One-Sided Joint Inversion of Shear Velocity and Resistivity from the PI-LAB Experiment at the Equatorial Mid-Atlantic Ridge

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Abstract

The lithosphere-asthenosphere system is fundamental to our understanding of mantle convection and plate tectonics. Seismic and electromagnetic methods are our primary means of determining its structure and physical properties. These independent constraints with different sensitivities to Earth's properties hold promise for understanding the system. Here we use the shear velocity model from Rayleigh waves and the MT based resistivity model from near the equatorial Mid-Atlantic Ridge. Crossplots of the models suggest a linear or near-linear trend that is also in agreement with petrophysical predictions. We therefore map the MT model to a new shear-wave starting model using the petrophysical relationship, which is then used to re-invert for shear-wave velocity. The resulting shear-wave velocity model fits the phase velocity data, and the correlation coefficient between the shear velocity and resistivity models is increased. Much of the model can be predicted by expectations for a thermal half-space cooling model, although some regions require a combination of higher temperatures, volatiles, or partial melt. We use the petrophysical predictions to estimate the melt fraction, melt volatile content, and temperature structure of the asthenospheric anomalies. We find up to 4% melt, with the lowest resistivities and shear velocities explained by up to 20% water or 20% CO₂ in the melt or $^{-1}$ % nearly pure sulfide melt, depending on the set of assumptions used. Melt is required in punctuated anomalies over broad depth ranges, and also in channels at the base of the lithosphere. Melt in the asthenosphere is dynamic, yet persistent on geologic time scales.

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11 Abstract

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- 13 convection and plate tectonics. Seismic and electromagnetic methods are our primary
- 14 means of determining its structure and physical properties. These independent constraints
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- 17 model from near the equatorial Mid-Atlantic Ridge. Cross-plots of the models suggest a
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- 21 wave velocity model fits the phase velocity data, and the correlation coefficient between
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- 25 predictions to estimate the melt fraction, melt volatile content, and temperature structure
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- shear velocities explained by up to 20% water or 20% CO_2 in the melt or ~1% nearly pure
- 28 sulfide melt, depending on the set of assumptions used. Melt is required in punctuated
- anomalies over broad depth ranges, and also in channels at the base of the lithosphere.
- 30 Melt in the asthenosphere is dynamic, yet persistent on geologic time scales.
- 31

32 Introduction

- 33 Plate tectonic theory is predicated on the idea of a rigid lithosphere that overrides a weaker
- 34 underlying asthenosphere (McKenzie & Parker, 1967), but the nature of the lithosphere-
- 35 asthenosphere system remains the subject of vigorous debate. The oceanic lithosphere
- 36 comprises the majority of the surface of the Earth and has the simplest evolution and
- history. It is classically thought to be thermally defined as a boundary layer in a simple
- thermal model (Parker & Oldenburg, 1973). In this model, increasing temperature with
- depth causes mantle rocks to weaken, creating the asthenosphere (e.g., Goetze et al., 1978).
- 40 However, a host of observations, including sharp seismic velocity discontinuities (Gaherty et
- 41 al., 1996; Rychert et al., 2020; Rychert et al., 2018; Rychert & Shearer, 2011; Tan &
- 42 Helmberger, 2007), low velocity zones (Forsyth et al., 1998; N. Harmon et al., 2020), and low
- 43 resistivity zones (Baba et al., 2006; Wang et al., 2020) in the asthenosphere, suggest that in

44 addition to temperature other factors are likely required to explain the observations. Many

- 45 potential explanations of these observations have been proposed including an increased
- 46 effect of hydration (Karato, 2012), the presence of partial melt (Anderson & Sammis, 1970),
- 47 and/or the enhanced effects at near sub-solidus conditions on seismic waves (Yamauchi &
- Takei, 2016). The debate centers around which of these explanations might be in operation
- 49 and how widely they apply.
- 50

51 Partial melt is likely to exist in the asthenosphere, in particular near mid-ocean ridges and 52 volcanic arcs where the volcanic systems must be fed by mantle melting (Anderson & 53 Sammis, 1970). However, further away from volcanic plate boundaries its presence is more 54 debated (Kawakatsu et al., 2009; Priestley & McKenzie, 2006; Rychert et al., 2005). The 55 amount of melt and its location is vital to our understanding of how the lithosphere-56 asthenosphere works, as the presence of partial melt is predicted to reduce the viscosity of 57 the asthenosphere (Hirth & Kohlstedt, 1995; Jackson et al., 2006) and could also facilitate 58 plate tectonics (Rychert et al., 2005; Rychert et al., 2007). However, different geophysical 59 techniques with different sensitivities and resolutions have imaged anomalies that have 60 been interpreted as melt in many forms (Rychert et al., 2020). For instance beneath mid-61 ocean ridges, seismic surface wave studies have interpreted a broad, hundreds of kilometers 62 wide, melt triangle beneath the ultrafast spreading East Pacific Rise at 17 °S (Dunn &63 Forsyth, 2003; Forsyth et al., 1998) and the intermediate spreading Juan De Fuca Ridge (Bell 64 et al., 2016; Gao, 2016), while other studies have imaged smaller scale and discrete melt 65 zones beneath the slow spreading equatorial Mid-Atlantic Ridges on the order of 100-200 66 km wide (N Harmon et al., 2020). Magnetotelluric methods have typically imaged smaller 67 and more discrete low resistivity zones interpreted as focused melt regions beneath the fast 68 spreading East Pacific Rise at 9 °N and the ultra-slow spreading Mohns Ridge (Johansen et 69 al., 2019; Key et al., 2013) that are typically < 100 km wide, although a broader region >200 70 km was inferred beneath the East Pacific Rise at 17 °S (Evans et al., 1999). Further off-axis, 71 layered and/or pervasive melt in the asthenosphere has been inferred based on the imaging 72 of discontinuities by scattered waves that require sharp drops in seismic velocity with depth 73 (Kawakatsu et al., 2009; Rychert & Shearer, 2011; Rychert & Shearer, 2009; Tharimena et 74 al., 2017). Active source seismic studies also find strong reflectors near the expected base of 75 the tectonic plate, that have been interpreted as channelized melt (Mehouachi & Singh, 76 2018; Stern et al., 2015). Similar channelized structures have also been interpreted from 77 thin low resistivity zones at 60-80 km depth (Naif et al., 2013; Wang et al., 2020). These 78 interpretations are intriguing but originate from methods with a variety of resolutions and 79 sensitivities in different locations. Therefore, whether or not differences are an artefact of 80 resolution and sensitivities of the individual methodologies or representative of real Earth 81 structure has remained unclear. 82 83 The complementary resolution and sensitivities of MT and seismic imaging techniques offer 84 a promising means of probing the Earth's physical properties to examine the thermal

structure and the presence of partial melt. The Earth's mantle is primarily composed of

- 86 olivine and pyroxene, and the conductivity of these minerals has a strong temperature
- 87 dependence (Gardés et al., 2014; Naif et al., 2021), enhanced by the presence of conducting
- 88 fluids such as partial melt (Naif et al., 2021; Ni et al., 2011) and the presence of water and
- 89 other crystallographic defects in the olivine mineral lattice (Gardés et al., 2014; Naif et al.,
- 2021). Water and other volatiles such as CO_2 are also thought to significantly increase the

91 conductivity of the fluid and therefore the overall conductivity of the mantle if present (Ni et

- al., 2011; Sifre et al., 2014). On the other hand, seismic velocities are dependent on
- temperature and pressure (e.g., Stixrude & Lithgow-Bertelloni, 2005), followed by the
- 94 presence of partial melt (Clark & Lesher, 2017; Hammond & Humphreys, 2000), particularly
- 95 for shear velocity, and are relatively insensitive to the presence of water as a
- 96 crystallographic defect (Abers et al., 2014) or as a component of the partial melt. These
- 97 differences mean that the two methods together have the potential to better constrain the
- 98 thermal properties of the mantle, the presence and amount of partial melt, and the amount
- of hydration in the melt.
- 100
- 101 There have been two main approaches to joint inversion of electromagnetic and seismic
- 102 data: 1) inversion based on underlying petrophysical or empirical relationships between
- 103 velocity and conductivity (Jegen et al., 2009) and 2) inversion based on a cross-gradient
- approach, e.g. forcing model changes in velocity and resistivity together (Bennington et al.,
- 2015; Gallardo & Meju, 2004; Haber & Oldenburg, 1997; Moorkamp et al., 2011). The
- 106 petrophysical or empirical approach requires either accurate models of the physical
- properties of the rocks (Gardés et al., 2014), or ideally a relatively simple system that can be
- 108 captured with simple linear or polynomial fits to data (Jegen et al., 2009), which is more
- 109 likely the case in locations with limited compositional and thermal variation. The cross-
- 110 gradient approach, on the other hand, presumes that low resistivity features should be
- associated with low or high velocity features, in other words, that the two are positively or
- 112 negatively correlated. However, this may not necessarily be the case in the presence of
- small amounts of certain minerals such as magnetite in serpentine (Stesky & Brace, 1973) or
- 114 graphite (Frost et al., 1989) and other highly conductive minerals which may not be
- volumetrically significant enough to have a strong seismic signature. Choosing between these two approaches or other approaches using Monte Carlo inversions (Moorkamp et al.)
- these two approaches or other approaches using Monte Carlo inversions (Moorkamp et al.,2010) is dependent on the details of the particular datasets and the structure involved.
- 118
- 119 The I-LAB (Imaging the Lithosphere-Asthenosphere Boundary) experiments including: 1)
- 120 Passive Imaging of the Lithosphere Asthenosphere Boundary (PI-LAB) experiment, 2)
- 121 Experiment to Unearth the Rheological Lithosphere-Asthenosphere Boundary (EURO-LAB),
- and 3) the Central Atlantic Lithosphere-Asthenosphere Boundary (CA-LAB) experiment
- 123 presented a unique opportunity to examine joint inversion and interpretation of MT and
- seismic data in order to understand the oceanic lithosphere-asthenosphere system at the
- equatorial Mid-Atlantic Ridge. We deployed 39 ocean bottom seismometers (OBS) and 39
- 126 ocean bottom magnetotelluric (OBMT) instruments from 0-80 Myr seafloor across the Chain
- and Romanche fracture zones (Agius et al., 2018; Harmon et al., 2018). The OBS and OBMT
- were co-located (within 1-2 km), in three lines perpendicular to the ridge (Fig. 1). The
- experiment was designed to image the uppermost mantle beneath the ridge system andexamine the evolution of the oceanic lithosphere-asthenosphere system and the nature of
- examine the evolution of the oceanic lithospherthe lithosphere-asthenosphere boundary.
- 132
- 133 Here we focus on two results for joint inversion, the 3-D shear-wave velocity model from
- 134 Rayleigh wave tomography and the 2-D MT inversion from the two southernmost lines (Fig.
- 135 1, 2). The shear velocity model images a high velocity lithosphere, and several punctuated
- 136 low velocity zones (<4.2 km/s) in the asthenosphere, that were interpreted as melt (N
- 137 Harmon et al., 2020). Near the ridge axis, asthenospheric low velocity zones are attributed

138 to sub-ridge upwelling (Anomalies A and E in line I and line II, respectively in Fig. 2), while 139 further off-axis the low velocity anomalies are attributed to melting due to upwelling caused 140 by small scale convection (Anomalies B, C, and F in Fig. 2) (N Harmon et al., 2020; Wang et 141 al., 2020). The MT result images similar structures to the surface wave model, e.g., a high 142 resistivity lithospheric lid ($\log_{10}(\rho) > 2$) and several low resistivity anomalies ($\log_{10}(\rho) < 1$) in 143 the asthenosphere (anomalies A, B, C, D, E, and F in Fig. 2) (Wang et al., 2020). In Line I there 144 is good agreement with the depth (50-80 km) and lateral extent (~100-200 km) of the low 145 resistivity anomaly and low seismic velocities (Anomalies B and C) as well as evidence for a 146 high resistivity, high velocity lithospheric drip (anomaly D) that extended from 50 to 150 km 147 depth. However, in line II (Fig. 2b and 2d) the agreement in terms of the shapes of the 148 anomalies is less remarkable, specifically anomaly F, where the conductive anomalies 149 suggest a channel structure < 20 km thick extending from the ridge to 30 Myr seafloor, while 150 the surface wave anomaly resembles a simple oval ~200 km wide from 50-80 km depth. In 151 addition, anomaly E is deeper in the resistivity model, >100 km depth, than in the shear 152 velocity model, where it extends from 50 to 100 km depth. While in line I, anomaly A is 153 shallower at ~30 km depth and smaller, <50 km wide, in the resistivity model than in the 154 shear velocity model, where it is located at 50-80 km depth and 150 km wide. In other 155 words, while there is some similarity in the lateral locations of the anomalies, the depth and 156 morphologies are a bit different.

157

158 Subsequent studies support the existence of these anomalies and suggest that apparent 159 discrepancies may be artefacts of resolution. For example, S-to-P receiver functions support 160 the existence of the anomalies. The receiver functions image discontinuities associated with 161 sharp velocity decreases with depth above the locations of the low shear velocity anomalies 162 E, C, and F in the asthenosphere and also the locations where the low resistivity anomalies 163 gradually decrease with depth in the asthenosphere (near anomaly E and directly beneath F) 164 (Rychert et al., 2021). In addition, a short period Rayleigh wave tomography study, which 165 had better resolution in the upper 60 km than Harmon et al. (2020), imaged a shallower 166 anomaly for anomaly A beneath line I, more consistent with the resistivity model (Saikia et 167 al., 2021). The differences between the surface wave models suggest that there are several 168 possibilities for shear-wave velocity models that will fit the Rayleigh wave data that could 169 also be consistent with the anomaly structure of the resistivity model as well. The primary 170 motivation of this study is to find a satisfactory shear velocity model that is also consistent 171 with structural information from the resistivity models.

172

173 Here we jointly consider the Rayleigh wave phase velocities and the MT data to evaluate 174 differences and similarities between the seismic and MT anomaly structures, in particular to 175 determine an Earth structure that can satisfy both datasets within data errors. We compare 176 the models one-to-one to develop an empirical relationship between the two and also 177 consider laboratory-based predictions for shear velocity and resistivity. We use the 178 relationship to translate the MT resistivity to shear-wave velocity and use this as the new 179 starting model for the surface wave tomography inversion. This approach assumes that the 180 structure within the resistivity model is closer to the true earth structure, which may be the 181 case, for example, if a thin channel structure exists, which surface waves would not be able 182 to resolve without prior knowledge. We evaluate the validity of this assumption. Finally, we 183 compare to petrophysical predictions for Earth properties to constrain temperature, the

184 amount of partial melt, and the amount of hydration, carbonization or sulfide weight

185 percentage of the partial melt in the asthenosphere.

186

187 Methods

188 MT data were inverted by Wang et al. (2020), which we briefly summarize here. The 189 determinant of the MT impedance tensor was used to invert logarithmic apparent resistivity 190 and linear phase along two wo-dimensional transects (line I and line II). The approach was 191 chosen to minimize strong 3-D coast effects from the nearby African Coast (Wang et al., 192 2019). Forward calculations and inversion were performed using the MAR2DEM code (Key, 193 2016), modified to accept determinant data as an input (Wang et al., 2021). Inversion of MT 194 data with this approach is less dependent on the starting model than surface wave inversion 195 due to the diffusive nature of electromagnetic fields and the smoothness and regularization 196 of the inversion problem. So here we focus on varying the starting model for shear velocity 197 inversion based on structural information from the resistivity data, but not vice versa. We 198 refer to this as a "one-sided" joint inversion.

199

200 We first establish a relationship between shear velocity and resistivity in our study area. We 201 use two transects through the three-dimensional shear-wave velocity model of Harmon et 202 al., (2020) in the same locations of the two two-dimensional resistivity model transects of 203 Wang et al., (2020). We make cross-plots separately for the two lines. Cross-plots of the 204 data suggest a linear relationship between the two data sets, but with scatter (Fig. 3). The 205 correlation coefficients of these cross-plots for line I is 0.43 and line II 0.39. A linear 206 regression of line I between shear-wave velocity (km/s) and resistivity ($\log_{10}(\rho)$) with the 207 highest correlation coefficient, yields a solution of Vs=4.19+0.10* $\log_{10}(\rho)$. We did not fit a 208 line to line 2 given the lower correlation and the fact that a linear relationship is less 209 apparent. We also consider predictions from laboratory petrophysical relationships between 210 shear velocity and resistivity for a half-space cooling model with a mantle potential 211 temperature of 1350 °C from 0-40 Myr, the approximate range in of ages along lines I and II 212 (Fig. 4). To model the predicted shear velocity for a given temperature, pressure, and melt 213 fraction we use the Very Broadband Rheology calculator (Havlin et al., 2021), assuming a 214 peridotite mantle composition. We use the attenuation parameterization of (Jackson & Faul, 215 2010) that is included in the calculator and use an average across the surface wave 216 frequency range used here. In this model, the addition of melt primarily affects shear 217 velocity with ~2-4% velocity reduction for 1% melt volume fraction depending on the 218 dihedral angle (Takei, 1998). The model of Takei (1998) assumes that melt is interconnected, 219 without necessarily proscribing a melt geometry. The associated predicted velocity 220 reduction depends on wetness, which is a measure of the amount of grain to grain contact 221 relative to the melt (Takei, 1998). Other models for the effect of melt on velocity exist based 222 on different assumptions of melt geometry (Clark & Lesher, 2017; Hammond & Humphreys, 223 2000; Schmeling, 1985) which we evaluate in the discussion section. For resistivity we use 224 the relationship for hydrated mantle peridotite (Gardés et al., 2014) and a model for the 225 conductivity of hydrous mantle melts (Ni et al., 2011). We then use the Hashin-Shtrikman 226 upper bound to calculate the total resistivity of a melt bearing peridotite mantle (Ni et al., 227 2011). The predictions for an example case with 100 ppm water content in the background 228 mantle and 1% melt in the melt triangle and variable amounts of water in the melt from 4-229 20 weight % are shown in Figure 4. We perform a linear regression on the melt-free mantle

- data points (black dots, Fig. 4), and find a relationship of $V_s = 4.14 + 0.11^* \log_{10}(\rho)$. This
- 231 relationship is very similar to the one derived for the cross-plot in line I, but the velocity
- intercept is 0.05 km/s lower than in our cross-plot, and the slope is only 0.01 km/s/ $\log_{10}(\rho)$
- higher than in the cross-plot. Given the similarity between the two and that the
- petrophysical line visually fits the data from the shear velocity and resistivity inversions, we
- 235 opt to use the relationship from the petrophysical modelling.
- 236

We use the aforementioned petrophysical relationship to translate the resistivity model (Fig. 2a, b) to shear-wave velocity, creating a new starting model (Fig. 5c, d) for the shear velocity inversion. We then invert the phase velocities from 18-143 s period from Harmon et al. (2020) sampled along lines I and II, for shear velocity as a function of depth, sampling at every point, 0.1°. We calculate the partial derivatives relating Rayleigh wave phase velocity to shear velocity using the Computer Programs in Seismology package (Herrmann, 2013), and we assume a fixed Vp/Vs ratio of 1.8, which is consistent with the Preliminary Earth

- 244 Reference Model (PREM), a global 1-D seismic velocity model (Dziewonski & Anderson,
- 245 1981). We include a seawater layer along lines I and II in the model based on the local
- bathymetry. We use a damped least-squares inversion and assume an *a priori* model error of 0.2 km/s following choices from previous work (Forsyth & Li, 2005; N Harmon et al.,
- 248 2020). We replace the upper 5 km of the model beneath the water layer with average
- crustal values (3.5 km/s) from the 1-D model of Harmon et al. (2020). The model is
- 250 parameterized every 5 km in depth down to 400 km. This parameterization is finer than that
- 251 presented in Harmon et al. (2020) (Fig. 2). Therefore, we also present an inversion using the
- 252 1-D model used in Harmon et al. (2020), but with the 5 km thick layers down to 400 km
- 253 depth used here for comparison purposes (Fig. 5).
- 254

255 We next determine the physical properties that explain the resulting anomalies. We 256 calculate the thermal structure for seafloor from 0 to 40 Myr age in 1 Myr intervals. The 257 thermal models have an adiabatic gradient added to them, and we assume a mantle 258 potential temperature of 1350 °C. Although seafloor age is known at each profile, we do not 259 proscribe age, given that our previously published models suggest that the age progression 260 of the lithosphere might not be monotonic everywhere. For each thermal structure from 0 261 to 40 Myr seafloor, we calculate the predicted shear velocity and resistivity for melt 262 fractions from 0.00 to 0.07 at 0.001 increments below 0.01 and 0.005 increment above 0.01 263 and melt water contents from 0 to 30 weight % in 1% increments for all temperatures > 264 1100 °C at the corresponding depth/pressure values using the relationships described above 265 for the half-space cooling model presented in Fig 4. We then examine the regions that 266 cannot be explained by temperature alone, specifically, where the shear velocity is <4.4 km/ 267 s and \log_{10} resistivity is < 1.5 (< 30 Ω m), which are the nominal limits of the melt free 268 predictions of the half-space cooling model (black dots, Fig 4). We perform a grid search 269 over melt fraction, melt water content, and apparent seafloor age/temperature for each 270 point in lines I and II. We then determine the chi-squared residual between the observed 271 resistivity and shear velocity with the predicted resistivity and shear velocity at the same 272 depth in each thermal structure from 0 to 40 Myr. The chi-squared residual is used to 273 determine goodness of fit assuming an *a priori* standard deviation of 0.05 km/s for the shear 274 velocity model and 0.10 $\log_{10}(\Omega m)$ for the resistivity model. A value of melt, melt hydration 275 and temperature is considered acceptable if the chi-squared value is < 1 for both the shear 276 velocity and resistivity data. The optimum value is the minimum summed value of the chi-

- 277 squared values for resistivity and shear velocity. We present the error as the maximum
- 278 value minus the minimum acceptable value divided by 2 for melt, melt water content and
- 279 temperature, which is the 95% confidence limit assuming symmetric error surfaces. We
- 280 acknowledge that this choice of reporting does not give a sense of the trade-offs in these
- 281 parameters.
- 282

Results 283

- 284 The shear-wave velocity structure derived from translating the MT models to seismic
- 285 velocity according to the petrophysical predictions (Fig. 5c, d) closely resembles the MT
- 286 models (Fig. 2 a, b), which is to be expected. We impose a water layer of 0.0 km/s in the
- 287 model, which results in the white area near the top of the model. The seismic velocities
- 288 range from 4.5 km/s in the upper 20-50 km of the Earth, with a minimum of 4.03 km/s
- 289 associated with the lowest resistivity regions. Strong lateral gradients are also visible in the
- 290 starting model, with changes of 0.4 km/s over less than 50 km, particularly near anomaly C.
- 291 The line II model has low velocity channels across the transect at 20-70 km depth and
- 292 several high velocity regions in the asthenospheric mantle.
- 293
- 294 When we use the shear-velocity model derived from MT (Fig. 5c, d) as the starting model for
- 295 the surface wave inversion we find a new shear-wave velocity model (Fig. 5 e, f) that more
- 296 closely resembles the MT models than the previously published model (Fig. 2). The highest
- 297 velocities are up to 4.81 km/s found in the fast lid, while the minimum velocity is 4.00 km/s,
- 298 found in anomaly B. The high velocity lid is more continuous than in the starting model but 299 follows a similar pattern of increasing thickness away from the ridges in both lines I and II. In
- 300 the asthenosphere, low velocity structures from the starting model are also retained.
- 301 Specifically, the channel structures in line II, anomaly E and F, are retained throughout much
- 302 of the model, particularly in the east near anomaly F, with similar velocities (~4.0 km/s). In
- 303 line I anomalies B and C are preserved i.e., ~4.0 km/s from the starting model. Anomaly A is
- 304 more pervasive beneath the ridge than in the MT starting model. Anomaly D is also
- 305 enhanced in the shear velocity model, with a high velocity of 4.56 km/s relative to the
- 306 starting model of 4.31 km/s at 100 km depth. The chi-squared values indicating goodness of
- 307 fit to the data is shown in Fig. 5a and 5b and are ~1 or less for most of the profile indicating
- 308 a fit this is within error. This goodness of fit is similar to the values from Harmon et al. (2020).
- 309
- 310
- 311 When we use the 1-D starting model from Harmon et al. (2020) for the surface wave
- 312 inversion, and the parameterisation and damping used here we find similarities and also
- 313 differences in comparison to Harmon et al. (2020) that illustrate the range of possible
- 314 models that fit the data (Fig. 5g, h). A high velocity lid is visible beneath the ridge and across
- 315 the region that ranges from 20-60 km in thickness. It shows low velocities beneath the ridge,
- with a stronger and shallower low velocity region beneath the ridge than in the model of 316
- 317 Harmon et al. (2020), although in general the features are similar and the velocity anomalies
- 318 are similar < 4.2 km/s but > 4.0 km/s. The normalized chi-squared fit to the data is shown in
- 319 Fig. 5 a,b for Line I and II, respectively. The chi-squared values are generally ~1 or less
- 320 indicating that the model fits the data within error and have a similar fit to the model with
- 321 the MT starting model. The new shear-velocity model with 1-D starting model presented in
- 322 Figure 5g, h is primarily for demonstrative purposes. The goal of the paper is to align the

323 previously published shear-wave velocity and resistivity models, and so we do not discuss

- 324 the model of Figure 5g, h further except for the purposes of resolution discussions.
- 325

The correlation between resistivity and shear-velocity is higher when the MT derived starting model is used in comparison to when the 1-D starting model is used in the shearvelocity inversion. For the 1-D starting model, there is a slope visible in line I (Fig. 6a), but

329 there is less of a visible relationship in line II (Fig 6b). Visually, the cross-plots for the MT

- derived starting model are more linear, with more of a slope visible in both lines I and II (Fig.
- 6 c, d). The correlation coefficients between the resistivity model and the shear velocity
- model assuming 1-D start model presented here are 0.41 and 0.29 for lines I and II, in other
- 333 words similar to that between the resistivity and the original shear velocity model presented
- in Harmon 2020 above (0.43 and 0.39 respectively). The correlation coefficients are higher,
- 0.56 and 0.62 for lines I and II respectively shear-wave model resulting from the MT-derived
 starting model. With the two lines combined the correlation coefficient is 0.60 (Fig. 7).
- 337
- 338 We illustrate the behaviour of the effect of varying the amounts of melt and water in the
- partial melt and compare it to the V_s and resistivity histogram for both line I and II (Fig. 7).
- 340 We use the thermal structure from the half-space cooling model shown in Fig. 4 but now
- allow partial melt at 0.1%, 1.0% and 3.0% where the mantle temperature exceeds 1100 °C.
- We also vary the amount water in the partial melt between 4-20%. The smallest amount of
- partial melt reduces the seismic velocity by << 1% in most cases, while the resistivity is 244 reduced by 264 log (Om) even the reason of vector contents presented here. At 1% melt bl
- reduced by ~0.6 $\log_{10}(\Omega m)$ over the range of water contents presented here. At 1% melt the shear velocity is reduced by ~2%, and the effect of increased water content is stronger,
- 2% reducing the resistivity up to ~1.5 log₁₀(Ω m) at the highest water contents. Finally, at 3%
- melt, the velocity is reduced by 4-5% and the resistivity reduction is up to ~2.1 $\log_{10}(\Omega m)$.
- 348 The span of partial melt and melt water contents considered here also generally spans the
- 349 range of most of the Vs/resistivity modelled values from our inversions, i.e., the
- 350 petrophysical values overlie the peak in the histogram. There is a slight bias in the seismic
- 351 velocities with a longer tail towards higher values.
- 352

353 Given the good general agreement between the petrophysical modelling and the shear 354 velocity and resistivity model values, we map the amount of partial melt, water content of 355 the melt, and temperature relative to the half-space cooling model onto the transects of 356 lines I and II (Fig. 8). We only perform this mapping where shear velocity is < 4.4 km/s and 357 $\log_{10}(\rho) < 1.5 \log_{10}(\Omega m)$, which is the nominal lower limit of the melt free half-space cooling 358 model (Fig. 4 and Fig. 7). In line I we find partial melt contents up to 4-4.5% near anomalies 359 B and C and similar maximum values in line II for anomalies E and F. Lower values of partial 360 melt < 2% are needed near anomaly A and for most of the other regions, typically requiring <361 1%. The water content of the melts is typically < 10 weight % for most (~60 %) of the total 362 anomaly area (colored regions in Fig. 8), with the notable exception of anomaly C which 363 requires up to 24 weight % water content to account for the low resistivity found in this 364 region and anomaly B which requires up to 15 weight %. There are other smaller patches of 365 high-water content visible near the edges of some of the regions, and within the channel of 366 anomaly E. The temperature structure generally has cooler temperatures 1100-1200 °C at depths < 100 km and temperatures > 1300 °C at greater depth. The grid search allows us to 367 368 assign error bounds corresponding to our presumed data errors (Fig. 9). The errors for the

- 369 melt percentages are typically <1%, while error for water content of the melt is on average 4
- 370 weight %, and the average errors for temperature are 26 °C.

371 Discussion

372 Our linear relationship between shear velocity and resistivity produced a reasonable starting

- 373 model for the shear-wave velocity inversion from Rayleigh waves. The inversion with the
- 374 MT-derived starting model was able to fit the phase velocity data within error for most of
- the two profiles shown here and fit the data as well as the 1-D model and the work of
- 376 Harmon et al., (2020). The MT-derived shear-velocity model improved the visual agreement
- and correlation coefficient between the resistivity and shear velocity model.
- 378

379 Overall, many of the in common features of the original works are retained and several of 380 the anomalies come into better agreement. For example, the MT-derived shear velocity 381 model retains the thickening of the lithosphere and the drip feature at anomaly D observed 382 in the Harmon et al. (2020) model. The lithospheric thickening with distance from the ridge 383 is more pronounced in the MT-derived shear velocity model in comparison to that of 384 Harmon et al. (2020). Anomalies B and C are also retained in the MT-derived model, 385 although anomaly B is more prominent than in the Harmon et al., (2020). In the 386 asthenosphere, better agreement between the resistivity model and the MT-derived shear 387 velocity model is achieved for the channel features in line II associated with anomaly F. 388 Anomaly C in the MT-derived shear velocity model has a morphology more similar to the MT 389 model than in the Harmon et al., (2020) model. Other anomalies such anomaly A shifts 390 shallower than the Harmon et al., (2020) model, and aligns better with a weak shallow 391 anomaly directly beneath the ridge in the resistivity model. Anomaly E is deeper than that in 392 the Harmon et al. (2020) model, again in better agreement with the resistivity model. 393

The differences in the shear velocity models here highlight some of the limitations of the

- approach. Specifically, inversion of Rayleigh wave phase velocities for shear velocity
 structure is non-unique, and this is well-known (Rychert et al., 2020) as many previous
- 397 works have demonstrated that a variety of models can fit a given dispersion curve. The
- differences between Harmon et al. (2020) (Fig. 2), the 1-D starting model with smoothing,
- damping, and parameterization of this study (Fig. 5e, f) and the MT-derived starting model
- 400 (Fig. 5g, h) illustrate this fact again and highlight that the strength of an anomaly can vary401 from model to model depending on the starting model, even if similar damping is used and
- 402 same fit is achieved as was the case here. For instance, the MT-derived shear-wave velocity
- 403 model includes velocities in Anomalies B, C by up to 1% slower in comparison with Harmon
- 404 et al. (2020), which impacts interpretation in terms of the presence of partial melt. Suitable
- 405 additional constraints are needed to determine which structure is the most likely, such as
- 406 information from receiver functions or resistivity.
- 407
- 408 The cross-plots indicate that the shear-wave velocity model and resistivity are in good
- 409 agreement with the petrophysics predictions for the half-space cooling model and variable
- 410 partial melt concentrations and melt water contents. About 80% of the shear velocity data
- 411 lie within 0.1 km/s of the petrophysical predictions for reasonable temperature structure,
- 412 melt and melt water contents (Fig. 7). The resistivity model is completely spanned by the
- 413 petrophysical predictions. Shear velocity appears to be biased towards higher values, which
- 414 may be a result of either the inversion process or a physical process. Shear-wave velocity

- 415 inversions can trade off velocities at shallow depths with deeper asthenospheric anomalies,
- 416 by compensating low asthenospheric values with higher lithospheric values. On the other
- 417 hand, other physical effects such as depletion (Schutt & Lesher, 2006) of peridotite through
- 418 ridge melting toward more harzburgitic compositions (Hacker & Abers, 2004) could cause
- 419 higher velocities by ~1-2%. In addition, anisotropy could also enhance the apparent velocity
- 420 by up to 1-3% (Rychert & Harmon, 2017; Saikia et al., 2021). In reality, it is likely some
- 421 combination of these physical effects which are not accounted for in the calculations used422 for predicting shear velocities.
- 423
- In this work we chose to force the shear velocity structure towards a closer match to the
- resistivity model, because the MT method has better resolution for certain features such as
- 426 thin channels, which is an assumption that is worth examination. We presumed the
- resistivity model has better structural resolution, but this assumption has limitations, since
 the 2-D assumption for the resistivity model may break down. For instance, anomaly E is
- 429 part of a larger 3-D anomaly visible that extends to the south along the Mid-Atlantic Ridge in
- 430 Harmon et al. (2020), and the depth of the anomaly is much greater in the resistivity
- 431 anomaly, perhaps owing to issues of dimensionality. Other observations, such as S-to-p
- 432 receiver functions, suggest there may be a shallower shear velocity anomaly associated with
- 433 anomaly E, which is necessary in order to produce a sharp velocity contrast in these regions
- 434 (Rychert et al., 2019). However, given that we prefer the MT-derived shear-wave velocity
- 435 structure for some of the major anomalies (A, B, C, D, and F), we proceed interpreting our
- 436 estimates for mantle melting and melt water content, bearing the limitations of the
- 437 inversions in mind.
- 438

439 The thermal structure predicted from our grid search (Fig. 8e, f) suggests relatively warm 440 temperatures beneath Anomalies B and C as well as the deeper parts of E (>1300 °C), while 441 Anomalies A and F have relatively low temperatures (1100-1200°C). This variability is likely a 442 result of the pressure dependence of the seismic waves. The low temperatures are generally 443 consistent with the interpretation that the shallow anomalies, particularly the channel 444 structures in F, are interacting with the base of the lithosphere (N Harmon et al., 2020; 445 Wang et al., 2020). The deeper, hotter anomalies (anomaly B and C) are also generally 446 consistent with the interpretation of upwelling from depth associated with small scale 447 convection.

448

449 The predicted melt fractions are in general agreement with our previous work from the 450 region, taking into account the various assumptions. Our melt fraction of up to 0.04 agrees 451 with the 0.01 - 0.07 previously reported based on the resistivity model alone (Wang et al., 452 2020). It is higher than the 0.005 to 0.015 reported by the previous shear-wave velocity 453 model (N Harmon et al., 2020). However, this can be explained by two main differences: 1) 454 The anomalies in the new shear velocity model presented here are up to 1 % slower than 455 those of the previous study (Harmon et al., 2020) and 2) We used the Takei, (1998) 456 relationship between melt and velocity here, which corresponds to about a 2 % velocity 457 reduction for 0.01 melt fraction in comparison to the 7.9% reduction for 0.01 melt fraction 458 from the work of (Hammond & Humphreys, 2000) used by Harmon et al. (2020). Our melt 459 fraction result of up to 0.04 is also consistent with the 6 – 11 % velocity drop with depth 460 velocity reduction required by receiver functions after correcting for the maximum effect of 461 temperature (Rychert et al., 2021), which would require melt fractions of 0.03 – 0.06

- 462 assuming the same melt-velocity relationship from Takei (1998) that we used here.
- 463

464 A different parameterization choice for the effects of melt on velocity due to different 465 assumptions on the melt geometry could yield lower melt fraction requirements by the 466 seismic constraints and still satisfy the resistivity model. Unconnected melt geometries such 467 as for isolated pockets or tubes (Schmeling, 1985) do not affect resistivity and so we can rule 468 those out (Naif et al., 2021). Assuming interconnected films and organized cuspate tubules 469 (Hammond & Humphreys, 2000), as used in Harmon et al., (2020), reduces the maximum 470 amount of partial melt fraction to < 0.02. Melt in the form of interconnected tubules and 471 cuspate geometries (Hammond & Humphreys, 2000), which have a velocity reduction of 472 14.5% per 0.01 melt fraction would suggest even lower melt fractions (< 0.01). Resistivity 473 does not depend on melt geometry. This is mostly due to the fact that the greatest 474 resistivity reduction occurs at melt fractions < 0.03, with a more gradual reduction in 475 resistivity at higher melt fractions (Fig. 10). However, since resistivity also has a strong 476 dependence on the volatile content in the melt, the lower melt fractions predicted for the 477 interconnected tubules and cuspate geometries could also satisfy the resistivity anomalies 478 with additional volatiles. More work would be required to determine the most likely partial 479 melt geometry and relationship for shear velocity reduction place better constraints on the 480 3-fold variation predicted from differing assumptions. 481 482 Predicted water contents are typically < 10 weight % for the melt but are surprisingly high, 483 up to weight 24%, in the centers of anomaly C, and F. Simple fractional or batch melting 484 calculations suggest that for a typical MORB mantle source with 100 ppm and an average 485 6% melting of the mantle suggest water contents of the melt should be ~0.2 weight % 486 (Workman & Hart, 2005). Higher water melt contents are possible for low degrees of partial 487 melting, for example <0.005 melt fraction yields > 1% weight water for 100 ppm in the 488 mantle source, and >7% weight water for 800 ppm in the mantle source. One possible

- 489 explanation is that these off-axis anomalies represent coalesced low-degree melts of a
 490 moderately wet mantle with high water content. There is some geochemical evidence
- 490 moderately wet mantle with high water content. There is some geochemical evidence for a
 491 moderately wet mantle from basalts collected from the ridge segments in the study area,
 492 is the study of the stud
- with estimated water contents that range from 110-770 ppm (~ 0.01-0.08 weight %) for the
 mantle source (Le Voyer et al., 2015). The advantage of this model is that wet melts are
- 494 stable and can persist in the mantle for long periods of time (Mehouachi & Singh, 2018).
 495

496 High CO₂ in the mantle melts is another possible explanation for the low resistivities 497 observed in region (Sifre et al., 2014), i.e., instead of high-water contents. Carbonated 498 peridotite is thought to exist in the mantle, although the abundance of carbon is relatively 499 low, likely < 100 ppm, as it is present in ancillary phases, rather than being hosted in olivine 500 or pyroxene (Dasgupta & Hirschmann, 2010). Carbonated melts are generated and stable at 501 greater depths, and only small degrees of partial melt are likely to be generated (<0.001 502 melt fraction) (Dasgupta & Hirschmann, 2010; Hirschmann, 2010). However, the melts could percolate upwards and coalesce generating higher CO₂ contents in the melt (Hirschmann, 503 504 2010). Fig. 10 shows the trade off in effective resistivity for 1 weight % water in the melt, 505 and 10% and 30% CO_2 by weight in the melt as a function of disequilibrium melt fraction 506 assuming 100 ppm in the un-melted mantle background for a depth of 80 km and a

507 temperature of 1350 °C. The figure is for demonstrative purposes since, melt fraction is

imposed rather than generated using batch melting or fractional melting, we did not vary
temperature as we did in the silicate case, and the melt may not necessarily be stable. At
30% CO₂ weight percent the resistivity is similar to the high-water content (20 weight %)

511 case. However, geochemical estimates of CO_2 in the primary ridge basalts range from 104

512 ppm to 1.9 weight % (Le Voyer et al., 2019), which is much lower than the >30 CO_2 weight %

needed to explain our results. To reach our high values, again aggregation of extremely low

- 514 degree partial melts would be required, and this also cannot be the melt that directly erupts
- 515 at the ridge.
- 516

517 Another possible explanation for the observed anomalies besides high water contents

518 (>10%) is sulfide melts, which are extremely conductive, >10⁴ S/m (Ducea & Park, 2000).

519 Small amounts of sulfide melts can rapidly reduce the effective resistivity of the aggregate.

520 To illustrate this we follow the parameterization of Ducea and Park (Ducea & Park, 2000), 521 using the (Gardés et al., 2014) parameterizations for the solid olivine and the Ni et al.,

- using the (Gardés et al., 2014) parameterizations for the solid olivine and the Ni et al.,
 (2011) parameterization for the silicate melt. We assume a conductivity of 10⁴ S/m. Fig. 10
- 523 shows a comparison between the effective resistivity for an olivine matrix with wet
- 524 disequilibrium melts and also for sulfide/wet disequilibrium melt mixtures with
- 525 predominately sulfide melt. Like the CO₂ case, this is for demonstrative purposes, without
- 526 varying a full suite of parameters. A nearly pure sulfide melt has a similar resistivity as a

527 silicate melt with 20% water, reaching values below 1 Ω m at < 0.01 melt fraction. So, in this

528 case, regions of high melt water contents in Fig. 8, e.g., anomaly C, could also be regions of

- 529 high sulfide melt content. Given the bulk abundance of sulphur measured in basaltic glasses
- in the region typically < 0.1 weight % (Le Voyer et al., 2015) and in <0.3 weight % in xenoliths
 from continents (Ducea & Park, 2000), it is unlikely that 0.04-0.05 sulfide melt fraction exists
- 532 in the mantle. However, a more conservative sulfide melt fraction of ~0.01 could at least
- 533 partially explain anomaly C (Hammond & Humphreys, 2000). There is also some evidence
- that melts from the nearby ridge segments are sulphur saturated (Le Voyer et al., 2015), and
- this may therefore suggest that sulfide melts may exist in higher abundance away from the

536 ridge melt triangle where silicate melts are in high abundance. Sulfide melts have also been

537 proposed to explain low seismic wave speeds in the asthenosphere (Helffrich et al., 2011).

538 Further work is needed to test whether sulfide melts would be compatible with small scale

- 539 convection and explain our off-axis anomalies, as they have a higher density than silicate 540 melts.
- 541

The melt anomalies inferred here extend to the base of our well-resolved region, ~150 km depth, which is greater than the 60 – 80 km predictions of a dry melting curve (Katz et al., 2003). This suggests that water or CO_2 induced melting is occurring at depth or the presence of sulfide melts or some combinations are active to produce melts so deep. In addition, the

546 largest melt fractions are associated with anomalies B, C, E and F, which are far from the

547 ridge axis. This suggests melt generation occurs away from the ridge either owing to small

548 scale upwellings, the presence of volatiles, or the combination of the two. Persistent melt

549 near the base of the lithosphere and apparent channelization near anomaly F also suggests

a role for water or other volatiles in the melts to stabilize them at relatively cool

temperatures near the base of the lithosphere (Mehouachi & Singh, 2018).

- 552
- 553 Our joint seismic -MT constraints require melt fractions (> 0.01) over large swaths of the
- asthenosphere mantle, several hundred kilometers, and hundreds of kilometers off the

- ridge axis. Such high percentages are not expected to persist over time and length scales
- that would enable seismic imaging (Spiegelman & Elliott, 1993). Melt fractions > 0.01 could
- be explained by a lack of a drainage route for the melt. Melt may coalesce at a permeability
- boundary at the lithosphere-asthenosphere boundary, as suggested by recent numerical
- 559 models that include 2-phase flow (Sim et al., 2020). Asthenospheric porosity in these models
- at a given snapshot in time can reach up to 10-20%, which could explain our melt fraction
- observations in the channels (Sim et al., 2020). The melt may also reduce the
- asthenospheric viscosity (Hirth & Kohlstedt, 1995; Jackson et al., 2006) potentially further
- 563 promoting small scale convection. Our observations in light of these geodynamic models
- 564 suggests that melt is dynamic but may be persistent on geological timescales.

565 Conclusions

- 566 We developed a simple relationship for shear velocity and resistivity of the oceanic
- 567 lithosphere and asthenosphere that can be used to link these quantities for joint inversions
- based on data from the I-LAB experiments and petrophysical modelling. We used the
- relationship to create a shear-wave starting model that we used to re-invert the phase
- 570 velocities. The new shear-wave velocity model more closely resembles the resistivity
- 571 models, in particular by including a low velocity channel and also in terms of the location
- 572 and shape of slow velocity anomalies. The apparent lithospheric drip was also enhanced.
- 573 Overall, the correlation between the surface wave and MT data sets increased. This suggests
- 574 that apparent discrepancies between the original models are more likely an artefact of
- 575 resolution and inversion schemes. Surface waves cannot resolve thin channel structures
- 576 unless significant prior knowledge is used in the starting model in the inversion. We also
- 577 demonstrate the utility for one-sided joint inversion of resistivity and shear velocity for
- 578 mantle melting and thermal structure based on petrophysical modelling. We show that
- shear velocity can place good constraints on melt volume, while resistivity can place good
- constraints on melt water content, CO_2 content or presence of sulfide melt given a simple
- 581 thermal structure such as the half-space cooling model.
- 582

583 Our estimates of melt, melt water content and temperature are in general reasonable and 584 within the expectations given geochemical outputs from the nearby ridge segments. The

- one exception is very high water or CO_2 contents (>15%) estimated in the slowest and least
- 586 resistive anomalies. These high melt water or CO₂ contents could be real but would require
- 587 coalescing low degree partial melts of moderately wet or carbon-rich mantle sources.
- 588 Alternatively, nearly pure sulfide melts at small fractions could potentially partially explain
- 589 these anomalies. Overall, joint interpretation and/or inversion of resistivity and shear
- 590 velocity models holds promise for resolving debates about the lithosphere-asthenosphere
- 591 system and the presence and character of partial melt in the mantle.
- 592

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- the IRIS DMC, as 2016-2017 network XS <u>https://doi.org/10.7914/SN/XS_2016</u>. (Rychert et al., 2016).









612 in inset map indicates study area.



613 614 Figure 2. Resistivity model and shear-wave velocity model from previous work. Panels (a) 615 and (b) show contoured resistivity transects from line I and line II, respectively, from Wang 616 et al. (2020). Contour interval is 0.5 log units. Panels (c) and (d) show contoured shear 617 velocity transects for line I and II, respectively, from Harmon et al. (2020). Contour interval is 618 0.05 km/s. Anomalies A, B, C, D, E and F from Harmon et al. (2020) are indicated.





620 621 Figure 3. Cross-plot histograms of resistivity and shear-wave velocity from previous work.





624 625

Figure 4. Petrophysical predictions for resistivity and shear-wave velocity for half-space 626 cooling model. Panel a shows the thermal structure for the half-space cooling model, b 627 shows the predicted shear-wave velocity structure, and c shows the predicted resistivity 628 structure predicted from petrophysics calculated as described in the text. White line in 629 panel a indicates the predicted melt triangle for 100 ppm water in a background mantle

630 (Katz et al., 2003). Panel d shows the cross-plot of predicted resistivity and shear velocity

631 without melt from panel b and c (black circles) and with a presumed melt fraction (0.01)

632 containing different amounts of water (4-20%), within the predicted melt triangle (yellow

633 and brown circles). Grey line in Panel d shows preferred linear relationship between

634 resistivity and shear velocity based petrophysical modelling presented here and consistent

635 with the cross-plot histograms presented in Figure 3.



636 637 Figure 5. Shear-wave velocity inversions based on resistivity predictions. Panel a and b 638 show misfit along line I and II using normalized chi-squared. Panel c and d show the shear-639 wave velocity models that result from translating the resistivity model shown in Fig. 2 to 640 velocity using the linear relationships based on petrophysical modelling. Panels e and f show 641 the shear-wave velocity inversion results using panels c and d, respectively, as starting 642 models. Panels g and h the show shear velocity inversion results using the 1-D starting 643 model from Harmon et al. (2020) and the smoothing, damping, and model parameterisation 644 used here. Contour interval is 0.05 km/s. Asthenospheric anomalies A, B, C, D, E and F from 645 Harmon et al. (2020) and Wang et al. (2020) are shown for reference.







656

Figure 7. Cross-plot histogram of resistivity and shear-wave velocity from the MT-derived

shear-wave velocity model for both lines I II and petrophysical predictions. Purely thermal

predictions are shown as black dots. Colored dots show predictions for various melt

fractions and melt water contents. Legend indicates the amount of imposed disequilibrium

melt fraction (0.001, 0.01 and 0.03) and water content of the melt in weight % (4-20%).



665 Figure 8. Results of grid search for partial melt, melt water content and mantle

temperature. Panels a and b show results for partial melt fraction, panels c and d show

water content of the partial melt, and panels e and f show the result for temperature for
lines I and II, respectively. Anomalies A, B, C, D, E and F are plotted at the same locations as
in Figure 2 for reference.



672 673 Figure 9. Error estimates of grid search for partial melt, water content of the melt and

674 mantle temperature. Panels a and b show partial melt fraction error, panels c and d show
675 water content of the partial melt error, and panels e and f show temperature error for line I

and line II, respectively. Anomalies A, B, C, D, E and F are plotted at the same locations as in

- 677 Figure 2 for reference.
- 678



Figure 10. Effective resistivity predictions for water, CO₂ and sulfide in silicate melts as a

681 function of melt fraction. We assume a solid mantle with 100 ppm water and disequilibrium 682 melt at 1300°C. Legend indicates the respective water, CO₂ and sulfide concentrations. Melt

683 with water only is shown as blue lines. Melt that includes water and CO₂ is shown as cyan

684 lines and melt that includes water and sulfide is shown as red lines.

685

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