HF Induced Modifications of the Electron Density Profile in the Earth's Ionosphere Using the Pump Frequencies Near the Fourth Electron Gyroharmonic

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Abstract

We discuss results on plasma density profile modifications in the F-region ionosphere caused by HF heating with the frequency f_0 in the range [(-150 kHz)-(+75 kHz)] around the 4th electron gyroharmonic $4f_c$. The experiments were conducted at the HAARP facility on June 2014. The multi-frequency Doppler Sounder (MDS) which measures the phase and amplitude of reflected sounding radio waves complemented by the observations of the Stimulated Electromagnetic Emission (SEE) were used for the diagnostics of the plasma perturbations. We detected noticeable plasma expulsion from the reflection region of the pumping wave, and from the upper hybrid region, the expulsion from the latter one had been strongly suppressed for f_0 [?] $4f_c$. The plasma expulsion from the upper hybrid region was accompanied with the sounding wave's anomalous absorption slower developing for f_0 [?] $4f_c$. Also, slower developing and weaker expulsion was detected for the height region between the pump wave reflection and upper hybrid altitudes.





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Key Points:

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11	•	plasma profile modifications in the HF-pumped ionosphere with frequencies near
12		the fourth electron gyroharmonic are studied at HAARP
13	•	plasma expulsion from the reflection height and the upper hybrid height of the pump
14		wave were detected
15	•	The expulsion from the upper hybrid height had been strongly suppressed for the
16		pump frequency close to fourth electron gyroharmonic

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17 Abstract

We discuss results on plasma density profile modifications in the F-region ionosphere caused 18 by HF heating with the frequency f_0 in the range [(-150 kHz)–(+75 kHz)] around the 19 4^{th} electron gyroharmonic $4f_c$. The experiments were conducted at the HAARP facil-20 ity on June 2014. The multi-frequency Doppler Sounder (MDS) which measures the phase 21 and amplitude of reflected sounding radio waves complemented by the observations of 22 the Stimulated Electromagnetic Emission (SEE) were used for the diagnostics of the plasma 23 perturbations. We detected noticeable plasma expulsion from the reflection region of the 24 pumping wave, and from the upper hybrid region, the expulsion from the latter one had 25 been strongly suppressed for $f_0 \approx 4 f_c$. The plasma expulsion from the upper hybrid 26 region was accompanied with the sounding wave's anomalous absorption slower devel-27 oping for $f_0 \approx 4 f_c$. Also, slower developing and weaker expulsion was detected for the 28 height region between the pump wave reflection and upper hybrid altitudes. 29

³⁰ Plain Language Summary

We report first results of systematic studying electron density profile modifications 31 in the F-region of the ionosphere caused by HF heating with the pumping frequencies 32 in the range [(-150 kHz)–(+75 kHz)] near the 4th electron gyroharmonic at the HAARP 33 heating facility. We measured phase and amplitude of ionospherically reflected sound-34 ing waves simultaneously with Stimulated Electromagnetic Emission. The phase data 35 allowed determining electron density profile evolution, the amplitude data provide in-36 formation of the pump-induced sounding waves' absorption, while SEE allows estimat-37 ing offset between the pump frequency and gyroharmonic. We detected noticeable plasma 38 expulsion from the reflection height and slower developing expulsion from the upper hy-39 brid height of the pump wave. Even slower developing and weaker expulsion was detected 40 for the height region between the pump wave reflection and upper hybrid heights. The 41 plasma expulsion from the upper hybrid region was accompanied with the sounding wave's 42 anomalous absorption, but was strongly suppressed for the pump wave frequency close 43 to fourth electron gyroharmonic, when the prominent "Downshifted Maximum" in SEE 44 was quenched. The combined phase sounding and SEE measurements allowed obtain-45 ing interconnection between different manifestations of the HF-induced turbulence and 46 determining altitude of the most effective pump wave energy input to ionospheric plasma. 47

48 1 Introduction

A powerful O-mode electromagnetic pump wave transmitted from the ground into 49 the bottom-site ionospheric F region excites a wide range of plasma processes leading 50 to the appearance of artificial ionospheric turbulence (AIT), i.e., generation of different 51 HF and LF plasma modes, plasma density inhomogeneities of scales from tens of cen-52 timeters to kilometers, causes electron heating, electron acceleration, ionization, gener-53 ation of ionospheric airglow etc. (Gurevich, 2007; Grach et al., 2016; Streltsov et al., 2018). 54 Diverse diagnostic methods and tools are used for the AIT studying, particularly, sound-55 ing of the heated volume of the ionosphere by diagnostic waves and a registration of sec-56 ondary, or stimulated, emission (SEE) in different frequency ranges. 57

The pump-plasma interaction is known to be strongest near the pump reflection height z_{r_0} at which $fp(z_{r_0})$ equals the pump frequency f_0 , and near the upper hybrid (UH) resonance height z_{UH} where $f_p(z_{UH}) = (f_0^2 - f_c^2)^{1/2}$ [here $fp = (e^2 N/\pi m)^{1/2}$, and $f_c = eB/(2\pi mc)$ are the electron plasma frequency and the electron cyclotron frequency, e and m are the electron charge and mass, N the electron density, c the speed of light, B the geomagnetic field strength]. This corresponds to existing theoretical concepts and is confirmed by investigations of the HF-pumped ionospheric volume by multifrequency Doppler sounding (MDS) at SURA and EISCAT heating facilities (Vaskov et al., 1986; Berezin et al., 1991; Lobachevsky et al., 1992; Grach et al., 1997; A. Shindin et al., 2012), which have revealed plasma expulsion from the resonance regions.

Stimulated electromagnetic emission (SEE) with frequencies f_{SEE} close to the pump 68 wave frequency f_0 occurs due to conversion of HF pump-driven electrostatic plasma modes, 69 most notably Langmuir (L) and upper hybrid (UH), into electromagnetic waves smaller 70 than the reflected pump wave (PW) by 50-90 dB (Thidé et al., 1982; Stubbe et al., 1984; 71 Leyser et al., 1989) and provides rich information about AIT (Grach et al., 2016). SEE 72 spectral characteristics depend on $\Delta f_c = f_0 - nf_c$, the offset of the pumping wave (PW) 73 frequency f_0 from the multiple electron gyroresonance nf_c , the most dramatic changes 74 occur during transition of f_0 via nf_c (e.g. from $f_0 < nf_c$ to $f_0 > nf_c$) and allows one 75 to estimate Δf_c during the experiment (Leyser et al., 1993; Leyser, 2001; Frolov et al., 76 2001; Sergeev et al., 2006; Grach et al., 2008, 2016). 77

We report results of the first experiments using MDS (phase) sounding of the HF-78 pumped ionosphere at the HAARP heating facility, located near Gakona, Alaska, USA 79 (62.40°N, 145.15°W) performed on June 2014. Simultaneously the SEE was monitored, 80 and the anomalous absorption (AA) of sounding waves was measured. The heating fa-81 cility was used both for the pump wave radiation and as pulsed Doppler HF radar. The 82 main purpose of the experiments was to study the dependence of HF-pump-induced elec-83 tron density expulsion from the resonance regions (in correlation with the AA and SEE) 84 on the offset of the pump wave frequency from the fourth gyroharmonic, $\Delta f_c = f_0 - 4f_c$, 85 the experiment was performed for.

⁸⁷ 2 Experimental Setup

Experiments on phase-amplitude sounding of the ionosphere heated volume were 88 performed on June 4. During time interval 14:55-16:25 AKDT the HAARP transmit-89 ter radiation schedule was organized as a sequence cycles at different PW frequencies f_0 90 varied from 5540 till 5730 kHz. The choice of the PW frequency nominals was conditioned 91 by covering available range around $4f_c$. Each cycle was organized as follows. During 4.5 92 min the HAARP operated as pulsed Doppler HF radar. The transmitters radiated ver-93 tically low-duty cycle diagnostic waves (DW): short (20 µs) pulses with the interpulse 94 period T = 100 ms at two carrier frequencies $f_{\rm DW} = f_0$ and $f_{\rm DW} = f_0$ -200 kHz with 95 effective radiated power (ERP) $P_{\rm ef} \sim 400$ MW each. After 30 seconds the radar mode 96 was combined for 45 seconds with quasi-continuous wave (QCW) pumping mode, i.e. high 97 duty cycle pump wave (pulse width $\tau = 70$ ms, the same interpulse period T = 10098 ms) at a frequency f_0 with the same ERP. During the QCW, the short pulses were ra-99 diated at 20th ms of the 30 ms pauses. After switching off the radar mode, the 30 s pause 100 of the HAARP transmitters was used for taking ionograms and changing the PW and 101 DW frequencies. Then the 5-min cycle was repeated at the new f_0 and f_{DW} . The power 102 of the 20 µs diagnostic pulses was sufficient to create a wide spectrum of DW (up to 300 103 kHz near each carrier frequency). An average DW power $\langle P \rangle = P/Q = 80$ kW was 104 far below the thresholds of the generation and maintenance of the pump-induced ther-105 mal plasma instabilities in the ionosphere (Grach et al., 1981). Also, the diagnostic pulses 106 were too short to excite ponderomotive instability in the ionosphere (Frolov et al., 2007; 107 Sergeev et al., 2007). 108

Under the combined radiation mode, the QCW created a perturbation in the iono-109 sphere, particularly at the plasma resonance (reflection altitude of the pump wave) and 110 at the UH resonance, while the DW simultaneously provided diagnostics of AIT (Vaskov 111 et al., 1986; Berezin et al., 1991; Grach et al., 1997; Sergeev et al., 2007, 2016; A. Shindin 112 et al., 2012). The high power of the HAARP transmitters used for MDS, applying broad-113 band radio receiver and specially developed signal processing algorithms have allowed 114 studying evolution of amplitude (A_i) and phase (φ_i) of the various spectral components 115 of the reflected DW (with frequencies f_i), which passed the perturbed region twice, in 116

a wide frequency range $\Delta f_{\text{total}} \sim 600 \text{kHz}$, (-450 kHz $\lesssim f_i - f_0 \lesssim +150 \text{ kHz}$), and 117 therefore, in a wide (till 25-35 km) altitude interval. Sometimes during QCW we could 118 analyze spectral components of the radar pulses in smaller range $\Delta f_{\rm QCW} \sim 360$ kHz 119 (-280 kHz $\lesssim f_i - f_0 \lesssim +80$ kHz. Outside of this range reflected signal fell till the back-120 ground noise due to the anomalous absorption (AA, see below), and signal amplitude 121 and phase analysis gave condemned results. The frequency resolution used for the anal-122 ysis was 1 kHz; the temporal resolution was determined by the interpulse period T =123 100 ms. The observational site were located, nearly under the heated region during in-124 jections at vertical. A 30-m folded-dipole BWDS receiver antenna was used in measure-125 ments. The receiver digitized a band 850 kHz, the dynamic range of the instruments af-126 ter spectral processing is estimated to be better than 90 dB. 127

Figures 1 and 2 present results obtained during the cycles at $f_0 = 5540, 5600, 5660$ 128 and 5730 kHz. Figure 1 shows the temporal evolution of Doppler frequency shifts $f_{di}(f_i, t) =$ 129 $d\varphi_i/dt$ (column a, φ_i is the phase incursion), normalized intensities $G_i(f_i, t)$ of the re-130 flected DW spectral components (column b), and SEE spectrograms (column c). The DW 131 intensities $A_i^2(f_i, t)$ are normalized to the $\langle A_i^2 \rangle$, the intensity averaged over 30 s before 132 the QCW switching on, $G_i[dB] = 10 \log(A_i^2/\langle A_i^2 \rangle)$) characterizes the anomalous ab-133 sorption (AA) due to scattering of DW into plasma (UH) waves on small-scale plasma 134 density irregularities (striations). In Fig. 2b we used running averaging over 5 pulses (0.5)135 s). In the spectrograms, prominent spectral SEE features are marked. There are the L-136 related ponderomotive narrow continuum (NCp) and UH-related downshifted maximum 137 (DM), 2DM, upshifted maximum (UM), thermal narrow continuum (NCt), broad con-138 tinuum (BC) and broad upshifted maximum (BUM). Stationary peculiarities and tem-139 poral evolution of different SEE features are well described in numerous papers, see, e.g. 140 (Frolov, 2004; Leyser et al., 1993, 1994; Leyser, 2001; Carozzi, 2002; Thidé et al., 2005; 141 Sergeev et al., 2006; Kotov et al., 2008; Grach et al., 2008, 2016), etc. The SEE can be 142 used for rough estimations of the offsets between f_0 and $4f_c$, $\Delta f_c = f_0 - 4f_c$, in the 143 range between successive gyroharmonics. Most precisely this can be done for $f_0 \approx 4 f_c$ 144 and $f_0 \gtrsim 4f_c$ (see formulas (7) and (8) below). 145

¹⁴⁶ 3 Phase Data Processing

For all f_0 , when heating (QCW) is turned on, as a rule, an increase of all f_i (pos-147 itive Doppler frequency shifts f_{di} is often observed. It is less pronounced for narrow ranges 148 of the diagnostic (sounding) waves frequencies with reflection heights near PW reflec-149 tion and UH heights z_{r0} and zUH, i.e. for $f_i \sim f_p(z_{r0})$ and $f_i \sim f_p(z_{UH})$. Negative 150 f_{di} are observed after the heating turns off. Such pump-induced changes of f_{di} occur on 151 the background of natural variations of the reflection heights (and f_{di}), determined by 152 the motion of the ionosphere. The measured temporal evolution of Doppler frequency 153 shifts $f_{di}(t) = \mathrm{d}\varphi_i/\mathrm{d}t$ for different f_i (Figure 1, panels a) provided data for the recon-154 struction of the electron density profile temporal evolution N(z,t) in the HF-pumped 155 ionosphere by solving the inverse problem of the phase sounding. 156

¹⁵⁷ Under geometric optics approximation, each of the sounding waves at the angu-¹⁵⁸ lar frequency $\omega_i = 2\pi f_i$ propagating from the ground up to the reflection points $z_r(f_i)$ ¹⁵⁹ and back to the ground, experiences the phase incursion (Ginzburg, 1970):

$$\varphi(\omega,t) = \frac{2\omega}{c} \int_0^{z_r} n(\omega,\omega_p(z,t)) dz - \frac{\pi}{2}$$
(1)

where $\omega_p(z,t) = 2\pi f_p(z,t)$ is the plasma frequency, $n(\omega, \omega_p(z,t))$ is the wave refractive index. The reflection altitude z_r is determined by the condition n = 0, and t is the time. This can be translated in the following expression for an additional phase change $\Delta \varphi(\omega) = \varphi(\omega, t_0) - \varphi(\omega, t)$ in the time interval $[t_0, t]$ associated with perturbation of the profile N(z,t) due to ionosphere HF-pumping or to natural reasons:

$$y(\omega) = \frac{c}{2\omega} \Delta \varphi(\omega) = \int_{\omega_{\min}}^{g(\omega)} K(\omega, \omega_p) \Delta z(\omega_p) d\omega_p$$
(2)

Here $K(\omega, \omega_p) = dn(\omega, \omega_p)/d\omega_p$ is a kernel of the integral equation (2), $g(\omega)$ is the an-167 gular plasma frequency at the reflection point, which is $g(\omega) = \omega$ for an ordinary wave, 168 t_0 is initial time, $\Delta z(\omega_p, t) = z(\omega_p, t) - z(\omega_p, t_0)$ is the altitude shift, i.e., the difference 169 between the sounding radio wave reflection altitude at the current (t) and initial (t_0) times. 170 Here the variable of the integration is replaced from the altitude z in (1) to the plasma 171 frequency ω_p in (2). It is taken into account that at the reflection point $n(\omega, g(\omega)) =$ 172 0, and at the entrance to the plasma layer $\Delta z = 0$. The left hand side $y(\omega)$ in (2) is 173 to be determined from the experimental data as $\Delta \varphi_i = \int_{t_0}^t f_{di}(t') dt'$. On the base of 174 the data obtained, an array $\Delta \Phi(\omega, t) = \Delta \varphi(\omega, t) - \Delta \varphi(\omega_{\min}, t)$ was created for the phase 175 shifts $\Delta \varphi(\omega, t)$ of the diagnostic waves. Here ω_{\min} is least of the probe wave frequencies, 176 in our experiment we have taken $\omega_{\min} = \omega_0 - 2\pi \cdot 280$ kHz. The probe wave at this frequency was reflected noticeably below (~ 25 - 35 km) z_{UH} , and we assumed that the 177 178 phase evolutions for this and lower frequencies do not depend on the pump-induced pro-179 cesses in the plasma resonance regions. This is confirmed, in particular, by measurements 180 of the AA bandwidth of the DW (Fig. 1b). Also, the subtraction of $\Delta \varphi(\omega_{\min}, t)$ allows 181 to exclude processes responsible for plasma density variations at lower altitudes caused 182 by violation of the balance between ionization and recombination in the lower ionosphere. 183

For the exact expression of the refractive index for magnetized plasma, the equation (2) cannot be reduced to the analytically solvable integral equation. In this case, we had applied in (A. Shindin et al., 2012; Sergeev et al., 2016) the regularization algorithms by Tikhonov (1995) and solved the inverse problem numerically. In this paper, instead of numerical Tikhonov algorithm, we used an approximate expression for well describing ordinary wave refractive index near the reflection point for HAARP experimental conditions,

$$n(\omega, \omega_p) \approx \left(1 - \frac{\omega_p^2}{\omega^2}\right)^{\beta}, \ \beta \approx 0.29.$$
 (3)

¹⁹² Substituting (3) into (2), we obtain the generalized Abel equation (Korn & Korn, 2000):

$$\int_{\omega_{\min}^2}^{\omega^2} \frac{\Delta z}{(\omega^2 - \omega_p^2)^{1-\beta}} d\omega_p^2 = F(\omega), \ F(\omega) = -c \frac{\omega^{2\beta-1}}{2\beta} \Delta \varphi(\omega)$$
(4)

¹⁹⁴ with the solution

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$$\Delta z = \frac{\sin\{\pi(1-\beta)\}}{\pi(1-\beta)} \int_{\omega_{\min}^2}^{\omega^2} \frac{F'(\omega_p^2)}{(\omega^2 - \omega_p^2)^\beta} \mathrm{d}\omega_p^2.$$
(5)

¹⁹⁶ Dynamics of the reflection altitude shifts of different DW spectral components, $\Delta z_r(\omega_i = 2\pi f_i, t)$ is displayed in Figure 2, panels a. In this figure t = 0 corresponds to QCW switch-¹⁹⁷ ing on, blue and red lines correspond to the reflection altitude of the DW at $f_i = f_0$, ¹⁹⁹ i.e. the PW reflection altitude, $(z(f_i) = z_r(f_0))$ and to the reflection altitude of the DW ²⁰⁰ at $f_i = f_p(z_{\rm UH}) = (f_0^2 - f_c^2)^{1/2}$, i.e. reflection from the pump wave UH resonance height, ²⁰¹ respectively. Chosen frequency step between neighbor spectral components displayed in ²⁰² the Figure 2a is $\Delta f = 30$ kHz. For clarity we introduce additional height shift 300 m ²⁰³ between reflection altitude shifts of the successive DW spectral components at t = 0.

Temporal variations of the reflection heights $\Delta z(f_i, t)$ allows to calculate velocities of the vertical motion of plasma density at a certain magnitude $Ni = \pi f_i^2 m/e^2$ as $V_v = \partial \Delta z_r(f_i, t)/\partial t$ at different f_i . The velocities of the sounding waves reflection height displacement vs. time and sounding wave frequency (pulse spectral component) $V_v(f_i, t)$ are presented in Figure 2b. Positive (red) velocity values correspond to upward motion of a certain plasma density level while negative (blue) values correspond to downward motion. Calculating the altitude shifts $\Delta z_r(f_i, t)$ allows us observing the evolution of the electron density profile N(z,t) from the reference one. The reference profile $N_0 = N(z,t_0)$ is taken from ionogram registered prior to the QCW pumping session. For this we shall transform $\Delta z_r(f_i, t)$ to $\Delta z_r(N, t)$ by using the univocal relation between the plasma frequency at the radio wave reflection point and electron density. Therefore,

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$$z(N, N_0, t) = z(N_0) + \Delta z(N, t)$$
(6)

is the dependence of the reflection height of the radio wave on the density. Then we find the required distribution N(z,t) by calculating the inverse of (6). Relative differences between reconstructed and reference electron density profiles $\delta N(z,t) = [N(z,t)-N_0]/N(z,t_0)$ vs. altitude z for all f_0 at 5th, 15th, 30th and 45th seconds of pumping are presented in Figure 2c.

Figure 2 illustrates the evolution of the electron density in the HF-pumped ionosphere in the altitude range from region above the reflection height till the region below the UH height.

4 Analysis of the Combined Phase sounding, AA, and SEE Data

Let's start an analysis from $f_0 = 5540$ kHz, the upper row of figures 1 and 2. During this cycle the pump frequency $f_0 = 5540$ kHz belongs to the "weak emission range" between 3rd and 4th gyroharmonic, where SEE spectrum contains weak BC, DM and UM (Fig. 1c). The offset $\Delta f_c = f_0 - 4f_c$ here can be roughly estimated from ionograms and geomagnetic field IGRF model as $\Delta f_c \sim -(130 - 150)$ kHz.

From Figure 1a, it is seen that immediately after QCW is switched on the Doppler 231 frequency shifts were mainly positive $(f_{di} > 0)$ for all spectral components except of 232 ones close to the pump frequency $(f_i \approx f_0)$ where $f_{di} < 0$. Such a behavior of f_{di} is 233 translated to increasing reflection altitudes of the sounding waves z_{ri} with the frequen-234 cies close to the pump wave frequency f_0 (Fig. 2a) with velocities V_v up to 100 m/sec, 235 and, to the contrary, slighter decreasing z_r for both $f_i > f_0$ and $f_i < f_0$ with veloci-236 ties V_v up to 60 m/s (Fig. 2b). This corresponds to electron density decrease in the vicin-237 ity of the pump wave reflection height $z_{r_0} = z_r(f_0)$ and increase on other heights, i. e. 238 to the plasma expulsion from the reflection point vicinity. The relative density deple-239 tion (hereafter reflection depletion, RD) reaches up to $\sim 0.1 - 0.2\%$ at 5th second of 240 the QCW pumping (Fig. 2c). According to Fig. 2a, the uplifting $\Delta z_r(f_i \approx f_0)$ grows 241 from 0.1 s till (3-5) s, reaching 100-300 m, and then slows down till t = 10-15 s. The 242 SEE feature NCp appears immediately after QCW switching on, simultaneously with 243 the start of the plasma expulsion from the vicinity of z_{r_0} and then exhibits strong overshoot-244 effect: its spectral width and intensity noticeably drops during the raise of $z_r(f_0)$ and 245 growth of the AA and UH-related SEE features (Fig. 1b, 1c). Such initial behavior of 246 the $f_d(f_i, t)$, $\Delta z_r(f_i, t)$, $V_v(f_i, t)$, NCp and AA is qualitatively similar for all f_0 and does 247 not depend on the offset Δf_c . 248

Few seconds later, in t = 1-3 s, phenomena related to excitation of the UH waves 249 and striations, resulting in the phenomena such as the AA, and the UH related SEE fea-250 tures (DM, UM, and BC), and the plasma expulsion from the UH height region ($z \sim$ 251 $z_{\rm UH}$, the "UH depletion", hereafter UHD) developed simultaneously. The expulsion cor-252 responded to appearance of the expanding range with $f_{di} < 0$ in Fig. 1a. At 3-5 s the 253 UHD became deeper than the RD, and then the UHD developed monotonously till QCW 254 pumping switchies off. At t 15 s the uplifting near the reflection point accelerated again 255 and continued till the QCW stops, for s deepening of the UHD and RD is accompanied 256 by plasma density decrease in the whole altitude range (Fig. 2c). 257

For the cycle at $f_0 = 5600$ kHz (Figs. 1, 2, second row) the results of AA measurements (temporal development and magnitude) and the phase sounding analysis ($f_d(f_i, t)$, ²⁶⁰ $\Delta z_r(f_i, t), V_v(f_i, t)$ and $\delta N(z, t)$) were similar to ones for $f_0 = 5540$ kHz, but their SEE ²⁶¹ were fairly differed. According to the SEE spectrogram (Fig. 1c), the PW frequency be-²⁶² longs to "below harmonic" range. It shows strong DM and resolved 2DM. The offset can ²⁶³ be roughly estimated as $\Delta f_c \sim -(60 - 70)$ kHz. According to the ionograms, $z_{\rm UH}$ in ²⁶⁴ the cycles at 5540 and 5600 kHz were close, $z_{\rm UH} \sim 225$ km.

Note, during these cycles the average anomalous absorption $\langle G \rangle$ in the range 170 $< f_0 - f_i < 220$ kHz decreased by $\sim 10 - 15$ dB (Figs 1b) in comparison with the range $-30 < f_0 - f_i < 150$ kHz. The most probably it can be attributed to focusing of HF diagnostic radio waves reflected from the UH altitudes with density depletion (Vaskov et al., 1986).

In the cycle at $f_0 = 5660$ kHz (Figs. 1,2, 3rd row) the behavior of the bulk of investigated parameters differed noticeably from cycles at $f_0 = 5540$ and 5600 kHz.

The RD developed similar to cycles with other f_0 during first 10 s of QCW pump-272 ing, but did not slow down after 10 s of QCW pumping. Negative f_{di} appeared near $f_i(z_{r_0})$ 273 than the range of $f_{di} < 0$ expanded for mainly to lower f_i . Dependence looked like shallow 274 low quasi-periodic structure with a period $\sim 3-4$ km, amplitude growing in time and 275 decreasing downward from z_{r_0} , and occupying a height interval exceeding spacing be-276 tween z_r and $z_{\rm UH}$. The UHD and zUH uplifting did not resolving in this cycle, more-277 over, the weak descending of the UH height at surrounded frequencies were observed. 278 Then the AA developed slower than in other cycles, while attained the same magnitudes 279 till 45th s. The difference is due to the fact that during this cycle at $f_0 = 5660$ kHz the 280 pump wave frequency was much closer to the double resonance where $f_{\rm UH} \approx 4 f_c$. This 281 is seen from the SEE spectrogram (Fig. 1c). Here the DM is not resolved till t $\sim 10-$ 282 15 s which means that 283

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$$_{\rm DM} = f_0 - \Delta f_{\rm DM} \approx f_{\rm UH} \approx 4f_c \tag{7}$$

at the UH height (the "resonance range"), $f_{\rm DM}$ is the DM peak frequency. Then the DM 285 appeared which was probably attributed to amplification of the $\delta N(z,t)$, and changing 286 of the UH height, and therefore, a violation of (7). This allows to estimate the offset $\Delta f_c =$ 287 f_0-4f_c during the cycle as 7-15 kHz: initially $\Delta f_c \approx \Delta f_{\rm DM}$ and then changed. Detailed 288 analysis of the SEE peculiarities (DM, UM and BUM) near the double resonance can 289 be found in (Levser et al., 1994; Carozzi, 2002; Sergeev et al., 2006; Kotov et al., 2008; 290 Grach et al., 2008). Note that the altitude of the double resonance obtained from (7) 291 and IGRF model ~ 240 km exceeded $z_{\rm UH}$ obtained from the ionogram by 10 km. This 292 value fits into the error in determining the heights when processing ionograms. 293

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The cycle at $f_0 = 5730$ kHz (Figs. 1,2, row 3) was the only cycle with $f_0 > 4f_c$, the offset $\Delta f_c = f_0 - 4f_c$ can be estimated from the SEE spectrogram as

$$\Delta f_c = f_0 - 4f_c \approx f_{\rm BUM} - f_0 = \Delta f_{\rm BUM},\tag{8}$$

 $\Delta f_c \sim +75$ kHz (Fig. 1c), fBUM is the BUM peak frequency. From Fig. 3 it can be 297 concluded the following. After QCW pumping switching on the plasma expulsion from 298 the vicinity of the reflection point (RD) and the NCp development in SEE spectrum be-299 haved similarly to all f_0 . Slowing down of the RD deepening after development of the 300 UH-related effects is were not observed. Like for $f_0 < 4f_c$, for $f_0 > 4f_c$ the UH-related 301 effects, namely AA; DM, 2DM, UM, SEE features; and UHD developed few seconds later 302 than L-related processes near the reflection height. The characteristics of this processes 303 during this cycle, however, differed from ones for $f_0 < 4f_c$. 304

First, the AA attained a saturation at $t \sim 10-15$ s, but for $f_0 > 4f_c$ at the saturation stage $\langle G \rangle \sim 25 - 30$ dB, while at $f_0 < 4f_c \langle G \rangle \sim 18 - 20$ dB, and the frequency range with strong AA is wider at $f_0 > 4f_c$ than at $f_0 < 4f_c$ (Figs 1b). Due to the strong AA, in the range 5590 $< f_i < 5650$ kHz the DW signal intensity fell below the background noise, and measurements of Doppler shifts/phase incursions and AA

were contaminated. This range is shown in Figures 1,2, row 4 by double arrows and dashed 310 parts of lines and shall be excluded from the analysis for t > 15 s. Second, the strong 311 NCt showed up in the SEE spectrogram at $\Delta f \sim 0 - (-7)$ kHz with a temporal be-312 havior similar to DM, 2DM and UM; after $t \sim 1.5$ s the NCt covered NCp. The DM, 313 2DM and UM developed concurrently with AA, exhibited the overshoot effect with max-314 ima at $t \sim 6-11$ s, and were more intensive then in other cycles. Such SEE behavior 315 is typical for the "above harmonic range" close to the "strong emission" range (Sergeev 316 et al., 2006). Third, the UHD started to develop ~ 2 km above the nominal upper hy-317 brid resonance height and occupied quite wide (> 5 km) altitude interval. Later, in $t \sim$ 318 10-15 s, the interval expanded (till ~ 8 km) and descended below the UH height. Dur-319 ing the whole 45 s of the QCW pumping the UHD remained shallower than the RD (Fig. 320 2c). 321

After the termination of the QCW, the signs of Doppler frequency shifts fdi and velocities $V_v = \partial \Delta z_r(f_i, t)/\partial t$ changed to the opposite ones (see Figures 1a and 2b). This leads to a reduction of the plasma density depletions around the reflection and UH heights, the depletions relaxed and disappeared in ~ 15-40 s after pump wave turns off.

327 5 Conclusions

The results can be briefly summarized as follows. It was obtained that during all 328 cycles the pump wave-plasma interaction developed most quickly (in a few milliseconds) 329 after QCW switching on in the vicinity of the pump wave reflection height z_{r_0} . It is ac-330 companied by the plasma expulsion from the interaction region (RD appearance) and 331 by the NCp SEE feature generation. At this time there were no essential differences be-332 tween cycles with different f_0 . Both the expulsion and NCp shall be attributed to the 333 excitation of L waves due to the ponderomotive parametric instability near the PW re-334 flection height z_{r_0} . 335

Later, in a 1-5 s after the QCW was switched on, for $f_0 < 4f_c$, the plasma expul-336 sion from a vicinity of the upper hybrid height $z_{\rm UH}$ (UHD development) began along with 337 development of the AA, UH-related SEE features such as DM, 2DM, UM and BC, as 338 well as with suppression (overshoot) of the NCp feature and slowing down of expulsion 330 from the vicinity of z_{r_0} . At 3-10 s the UHDs became deeper than the RDs. The expul-340 sion from upper hybrid height continues until the end of 45-s-long QCW-pumping. All 341 these effects are related to excitation of the striations and UH plasma waves. The slow-342 ing down of the z_{r_0} uplifting, RD deepening and NCp suppression indicated that the UH-343 related processes leaded to the noticeable shielding of the reflection point from the pump 344 wave energy. A sequence of the described effects is consistent with a general scenario of 345 the phenomena developing in the HF-pumped ionosphere (Frolov et al., 1997; Thidé et 346 al., 2005; Grach et al., 2016) and is clearly illustrated by Sergeev et al. (2016), where the 347 results of 3 successive 2-minute cycles at the frequency $f_0 = 5500$ kHz obtained on the 348 same experiment (2014/06/04, 15:40-15:46 AKDT) were presented. Similar results were 349 obtained also at the SURA facility (A. Shindin et al., 2012). In the described experiment 350 during the cycles at $f_0 < 4f_c$ the RD and NCp develop similarly to the cycles at $f_0 =$ 351 5500 kHz till t= 5-10 s, but later the RD started to deepen again, concurrently with UHD 352 deepening on the background of plasma density decrease in the interval $z_{\rm UH} \lesssim z \lesssim z_{r_0}$. 353 A dependence $\delta N(z)$ looked as two isolated minima, close, respectively, to the reflection 354 and UH heights. Note that for presented cycles at $f_0 < 4f_c$ (with approximately same 355 initial UH heights ~ 225 km) the expulsion parameters $\Delta z(f_i)$ and $\Delta N(z)$ as well the 356 AA behaved similarly, even quantitatively, while the SEE spectra are different. This points 357 on a weak dependence on the AIT peculiarities on in the range $-150 \lesssim \Delta f_c \lesssim -60$ 358 kHz. 359

In the single cycle with $f_0 > 4f_c$ ($f_0 = 5730$ kHz, $\Delta f_c \sim 75$ kHz) the $\delta N(z)$ de-360 veloped, again, as two isolated minima, RD and UHD. In this cycle the UHD depth re-361 mained less than the RD depth during the whole OCW pumping interval; the altitude 362 ranges, occupied by depletions, were wider, the AA of the diagnostic pulses (DW) was 363 stronger (by ~ 10 dB) and occupied larger frequency range than at $f_0 < 4f_c$. The up-364 lifting of the diagnostic waves (DW) reflection heights z_{ri} near the PW UH height started 365 after ~ 15 s delay (Fig. 3c, row 4) after QCW switch on. This points to the essential 366 difference in the AIT excitation for $f_0 > 4f_c$ and $f_0 < 4f_c$. 367

For the cycle at $f_0 = 5660$ kHz the PW frequency was close to $4f_c$, i.e. got into 368 the "resonance range", where the DM was totally suppressed or very weak, but the BUM 369 and UM were present in the SEE spectra. Taking $f_{\rm DM} = f_0 - \Delta f_{\rm DM} \approx 4 f_c(z_{\rm UH}) \approx$ 370 5650 kHz and using IGRF model we obtain $z_{\rm UH} \approx 239$ km in the beginning of the cy-371 cle. In this cycle the DW frequency range with negative f_{di} expanded from the PW fre-372 quency f_0 (Fig. 1a, 3rd row), the temporal behavior of $\delta N(z,t)$ demonstrated deepen-373 ing quasi-periodic structure with a period $\sim 3-4$ km, with amplitude growing in time 374 and decreasing downward from z_{r_0} , and occupying height interval exceeding spacing be-375 tween z_{r_0} and z_{UH} . Unfortunately, the total range of the diagnostic signals available for 376 the phase data processing was too narrow to estimate lower boundary of the height in-377 terval. "Independent" z_{ri} uplifting for $z_{ri} \sim z_{\rm UH}$ wasn't resolved for this cycle. There-378 fore, for $f_0 \approx 4 f_c$, the plasma expulsion from the UH region is quenched as well as the 379 DM generation. The AA, and hence the striations developed slower, than at f_0 far from 380 gyroharmonic, but attained quite large values till the end of QCW pumping, the AA value 381 G achieved ~ 20 dB in DW frequency range $-150 \text{ kHz} < f_i - f_0 < 50 \text{ kHz}.$ 382

Assuming that the height of greatest plasma expulsion corresponds to maximum 383 PW energy consumption by ionospheric plasma we can conclude that for $f_0 < 4f_c$ the 384 most effective PW energy input occurs initially near the PW reflection height z_{r_0} . Then 385 the energy input becomes more effective near the UH height $z_{\rm UH}$. Later, when the plasma 386 expulsion develops in the whole interval $z_{\rm UH} \lesssim z \lesssim z_{r_0}$, some kind of mutual influence of these two separated regions is, presumably, observed. For $f_0 > 4f_c$ these two 387 388 isolated regions remained independent, although the UHD occupied greater altitude range. 389 For $f_0 \approx 4f_c$ the PW energy contributed mainly near z_{r_0} . The AIT excitation near the 390 UH height is suppressed. 391

Finally, combined investigations of the HF heated volume by MDS, SEE and AA allow establishing interconnection between different manifestations of the AIT, and determining position (altitude) of the most effective pump wave energy input in the HFpumped ionosphere as dependent on the offset between f_0 and nf_c .

The reasons for observed plasma expulsion is the enhancement of the gas-kinetic pressure due to electron heating and of the averaged high-frequency (ponderomotive) pressure, the enhancement is conditioned by the excitation of plasma (L and UH) waves by the pump wave near its reflection and UH heights (Dimant Ya., 1989; Grach et al., 1989; Vas'kov & Dimant, 1989). However, for adequate description of the dynamics of the electron density profile modification near PW resonances, further theoretical efforts are required.

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Figure 1. Column (a): the Doppler frequency shifts (colors) vs. time t and diagnostic wave spectral component frequency shift f_i-f_0 . Column (b): normalized amplitude of diagnostic waves G (color) vs. t and frequency shift f_i-f_0 . Column (c): the SEE spectrograms. The PW frequencies f_0 are shown above b-panels. Double arrows show contaminated frequency range for phase data processing.



Figure 2. Column (a): temporal variations of the reflection heights of different spectral components of sounding pulses $\Delta z(f_i, t)$ for different cycles. Frequency step between shown spectral components $\Delta f_i = 60$ kHz starting from f_0 . Additional height shift 300 m between reflection heights of the successive shown frequencies f_i at t = 0 is added for clarity. Red and blue lines correspond the spectral components reflected from $z_{\rm UH}$ and z_r . Column (b): velocity of the sounding waves reflection height displacement $V_v = \partial \Delta z_r(f_i, t)/\partial t$ vs. time and frequency. Column (c): relative variations of electron density [N(t)-N(0)]/N(0) vs. height at t = 5 s (blue), 15 s (orange) 30 s (green) and 45 s (red) in the same cycles. Double arrow and dashed parts of lines show contaminated frequency range.

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HF Induced Modifications of the Electron Density Profile in the Earth's Ionosphere Using the Pump Frequencies Near the Fourth Electron Gyroharmonic

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Key Points:

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11	•	plasma profile modifications in the HF-pumped ionosphere with frequencies near
12		the fourth electron gyroharmonic are studied at HAARP
13	•	plasma expulsion from the reflection height and the upper hybrid height of the pump
14		wave were detected
15	•	The expulsion from the upper hybrid height had been strongly suppressed for the
16		pump frequency close to fourth electron gyroharmonic

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17 Abstract

We discuss results on plasma density profile modifications in the F-region ionosphere caused 18 by HF heating with the frequency f_0 in the range [(-150 kHz)–(+75 kHz)] around the 19 4^{th} electron gyroharmonic $4f_c$. The experiments were conducted at the HAARP facil-20 ity on June 2014. The multi-frequency Doppler Sounder (MDS) which measures the phase 21 and amplitude of reflected sounding radio waves complemented by the observations of 22 the Stimulated Electromagnetic Emission (SEE) were used for the diagnostics of the plasma 23 perturbations. We detected noticeable plasma expulsion from the reflection region of the 24 pumping wave, and from the upper hybrid region, the expulsion from the latter one had 25 been strongly suppressed for $f_0 \approx 4 f_c$. The plasma expulsion from the upper hybrid 26 region was accompanied with the sounding wave's anomalous absorption slower devel-27 oping for $f_0 \approx 4 f_c$. Also, slower developing and weaker expulsion was detected for the 28 height region between the pump wave reflection and upper hybrid altitudes. 29

³⁰ Plain Language Summary

We report first results of systematic studying electron density profile modifications 31 in the F-region of the ionosphere caused by HF heating with the pumping frequencies 32 in the range [(-150 kHz)–(+75 kHz)] near the 4th electron gyroharmonic at the HAARP 33 heating facility. We measured phase and amplitude of ionospherically reflected sound-34 ing waves simultaneously with Stimulated Electromagnetic Emission. The phase data 35 allowed determining electron density profile evolution, the amplitude data provide in-36 formation of the pump-induced sounding waves' absorption, while SEE allows estimat-37 ing offset between the pump frequency and gyroharmonic. We detected noticeable plasma 38 expulsion from the reflection height and slower developing expulsion from the upper hy-39 brid height of the pump wave. Even slower developing and weaker expulsion was detected 40 for the height region between the pump wave reflection and upper hybrid heights. The 41 plasma expulsion from the upper hybrid region was accompanied with the sounding wave's 42 anomalous absorption, but was strongly suppressed for the pump wave frequency close 43 to fourth electron gyroharmonic, when the prominent "Downshifted Maximum" in SEE 44 was quenched. The combined phase sounding and SEE measurements allowed obtain-45 ing interconnection between different manifestations of the HF-induced turbulence and 46 determining altitude of the most effective pump wave energy input to ionospheric plasma. 47

48 1 Introduction

A powerful O-mode electromagnetic pump wave transmitted from the ground into 49 the bottom-site ionospheric F region excites a wide range of plasma processes leading 50 to the appearance of artificial ionospheric turbulence (AIT), i.e., generation of different 51 HF and LF plasma modes, plasma density inhomogeneities of scales from tens of cen-52 timeters to kilometers, causes electron heating, electron acceleration, ionization, gener-53 ation of ionospheric airglow etc. (Gurevich, 2007; Grach et al., 2016; Streltsov et al., 2018). 54 Diverse diagnostic methods and tools are used for the AIT studying, particularly, sound-55 ing of the heated volume of the ionosphere by diagnostic waves and a registration of sec-56 ondary, or stimulated, emission (SEE) in different frequency ranges. 57

The pump-plasma interaction is known to be strongest near the pump reflection height z_{r_0} at which $fp(z_{r_0})$ equals the pump frequency f_0 , and near the upper hybrid (UH) resonance height z_{UH} where $f_p(z_{UH}) = (f_0^2 - f_c^2)^{1/2}$ [here $fp = (e^2 N/\pi m)^{1/2}$, and $f_c = eB/(2\pi mc)$ are the electron plasma frequency and the electron cyclotron frequency, e and m are the electron charge and mass, N the electron density, c the speed of light, B the geomagnetic field strength]. This corresponds to existing theoretical concepts and is confirmed by investigations of the HF-pumped ionospheric volume by multifrequency Doppler sounding (MDS) at SURA and EISCAT heating facilities (Vaskov et al., 1986; Berezin et al., 1991; Lobachevsky et al., 1992; Grach et al., 1997; A. Shindin et al., 2012), which have revealed plasma expulsion from the resonance regions.

Stimulated electromagnetic emission (SEE) with frequencies f_{SEE} close to the pump 68 wave frequency f_0 occurs due to conversion of HF pump-driven electrostatic plasma modes, 69 most notably Langmuir (L) and upper hybrid (UH), into electromagnetic waves smaller 70 than the reflected pump wave (PW) by 50-90 dB (Thidé et al., 1982; Stubbe et al., 1984; 71 Leyser et al., 1989) and provides rich information about AIT (Grach et al., 2016). SEE 72 spectral characteristics depend on $\Delta f_c = f_0 - nf_c$, the offset of the pumping wave (PW) 73 frequency f_0 from the multiple electron gyroresonance nf_c , the most dramatic changes 74 occur during transition of f_0 via nf_c (e.g. from $f_0 < nf_c$ to $f_0 > nf_c$) and allows one 75 to estimate Δf_c during the experiment (Leyser et al., 1993; Leyser, 2001; Frolov et al., 76 2001; Sergeev et al., 2006; Grach et al., 2008, 2016). 77

We report results of the first experiments using MDS (phase) sounding of the HF-78 pumped ionosphere at the HAARP heating facility, located near Gakona, Alaska, USA 79 (62.40°N, 145.15°W) performed on June 2014. Simultaneously the SEE was monitored, 80 and the anomalous absorption (AA) of sounding waves was measured. The heating fa-81 cility was used both for the pump wave radiation and as pulsed Doppler HF radar. The 82 main purpose of the experiments was to study the dependence of HF-pump-induced elec-83 tron density expulsion from the resonance regions (in correlation with the AA and SEE) 84 on the offset of the pump wave frequency from the fourth gyroharmonic, $\Delta f_c = f_0 - 4f_c$, 85 the experiment was performed for.

⁸⁷ 2 Experimental Setup

Experiments on phase-amplitude sounding of the ionosphere heated volume were 88 performed on June 4. During time interval 14:55-16:25 AKDT the HAARP transmit-89 ter radiation schedule was organized as a sequence cycles at different PW frequencies f_0 90 varied from 5540 till 5730 kHz. The choice of the PW frequency nominals was conditioned 91 by covering available range around $4f_c$. Each cycle was organized as follows. During 4.5 92 min the HAARP operated as pulsed Doppler HF radar. The transmitters radiated ver-93 tically low-duty cycle diagnostic waves (DW): short (20 µs) pulses with the interpulse 94 period T = 100 ms at two carrier frequencies $f_{\rm DW} = f_0$ and $f_{\rm DW} = f_0$ -200 kHz with 95 effective radiated power (ERP) $P_{\rm ef} \sim 400$ MW each. After 30 seconds the radar mode 96 was combined for 45 seconds with quasi-continuous wave (QCW) pumping mode, i.e. high 97 duty cycle pump wave (pulse width $\tau = 70$ ms, the same interpulse period T = 10098 ms) at a frequency f_0 with the same ERP. During the QCW, the short pulses were ra-99 diated at 20th ms of the 30 ms pauses. After switching off the radar mode, the 30 s pause 100 of the HAARP transmitters was used for taking ionograms and changing the PW and 101 DW frequencies. Then the 5-min cycle was repeated at the new f_0 and f_{DW} . The power 102 of the 20 µs diagnostic pulses was sufficient to create a wide spectrum of DW (up to 300 103 kHz near each carrier frequency). An average DW power $\langle P \rangle = P/Q = 80$ kW was 104 far below the thresholds of the generation and maintenance of the pump-induced ther-105 mal plasma instabilities in the ionosphere (Grach et al., 1981). Also, the diagnostic pulses 106 were too short to excite ponderomotive instability in the ionosphere (Frolov et al., 2007; 107 Sergeev et al., 2007). 108

Under the combined radiation mode, the QCW created a perturbation in the iono-109 sphere, particularly at the plasma resonance (reflection altitude of the pump wave) and 110 at the UH resonance, while the DW simultaneously provided diagnostics of AIT (Vaskov 111 et al., 1986; Berezin et al., 1991; Grach et al., 1997; Sergeev et al., 2007, 2016; A. Shindin 112 et al., 2012). The high power of the HAARP transmitters used for MDS, applying broad-113 band radio receiver and specially developed signal processing algorithms have allowed 114 studying evolution of amplitude (A_i) and phase (φ_i) of the various spectral components 115 of the reflected DW (with frequencies f_i), which passed the perturbed region twice, in 116

a wide frequency range $\Delta f_{\text{total}} \sim 600 \text{kHz}$, (-450 kHz $\lesssim f_i - f_0 \lesssim +150 \text{ kHz}$), and 117 therefore, in a wide (till 25-35 km) altitude interval. Sometimes during QCW we could 118 analyze spectral components of the radar pulses in smaller range $\Delta f_{\rm QCW} \sim 360$ kHz 119 (-280 kHz $\lesssim f_i - f_0 \lesssim +80$ kHz. Outside of this range reflected signal fell till the back-120 ground noise due to the anomalous absorption (AA, see below), and signal amplitude 121 and phase analysis gave condemned results. The frequency resolution used for the anal-122 ysis was 1 kHz; the temporal resolution was determined by the interpulse period T =123 100 ms. The observational site were located, nearly under the heated region during in-124 jections at vertical. A 30-m folded-dipole BWDS receiver antenna was used in measure-125 ments. The receiver digitized a band 850 kHz, the dynamic range of the instruments af-126 ter spectral processing is estimated to be better than 90 dB. 127

Figures 1 and 2 present results obtained during the cycles at $f_0 = 5540, 5600, 5660$ 128 and 5730 kHz. Figure 1 shows the temporal evolution of Doppler frequency shifts $f_{di}(f_i, t) =$ 129 $d\varphi_i/dt$ (column a, φ_i is the phase incursion), normalized intensities $G_i(f_i, t)$ of the re-130 flected DW spectral components (column b), and SEE spectrograms (column c). The DW 131 intensities $A_i^2(f_i, t)$ are normalized to the $\langle A_i^2 \rangle$, the intensity averaged over 30 s before 132 the QCW switching on, $G_i[dB] = 10 \log(A_i^2/\langle A_i^2 \rangle)$) characterizes the anomalous ab-133 sorption (AA) due to scattering of DW into plasma (UH) waves on small-scale plasma 134 density irregularities (striations). In Fig. 2b we used running averaging over 5 pulses (0.5)135 s). In the spectrograms, prominent spectral SEE features are marked. There are the L-136 related ponderomotive narrow continuum (NCp) and UH-related downshifted maximum 137 (DM), 2DM, upshifted maximum (UM), thermal narrow continuum (NCt), broad con-138 tinuum (BC) and broad upshifted maximum (BUM). Stationary peculiarities and tem-139 poral evolution of different SEE features are well described in numerous papers, see, e.g. 140 (Frolov, 2004; Leyser et al., 1993, 1994; Leyser, 2001; Carozzi, 2002; Thidé et al., 2005; 141 Sergeev et al., 2006; Kotov et al., 2008; Grach et al., 2008, 2016), etc. The SEE can be 142 used for rough estimations of the offsets between f_0 and $4f_c$, $\Delta f_c = f_0 - 4f_c$, in the 143 range between successive gyroharmonics. Most precisely this can be done for $f_0 \approx 4 f_c$ 144 and $f_0 \gtrsim 4f_c$ (see formulas (7) and (8) below). 145

¹⁴⁶ 3 Phase Data Processing

For all f_0 , when heating (QCW) is turned on, as a rule, an increase of all f_i (pos-147 itive Doppler frequency shifts f_{di} is often observed. It is less pronounced for narrow ranges 148 of the diagnostic (sounding) waves frequencies with reflection heights near PW reflec-149 tion and UH heights z_{r0} and zUH, i.e. for $f_i \sim f_p(z_{r0})$ and $f_i \sim f_p(z_{UH})$. Negative 150 f_{di} are observed after the heating turns off. Such pump-induced changes of f_{di} occur on 151 the background of natural variations of the reflection heights (and f_{di}), determined by 152 the motion of the ionosphere. The measured temporal evolution of Doppler frequency 153 shifts $f_{di}(t) = \mathrm{d}\varphi_i/\mathrm{d}t$ for different f_i (Figure 1, panels a) provided data for the recon-154 struction of the electron density profile temporal evolution N(z,t) in the HF-pumped 155 ionosphere by solving the inverse problem of the phase sounding. 156

¹⁵⁷ Under geometric optics approximation, each of the sounding waves at the angu-¹⁵⁸ lar frequency $\omega_i = 2\pi f_i$ propagating from the ground up to the reflection points $z_r(f_i)$ ¹⁵⁹ and back to the ground, experiences the phase incursion (Ginzburg, 1970):

$$\varphi(\omega,t) = \frac{2\omega}{c} \int_0^{z_r} n(\omega,\omega_p(z,t)) dz - \frac{\pi}{2}$$
(1)

where $\omega_p(z,t) = 2\pi f_p(z,t)$ is the plasma frequency, $n(\omega, \omega_p(z,t))$ is the wave refractive index. The reflection altitude z_r is determined by the condition n = 0, and t is the time. This can be translated in the following expression for an additional phase change $\Delta \varphi(\omega) = \varphi(\omega, t_0) - \varphi(\omega, t)$ in the time interval $[t_0, t]$ associated with perturbation of the profile N(z,t) due to ionosphere HF-pumping or to natural reasons:

$$y(\omega) = \frac{c}{2\omega} \Delta \varphi(\omega) = \int_{\omega_{\min}}^{g(\omega)} K(\omega, \omega_p) \Delta z(\omega_p) d\omega_p$$
(2)

Here $K(\omega, \omega_p) = dn(\omega, \omega_p)/d\omega_p$ is a kernel of the integral equation (2), $g(\omega)$ is the an-167 gular plasma frequency at the reflection point, which is $g(\omega) = \omega$ for an ordinary wave, 168 t_0 is initial time, $\Delta z(\omega_p, t) = z(\omega_p, t) - z(\omega_p, t_0)$ is the altitude shift, i.e., the difference 169 between the sounding radio wave reflection altitude at the current (t) and initial (t_0) times. 170 Here the variable of the integration is replaced from the altitude z in (1) to the plasma 171 frequency ω_p in (2). It is taken into account that at the reflection point $n(\omega, g(\omega)) =$ 172 0, and at the entrance to the plasma layer $\Delta z = 0$. The left hand side $y(\omega)$ in (2) is 173 to be determined from the experimental data as $\Delta \varphi_i = \int_{t_0}^t f_{di}(t') dt'$. On the base of 174 the data obtained, an array $\Delta \Phi(\omega, t) = \Delta \varphi(\omega, t) - \Delta \varphi(\omega_{\min}, t)$ was created for the phase 175 shifts $\Delta \varphi(\omega, t)$ of the diagnostic waves. Here ω_{\min} is least of the probe wave frequencies, 176 in our experiment we have taken $\omega_{\min} = \omega_0 - 2\pi \cdot 280$ kHz. The probe wave at this frequency was reflected noticeably below (~ 25 - 35 km) z_{UH} , and we assumed that the 177 178 phase evolutions for this and lower frequencies do not depend on the pump-induced pro-179 cesses in the plasma resonance regions. This is confirmed, in particular, by measurements 180 of the AA bandwidth of the DW (Fig. 1b). Also, the subtraction of $\Delta \varphi(\omega_{\min}, t)$ allows 181 to exclude processes responsible for plasma density variations at lower altitudes caused 182 by violation of the balance between ionization and recombination in the lower ionosphere. 183

For the exact expression of the refractive index for magnetized plasma, the equation (2) cannot be reduced to the analytically solvable integral equation. In this case, we had applied in (A. Shindin et al., 2012; Sergeev et al., 2016) the regularization algorithms by Tikhonov (1995) and solved the inverse problem numerically. In this paper, instead of numerical Tikhonov algorithm, we used an approximate expression for well describing ordinary wave refractive index near the reflection point for HAARP experimental conditions,

$$n(\omega, \omega_p) \approx \left(1 - \frac{\omega_p^2}{\omega^2}\right)^{\beta}, \ \beta \approx 0.29.$$
 (3)

¹⁹² Substituting (3) into (2), we obtain the generalized Abel equation (Korn & Korn, 2000):

$$\int_{\omega_{\min}^2}^{\omega^2} \frac{\Delta z}{(\omega^2 - \omega_p^2)^{1-\beta}} d\omega_p^2 = F(\omega), \ F(\omega) = -c \frac{\omega^{2\beta-1}}{2\beta} \Delta \varphi(\omega)$$
(4)

¹⁹⁴ with the solution

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$$\Delta z = \frac{\sin\{\pi(1-\beta)\}}{\pi(1-\beta)} \int_{\omega_{\min}^2}^{\omega^2} \frac{F'(\omega_p^2)}{(\omega^2 - \omega_p^2)^\beta} \mathrm{d}\omega_p^2.$$
(5)

¹⁹⁶ Dynamics of the reflection altitude shifts of different DW spectral components, $\Delta z_r(\omega_i = 2\pi f_i, t)$ is displayed in Figure 2, panels a. In this figure t = 0 corresponds to QCW switch-¹⁹⁷ ing on, blue and red lines correspond to the reflection altitude of the DW at $f_i = f_0$, ¹⁹⁹ i.e. the PW reflection altitude, $(z(f_i) = z_r(f_0))$ and to the reflection altitude of the DW ²⁰⁰ at $f_i = f_p(z_{\rm UH}) = (f_0^2 - f_c^2)^{1/2}$, i.e. reflection from the pump wave UH resonance height, ²⁰¹ respectively. Chosen frequency step between neighbor spectral components displayed in ²⁰² the Figure 2a is $\Delta f = 30$ kHz. For clarity we introduce additional height shift 300 m ²⁰³ between reflection altitude shifts of the successive DW spectral components at t = 0.

Temporal variations of the reflection heights $\Delta z(f_i, t)$ allows to calculate velocities of the vertical motion of plasma density at a certain magnitude $Ni = \pi f_i^2 m/e^2$ as $V_v = \partial \Delta z_r(f_i, t)/\partial t$ at different f_i . The velocities of the sounding waves reflection height displacement vs. time and sounding wave frequency (pulse spectral component) $V_v(f_i, t)$ are presented in Figure 2b. Positive (red) velocity values correspond to upward motion of a certain plasma density level while negative (blue) values correspond to downward motion. Calculating the altitude shifts $\Delta z_r(f_i, t)$ allows us observing the evolution of the electron density profile N(z,t) from the reference one. The reference profile $N_0 = N(z,t_0)$ is taken from ionogram registered prior to the QCW pumping session. For this we shall transform $\Delta z_r(f_i, t)$ to $\Delta z_r(N, t)$ by using the univocal relation between the plasma frequency at the radio wave reflection point and electron density. Therefore,

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$$z(N, N_0, t) = z(N_0) + \Delta z(N, t)$$
(6)

is the dependence of the reflection height of the radio wave on the density. Then we find the required distribution N(z,t) by calculating the inverse of (6). Relative differences between reconstructed and reference electron density profiles $\delta N(z,t) = [N(z,t)-N_0]/N(z,t_0)$ vs. altitude z for all f_0 at 5th, 15th, 30th and 45th seconds of pumping are presented in Figure 2c.

Figure 2 illustrates the evolution of the electron density in the HF-pumped ionosphere in the altitude range from region above the reflection height till the region below the UH height.

4 Analysis of the Combined Phase sounding, AA, and SEE Data

Let's start an analysis from $f_0 = 5540$ kHz, the upper row of figures 1 and 2. During this cycle the pump frequency $f_0 = 5540$ kHz belongs to the "weak emission range" between 3rd and 4th gyroharmonic, where SEE spectrum contains weak BC, DM and UM (Fig. 1c). The offset $\Delta f_c = f_0 - 4f_c$ here can be roughly estimated from ionograms and geomagnetic field IGRF model as $\Delta f_c \sim -(130 - 150)$ kHz.

From Figure 1a, it is seen that immediately after QCW is switched on the Doppler 231 frequency shifts were mainly positive $(f_{di} > 0)$ for all spectral components except of 232 ones close to the pump frequency $(f_i \approx f_0)$ where $f_{di} < 0$. Such a behavior of f_{di} is 233 translated to increasing reflection altitudes of the sounding waves z_{ri} with the frequen-234 cies close to the pump wave frequency f_0 (Fig. 2a) with velocities V_v up to 100 m/sec, 235 and, to the contrary, slighter decreasing z_r for both $f_i > f_0$ and $f_i < f_0$ with veloci-236 ties V_v up to 60 m/s (Fig. 2b). This corresponds to electron density decrease in the vicin-237 ity of the pump wave reflection height $z_{r_0} = z_r(f_0)$ and increase on other heights, i. e. 238 to the plasma expulsion from the reflection point vicinity. The relative density deple-239 tion (hereafter reflection depletion, RD) reaches up to $\sim 0.1 - 0.2\%$ at 5th second of 240 the QCW pumping (Fig. 2c). According to Fig. 2a, the uplifting $\Delta z_r(f_i \approx f_0)$ grows 241 from 0.1 s till (3-5) s, reaching 100-300 m, and then slows down till t = 10-15 s. The 242 SEE feature NCp appears immediately after QCW switching on, simultaneously with 243 the start of the plasma expulsion from the vicinity of z_{r_0} and then exhibits strong overshoot-244 effect: its spectral width and intensity noticeably drops during the raise of $z_r(f_0)$ and 245 growth of the AA and UH-related SEE features (Fig. 1b, 1c). Such initial behavior of 246 the $f_d(f_i, t)$, $\Delta z_r(f_i, t)$, $V_v(f_i, t)$, NCp and AA is qualitatively similar for all f_0 and does 247 not depend on the offset Δf_c . 248

Few seconds later, in t = 1-3 s, phenomena related to excitation of the UH waves 249 and striations, resulting in the phenomena such as the AA, and the UH related SEE fea-250 tures (DM, UM, and BC), and the plasma expulsion from the UH height region ($z \sim$ 251 $z_{\rm UH}$, the "UH depletion", hereafter UHD) developed simultaneously. The expulsion cor-252 responded to appearance of the expanding range with $f_{di} < 0$ in Fig. 1a. At 3-5 s the 253 UHD became deeper than the RD, and then the UHD developed monotonously till QCW 254 pumping switchies off. At t 15 s the uplifting near the reflection point accelerated again 255 and continued till the QCW stops, for s deepening of the UHD and RD is accompanied 256 by plasma density decrease in the whole altitude range (Fig. 2c). 257

For the cycle at $f_0 = 5600$ kHz (Figs. 1, 2, second row) the results of AA measurements (temporal development and magnitude) and the phase sounding analysis ($f_d(f_i, t)$, ²⁶⁰ $\Delta z_r(f_i, t), V_v(f_i, t)$ and $\delta N(z, t)$) were similar to ones for $f_0 = 5540$ kHz, but their SEE ²⁶¹ were fairly differed. According to the SEE spectrogram (Fig. 1c), the PW frequency be-²⁶² longs to "below harmonic" range. It shows strong DM and resolved 2DM. The offset can ²⁶³ be roughly estimated as $\Delta f_c \sim -(60 - 70)$ kHz. According to the ionograms, $z_{\rm UH}$ in ²⁶⁴ the cycles at 5540 and 5600 kHz were close, $z_{\rm UH} \sim 225$ km.

Note, during these cycles the average anomalous absorption $\langle G \rangle$ in the range 170 $< f_0 - f_i < 220$ kHz decreased by $\sim 10 - 15$ dB (Figs 1b) in comparison with the range $-30 < f_0 - f_i < 150$ kHz. The most probably it can be attributed to focusing of HF diagnostic radio waves reflected from the UH altitudes with density depletion (Vaskov et al., 1986).

In the cycle at $f_0 = 5660$ kHz (Figs. 1,2, 3rd row) the behavior of the bulk of investigated parameters differed noticeably from cycles at $f_0 = 5540$ and 5600 kHz.

The RD developed similar to cycles with other f_0 during first 10 s of QCW pump-272 ing, but did not slow down after 10 s of QCW pumping. Negative f_{di} appeared near $f_i(z_{r_0})$ 273 than the range of $f_{di} < 0$ expanded for mainly to lower f_i . Dependence looked like shallow 274 low quasi-periodic structure with a period $\sim 3-4$ km, amplitude growing in time and 275 decreasing downward from z_{r_0} , and occupying a height interval exceeding spacing be-276 tween z_r and $z_{\rm UH}$. The UHD and zUH uplifting did not resolving in this cycle, more-277 over, the weak descending of the UH height at surrounded frequencies were observed. 278 Then the AA developed slower than in other cycles, while attained the same magnitudes 279 till 45th s. The difference is due to the fact that during this cycle at $f_0 = 5660$ kHz the 280 pump wave frequency was much closer to the double resonance where $f_{\rm UH} \approx 4 f_c$. This 281 is seen from the SEE spectrogram (Fig. 1c). Here the DM is not resolved till t $\sim 10-$ 282 15 s which means that 283

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$$_{\rm DM} = f_0 - \Delta f_{\rm DM} \approx f_{\rm UH} \approx 4f_c \tag{7}$$

at the UH height (the "resonance range"), $f_{\rm DM}$ is the DM peak frequency. Then the DM 285 appeared which was probably attributed to amplification of the $\delta N(z,t)$, and changing 286 of the UH height, and therefore, a violation of (7). This allows to estimate the offset $\Delta f_c =$ 287 f_0-4f_c during the cycle as 7-15 kHz: initially $\Delta f_c \approx \Delta f_{\rm DM}$ and then changed. Detailed 288 analysis of the SEE peculiarities (DM, UM and BUM) near the double resonance can 289 be found in (Levser et al., 1994; Carozzi, 2002; Sergeev et al., 2006; Kotov et al., 2008; 290 Grach et al., 2008). Note that the altitude of the double resonance obtained from (7) 291 and IGRF model ~ 240 km exceeded $z_{\rm UH}$ obtained from the ionogram by 10 km. This 292 value fits into the error in determining the heights when processing ionograms. 293

f

The cycle at $f_0 = 5730$ kHz (Figs. 1,2, row 3) was the only cycle with $f_0 > 4f_c$, the offset $\Delta f_c = f_0 - 4f_c$ can be estimated from the SEE spectrogram as

$$\Delta f_c = f_0 - 4f_c \approx f_{\rm BUM} - f_0 = \Delta f_{\rm BUM},\tag{8}$$

 $\Delta f_c \sim +75$ kHz (Fig. 1c), fBUM is the BUM peak frequency. From Fig. 3 it can be 297 concluded the following. After QCW pumping switching on the plasma expulsion from 298 the vicinity of the reflection point (RD) and the NCp development in SEE spectrum be-299 haved similarly to all f_0 . Slowing down of the RD deepening after development of the 300 UH-related effects is were not observed. Like for $f_0 < 4f_c$, for $f_0 > 4f_c$ the UH-related 301 effects, namely AA; DM, 2DM, UM, SEE features; and UHD developed few seconds later 302 than L-related processes near the reflection height. The characteristics of this processes 303 during this cycle, however, differed from ones for $f_0 < 4f_c$. 304

First, the AA attained a saturation at $t \sim 10-15$ s, but for $f_0 > 4f_c$ at the saturation stage $\langle G \rangle \sim 25 - 30$ dB, while at $f_0 < 4f_c \langle G \rangle \sim 18 - 20$ dB, and the frequency range with strong AA is wider at $f_0 > 4f_c$ than at $f_0 < 4f_c$ (Figs 1b). Due to the strong AA, in the range 5590 $< f_i < 5650$ kHz the DW signal intensity fell below the background noise, and measurements of Doppler shifts/phase incursions and AA

were contaminated. This range is shown in Figures 1,2, row 4 by double arrows and dashed 310 parts of lines and shall be excluded from the analysis for t > 15 s. Second, the strong 311 NCt showed up in the SEE spectrogram at $\Delta f \sim 0 - (-7)$ kHz with a temporal be-312 havior similar to DM, 2DM and UM; after $t \sim 1.5$ s the NCt covered NCp. The DM, 313 2DM and UM developed concurrently with AA, exhibited the overshoot effect with max-314 ima at $t \sim 6-11$ s, and were more intensive then in other cycles. Such SEE behavior 315 is typical for the "above harmonic range" close to the "strong emission" range (Sergeev 316 et al., 2006). Third, the UHD started to develop ~ 2 km above the nominal upper hy-317 brid resonance height and occupied quite wide (> 5 km) altitude interval. Later, in $t \sim$ 318 10-15 s, the interval expanded (till ~ 8 km) and descended below the UH height. Dur-319 ing the whole 45 s of the QCW pumping the UHD remained shallower than the RD (Fig. 320 2c). 321

After the termination of the QCW, the signs of Doppler frequency shifts fdi and velocities $V_v = \partial \Delta z_r(f_i, t)/\partial t$ changed to the opposite ones (see Figures 1a and 2b). This leads to a reduction of the plasma density depletions around the reflection and UH heights, the depletions relaxed and disappeared in ~ 15-40 s after pump wave turns off.

327 5 Conclusions

The results can be briefly summarized as follows. It was obtained that during all 328 cycles the pump wave-plasma interaction developed most quickly (in a few milliseconds) 329 after QCW switching on in the vicinity of the pump wave reflection height z_{r_0} . It is ac-330 companied by the plasma expulsion from the interaction region (RD appearance) and 331 by the NCp SEE feature generation. At this time there were no essential differences be-332 tween cycles with different f_0 . Both the expulsion and NCp shall be attributed to the 333 excitation of L waves due to the ponderomotive parametric instability near the PW re-334 flection height z_{r_0} . 335

Later, in a 1-5 s after the QCW was switched on, for $f_0 < 4f_c$, the plasma expul-336 sion from a vicinity of the upper hybrid height $z_{\rm UH}$ (UHD development) began along with 337 development of the AA, UH-related SEE features such as DM, 2DM, UM and BC, as 338 well as with suppression (overshoot) of the NCp feature and slowing down of expulsion 330 from the vicinity of z_{r_0} . At 3-10 s the UHDs became deeper than the RDs. The expul-340 sion from upper hybrid height continues until the end of 45-s-long QCW-pumping. All 341 these effects are related to excitation of the striations and UH plasma waves. The slow-342 ing down of the z_{r_0} uplifting, RD deepening and NCp suppression indicated that the UH-343 related processes leaded to the noticeable shielding of the reflection point from the pump 344 wave energy. A sequence of the described effects is consistent with a general scenario of 345 the phenomena developing in the HF-pumped ionosphere (Frolov et al., 1997; Thidé et 346 al., 2005; Grach et al., 2016) and is clearly illustrated by Sergeev et al. (2016), where the 347 results of 3 successive 2-minute cycles at the frequency $f_0 = 5500$ kHz obtained on the 348 same experiment (2014/06/04, 15:40-15:46 AKDT) were presented. Similar results were 349 obtained also at the SURA facility (A. Shindin et al., 2012). In the described experiment 350 during the cycles at $f_0 < 4f_c$ the RD and NCp develop similarly to the cycles at $f_0 =$ 351 5500 kHz till t= 5-10 s, but later the RD started to deepen again, concurrently with UHD 352 deepening on the background of plasma density decrease in the interval $z_{\rm UH} \lesssim z \lesssim z_{r_0}$. 353 A dependence $\delta N(z)$ looked as two isolated minima, close, respectively, to the reflection 354 and UH heights. Note that for presented cycles at $f_0 < 4f_c$ (with approximately same 355 initial UH heights ~ 225 km) the expulsion parameters $\Delta z(f_i)$ and $\Delta N(z)$ as well the 356 AA behaved similarly, even quantitatively, while the SEE spectra are different. This points 357 on a weak dependence on the AIT peculiarities on in the range $-150 \lesssim \Delta f_c \lesssim -60$ 358 kHz. 359

In the single cycle with $f_0 > 4f_c$ ($f_0 = 5730$ kHz, $\Delta f_c \sim 75$ kHz) the $\delta N(z)$ de-360 veloped, again, as two isolated minima, RD and UHD. In this cycle the UHD depth re-361 mained less than the RD depth during the whole OCW pumping interval; the altitude 362 ranges, occupied by depletions, were wider, the AA of the diagnostic pulses (DW) was 363 stronger (by ~ 10 dB) and occupied larger frequency range than at $f_0 < 4f_c$. The up-364 lifting of the diagnostic waves (DW) reflection heights z_{ri} near the PW UH height started 365 after ~ 15 s delay (Fig. 3c, row 4) after QCW switch on. This points to the essential 366 difference in the AIT excitation for $f_0 > 4f_c$ and $f_0 < 4f_c$. 367

For the cycle at $f_0 = 5660$ kHz the PW frequency was close to $4f_c$, i.e. got into 368 the "resonance range", where the DM was totally suppressed or very weak, but the BUM 369 and UM were present in the SEE spectra. Taking $f_{\rm DM} = f_0 - \Delta f_{\rm DM} \approx 4 f_c(z_{\rm UH}) \approx$ 370 5650 kHz and using IGRF model we obtain $z_{\rm UH} \approx 239$ km in the beginning of the cy-371 cle. In this cycle the DW frequency range with negative f_{di} expanded from the PW fre-372 quency f_0 (Fig. 1a, 3rd row), the temporal behavior of $\delta N(z,t)$ demonstrated deepen-373 ing quasi-periodic structure with a period $\sim 3-4$ km, with amplitude growing in time 374 and decreasing downward from z_{r_0} , and occupying height interval exceeding spacing be-375 tween z_{r_0} and $z_{\rm UH}$. Unfortunately, the total range of the diagnostic signals available for 376 the phase data processing was too narrow to estimate lower boundary of the height in-377 terval. "Independent" z_{ri} uplifting for $z_{ri} \sim z_{\rm UH}$ wasn't resolved for this cycle. There-378 fore, for $f_0 \approx 4 f_c$, the plasma expulsion from the UH region is quenched as well as the 379 DM generation. The AA, and hence the striations developed slower, than at f_0 far from 380 gyroharmonic, but attained quite large values till the end of QCW pumping, the AA value 381 G achieved ~ 20 dB in DW frequency range $-150 \text{ kHz} < f_i - f_0 < 50 \text{ kHz}.$ 382

Assuming that the height of greatest plasma expulsion corresponds to maximum 383 PW energy consumption by ionospheric plasma we can conclude that for $f_0 < 4f_c$ the 384 most effective PW energy input occurs initially near the PW reflection height z_{r_0} . Then 385 the energy input becomes more effective near the UH height $z_{\rm UH}$. Later, when the plasma 386 expulsion develops in the whole interval $z_{\rm UH} \lesssim z \lesssim z_{r_0}$, some kind of mutual influence of these two separated regions is, presumably, observed. For $f_0 > 4f_c$ these two 387 388 isolated regions remained independent, although the UHD occupied greater altitude range. 389 For $f_0 \approx 4f_c$ the PW energy contributed mainly near z_{r_0} . The AIT excitation near the 390 UH height is suppressed. 391

Finally, combined investigations of the HF heated volume by MDS, SEE and AA allow establishing interconnection between different manifestations of the AIT, and determining position (altitude) of the most effective pump wave energy input in the HFpumped ionosphere as dependent on the offset between f_0 and nf_c .

The reasons for observed plasma expulsion is the enhancement of the gas-kinetic pressure due to electron heating and of the averaged high-frequency (ponderomotive) pressure, the enhancement is conditioned by the excitation of plasma (L and UH) waves by the pump wave near its reflection and UH heights (Dimant Ya., 1989; Grach et al., 1989; Vas'kov & Dimant, 1989). However, for adequate description of the dynamics of the electron density profile modification near PW resonances, further theoretical efforts are required.

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Figure 1. Column (a): the Doppler frequency shifts (colors) vs. time t and diagnostic wave spectral component frequency shift f_i-f_0 . Column (b): normalized amplitude of diagnostic waves G (color) vs. t and frequency shift f_i-f_0 . Column (c): the SEE spectrograms. The PW frequencies f_0 are shown above b-panels. Double arrows show contaminated frequency range for phase data processing.



Figure 2. Column (a): temporal variations of the reflection heights of different spectral components of sounding pulses $\Delta z(f_i, t)$ for different cycles. Frequency step between shown spectral components $\Delta f_i = 60$ kHz starting from f_0 . Additional height shift 300 m between reflection heights of the successive shown frequencies f_i at t = 0 is added for clarity. Red and blue lines correspond the spectral components reflected from $z_{\rm UH}$ and z_r . Column (b): velocity of the sounding waves reflection height displacement $V_v = \partial \Delta z_r(f_i, t)/\partial t$ vs. time and frequency. Column (c): relative variations of electron density [N(t)-N(0)]/N(0) vs. height at t = 5 s (blue), 15 s (orange) 30 s (green) and 45 s (red) in the same cycles. Double arrow and dashed parts of lines show contaminated frequency range.

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