

True Gravity in Atmospheric Ekman Layer Dynamics

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Abstract

True gravity is a three-dimensional vector, $\mathbf{g} = ig\lambda + jg\phi + kgz$, with (λ, ϕ, z) the (longitude, latitude, height) and (i, j, k) the corresponding unit vectors. The vertical direction is along \mathbf{g} , not along \mathbf{k} , which is normal to the Earth spherical (or ellipsoidal) surface (called deflected-vertical). Correspondingly, the spherical (or ellipsoidal) surfaces are not horizontal surfaces (called deflected-horizontal surfaces). In the (λ, ϕ, z) coordinates, the true gravity \mathbf{g} has longitudinal-latitude component, $g_h = ig\lambda + jg\phi$, but it is neglected completely in meteorology through using the standard gravity $(-g_0\mathbf{k}, g_0 = 9.81 \text{ m/s}^2)$ instead. Such simplification on the true gravity \mathbf{g} has never been challenged. This study uses the atmospheric Ekman layer as an example to illustrate the importance of g_h . The standard gravity $(-g_0\mathbf{k})$ is replaced by the true gravity \mathbf{g} in the classical atmospheric Ekman layer equation with a constant eddy viscosity (K) and a height-dependent-only density $\rho(z)$ represented by an e -folding stratification. New formulas for the Ekman spiral and Ekman pumping are obtained. The second derivative of the gravity disturbance (T) versus z , also causes the Ekman pumping, w , in addition to the geostrophic vorticity with Δz the Ekman layer thickness and f the Coriolis parameter. With data from the EIGEN-6C4 static gravity model, the global mean strength of the Ekman pumping due to the true gravity is found to be 4.0 cm s^{-1} . Such evidently large value implies the urgency to include the true gravity \mathbf{g} into the atmospheric dynamics.

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Key Points:

- Vertical that meteorologists think is not vertical. True gravity $\mathbf{g}(\lambda, \varphi, z)$ represents vertical with latitudinal-longitudinal components
- Replacement of the standard gravity $-g_0\mathbf{k}$ ($g_0=9.81 \text{ m s}^{-2}$) by $\mathbf{g}(\lambda, \varphi, z)$ in the classical Ekman layer equation leads to a new solution
- With data from the EIGEN-6C4 static gravity model, the global mean strength of the Ekman pumping due to $\mathbf{g}(\lambda, \varphi, z)$ is 4.0 cm s^{-1} .

14 Abstract

15 True gravity is a three-dimensional vector, $\mathbf{g} = \mathbf{i}g_\lambda + \mathbf{j}g_\phi + \mathbf{k}g_z$, with (λ, ϕ, z) the (longitude, latitude,
 16 height) and $(\mathbf{i}, \mathbf{j}, \mathbf{k})$ the corresponding unit vectors. The vertical direction is along \mathbf{g} , not along \mathbf{k} ,
 17 which is normal to the Earth spherical (or ellipsoidal) surface (called deflected-vertical).
 18 Correspondingly, the spherical (or ellipsoidal) surfaces are not horizontal surfaces (called
 19 deflected-horizontal surfaces). In the (λ, ϕ, z) coordinates, the true gravity \mathbf{g} has longitudinal-
 20 latitudinal component, $\mathbf{g}_h = \mathbf{i}g_\lambda + \mathbf{j}g_\phi$, but it is neglected completely in meteorology through using
 21 the standard gravity $(-g_0\mathbf{k}, g_0 = 9.81 \text{ m/s}^2)$ instead. Such simplification on the true gravity \mathbf{g} has
 22 never been challenged. This study uses the atmospheric Ekman layer as an example to illustrate
 23 the importance of \mathbf{g}_h . The standard gravity $(-g_0\mathbf{k})$ is replaced by the true gravity \mathbf{g} in the classical
 24 atmospheric Ekman layer equation with a constant eddy viscosity (K) and a height-dependent-only
 25 density $\rho(z)$ represented by an e-folding stratification. New formulas for the Ekman spiral and
 26 Ekman pumping are obtained. The second derivative of the gravity disturbance (T) versus z , also
 27 causes the Ekman pumping, $(D_E / 2\pi f)\partial^2 T / \partial z^2$, in addition to the geostrophic vorticity with D_E
 28 the Ekman layer thickness and f the Coriolis parameter. With $\partial^2 T / \partial z^2$ data from the EIGEN-6C4
 29 static gravity model, the global mean strength of the Ekman pumping due to the true gravity is
 30 found to be 4.0 cm s^{-1} . Such evidently large value implies the urgency to include the true gravity
 31 \mathbf{g} into the atmospheric dynamics.

32 Plain Language Summary

33
 34 Meteorologists use the spherical (or ellipsoidal) surfaces represented by latitude (ϕ) and longitude
 35 (λ) as the horizontal and the direction normal to them represented by height (z) as the vertical. It is
 36 not correct since the vertical direction is represented by the true gravity $\mathbf{g}(\lambda, \phi, z)$; and the horizontal
 37 surfaces are the equipotential surfaces of $\mathbf{g}(\lambda, \phi, z)$ such as the geoid surface which is nearest to
 38 the Earth spherical (or ellipsoidal) surface ($z = 0$). In the (λ, ϕ, z) coordinates, the true gravity
 39 $\mathbf{g}(\lambda, \phi, z)$ has latitudinal and longitudinal components, which are neglected completely in
 40 meteorology. This study uses the atmospheric Ekman layer dynamics and the true gravity data
 41 from the EIGEN-6C4 static gravity model as an example to show the importance of using the true
 42 gravity $\mathbf{g}(\lambda, \phi, z)$ in the atmospheric dynamics.

43 1 Introduction

44 Meteorologists usually use the Earth-fixed coordinate system with (λ, ϕ, z) representing the
 45 longitude, latitude, and spherical normal (or height) with $(\mathbf{i}, \mathbf{j}, \mathbf{k})$ the corresponding unit vectors.
 46 The unit vector \mathbf{k} does not represent the vertical direction since the Earth true gravity $\mathbf{g} (= g_\lambda\mathbf{i} +$
 47 $g_\phi\mathbf{j} + g_z\mathbf{k})$ represents the vertical. We may call the direction of \mathbf{k} the deflected-vertical. The angle
 48 between $-\mathbf{k}$ and \mathbf{g} is the vertical deflection. The spherical (or ellipsoidal) surfaces are not the
 49 horizontal surfaces since the equipotential surfaces of \mathbf{g} such as the geoid surface represent the
 50 horizontal surfaces. We may call the spherical (or ellipsoidal) surfaces the deflected-horizontal
 51 surfaces. Appendix A provides difference between the oblate spheroid (ellipsoidal) coordinates
 52 versus polar spherical coordinates.

53 The turbulent mixing in atmospheric planetary boundary layer is treated as a diffusion
 54 process similar to molecular diffusion, with an eddy viscosity K , which is many orders of
 55 magnitude larger than the molecular viscosity. The turbulent mixing generates ageostrophic wind
 56 (called the Ekman spiral), decaying by an e-folding over a height as the wind vector rotate to the

57 right (left) in the northern (southern) hemisphere through one radian (Ekman, 1905). Along with
58 the Ekman spiral, several important processes such as Ekman pumping can be identified.

59 As in other atmospheric dynamics, the Ekman theory was established using the standard
60 gravity ($-g_0\mathbf{k}$, $g_0 = 9.81 \text{ m s}^{-2}$), rather than the true gravity \mathbf{g} (Pedlosky, 1987; Holton, 2004). Its
61 longitudinal-latitudinal component, $\mathbf{g}_h (= g_\lambda\mathbf{i} + g_\varphi\mathbf{j})$, is neglected completely. Use of the standard
62 gravity ($-g_0\mathbf{k}$) instead of the true gravity \mathbf{g} is based on the comparison that the strength of the
63 deflected-vertical component $|g_z|$ is 5-6 orders of magnitude larger than the strength of the
64 deflected-horizontal gravity $|\mathbf{g}_h|$. This comparison is unphysical because such a huge difference in
65 magnitude between the components in \mathbf{k} and in (\mathbf{i}, \mathbf{j}) also occurs in the pressure gradient force in
66 large-scale atmospheric dynamics. But, the pressure gradient force in (\mathbf{i}, \mathbf{j}) is never neglected
67 against the pressure gradient force in \mathbf{k} . Thus, the feasibility of using the standard gravity ($-g_0\mathbf{k}$) in
68 meteorology needs to be investigated. The Ekman dynamics provides a theoretical framework for
69 such a study.

70 The rest of the paper is outlined as follows. Section 2 presents the dynamic equation with
71 the true gravity for atmospheric Ekman layer. Section 3 shows the Ekman layer solution and a new
72 equation for the Ekman pumping due to the use of the true gravity. Section 4 describes the data
73 source (EIGEN-6C4 model) of the second derivative of the disturbing static gravity potential ($\partial^2 T / \partial z^2$), and
74 shows the Global Ekman pumping velocity due to $\partial^2 T / \partial z^2$ using the EIGEN-6C4
75 data. Section 5 shows the feasibility of using the (λ, φ, z) coordinates. Section 6 presents the
76 conclusions. Appendices A-D present the two Earth coordinate systems and the basic information
77 about the true gravity \mathbf{g} and related disturbing static gravity potential T .

78 2 Dynamic Equation with the True Gravity

79 Steady-state large-scale atmospheric dynamic equation with the Boussinesq approximation
80 (replacement of density ρ by a constant ρ_0 except ρ being multiplied by the gravity and
81 incompressibility) is given by (Chu, 2021)

$$82 \quad \rho_0 [2\boldsymbol{\Omega} \times \mathbf{U}] = -\nabla_3 p + \rho \mathbf{g} + \rho_0 \mathbf{F} \quad (1a)$$

$$83 \quad \nabla \cdot \mathbf{U} + \frac{\partial w}{\partial z} = 0 \quad (1b)$$

84 if the pressure gradient force, true gravity \mathbf{g} (see Appendices B and C), and friction are the only

85 real forces. Here, $\nabla_3 \equiv \mathbf{i} \frac{1}{R \cos \varphi} \frac{\partial}{\partial \lambda} + \mathbf{j} \frac{1}{R} \frac{\partial}{\partial \varphi} + \mathbf{k} \frac{\partial}{\partial z}$, and $\nabla \equiv \mathbf{i} \frac{1}{R \cos \varphi} \frac{\partial}{\partial \lambda} + \mathbf{j} \frac{1}{R} \frac{\partial}{\partial \varphi}$ are the 3D and

86 2D vector differential operators in the polar spherical coordinates; $\boldsymbol{\Omega} = \Omega(\mathbf{j} \cos \varphi + \mathbf{k} \sin \varphi)$, is the

87 Earth rotation vector with $\Omega = 2\pi/(86164 \text{ s})$ the Earth rotation rate; ρ is the density; $\rho_0 = 1.225$

88 kg/m^3 , is the characteristic density near the ocean surface; $\mathbf{U} = (u, v)$, is the deflected-horizontal

89 velocity vector; w is the deflected-vertical velocity; p is the pressure; and \mathbf{F} is the turbulent

90 diffusive force due to the vertical shear represented by

$$91 \quad \mathbf{F} = \frac{\partial}{\partial z} \left(K \frac{\partial \mathbf{U}}{\partial z} \right) \quad (2)$$

92 Let \mathbf{U}_g be the geostrophic wind

$$93 \quad \rho_0 [2\boldsymbol{\Omega} \times \mathbf{U}_g] = -\nabla p \quad (3)$$

94 After substitution of (3) into (1a), we get the dynamic equation for the Ekman layer

$$95 \quad \rho_0 [2\boldsymbol{\Omega} \times (\mathbf{U} - \mathbf{U}_g)] = \rho \mathbf{g}_h + \rho_0 \mathbf{F}, \quad \mathbf{g}_h = \mathbf{i} g_\lambda + \mathbf{j} g_\varphi \quad (4)$$

96 where \mathbf{g}_h is independent on z in the troposphere (see Appendix D).

97 Baroclinicity (i.e., non-zero latitudinal or longitudinal density gradient) and spatially
 98 varying eddy viscosity K affect the Ekman layer dynamics (Chu, 2018; Sun & Sun, 2020). To limit
 99 the study on the effect of \mathbf{g}_h , the eddy viscosity K is assumed constant and the density varies in the
 100 z -direction, i.e., the geostrophic wind does not depend on z ,

$$101 \quad \partial \mathbf{U}_g / \partial z = 0.$$

102 Furthermore, a special density stratification is selected for this study as the e-folding decreasing
 103 with height

$$104 \quad \frac{\rho}{\rho_0} = s(z), \quad s(z) \equiv \exp\left(-\frac{z}{H}\right), \quad H = 10.4 \text{ km} \quad (5)$$

105 where H is considered as the height of the troposphere. Substitution of (5) into (4) leads to

$$106 \quad 2\boldsymbol{\Omega} \times (\mathbf{U} - \mathbf{U}_g) = s(z)\mathbf{g}_h + K \frac{\partial^2 \mathbf{U}}{\partial z^2} \quad (6)$$

107 With the complex variables, the deflected-horizontal gravity (\mathbf{g}_h), Ekman velocity (\mathbf{U}), and
 108 geostrophic wind (\mathbf{U}_g) are defined by

$$109 \quad G_h = g_\lambda + i g_\varphi, \quad U = u_E + i v_E, \quad U_g = u_g + i v_g, \quad i \equiv \sqrt{-1} \quad (7)$$

110 Eq.(6) is converted into

$$111 \quad \frac{\partial^2 U}{\partial z^2} - i \frac{f}{K} (U - U_g) = -\frac{s(z)}{K} G_h. \quad (8)$$

112 Substitution of (5) into (8) leads to

$$113 \quad \frac{\partial^2 U}{\partial z^2} - i \frac{f}{K} (U - U_g) = -\frac{G_h}{K} \exp\left(-\frac{z}{H}\right) \quad (9)$$

114 The Ekman velocity U needs to be satisfied by the upper boundary condition,

$$115 \quad U \rightarrow U_g \quad \text{as } z \rightarrow \infty \quad (10)$$

116 and the surface's boundary condition

$$117 \quad U = 0 \quad \text{as } z = 0 \quad (11)$$

118 **3 Ekman Layer Solution**

119 Eq.(9) with the boundary conditions (10) and (11) has the exact solution

$$120 \quad U(z) = U_g - \left[U_g - \Gamma(\delta^2 + 2i)G_h \right] e^{-(1+i)\pi z / D_E} - \Gamma(\delta^2 + 2i)G_h e^{-z/H} \quad (12)$$

121 where

$$122 \quad D_E \equiv \pi \sqrt{\frac{2K}{|f|}}, \quad \delta \equiv \frac{D_E}{\pi H} = \frac{\sqrt{2K/|f|}}{H}, \quad \Gamma \equiv \frac{2}{f(\delta^4 + 4)} \quad (13)$$

123 Here, D_E is the Ekman layer depth; and δ is the ratio between the Ekman layer depth (D_E) and
 124 the height of troposphere (H). Converting the Ekman layer spiral (12) into the vector form

$$u = u_g - \left\{ \left[u_g - \Gamma(\delta^2 g_\lambda - 2g_\varphi) \right] \cos\left(\frac{\pi z}{D_E}\right) + \left[v_g - \Gamma(\delta^2 g_\varphi + 2g_\lambda) \right] \sin\left(\frac{\pi z}{D_E}\right) \right\} e^{-\pi z/D_E} - \Gamma(\delta^2 g_\lambda - 2g_\varphi) e^{-z/H} \quad (14)$$

$$v = v_g - \left\{ -\left[u_g - \Gamma(\delta^2 g_\lambda - 2g_\varphi) \right] \sin\left(\frac{\pi z}{D_E}\right) + \left[v_g - \Gamma(\delta^2 g_\varphi + 2g_\lambda) \right] \cos\left(\frac{\pi z}{D_E}\right) \right\} e^{-\pi z/D_E} - \Gamma(\delta^2 g_\varphi + 2g_\lambda) e^{-z/H}$$

The eddy viscosity K is taken as a constant ($K = 5 \text{ m}^2 \text{ s}^{-1}$) in the atmospheric planetary boundary layer (Holton, 2004). The parameter δ is estimated by

$$\delta = \frac{\sqrt{2K/|f|}}{H} = \frac{0.0252}{\sqrt{|\sin \varphi|}}, \text{ for } K = 5 \text{ m}^2 \text{ s}^{-1}, H = 10.4 \text{ km}, \Omega = \frac{2\pi}{86164 \text{ s}} \quad (15)$$

where the parameter δ varies from 0.0854 at $\varphi=5^\circ$ (N or S) to 0.0252 at $\varphi=90^\circ$ (N or S). The range of $\pi\delta$ and the maximum values of δ^2 (at $\varphi=5^\circ$ S or N) are estimated by

$$0.07917 \leq \pi\delta \leq 0.2683, \quad \delta^2 \leq 0.7286 \times 10^{-2} \quad (16)$$

It is reasonable to delete terms with δ^2 in (14). The Ekman profile (14) is simplified by

$$u = u_g - \left[\left(u_g + \frac{g_\varphi}{f} \right) \cos\left(\frac{\pi z}{D_E}\right) + \left(v_g - \frac{g_\lambda}{f} \right) \sin\left(\frac{\pi z}{D_E}\right) \right] e^{-\pi z/D_E} + \frac{g_\varphi}{f} e^{-z/H} \quad (17)$$

$$v = v_g - \left[-\left(u_g + \frac{g_\varphi}{f} \right) \sin\left(\frac{\pi z}{D_E}\right) + \left(v_g - \frac{g_\lambda}{f} \right) \cos\left(\frac{\pi z}{D_E}\right) \right] e^{-\pi z/D_E} - \frac{g_\lambda}{f} e^{-z/H}$$

where the parameter Γ is simplified as $\Gamma = 1/(2f)$. Substitution of (17) into the continuity equation (1b) and integration with respect to z from $z = 0$ to $z = D_E$ leads to

$$w(D_E) = \frac{1}{R} \int_0^{D_E} \left(\frac{1}{\cos \varphi} \frac{\partial u}{\partial \lambda} + \frac{\partial v}{\partial \varphi} \right) dz \quad (18)$$

Substitution of (17) into (18) gives the Ekman pumping velocity

$$w(D_E) = \frac{D_E}{2\pi} \zeta_g - \frac{D_E}{2\pi f} \nabla \bullet \mathbf{g}_h \quad (19)$$

where the following approximations in the definite integration (18) are used

$$e^{-\pi} = 0.04321 \ll 1, \quad e^{-\pi\delta} \approx 1 \quad (20)$$

Here, $\zeta_g = \mathbf{k} \bullet \nabla \times \mathbf{U}_g$ is the geostrophic vorticity. Eq.(19) clearly shows that $\nabla \bullet \mathbf{g}_h$ causes the Ekman pumping in addition to the geostrophic vorticity. Substitution of (D6) in Appendix D into (19) leads to

$$w(D_E) = \frac{D_E}{2\pi} \zeta_g + \frac{D_E}{2\pi f} \frac{\partial^2 T}{\partial z^2} \Big|_{z=0} \quad (21)$$

where T is the disturbing static gravity potential (see Appendices C and D). The second term in the righthand side of (21) is the Ekman pumping due to the use of the true gravity

$$w_{ig}(D_E) = \frac{D_E}{2\pi f} \frac{\partial^2 T}{\partial z^2} \Big|_{z=0} \quad (22)$$

148 Here, the second derivative of the disturbing static gravity potential ($\partial^2 T / \partial z^2$) is obtained from a
 149 gravity model from the geodetic community.

150 **4 Global Ekman Pumping Velocity Due to $\partial^2 T / \partial z^2$**

151 The global data of the second derivative of the disturbing static gravity potential $\partial^2 T / \partial z^2$
 152 is obtained from the global static gravity model EIGEN-6C4 (<http://icgem.gfz-potsdam.de/home>)
 153 (Kostelecký et al. 2015), which was developed jointly by the GFZ Potsdam and GRGS Toulouse
 154 up to degree and order 2190, on $1^\circ \times 1^\circ$ grids (Figure 1), with -603.6 Eotvos as the minimum and
 155 642.8 Eotvos as the maximum (1 Eotvos = 10^{-9} s^{-2}). With the openly available data of $\partial^2 T / \partial z^2$,
 156 the Ekman pumping velocity due the use of the true gravity is easily identified. The global Ekman
 157 pumping velocity due to the true gravity $w_{ig}(D_E)$ (Figure 2) is calculated using (22) with the
 158 EIGEN-6C4 $\partial^2 T / \partial z^2$ data. The equatorial region ($5^\circ\text{S} - 5^\circ\text{N}$) is not included since the geostrophic
 159 balance does not exist there. Histogram of $w_{ig}(D_E)$ shows the Gaussian type distribution (Figure
 160 3a), and histogram of $|w_{ig}(D_E)|$ shows the near Gamma distribution with the mean of 4.0 cm s^{-1} ,
 161 and standard deviation of 13.59 cm s^{-1} . The result that the global mean $|w_{ig}(D_E)|$ reaches an
 162 evidently large value of 4.0 cm s^{-1} indicates the importance of $\partial^2 T / \partial z^2$ in atmospheric Ekman
 163 pumping.

164 **5 True-Vertical Coordinate versus Deflected-Vertical Coordinate**

165 As mentioned in the Introduction section, the true vertical direction \mathbf{e}_3 (upward positive) is
 166 with the true gravity \mathbf{g} ,

$$167 \quad \mathbf{g}(\lambda, \varphi, z) = -|\mathbf{g}(\lambda, \varphi, z)|\mathbf{e}_3(\lambda, \varphi, z). \quad (23)$$

168 The true horizontal surfaces are the equipotential surfaces of the true gravity $[V(\lambda, \varphi, z)]$ (see
 169 Appendix C). The geoid is one of them. On a true horizontal surface, the orthogonal unit vectors
 170 are represented by $[\mathbf{e}_1(\lambda, \varphi, z), \mathbf{e}_2(\lambda, \varphi, z)]$, but not (\mathbf{i}, \mathbf{j}) . With such a true-vertical coordinate, the
 171 true gravity \mathbf{g} has the vertical component only with no true-horizontal component. This treatment
 172 seems attractive to meteorologist. However, it is not feasible at all since the unit vectors $[\mathbf{e}_1(\lambda, \varphi, z), \mathbf{e}_2(\lambda, \varphi, z), \mathbf{e}_3(\lambda, \varphi, z)]$
 173 vary at each point inside the troposphere, and it is almost
 174 impossible to convert any atmospheric model (theoretical or numerical) with the standard gravity
 175 ($-\mathbf{g}_0\mathbf{k}$) into the model with the true gravity \mathbf{g} using the reference coordinates with the unit vectors
 176 $[\mathbf{e}_1(\lambda, \varphi, z), \mathbf{e}_2(\lambda, \varphi, z), \mathbf{e}_3(\lambda, \varphi, z)]$. The alternative treatment is to keep the deflected-vertical
 177 direction \mathbf{k} and deflected-horizontal surface (\mathbf{i}, \mathbf{j}) as the same as the meteorologists use. With this
 178 treatment, the unit vectors $(\mathbf{i}, \mathbf{j}, \mathbf{k})$ are independent on (λ, φ, z) . It is easy to replace the standard
 179 gravity ($-\mathbf{g}_0\mathbf{k}$) by the true gravity $\mathbf{g} (= \mathbf{g}_h - \mathbf{g}_0\mathbf{k})$ into any atmospheric models.

180

181 **6 Conclusions**

182 Meteorologists use the deflected-vertical (i.e., normal to the Earth spherical/ellipsoidal
 183 surface) as the “vertical”, the deflected-horizontal (i.e., the Earth spherical/ellipsoidal surfaces) as
 184 the “horizontal”, and the standard gravity ($-\mathbf{g}_0\mathbf{k}$) instead of the true gravity \mathbf{g} . The true gravity \mathbf{g}
 185 has latitudinal-longitudinal (i.e., deflected-horizontal) component \mathbf{g}_h , which is neglected
 186 completely. This study demonstrates the importance of \mathbf{g}_h in atmospheric dynamics using the
 187 Ekman layer as an example. With the constant eddy viscosity K and the e-folding type height-
 188 decreasing density, new equation for the atmospheric Ekman layer dynamics was derived

189 including both geostrophic forcing and \mathbf{g}_h . The evident Ekman pumping velocity due to the true
 190 gravity [$w_{tg}(D_E)$] is identified using the openly available data of $\partial^2 T / \partial z^2$ (T is the disturbing static
 191 gravity potential) from the EIGEN-6C4 gravity model with the global mean of 4.0 cm s^{-1} (evidently
 192 large) and standard deviation of 13.59 cm s^{-1} for $|w_{tg}(D_E)|$. Note that the results in this study is only
 193 for the specially selected density field represented by the e-folding density stratification with one
 194 specific gravity model (i.e., EIGEN-6C4), not for the density in the real atmosphere. However, it
 195 demonstrates that \mathbf{g}_h is an important forcing term in the atmospheric dynamics. Finally, if the
 196 meteorological community wants to keep the traditional terminology about the vertical (normal to
 197 the Earth sphere/ellipsoid) and horizontal (Earth spherical/ellipsoidal surface), the direction along
 198 the true gravity vector $\mathbf{g} (= \mathbf{i}g_\lambda + \mathbf{j}g_\varphi + \mathbf{k}g_z)$ should be called the **true vertical**; and the equipotential
 199 surfaces such as the geoid should be called the **true horizontal**.

200 Appendix A. Oblate Spheroid Coordinates Versus Polar Spherical Coordinates

201 The oblate spheroid coordinates share the same longitude (λ) but different latitude (φ_{ob})
 202 and radial coordinate (representing vertical) (r_{ob}) with corresponding unit vectors (\mathbf{i} , \mathbf{j} , \mathbf{k}). The
 203 relationship between the oblate spheroid coordinates (λ , φ_{ob} , r_{ob}) and the polar spherical
 204 coordinates (λ , φ , r) is given by (Gill, 1982)

$$205 \quad r^2 = r_{ob}^2 + \frac{1}{2}d^2 - d^2 \sin^2 \varphi_{ob}, \quad r^2 \cos^2 \varphi = (r_{ob}^2 + \frac{1}{2}d^2) \cos^2 \varphi_{ob} \quad (\text{A1})$$

206 where d is the half distance between the two foci of the ellipsoid. For the normal Earth, $d = 521.854$
 207 km. The 3D vector differential operator in the oblate spheroid coordinates is represented by

$$208 \quad \nabla_3 = \mathbf{i} \frac{1}{h_\lambda^{ob}} \frac{\partial}{\partial \lambda} + \mathbf{j} \frac{1}{h_\varphi^{ob}} \frac{\partial}{\partial \varphi} + \mathbf{k} \frac{1}{h_r^{ob}} \frac{\partial}{\partial z}, \quad z = r - R \quad (\text{A2})$$

209 where $R = 6.3781364 \times 10^6$ m, is the semi-major axis of the normal Earth (Earth radius). The
 210 coefficients (or called Lamé numbers) (h_λ^{ob} , h_φ^{ob} , h_r^{ob}) are given by

$$211 \quad h_\lambda^{ob} = \sqrt{r^2 + \frac{1}{2}d^2} \cos \varphi, \quad h_\varphi^{ob} = \sqrt{r^2 - \frac{1}{2}d^2 + d^2 \sin^2 \varphi}, \quad h_r^{ob} = \frac{r \sqrt{r^2 - \frac{1}{2}d^2 + d^2 \sin^2 \varphi}}{\sqrt{r^4 - \frac{1}{4}d^4}}. \quad (\text{A3})$$

212 However, the 3D vector differential operator in the polar spherical coordinates is represented by

$$213 \quad \nabla_3 = \mathbf{i} \frac{1}{R \cos \varphi} \frac{\partial}{\partial \lambda} + \mathbf{j} \frac{1}{R} \frac{\partial}{\partial \varphi} + \mathbf{k} \frac{\partial}{\partial z}. \quad (\text{A4})$$

214 The difference between the two coordinates is 0.17% (Gill, 1982).

215

216 Appendix B. True Gravity and Its approximation

217

218 The true gravity \mathbf{g} is represented by a three-dimensional vector in the (λ , φ , z) coordinate
 219 system,

$$220 \quad \mathbf{g} = \mathbf{g}_h + \mathbf{k}g_z, \quad \mathbf{g}_h = \mathbf{i}g_\lambda + \mathbf{j}g_\varphi \quad (\text{B1})$$

221 where \mathbf{g}_h is its deflected-horizontal component, and $g_z \mathbf{k}$ the deflected-vertical component. It has
 222 two approximated forms. The first one is the normal gravity [$-g(\varphi)\mathbf{k}$] and usually represented in
 223 the oblate spheroid coordinates (see Appendix A) and associated with a mathematically modeled

224 Earth (i.e., a rigid and geocentric ellipsoid) called the normal Earth. The normal Earth is a spheroid
 225 (i.e., an ellipsoid of revolution), has the same total mass and angular velocity as the Earth, and
 226 coincides its minor axis with the mean rotation of the Earth (Vaniček & Krakiwsky, 1986). The
 227 normal gravity vector $[-g(\varphi)\mathbf{k}]$ is the sum of the gravitational and centrifugal accelerations exerted
 228 on the water particle by the normal Earth. Its intensity $g(\varphi)$ is determined analytically. For example,
 229 the World Geodetic System 1984 uses the Somigliana equation (National Geospatial-Intelligence
 230 Agency, 1984) to represent $g(\varphi)$

$$231 \quad g(\varphi) = g_e \left[\frac{1 + \kappa \sin^2 \varphi}{\sqrt{1 - e^2 \sin^2 \varphi}} \right], \quad e^2 = \frac{a^2 - b^2}{a^2}, \quad \kappa = \frac{bg_p - ag_e}{ag_e} \quad (\text{B2})$$

232 where (a, b) are the equatorial and polar semi-axes; a is used for the Earth radius, $R = a =$
 233 6.3781364×10^6 m; $b = 6.3567523 \times 10^6$ m; e is the spheroid's eccentricity; $g_e = 9.780$ m/s², is the
 234 gravity at the equator; and $g_p = 9.832$ m/s² is the gravity at the poles. The second one is the standard
 235 gravity vector, $-g_0\mathbf{k}$, with $g_0 = 9.81$ m/s². Meteorologist uses the standard gravity. Both normal and
 236 standard gravities don't have latitudinal and longitudinal components.

237

238 **Appendix C. True, Normal, and Standard Gravity Potentials**

239

240 Let (V, E, E_0) be the gravity potentials associated with the true gravity \mathbf{g} , the normal gravity
 241 $-g(\varphi)\mathbf{k}$, and the standard gravity $-g_0\mathbf{k}$. The potentials of the normal and standard gravities are given
 242 by

$$243 \quad E(\varphi, z) = -g(\varphi)z, \quad E_0(z) = -g_0z \quad (\text{C1})$$

244 Both V and E include the potential of the Earth's rotation (P_R)

$$245 \quad P_R = \Omega^2 r^2 \cos^2 \varphi / 2. \quad (\text{C2})$$

246 The gravity disturbance is the difference between the true gravity $\mathbf{g}(\lambda, \varphi, z)$ and the normal gravity
 247 $[-g(\varphi)\mathbf{k}]$ at the same point (Hackney & Featherstone 2003). The potential of the gravity
 248 disturbance (called the disturbing gravity potential) is given by

$$249 \quad T = V - E = V + g(\varphi)z \quad (\text{C3})$$

250 With the disturbing gravity potential T , the true gravity $\mathbf{g} (= \mathbf{g}_h + g_z\mathbf{k})$ and its components are
 251 represented by (Sandwell & Smith 1997)

$$252 \quad \mathbf{g} = \nabla_3 V, \quad \mathbf{g}_h = \nabla T, \quad g_z = -g(\varphi) + \frac{\partial T}{\partial z} \quad (\text{C4})$$

253 where $\nabla \equiv \mathbf{i} \frac{1}{R \cos \varphi} \frac{\partial}{\partial \lambda} + \mathbf{j} \frac{1}{R} \frac{\partial}{\partial \varphi}$ is the 2D vector differential operator. The disturbing static
 254 gravity potential (T) outside the Earth masses in the spherical coordinates with the spherical
 255 expansion is given by (Kostelecký et al. 2015)

$$256 \quad T(r, \lambda, \varphi) = \frac{GM}{r} \sum_{l=2}^{\infty} \sum_{m=0}^l \left(\frac{R}{r} \right)^l \left[(C_{l,m} - C_{l,m}^{el}) \cos m\lambda + S_{l,m} \sin m\lambda \right] P_{l,m}(\sin \varphi), \quad (\text{C5})$$

257 where $G = 6.674 \times 10^{-11} \text{m}^3 \text{kg}^{-1} \text{s}^{-2}$, is the gravitational constant; $M = 5.9736 \times 10^{24}$ kg, is the mass
 258 of the Earth; r is the radial distance with $z = r - R$; $P_{l,m}(\sin \varphi)$ are the Legendre associated functions
 259 with (l, m) the degree and order of the harmonic expansion; $(C_{l,m}, C_{l,m}^{el}, S_{l,m})$ are the harmonic
 260 geopotential coefficients (Stokes parameters with $C_{l,m}^{el}$ belonging to the reference ellipsoid. From
 261 Eqs. (C1) and (C3) the potential of the true gravity is given by

$$V = T - g(\varphi)z. \quad (C6)$$

From Eq.(C4) the true gravity is represented by

$$\mathbf{g}(\lambda, \varphi, z) = \nabla T + \left[\frac{\partial T}{\partial z} - g(\varphi) \right] \mathbf{k} \quad (C7)$$

265

266 **Appendix D. An Approximate 3D True Gravity Field for the Troposphere**

267 According to Eq.(C5) (i.e., the spectral of the disturbing static gravity potential T), the ratio
268 between $T(\lambda, \varphi, z)$ to $T(\lambda, \varphi, 0)$ through the troposphere can be roughly estimated by

$$\left| \frac{T(\lambda, \varphi, z)}{T(\lambda, \varphi, 0)} \right| \approx \frac{R}{R+z} \approx 1, \quad H \geq z \geq 0 \quad (D1)$$

270 where $H = 10.4$ km, is the height of the troposphere. Since R is the radius of the Earth and more
271 than 3 orders of magnitude larger than H . This leads to the first approximation that the disturbing
272 gravity potential $T(\lambda, \varphi, z)$ does not change with z in the whole troposphere (approximation of
273 thin layer for the troposphere)

$$T(\lambda, \varphi, z) \approx T(\lambda, \varphi, 0), \quad H \geq z \geq 0 \quad (D2)$$

275 which makes

$$\nabla^2 T(\lambda, \varphi, z) \approx \nabla^2 T(\lambda, \varphi, 0); \quad (D3)$$

277 The potential of the true gravity satisfies the Laplace equation outside the Earth surface (Vaniček
278 & Krakiwsky, 1986),

$$\nabla^2 V + \frac{\partial^2 V}{\partial z^2} = 0 \quad (D4)$$

280 Substitution of (C6) into (D4) leads to the Laplace equation for the disturbing static gravity
281 potential

$$\nabla^2 T + \frac{\partial^2 T}{\partial z^2} = 0 \quad (D5)$$

283 Use of the second formula in (C4) gives

$$\nabla \cdot \mathbf{g}_h = -\frac{\partial^2 T}{\partial z^2} \approx -\frac{\partial^2 T}{\partial z^2} \Big|_{z=0} \quad (D6)$$

285 where the approximation (D3) for the troposphere is used.

286

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288

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291 disturbing static gravity potential ($\partial^2 T / \partial z^2$) data from the EIGEN-6C4 model, which is publicly
292 available at the website: <http://icgem.gfz-potsdam.de/home>.

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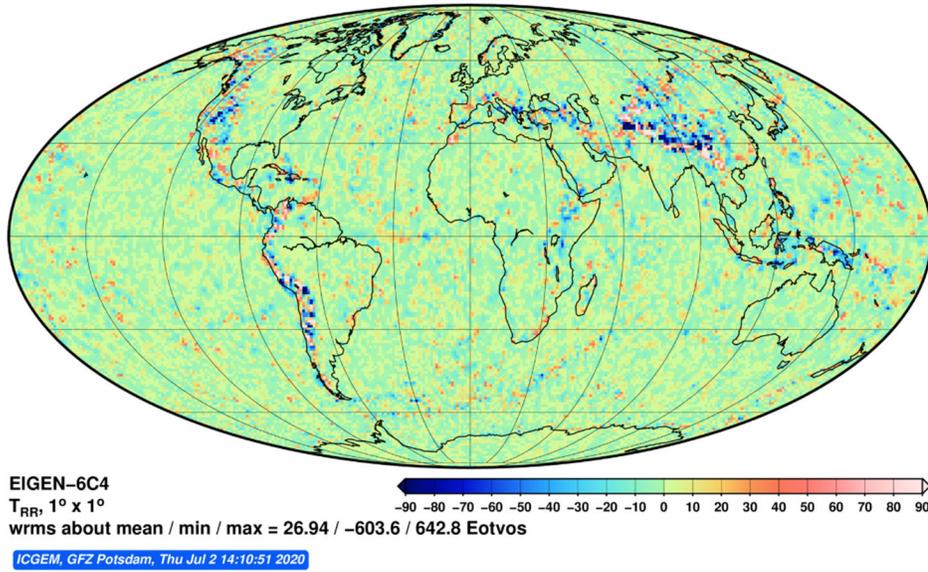
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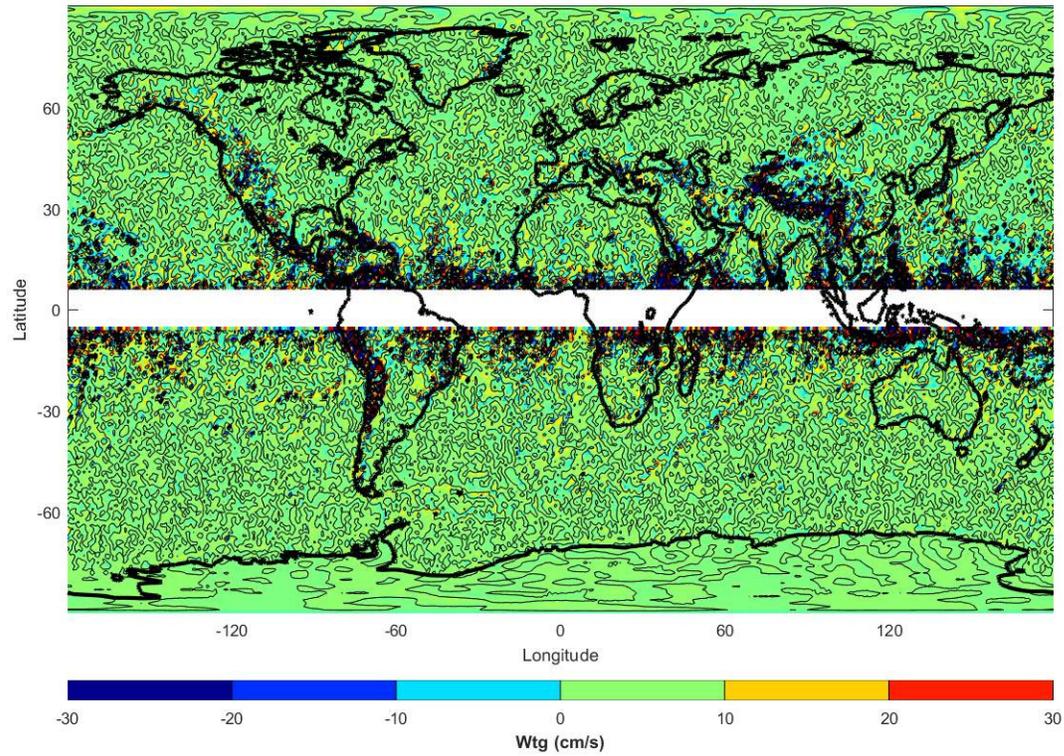
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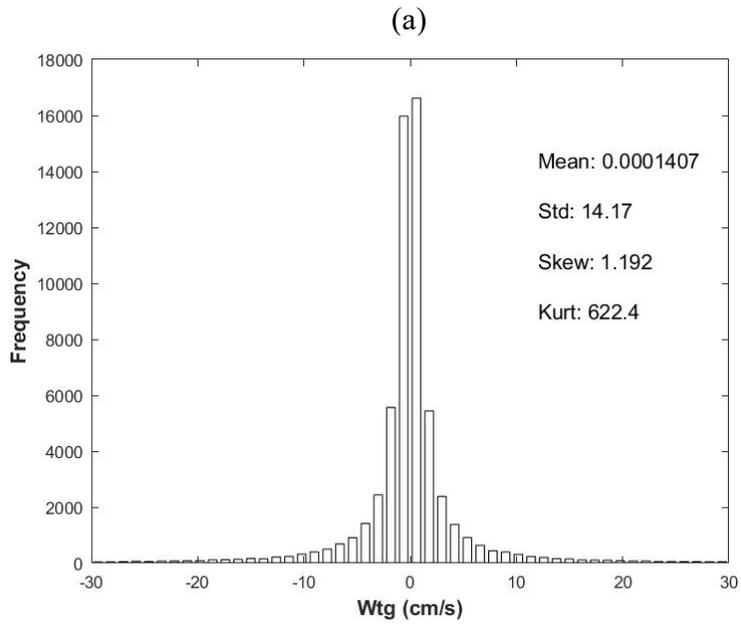
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Figure 1. Global second radial derivative of the disturbing static gravity potential ($\partial^2 T / \partial z^2$) (unit: Eotvos) at $z = 0$ from the EIGEN-6C4 with $1^\circ \times 1^\circ$, computed online at the website <http://icgem.gfz-potsdam.de/home>. Large positive (negative) values indicate evident upward (downward) Ekman pumping velocity due to the true gravity in the Northern Hemisphere, and opposite in the Southern Hemisphere.

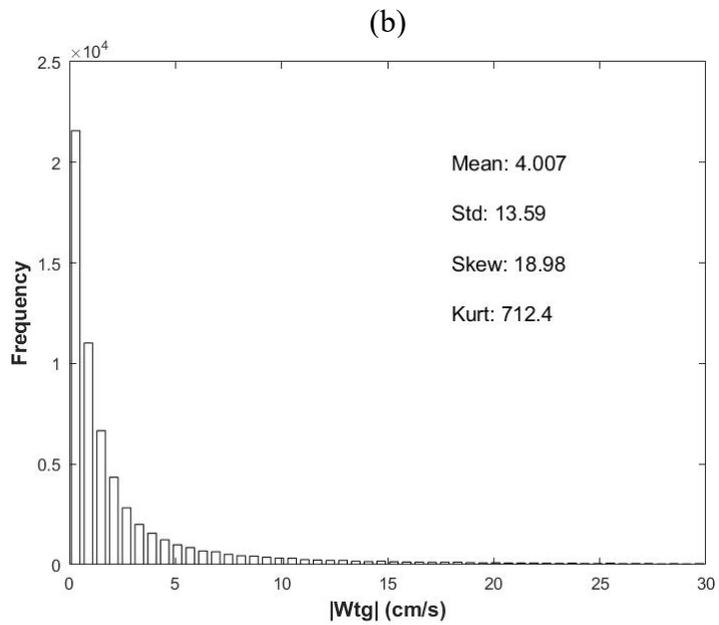
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345 **Figure 2.** Ekman pumping velocity (cm s^{-1}) due to the use of the true gravity $w_{tg}(D_E)$ calculated
346 with the EIGEN-6C4 $\partial^2 T / \partial z^2$ data.
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Figure 3. Histograms of global (a) $w_{tg}(D_E)$ and (b) $|w_{tg}(D_E)|$ (cm s^{-1}) with statistical parameters calculated using the EIGEN-6C4 $\partial^2 T / \partial z^2$ data.