 \rho'}{\sqrt{a^2 - b^2 + q + {\rho'}^2} \left(-a^2 + b^2 + q + {\rho'}^2\right) \sqrt{\frac{(b^2 - a^2)(a^2 + 2ar\cos(\theta) + r^2\cos(2\theta))}{a^2 + 2ar\cos(\theta) + r^2}}},$$
(16)

with the shorthand $q = \sqrt{\frac{2\rho'^2(b^2-a^2)(a^2+2ar\cos(\theta)+r^2\cos(2\theta))}{a^2+2ar\cos(\theta)+r^2}} + (a^2-b^2)^2 + {\rho'}^4$ and where ρ' is converted to ρ using equation (15). As eventually the $\gamma = 1$ equation is solved numerically, the dependence of r is not a problem as long as the resulting α_{eff} (which depends on the electric field) behaves correctly.

Text S3. Derivation of the distance r_{max}

The distance r_{max} between the tip of the ellipsoid and the position of the breakdown field E_k can be found from the relation $\rho_{ab}(r_{max}, \theta.\phi) = \rho_c$. Here, ρ_{ab} is the bisector in the direction of the electric field, and ρ_c is the distance (in terms of the ρ coordinate) to the position of the breakdown field $E(\rho_c) = E_k$. Thus, to solve for r_{max} , ρ_c needs to be found first, where ρ_c is defined by the equation $E(\rho_c) = E_k$ (the electric field is derived in the supplementary materials). However, due to the complicated nature of the electric field of a conducting ellipsoid, this equation cannot be solved explicitly for ρ_c . Nevertheless, it turns out that there is a fairly good approximation for r_{max} .

The surface of a conducting ellipsoid is an equipotential. One can mistakenly think that the electric field is constant on the surface. From symmetry it then follows that the points where the electric field has a certain constant value, such as E_k , will lie on an ellipsoid. Though this is based on a false assumption, it might prove useful to approximate the surface $E = E_k$ as forming an ellipsoid. Simulating the electric field in COMSOL, specifically for the semi-axes a = 5 cm and b = 2 cm, the field at the tip $E_0 = 95.2$ kV/cm and the breakdown field $E_k = 32$ kV/cm, results in the plot of Figure S2. This figure shows that near the tip of the ellipsoid in the (x, z)-plane, where x is the horizontal coordinate and z the vertical coordinate, the line of $E = E_k$ approximately follows the shape of an ellipse around the conducting ellipse. The validity of the approximation can be tested, by importing the data points where $E = E_k$ into MATLAB, and fitting these using the equation of an ellipse. This ellipse has unknown semi-axes a' and b', which are the fitting parameters, and is centered at the center of the conducting ellipsoid (here at (0,0) because of symmetry. The fitting equation is thus $z = a'\sqrt{1-(x/b')^2}$. The data points imported from COMSOL and the fit through these points is depicted in Figure S3, where (x, z) = (0, 5) is the position of the tip of the ellipsoid. It follows that the data points indeed approximately lie on an ellipse near the tip of the conducting ellipsoid, as the fit agrees very well with the data points.

It can thus be concluded that the surface where $E = E_k$ can be approximated as an ellipsoid. Because of symmetry, only an ellipse in the (x, z)-plane or (y, z)-plane needs to be considered. The next step is to find an analytical expression for r_{max} going from the origin at r = 0 at the tip of the conducting ellipsoid to the ellipse where $E = E_k$.

While r_{max} cannot be found from the electric field for arbitrary θ , it can be found for $\theta = 0$ and $\theta = \pi/2$. First, the electric field is reformulated in terms of only r and θ . Then, filling in $\theta = 0$ and solving $E(r, 0) = \frac{b^2 E_0}{2ar+b^2+r^2} = E_k$ for r gives:

$$r_{max}(\theta = 0) = \sqrt{a^2 + \frac{b^2(E_0 - E_k)}{E_k}} - a.$$
 (17)

Obtaining $r_{max}(\theta = \pi/2)$ proves more difficult. Filling in $\theta = \pi/2$ in the electric field expression yields:

$$E(r,\theta=\pi/2) = \frac{2b^2 E_0 \sqrt{\frac{(a^2+r^2)\left((b^2-a^2)\left(\frac{2a^2}{a^2+r^2}-1\right)+\sqrt{4a^2r^2+b^4-2b^2r^2+r^4}+a^2+r^2\right)}{\sqrt{4a^2r^2+b^4-2b^2r^2+r^4+2a^2-b^2+r^2}}}}{\frac{4}{\sqrt{4a^2r^2+b^4-2b^2r^2+r^4}\left(\sqrt{4a^2r^2+b^4-2b^2r^2+r^4}+b^2+r^2\right)}}.$$
 (18)

Setting the above expression equal to E_k and solving for r leads to a case known as 'Casus irreducibilis'. For an irreducible degree 3 polynomial with three real roots, it has been proven that complex numbers need to be introduced to express the solution in roots of any degree, even though the solution is real (Wantzel, 1843). Solving $E(r, \theta = \pi/2) = E_k$ for r, for example using software like Mathematica, leads to 6 solutions containing imaginary parts. Setting b close to a, it is found that one of these solutions approaches the solution of a sphere, accompanied by a very small imaginary part. For example, for a = 3 cm, b = 2.99 cm, $E_0 = 100$ kV/cm, $E_k = 32.75$ kV/cm a value of 10^{-12} cm is found for the imaginary part. These negligibly small imaginary contributions, which are found for any a and b and remain negligibly small, are a results of numerical noise in the machine number calculations in Mathematica (or other numerical software packages).

:

Possibly due to this 'Casus irreducibilits' issue, $r_{max}(\theta = \pi/2)$ has no solution at a = b, so for the reduction to a sphere. However, when b approaches a, $r_{max}(\theta = \pi/2)$ approaches the solution of a sphere. For example, taking again a = 3 cm, b = 2.99 cm, $E_0 = 100$ kV/cm, $E_k = 32.75 \text{ kV/cm}$, the real part of $r_{max}(\theta = \pi/2)$ is 4.27758 cm, while the r_{max} of a sphere at $\theta = \pi/2$ is 4.29894 cm. Instead taking b = 2.9999999 cm (seven decimals) gives $r_{max}(\theta = \pi/2) = 4.29894 \text{ cm}$ for the ellipsoid. It can thus be concluded that $r_{max}(\theta = \pi/2)$ approaches the correct solution for a sphere and can be safely used, as the goal is to implement the model for an ellipsoid and not a sphere, for which simpler expressions are already known. Writing out the found solution for r_{max} at $\theta = \pi/2$ for a conducting ellipsoid gives:

$$r_{max}(\theta = \pi/2) = \frac{1}{2\sqrt{3}}\sqrt{\frac{F}{G}},\tag{19}$$

where F and G are given by:

$$\begin{split} F &= 32 \sqrt[3]{2} (1-i\sqrt{3}) a^8 E_k^4 + 12 \sqrt[3]{2} i(\sqrt{3}+i) a^2 b^6 E_k^2 \Big(E_0^2 + 3E_k^2 \Big) \\ &+ 24a^2 b^2 E_k^2 \Big(\sqrt[3]{b^4} \Big(3b^4 E_k \Big(a^4 C_5 E_k - C_1 E_0^2 + 2C_1 E_k^2 \Big) + 6a^4 E_k^3 (2a^4 C_7 E_k + C_1) - 18a^2 b^6 C_4 E_k^2 + 3a^2 b^2 C_1 E_k \Big(E_0^2 - 4E_k^2 \Big) - b^8 C_3 \Big) + C_2 \\ &+ 4 \sqrt[3]{2} i(\sqrt{3}+i) a^4 E_k^2 \Big) \\ &+ b^4 \Big(4 \Big(E_0^2 - 2E_k^2 \Big) \sqrt[3]{b^4} \Big(12a^8 C_7 E_k^4 + 3a^4 b^4 C_5 E_k^2 - 18a^2 b^6 C_4 E_k^2 + 3a^2 C_1 C_6 E_k - b^8 C_3 - 3b^2 C_1 C_6 E_k \Big) + C_2 + 2 \sqrt[3]{2} (1 - i\sqrt{3}) a^4 E_k^2 \Big(4E_0^2 + 49E_k^2 \Big) \Big) \\ &- 16a^4 E_k^2 \sqrt[3]{b^4} \Big(12a^8 C_7 E_k^4 + 3a^4 b^4 C_5 E_k^2 - 18a^2 b^6 C_4 E_k^2 + 3a^2 C_1 C_6 E_k - b^8 C_3 - 3b^2 C_1 C_6 E_k \Big) + C_2 \\ &+ (1 + i\sqrt{3}) (256a^{12} E_k^6 - 1152a^{10} b^2 E_k^6 - 6b^6 \Big(6a^6 E_k^4 \Big(8E_0^2 + 49E_k^2 \Big) + C_1 C_6 E_k \Big) \\ &+ 6a^4 b^8 C_5 E_k^2 - 36a^2 b^{10} C_4 E_k^2 + 6b^4 \Big(4a^8 C_7 E_k^4 + a^2 C_1 C_6 E_k \Big) - 2b^{12} C_3 \Big)^{2/3} + 2 \sqrt[3]{2} (1 - i\sqrt{3}) b^8 \Big(E_0^2 + E_k^2 \Big)^2, \end{split}$$

$$G = E_k^2 \left(a^2 - b^2\right) \sqrt[3]{b^4} \left(12a^8C_7E_k^4 + 3a^4b^4C_5E_k^2 - 18a^2b^6C_4E_k^2 + 3a^2C_1C_6E_k - b^8C_3 - 3b^2C_1C_6E_k\right) + C_2,$$

and where $C_1, C_2, C_3, C_4, C_5, C_6$ and C_7 are defined as follows:

$$C_{1}^{2} = -12a^{8}E_{k}^{4} \left(8E_{0}^{2} + E_{k}^{2}\right) + 36a^{6}b^{2}E_{k}^{4} \left(7E_{0}^{2} + E_{k}^{2}\right) - 3a^{4}b^{4}E_{k}^{2} \left(13E_{0}^{4} + 72E_{0}^{2}E_{k}^{2} + 12E_{k}^{4}\right) + 6a^{2}b^{6}E_{k}^{2} \left(7E_{0}^{4} + 10E_{0}^{2}E_{k}^{2} + 2E_{k}^{4}\right) - 3b^{8}E_{0}^{4} \left(4E_{0}^{2} + E_{k}^{2}\right),$$

$$C_2 = 128a^{12}E_k^6 - 576a^{10}b^2E_k^6 - 18a^6b^6E_k^4 \left(8E_0^2 + 49E_k^2\right),$$

$$C_3 = 2E_0^6 - 21E_0^4 E_k^2 + 6E_0^2 E_k^4 + 2E_k^6,$$

$$C_4 = 2E_0^4 + 2E_0^2 E_k^2 + 3E_k^4,$$

$$C_5 = 5E_0^4 + 46E_0^2 E_k^2 + 122E_k^4,$$

$$C_6 = 2a^2 E_k^2 + b^2 \left(E_0^2 - 2E_k^2\right),$$

:

$$C_7 = 4E_0^2 + 85E_k^2.$$

Now that analytical expressions are found for $r_{max}(\theta = 0)$ and $r_{max}(\theta = \pi/2)$, the ellipse approximation for r_{max} at arbitrary θ can be applied. Noting that the origin is placed at the tip of the ellipsoidal conductor, the equation for the ellipse where $E = E_k$ is given by:

$$\frac{(z'+a)^2}{a'^2} + \frac{x'^2}{b'^2} = 1,$$
(20)

where (z',x') is a point on the ellipse $E = E_k$, a' is its major semi-axis and b' its minor semi-axis. This configuration is depicted in Figure S4, where the grey region represents the photon absorption region. Here it is seen that $a' = a + r_{max}(\theta = 0)$, $z' = r_{max} \cos \theta$, and $x' = r_{max} \sin \theta$.

Setting z' = 0 in equation (20), which corresponds to $x' = \pm r_{max}(\theta = \pi/2)$ as can be seen from Figure S4, gives

$$\frac{a^2}{a'^2} + \frac{r_{max}(\theta = \pi/2)^2}{b'^2} = 1.$$
(21)

which can be solved for the minor semi-axis b' of the ellipse $E = E_k$:

$$b' = \frac{r_{max}(\theta = \pi/2)}{\sqrt{1 - \frac{a^2}{a'^2}}}.$$
(22)

X - 12

Substituting the found expressions for a', b', z' and x' into equation (20) results in the final expression for $r_{max}(\theta)$ in the ellipse approximation:

$$r_{max}(\theta) = \frac{r_{max}(\theta = \pi/2) \Big(\sqrt{r_{max}(\theta = 0)^2 (2a + r_{max}(\theta = 0))^2 \sin^2(\theta) + r_{max}(\theta = \pi/2)^2 (a + r_{max}(\theta = 0))^2 \cos^2(\theta)}{r_{max}(\theta = 0) (2a + r_{max}(\theta = 0)) \sin^2(\theta) + r_{max}(\theta = \pi/2)^2 \cos^2(\theta)}.$$

(23)

Looking back at Figure S3, the ellipse fit of the data points gives $r_{max}(\theta = 0) = (5.743 - 5) \text{ cm} = 0.743 \text{ cm}$ in the z-direction, and $r_{max}(\theta = \pi/2) = 1.408 \text{ cm}$ in the z-direction. Equation (23) gives for the same input, a = 5 cm, b = 2 cm, $E_0 = 95.2 \text{ kV/cm}$ and $E_k = 32 \text{ kV/cm}$, the values of $r_{max}(\theta = 0) = 0.736 \text{ cm}$ and $r_{max}(\theta = \pi/2) = 1.397 \text{ cm}$. The discrepancy between these values is very small (relative error of about 1%) and caused by equation (23) being derived using the data points of $E = E_k$ at $\theta = 0$ and $\theta = \pi/2$ and basing the ellipse shape on that, while the ellipse in Figure S3 is based on more data points. Hence, the ellipse of equation (23) is formulated such that the two computed r_{max} at $\theta = 0$ and $\theta = \pi/2$ lie on the ellipse, while for the fit the optimal fit does not necessarily go through these two points precisely. Equation (23) can also be compared to Figure S3 for arbitrary $0 \le \theta \le \pi/2$. For example, $\theta = 0.9497$ rad gives $r_{max} = 0.928 \text{ cm}$ for the ellipse fit, and $r_{max} = 0.919 \text{ cm}$ for equation (23). Moreover, $\theta = 0.5120$ rad gives $r_{max} = 0.792 \text{ cm}$ and $r_{max} = 0.784 \text{ cm}$, respectively. We thus conclude that equation (23) is a fair approximation of the actual r_{max} of a conducting ellipsoid.

References

Curtright, T. L., Cao, Z., Huang, S., Sarmiento, J. S., Subedi, S., Tarrence, D. A.,
& Thapaliya, T. R. (2020, apr). Charge densities for conducting ellipsoids. *Eu-*

ropean Journal of Physics, 41(3), 035204. Retrieved from https://doi.org/ 10.1088%2F1361-6404%2Fab806a doi: 10.1088/1361-6404/ab806a

:

- Köhn, C., & Ebert, U. (2015). Calculation of beams of positrons, neutrons, and protons associated with terrestrial gamma ray flashes. *Journal of Geophysical Research: Atmospheres*, 120(4), 1620-1635. Retrieved from https://agupubs .onlinelibrary.wiley.com/doi/abs/10.1002/2014JD022229 doi: https://doi .org/10.1002/2014JD022229
- Wantzel, L. (1843). Classification des nombres incommensurables d'origine algébrique. Nouvelles annales de mathématiques : journal des candidats aux écoles polytechnique et normale, 1e série, 2, 117-127. Retrieved from http://www.numdam.org/item/ NAM_1843_1_2__117_1



Figure S1. Schematic of the bisector giving the electric field direction outside a conducting ellipsoid.



Figure S2. The line of constant $E = E_k$ (red) for an ellipse with semi-axes a = 5 cm (vertical) and b = 2 cm (horizontal). Here only half of the width of the ellipse is shown and the axis of symmetry is drawn.



Figure S3. Fit of $E = E_k$ data points using an ellipse fit of $z = a'\sqrt{1 - (x/b')^2}$. It is found that $a' = (5.743 \pm 0.002)$ cm and $b' = (2.86 \pm 0.01)$ cm.



Figure S4. The conducting ellipsoid (solid, black line), with positions of constant $E = E_k$ (solid, red line), the edge of the ionization region, approximated as an ellipse shape (dashed line) with coordinates (z',x') and semi-axes a' and b'.