# Modeling lithospheric thickness along the conjugate South Atlantic passive margins implies asymmetric rift initiation

Peter Haas<sup>1</sup>, R. Dietmar Muller<sup>2</sup>, Jörg Ebbing<sup>1</sup>, Nils-Peter Finger<sup>3</sup>, and Mikhail K Kaban<sup>4</sup>

<sup>1</sup>Kiel University <sup>2</sup>University of Sydney <sup>3</sup>GFZ German Research Centre for Geosciences <sup>4</sup>GeoForschungsZentrum Potsdam

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#### Abstract

The lithospheric architecture of passive margins is crucial for understanding the tectonic processes that caused the break-up of Gondwana. We highlight the evolution of the South Atlantic passive margins by a simple thermal lithosphere-asthenosphereboundary (LAB) model based on rifting time, crustal thickness, and stretching factors. We simulate the different rifting stages that caused the opening of the South Atlantic Ocean and pick the LAB as the T=1330 °C isotherm, which is calculated by 1D advection and diffusion. In a synthetic example, we demonstrate that the initial crustal thickness has the largest effect on the thermal LAB. For the South American passive margin, our modeled LAB shows a deep and smooth structure between 110-150 km depth at equatorial latitudes and a more variable LAB between 50-200 km along the southern part. This division reflects different stages of the South Atlantic opening: initial opening of the southern South Atlantic causing substantial lithospheric thinning, followed by rather oblique opening of the equatorial South Atlantic accompanied by severe thinning. The modeled LAB reflects a high variability associated with tectonic features on a small scale. Comparing the LAB of the conjugate South American and African passive margins in a Gondwana framework reveals a variable lithospheric architecture for the southern conjugate margins. Along selected conjugate margin segments stark differences up to 80 km of the LAB depths correlate with strong gradients in margin width. This mutual asymmetry suggests highly asymmetric melting and lithospheric thinning prior to rifting.

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7	<sup>1</sup> Institute of Geosciences, Kiel University, Kiel, Germany
8	<sup>2</sup> Earthbyte Group, School of Geosciences, The University of Sydney, Sydney, Australia
9	<sup>3</sup> GFZ Research Centre for Geosciences, Potsdam, Germany
10	<sup>4</sup> Free University of Berlin, Berlin, Germany
11	<sup>5</sup> Schmidt Institute of Physics of the Earth, Moscow, Russia
12	
13	Corresponding author: Peter Haas (peter.haas@ifg.uni-kiel.de)
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15	Key Points:
16	• A simple thermal LAB model for the South Atlantic passive margins has been developed
17 18	• The LAB model shows highly variable lithospheric thickness that reflects different rifting stages of the South Atlantic
19 20 21	• The difference of LAB depths for conjugate margin pairs correlates with asymmetry in margin width, implying asymmetric lithospheric thinning

#### 22 Abstract

The lithospheric architecture of passive margins is crucial for understanding the tectonic 23 processes that caused the break-up of Gondwana. We highlight the evolution of the South 24 Atlantic passive margins by a simple thermal lithosphere-asthenosphere-boundary (LAB) model 25 based on rifting time, crustal thickness, and stretching factors. We simulate the different rifting 26 27 stages that caused the opening of the South Atlantic Ocean and pick the LAB as the T=1330 °C isotherm, which is calculated by 1D advection and diffusion. In a synthetic example, we 28 demonstrate that the initial crustal thickness has the largest effect on the thermal LAB. For the 29 South American passive margin, our modeled LAB shows a deep and smooth structure between 30 110-150 km depth at equatorial latitudes and a more variable LAB between 50-200 km along the 31 southern part. This division reflects different stages of the South Atlantic opening: initial opening 32 33 of the southern South Atlantic causing substantial lithospheric thinning, followed by rather oblique opening of the equatorial South Atlantic accompanied by severe thinning. The modeled 34 LAB reflects a high variability associated with tectonic features on a small scale. Comparing the 35 LAB of the conjugate South American and African passive margins in a Gondwana framework 36 reveals a variable lithospheric architecture for the southern conjugate margins. Along selected 37 conjugate margin segments stark differences up to 80 km of the LAB depths correlate with 38 strong gradients in margin width. This mutual asymmetry suggests highly asymmetric melting 39 40 and lithospheric thinning prior to rifting.

#### 41 Plain Language Summary

42 Passive margins mark the transition zone from a continent to the ocean without being an active boundary of tectonic plates. They are typical for all continents on the globe. In the South 43 Atlantic, the passive margins are located adjacent to the eastern coastline of South America and 44 the western coastline of Africa. Studying the architecture of passive margins is essential for 45 understanding plate tectonic history of the earth because they define how the continents once 46 belonged together and how they broke apart. Passive margin segments on opposite sides of an 47 48 ocean form so called conjugate margin pairs. Most geophysical studies of passive margins focus on the near-surface architecture. However, their deeper extension to the base of the rigid shell of 49 50 the earth, known as lithospheric thickness, is to a large extent unknown. Based on a simple temperature model we find that the lithospheric thickness is highly variable and shows structural 51 variations along the South Atlantic passive margins. These differences are associated with the 52 extension of conjugate margin pairs: where one margin is narrower than the conjugate, its 53 lithospheric thickness is greater. This asymmetry indicates that the geodynamic processes, 54 causing the break-up of the two continents, must have been asymmetric as well. 55

### 56 **1 Introduction**

The architecture and evolution of passive margins have been extensively studied over the last decades (e.g., Duretz et al., 2016; Geoffroy, 2005; Lister et al., 1986; Reston, 2009). For a long time, this was predominantly motivated by hydrocarbon exploration because numerous oil and gas reservoirs are connected to passive margin formation. Nowadays, passive margins become the focus of attention for sequestering carbon dioxide (e.g., Ringrose and Meckel, 2019), as well

as for estimating the global carbon dioxide budget over deep time (e.g., Brune et al., 2017).

The McKenzie model of rifting is a widely accepted model that explains thinning of the continental lithosphere and subsequent stretching in pure-shear mode associated with tectonic

65 subsidence (McKenzie, 1978). Jarvis and McKenzie (1980) introduced a time-dependent

analytical model that relates variations in heat flow and subsidence history to the rate of
 extension. The formation of non-volcanic passive margins is the endmember of the McKenzie
 rifting model with passive upwelling of buoyant sublithospheric mantle material, driven by far
 field extension forces (Geoffroy 2005; Sengör and Burke, 1978)

69 field extension forces (Geoffroy, 2005; Sengör and Burke, 1978).

70 In many places, volcanic passive margins have been interpreted as the result of active upwelling

of a mantle plume, associated with a thick crust due to magmatic underplating and the formation  $G_{1}$ 

of Seaward Dipping Reflectors (SDR; e.g., Geoffroy, 2005; Mutter et al., 1982). The occurrence of volcanic rifted margins does not necessarily require a pronounced thermal anomaly in the

of volcanic rifted margins does not necessarily require a pronounced thermal anomaly in the mantle related to a plume (e.g., Bown and White, 1995), but can be explained with transient

small-scale mantle convection underneath the lithosphere (Nielsen, 2002; Simon et al., 2009) or

76 plume-rift interaction (Morgan et al., 2020).

The crustal architecture of passive margins shows a high diversity that cannot be characterized 77 only in terms of magmatic budget (e.g., Tugend et al., 2018). One important observation is 78 asymmetry of opposed margins, which requires other mechanisms than pure shear (Lister et al., 79 1986). A proposed mechanism is detachment faulting along low-angle normal faults, which cut 80 through the entire lithosphere (Lister et al., 1986). This concept is based on simple shear 81 (Wernicke, 1981). Brune et al. (2014) showed that margin asymmetry is caused by rift migration 82 of only upper crustal faults, associated with a large melt transfer between two rift sides. Tugend 83 et al. (2018) propose that timing of decompression melting may be more important than 84 85 estimates of the magmatic budget of passive margins to understand their evolution and variability. 86

All rifting models have in common that the initial lithospheric thickness is thinned prior to rifting, followed by subsidence and cooling of the lithosphere. The concept of half-space cooling predicts lithospheric thickening with time during the first ~80 million years after crustal accretion at a mid-ocean-ridge (e.g., Turcotte and Schubert, 2018). However, at passive margins this concept fails as the oceanic crust is older than 80 Ma in many places and the relation of passive and active upwelling of mantle material is not known. Due to this ambiguity, the amount of lithospheric thinning prior to margin formation is also often unknown.

In this study, we focus on the formation and lithospheric thickness of the South Atlantic passive 94 margins. Their architecture cannot be explained by uniform rifting models only and shows a 95 wide range of volcanic and non-volcanic passive margin types. In the Late Jurassic, rifting 96 started and caused the disintegration of Western Gondwana, leading to the opening of the 97 southern South Atlantic (e.g., Rabinowitz and LaBrecque, 1979). In the Late Aptian/Early 98 Albian, the equatorial part of the South Atlantic opened (Moulin et al., 2010), characterized by a 99 higher degree of oblique rifting (Brune et al., 2018). The equatorial opening of the South Atlantic 100 was dominated by far field forces (e.g., Heine et al., 2013; Moulin et al., 2010). 101

To what extent the Tristan Hotspot, which formed the Parana-Etendeka flood basalts (Granot and 102 Dyment, 2015; Renne et al., 1992), contributed to the initial break-up is still discussed. 103 104 Combined seismic imaging and potential field data analysis for the central and southern segments of the South Atlantic suggests that the Tristan Hotspot sourced the magmatism volume 105 but might not necessarily have altered the process of rifted margin formation (Blaich et al., 106 2011). Comparison between Large Igneous Provinces (LIP) and break-up age shows that rifting 107 occurred before LIP emplacement. This indicates that rifting might have been initiated by 108 tectonic forces and the Tristan plume only guides the mantle material towards the thinned 109

110 lithosphere (Buiter and Torsvik, 2014). Morgan et al. (2020) propose that along-rift flow of

plume material causes the formation of volcanic passive margins, associated with thinning of the initial lithosphere.

In recent years, several global and regional lithospheric thickness models have been published 113 that cover the South Atlantic passive margins. Global models are, for example, derived from 114 surface-wave dispersion maps (Pasyanos et al., 2014), conversion of seismic tomography to 115 thermal LAB (Steinberger and Becker, 2018) or multi-probabilistic joint inversion (Afonso et al., 116 2019). The global LAB models have a wide depth range, partly depending on the data sets and 117 regularizations used in establishing the models. Due to the narrow and elongated margin 118 geometry, many of these global models are not capable of mapping the LAB in this region. For 119 example, Finger et al. (2021) present a regional model for the South American continent based 120 121 on combined density, thermal, and compositional modeling, which can be converted to a LAB model. However, the thermal field, which can be converted to LAB depths, does not adequately 122 represent the passive margins as their modeling approach focusses on the continental platforms. 123

In this paper, we predict a thermal LAB depth of the stretched region along the passive margins 124 of the South Atlantic. We first introduce our method for the South American passive margin. The 125 thermal LAB depth is derived from three input parameters: stretching factors, rifting time, and 126 127 crustal thickness. We calculate the stretching factors by accounting for crustal thickness gradients across the deforming region. Together with rifting time and published crustal thickness 128 models, we then calculate a thermal model of the extended lithosphere. The LAB depth is 129 130 defined by extrapolating the linear geotherm from the crust throughout the lithosphere after rifting occurred. Next, we discuss the evolution of the thermal LAB for the conjugate South 131 Atlantic passive margins in a Gondwana framework using GPlates software (Müller et al., 2018). 132 We further evaluate differences between conjugate margin basins by correlating the predicted 133

134 LAB depth with margin width.

### 135 **2 Methods**

We calculate lithospheric thickness as a function of rifting time, crustal thickness, and stretching factor. For that, we use the python code *RiftSubsidence* based on software that was originally designed to calculate theoretical subsidence curves for rifting scenarios in 1D (White et al., pers.

com.). In *RiftSubsidence*, the subsidence is calculated based on the amount and timing of pure shear lithospheric extension, as well as on the thermal and density structure of the lithosphere.

140 In this approach, the lithospheric thickness is derived from the thermal structure after the

141 In this approach, the hubspheric uncertex is derived from the methal structure after the 142 lithosphere has been stretched. The temperature of the model is calculated by 1D advection and 143 diffusion using finite differences. The top of the model is defined at sea level, while the base of 144 the model is defined as the LAB. At these boundaries, the temperature is fixed throughout the 145 entire rifting period.

146 Prior to lithospheric stretching the LAB depth  $z_{LAB}$  and LAB temperature  $T_{LAB}$  must be defined.

147  $z_{LAB}$  is balanced isostatically against a reference Mid-Ocean-Ridge (MOR). Figure 1 shows the

isostatic model of the reference MOR on the left and passive margin on the right. Assuming that

149 the thickness of the mantle lithosphere  $h_m$  is the only unknown parameter, the isostatic equation

150 can be defined as:

151

$$h_m = \frac{\rho_w h_w + \rho_c (h_{c0} - h_c) + \rho_a (h_c - h_w - h_{c,ref})}{(\rho_m - \rho_a)} \tag{1}$$

The assumed values of density and thickness of each layer are listed in Table 1. The crustal thickness varies spatially for the passive margin. Accordingly, the initial LAB depth  $z_{LAB}$  is

individual for each point and is defined as:  $z_{LAB} = h_c + h_m$ . For crustal density, we calculate the

average value over the entire deformable region (see Table 1). The value of  $\rho_c = 2.809$  g/cm<sup>3</sup> is 155

obtained by balancing isostatically thicknesses and densities of crystalline crust and sediments of 156 the crustal model of Finger et al. (2021). Thus, it represents a mean value of the entire crust. 157

158

Variable	Name	Value
$\rho_{w}$	Density of sea water	1.03 g/cm <sup>3</sup>
ρ <sub>c</sub>	Density of crust	$2.809 \text{ g/cm}^{3}$
$\rho_{\rm m}$	Density of lithospheric mantle	Unknown
$\rho_{a}$	Density of asthenosphere	$3.3 \text{ g/cm}^{3}$
h <sub>w</sub>	Height of sea water	$2.5 \text{ km}^2$
h <sub>c,ref</sub>	Thickness of crust at MOR	7 km <sup>2</sup>
h <sub>c</sub>	Thickness of crust	Variable <sup>1</sup>
h <sub>m</sub>	Thickness of lithospheric mantle	Unknown

**Table 1.** Layers of isostatic model at passive margin with respective densities. <sup>1</sup>Thickness and 159

- density of the crust are taken from Finger et al. (2021). <sup>2</sup>Thicknesses of MOR reference column 160
- are taken from Afonso et al. (2019). <sup>3</sup>Density of asthenosphere taken from Zoback and Mooney 161 (2010).
- 162
- 163

$ ho_{\scriptscriptstyle W}$	h <sub>w</sub>	$\rho_c$	$h_c$
$ ho_c$	h <sub>c,ref</sub>	PC	
			Î
$ ho_a$		$\rho_m$	h <sub>m</sub>
			Ļ
		$ ho_a$	

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Figure 1. Isostatic balance of MOR (left) and passive margin (right). Mantle density and 165 thickness of mantle lithosphere are the unknown parameters. Full names of the parameters are 166 given in Table 1. 167

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The isostatic balance in Eq. 1 assumes a constant mantle density. As the South American passive 169 margin is bounded by different continental and oceanic tectonic regimes, its mantle density is 170 heterogeneous. Estimates from global and regional inversions indicate varying mantle densities 171 along the South American passive margin (e.g., Afonso et al., 2019; Finger et al., 2021). 172 Therefore, the mantle density is the second unknown parameter of the isostatic column. 173

Prior to rifting the thermal state of the lithosphere can be regarded as purely conductive with a 174

linear geotherm. Assuming a constant linear geotherm for the entire lithosphere (blue lines in 175

Figure 2), we calculate the mantle density and thickness of the mantle lithosphere in an iterative 176

scheme: 177

178				
179	1. Select a starting	ng value of $h_{m,0}$		
180		-	out the lithosphere, the mantle temperature $T_m$ can	be
181	-		temperature at the Moho $T_{Moho}$ :	
	-			
	$T_m = \frac{T_{LAB} + T_{LAB}}{2}$	T <sub>Moho</sub>	(2)	
	<sup>1</sup> <sup>m</sup> – 2			
182				
183			tient of crustal and lithospheric thickness:	
	$T_{IAR}(1$	$+\frac{h_c}{h_c+h_{m,i}}\Big)$	(3)	
184	T -	$n_c + n_{m,i}$		
185	m = 1	2		
186	with $n_{m,i} = \tan \theta$	ickness of mantle lith	osphere at iteration step <i>i</i> .	
187	2 1	la danaitar an a francti	on of tommentum based on volumetric coefficient	.f
188		is density as a function sion $\alpha$ (e.g., Turcotte	on of temperature, based on volumetric coefficient	01
189	inermai expan	ision $\alpha$ (e.g., 1 urcolle		
	······································		and Schubert 2010).	
	$d\rho = -\rho \alpha dT$		(4)	
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190	$d\rho = -\rho\alpha dT$	-		
190 191	$d\rho = -\rho \alpha dT$ Solve the diffe	erential:	(4)	
191	$d\rho = -\rho\alpha dT$	erential:		
191 192	$d\rho = -\rho \alpha dT$ Solve the diffe	erential:	(4)	
191 192 193	$d\rho = -\rho \alpha dT$ Solve the diffe $\rho_a - \rho_{m,i} = -\rho_{m,i}$	erential: $_{i}\alpha(T_{LAB} - T_{m})$	(4)	
191 192	$d\rho = -\rho \alpha dT$ Solve the difference $\rho_a - \rho_{m,i} = -\rho_{m,i}$ Solve Eq. 5 for	erential: $_{i}\alpha(T_{LAB} - T_m)$ or mantle density:	(4) (5)	
191 192 193 194	$d\rho = -\rho \alpha dT$ Solve the difference $\rho_a - \rho_{m,i} = -\rho_{m,i}$ Solve Eq. 5 for	erential: $_{i}\alpha(T_{LAB} - T_m)$ or mantle density:	(4)	
191 192 193 194 195	$d\rho = -\rho \alpha dT$ Solve the diffe $\rho_a - \rho_{m,i} = -\rho_{m,i}$	erential: $_{i}\alpha(T_{LAB} - T_m)$ or mantle density:	(4) (5)	
191 192 193 194 195 196	$d\rho = -\rho\alpha dT$ Solve the difference of the second s	erential: $_{i}\alpha(T_{LAB} - T_m)$ or mantle density: $\overline{T_{3} - T_m}$	(4) (5) (6)	t
191 192 193 194 195 196 197	$d\rho = -\rho \alpha dT$ Solve the difference of the second	erential: $_{i}\alpha(T_{LAB} - T_m)$ or mantle density: $\overline{T_{3} - T_m}$	(4) (5)	ent
191 192 193 194 195 196 197 198	$d\rho = -\rho\alpha dT$ Solve the difference of the second s	erential: $_{i}\alpha(T_{LAB} - T_m)$ or mantle density: $\overline{T_{3} - T_m}$	(4) (5) (6)	ent
191 192 193 194 195 196 197 198 199	$d\rho = -\rho\alpha dT$ Solve the difference of the second s	erential: $_{i}\alpha(T_{LAB} - T_m)$ or mantle density: $\overline{g_{3} - T_m}$ antle density at the i	(4) (5) (6) iteration step <i>i</i> and $\alpha$ =thermal expansivity coefficient	ent
191 192 193 194 195 196 197 198	$d\rho = -\rho\alpha dT$ Solve the difference of the second s	erential: $_{i}\alpha(T_{LAB} - T_m)$ or mantle density: $\overline{T_{3} - T_m}$	(4) (5) (6) iteration step <i>i</i> and $\alpha$ =thermal expansivity coefficient	ent

202 5. Repeat Step 2-4

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The process is iterated until the density change reaches the threshold  $\|\rho_{m,i+1} - \rho_{m,i}\| < 0.001$ g/cm<sup>3</sup>. For  $T_{LAB}$  and  $\alpha$  we choose standard values of  $T_{LAB}=1333$  °C and  $\alpha=3.28*10^{-5}$  1/K (e.g., Jarvis and McKenzie, 1980; Parsons and Sclater, 1977). More specific values of  $\alpha$  for oceanic lithosphere exist ( $\alpha=3.45$ ), which are derived from mineral physics (Afonso et al., 2005). However, the effect on the modeled LAB is only minor.

 $z_{LAB}$  represents the depth of the LAB prior to rifting. If the lithosphere is not stretched, a linear geotherm for the entire lithosphere can be assumed. However, rifting causes lithospheric extension with subsequent non-linear displacement of the geotherm. After extension ceases, the lithosphere cools and the geotherm relaxes back to the linear state at infinite time. Depending on initiation and end of rifting and the amount of stretching, the geotherm will deviate from its initial linear state (Figure 2). If we assume that conductive heat transport in the crust is the

- 215 dominant heat source and that thermal expansion is constant both for the crust and lithosphere,
- we can extrapolate the geotherm of the crust throughout the mantle lithosphere. As a result,  $z_{LAB}$ is shifted upward (Figure 2). The difference between  $z_{LAB,init}$  based on the isostatic model and
- 218  $z_{LAB,rifted}$  based on the extrapolated crustal geotherm, is defined as  $\Delta z_{LAB}$ .
- 219

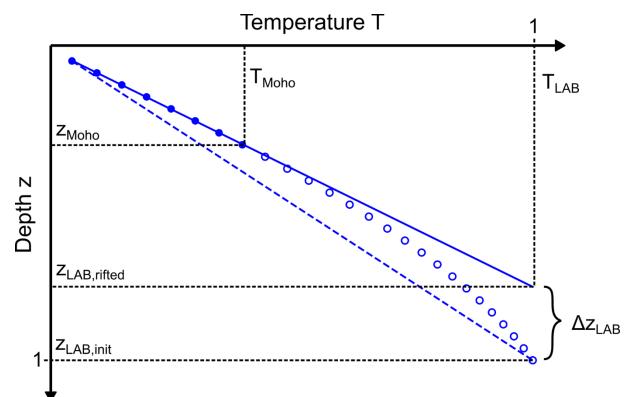


Figure 2. Geotherms and the LABs. Dashed line: Initial geotherm before rifting, circles: uplifted geotherm after rifting and cooling, solid blue line: extrapolated linear geotherm based on thermal structure of crust (filled circles). The depth at which the extrapolated geotherm reaches  $T_{LAB}$ provides an approximate estimate of  $z_{LAB}$ .

225

The amount of non-linearity of the distorted geotherm in Figure 2 depends on the elapsed time and on the amount of stretching. Based on the concepts described by McKenzie (1978), we define the stretching factor as the thickness of unthinned crust divided by thinned crust:

$$\beta = \frac{h_i}{h_s} \tag{7}$$

230  $h_i$  defines the initial crustal thickness, whereas  $h_s$  is the crustal thickness after stretching.

Müller et al. (2019) used this approach to derive stretching factors globally for all deforming regions. However, their values are calculated based on uniform stretching and do not consider crustal thickness gradients. We derive new stretching factors for the South American passive margin, by selecting an updated crustal thickness model of South America by Finger et al. (2021), which is based on available seismic determinations.

The Continent-Ocean-Boundary (COB) can be divided in an inner part, defined as the landward limit of stretched continental crust, and an outer part, defined as the oceanward limit of stretched

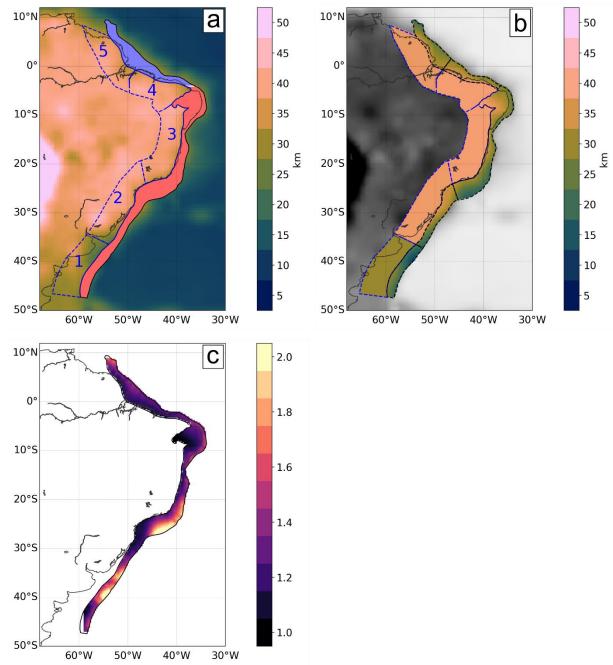
- continental crust (Müller et al., 2019). The area between the inner and outer COB is defined as the deforming region and represents the thinned error h (see Figure 2). In the following, we will
- the deforming region and represents the thinned crust  $h_s$  (see Figure 3). In the following, we will

treat the term 'deforming region' as 'passive margin'. To obtain the unthinned crust  $h_i$ , the inner 240 margin of the COB is extended 500 km towards the continent. Figure 3a shows the crustal 241 thickness of Finger et al. (2021), where the South American passive margin is defined by its 242 inner and outer COB. For the equatorial and southern margin different rifting stages are assumed. 243 We define five different segments of extended COB, based on adjacent onshore crustal thickness 244 and surface geology. For each segment, we calculate the mean value of crustal thickness, which 245 represents unthinned crust  $h_i$  (Figure 3b). Table 2 lists the average values of the different 246 segments. Note that the average crustal thickness of units 2-5 is almost identical. Figure 3c 247 shows the resulting stretching factor. 248

249

Segment	Geology	Crustal Thickness [km]
1	Colorado Basin (offshore)	29,9
2	Parana Basin	38,2
3	Sao Francisco	38,8
4	Amazonia South	38,6
5	Amazonia North	37,5

Table 2. Average crustal thickness values for the different geological segments shown in
 Figure 3b.



253

Figure 3. Crustal thickness to calculate stretching factors. a: Crustal thickness model of Finger et 254 al. (2021). Masked area: Geometry of the South American passive margin. The coral polygon 255 corresponds to the early opening of the Atlantic from 141-120 Ma, the blue polygon to the later 256 stage from 121-107 Ma (rifting time taken from the GPlates data base, as published in Müller et 257 al., 2019). Blue dashed contours and numbers represent different segments of crustal thickness, 258 extended 500 km towards the inner continent. b: Crustal thickness values for deformable region 259 and inner extension area. Values for inner area are averaged. c: Stretching factors of deforming 260 region of passive margin. 261 262

#### **3 Synthetic example – Varying input parameters of** *RiftSubsidence*

Originally, *RiftSubsidence* was designed to calculate theoretical subsidence curves based on the input parameters rifting time, stretching factor, and crustal thickness (White et al., pers. com.). Within the following synthetic example, we show the effect of each parameter on the resulting lithospheric thickness after rifting.

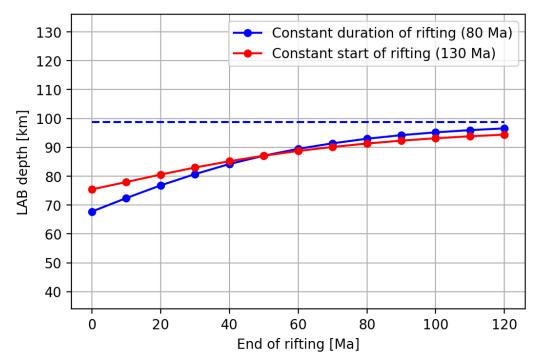
First, we select default values of the initial parameters. For rifting time  $t_{rift}$  we choose 130-50 Ma, stretching factor  $\beta = 2$ , crustal thickness  $h_c = 30$  km. Each parameter is varied in a certain range, while the other parameters are fixed at the respective default values. Densities and thicknesses of the isostatic reference model are taken from Table 1. Here, we assume a constant mantle density of  $\rho_m = 3.35$  g/cm<sup>3</sup>. According to Eq. 1, the only unknown parameter is now  $h_m$ . In another synthetic example, we vary the densities of the reference isostatic column and investigate to which extent this contributes to the LAB depth after rifting (supporting information S1).

We investigate two different scenarios of  $t_{rift}$ : first, we set the initiation of rifting at 130 Ma and vary the end from 120 to 0 Ma. Second, we set a constant duration of rifting of 80 Ma and start at 200 Ma Eigenvalue the and time of victime scenario the LAD doubt of second scenario from the second scenario the second scenario from the second scenario the second scenario from the secon

200 Ma. Figure 4 shows the end time of rifting versus the LAB depth after rifting  $z_{LAB,rifted}$ . For 277 constant time of rifting initiation at 130 Ma zLAB, rifted decreases towards the present time. The 278 same trend can be observed for a constant duration of rifting, but  $z_{LAB,rifted}$  is even shallower than 279 before. *z<sub>LAB,rifted</sub>* is progressively decreasing towards the present day. The shorter the rifting time, 280 the less time the lithosphere has to cool down, resulting in an uplifted geotherm and a stronger 281 deviation from the initial isostatic LAB z<sub>LAB,init</sub>. Comparing both scenarios, the advection rate is 282 the same only if the period of extension and the stretching factor are the same. z<sub>LAB,rifted</sub> becomes 283 shallower for the model with constant duration of rifting because the extension has occurred until 284 285 more recent times (For example, 130-0 Ma for the red curve vs. 80-0 Ma for the blue curve).

By varying the stretching factor  $\beta$  between 1-3 and taking the default values of the other 286 parameters, we observe an almost linear shallowing of  $z_{LAB,rifted}$  towards higher stretching factor 287 (Figure 5). The difference of ~20 km between  $z_{LAB,rifted}$  and  $z_{LAB,init}$  is of the same magnitude as 288 289 with the variable rifting time in Figure 4. Figure 5 also shows the LAB depth after varying the initial crustal thickness between 20-40 km. Note that in this case  $z_{LAB.init}$  is not constant. 290 Increasing the Moho depth causes increasing the LAB depth. Isostasy shows that the mass deficit 291 of a deeper Moho is compensated by a thicker mantle lithosphere (right column of Figure 1). 292 293 This trend is not linear because the crustal thickness  $h_c$  controls two coefficients  $\rho_c$  and  $\rho_a$  in Eq. 1. At the of Moho depth ~26 km and LAB depth of ~60 km,  $z_{LAB,rifted}$  starts to deviate from 294  $z_{LAB,init}$ . The effect becomes stronger with increasing the Moho depth. Here, the effect of rifting 295 and stretching becomes significant. For this configuration, a ratio of the Moho and LAB depths 296 297 of 0.43 is the threshold for significant difference between  $z_{LAB,init}$  and  $z_{LAB,rifted}$ .

This synthetic example demonstrates that the initial crustal thickness has the strongest effect on  $z_{LAB,rifted}$ . In the supporting information we show how  $z_{LAB,rifted}$  varies for spatially variable crustal thickness, using two recently published crustal thickness models of South America (Finger et al., 2021; Haas et al., 2020).



**Figure 4.** Rifting time vs. LAB depth for different initiation time and duration of rifting. The red curve shows the  $z_{LAB,rifted}$  for constant start of rifting at 130 Ma with variable end between 120 to 0 Ma. The blue curve indicates  $z_{LAB,rifted}$  for constant duration of rifting of 80 Ma with variable onset of rifting. The first red dot on the right side cuts the x-axis at 120 Ma and corresponds to a rifting period from 130-120 Ma. The first blue dot corresponds to a rifting period from 200-120 Ma. The dashed blue line shows  $z_{LAB,init}$  based on isostatic balance of Figure 1.

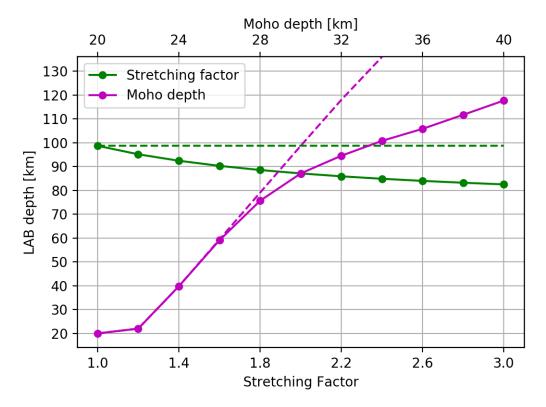


Figure 5. Stretching factor and Moho depth vs. LAB depth. The solid green line indicates  $z_{LAB,rifted}$  for varying stretching factors  $\beta$  between 1-3, while the solid magenta line shows  $z_{LAB,rifted}$  for variable initial crustal thickness between 20-40 km. The dashed lines represent LAB depths of the initial isostatic balance.

#### 316 **4 Results**

#### 317 4.1 LAB for the South American passive margin

Figure 6 shows  $z_{LAB,rifted}$  for the South American passive margin. The South American passive margin can be split into an equatorial and a southern part, separated by the Chain Fracture Zone.  $z_{LAB,rifted}$  varies from values locally lower than 50 km to deeper than 200 km. Remarkably, there are also large gradients of  $z_{LAB,rifted}$  between the inner and outer COB. The magnitude depends on the distance between inner and outer COB. For example, in the Pernambuco-Parnaiba Basin  $z_{LAB,rifted}$  varies ~100 km between inner and outer COB as the inner COB extends deeply into the Borborema Province.

The sedimentary basins south of the Chain Fracture Zone have been formed during early opening 325 of the South Atlantic. In this area,  $z_{IAB,rifted}$  is characterized by a rather heterogeneous structure 326 327 with values mostly lower than 100 km. Locally, there are large horizontal gradients, causing strong deviations of the lithospheric thickness even inside a single sedimentary basin. Very 328 329 shallow lithosphere can be observed in the Campos, Santos, Punta del Este, and Colorado Basins with values lower than 50 km, indicating strong deformation due to high stretching factors. The 330 shallow lithosphere of the Campos and Santos Basins is intersected by a thick lithospheric piece 331 in the southern Santos and northern Pelotas Basin. Here, z<sub>LAB,rifted</sub> reaches values up to 160 km 332 333 even towards the outer COB.

In contradiction to the southern part,  $z_{LAB,rifted}$  reveals a rather smooth structure between 100-150 km in the equatorial part of the South American passive margin. Even though the deforming region of the equatorial segment is narrower than for the southern segment, for most basins a high gradient of lithospheric thickness between inner and outer COB is present with thinning towards the outer COB.

Most features of  $z_{LAB,rifted}$  are already present in the isostatically derived  $z_{LAB,init}$  (Figure 7a).

However, in the equatorial part, the gradient of lithospheric thickness between the inner and

outer COB is higher than for  $z_{LAB,rifted}$ . The difference between  $z_{LAB,init}$  and  $z_{LAB,rifted}$ ,  $\Delta z_{LAB}$ , shows

that in this area rifting causes more variety than for other parts of the margin (Figure 7b).  $\Delta z_{LAB}$ has a maximum of 12 km, with the highest values in the Borborema Province, which is the onshore continuation of the Pernambuco-Parnaiba and Sergipe-Alagoas Basins, and in the

Pelotas and Santos Basins offshore of Brazil. For other regions  $\Delta z_{LAB}$  is mostly lower than 7 km.

A cumulative plot of  $z_{LAB,init}$  versus  $\Delta z_{LAB}$  follows a Gaussian distribution with a standard deviation of 22.5 km (Figure 8).  $\Delta z_{LAB}$  reflects the behavior of the geotherms as shown in Figure

deviation of 22.5 km (Figure 8).  $\Delta z_{LAB}$  reflects the behavior of the geotherms as shown in Figure 2. The isostatic equation (Eq. 1) shows that thin lithosphere correlates with thin crust. If both

layers are thin,  $\Delta z_{IAB}$  has no time to propagate and is negligible for lower values of  $z_{IAB,init}$ . Thick

crust correlates with thick lithosphere. In this case, the geotherms deviate from each other in

deeper levels which cannot be compensated by the mantle lithosphere. Consequently,  $\Delta z_{IAB}$  gets

small for higher values of  $z_{LAB,init}$  (Figure 8). Selecting  $\Delta z_{LAB} = 5$  km as a benchmark shows that

only a range of 110-170 km for  $z_{LAB,init}$  cause deviations of  $z_{LAB,rifted}$  during the rifting process.

The maximum of the curve corresponds to  $z_{LAB,init}$ =135 km and  $\Delta z_{LAB}$ =10 km.

Figure 9a displays the mantle density of the isostatic column prior to rifting. Notably,  $\rho_m$  is 3.35

 $g/cm^3$  for most areas. The difference to  $\rho_a$  (3.3 g/cm<sup>3</sup>) is 0.05 g/cm<sup>3</sup> and reflects the temperature-

dependent decrease of density with depth from lithosphere to asthenosphere. The outermost

Campos and Santos Basins show anomalous minima of  $\rho_m$  around 3.31-3.32 g/cm<sup>3</sup>. These structures partly require a higher number of iterations that are needed to satisfy the density

360 threshold criterium (Figure 9b).

Most of the points in Figure 9b reach the density threshold after  $i \le 4$  iterations. Only where lithospheric thickness is very shallow, the algorithm needs more time to converge. This shows an inherent stability between mantle density and thickness of the mantle lithosphere, which are the

variable parameters of our inverse approach.

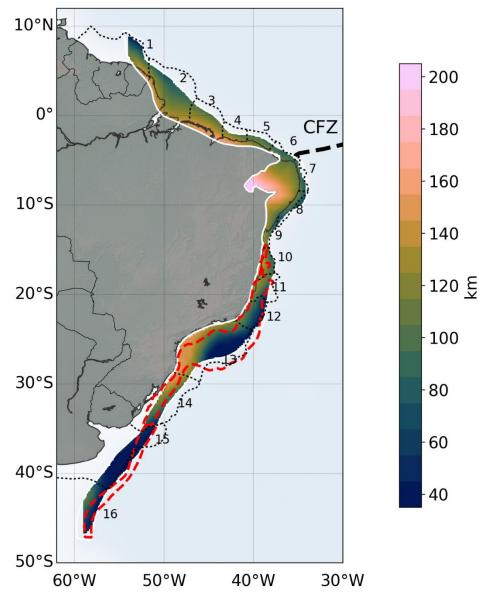
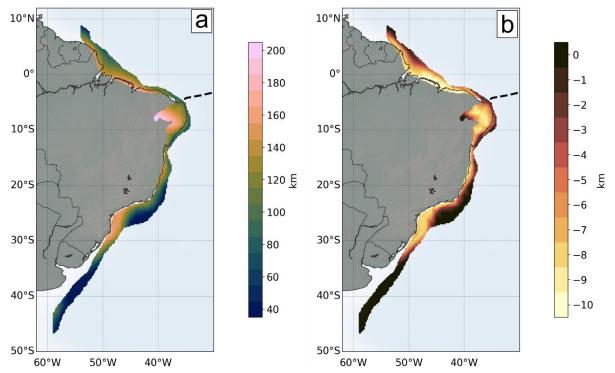


Figure 6. LAB after rifting at the passive margin of South America. Thick dashed black line
indicates Chain Fracture Zone (CFZ). Thick dashed red polygon marks the extension of Large
Igneous Provinces (LIPs) offshore South America, taken from the Johansson et al. (2018) data
base. Dashed contours and numbers indicate offshore sedimentary basin locations, taken from
Wen et al. (2019). 1: Guyana, 2: Foz do Amazonas-Marajo, 3: Para-Maranhao, 4: Barreirinhas,
5: Ceara, 6: Potiguar, 7: Pernambuco-Parnaiba, 8: Sergipe-Alagoas, 9: Bahia Norte, 10: Bahia
Sul, 11: Espirito Santo, 12: Campos, 13: Santos, 14: Pelotas, 15: Punta del Este, 16: Colorado.



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**Figure 7. a**: Initial isostatic LAB before rifting, **b**: Difference of LAB before and after rifting. Dashed line indicated Chain Fracture Zone.

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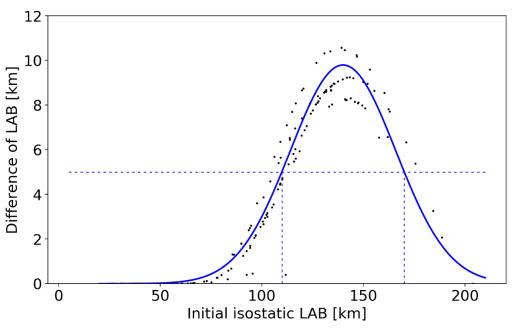




Figure 8. Distribution of Initial Isostatic LAB vs. Difference of LAB before and after rifting. The horizontal dashed line indicates the difference benchmark of  $\Delta z_{LAB} = 5$  km.

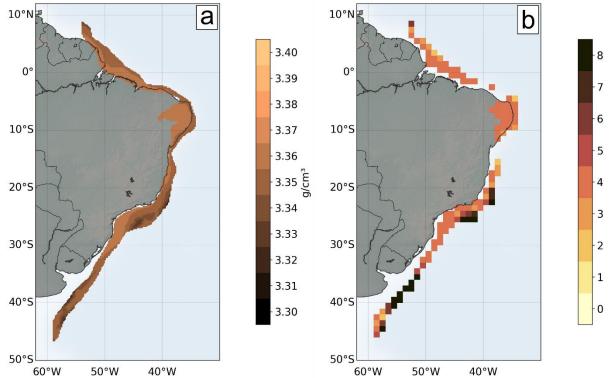


Figure 9. a: Mantle density before rifting, b: Number of iterations needed to satisfy density fit.
Note that the resolution is kept to 1 degree to show the convergence at each point of the grid.
4.2 LAB of the conjugate South Atlantic passive margins in a Western Gondwana

4.2 LAB of the conjugate South Atlantic passive margins in a Western Gondwana framework

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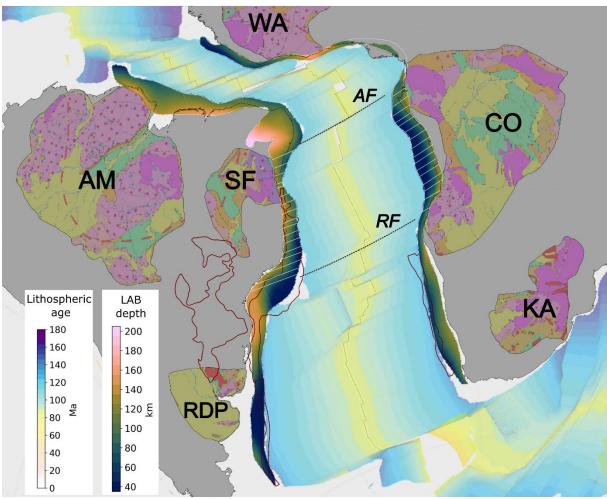
We used the same approach to calculate  $z_{LAB,rifted}$  for the African passive margin (see supporting information S2 for the crustal model, segments of crustal thickness, and stretching factors). Figure 10 shows  $z_{LAB,rifted}$  in a Gondwana framework, rotated to 83 Ma while the entire South Atlantic had been opened. The conjugate margins between Ascension and Rio de Janeiro Fracture Zone are connected by flowlines, which have been calculated by a set of seed points located on the Mid Atlantic Ridge. The flowlines are used to calculate the width of the passive margins.

In many parts of the African passive margin, the width is less than on the conjugate South American side, especially in the African equatorial segment. However, the general trend of deeper lithosphere in the equatorial segment and shallower lithosphere in the southern segment is also observed at the African passive margin. A notable difference appears at the northernmost tip of the African deforming region, where the lithospheric thickness is shallower than 50 km.

For the central part of the African passive margin, the lithospheric thickness is less variable than for South America. The thickness is mostly lower than 100 km with lowest values offshore Congo and Kwanza Basin. In this area, the passive margin appears to be wider than for the South American counterpart, suggesting more constant lithospheric thinning. The across-margin gradients are not as pronounced as for the South American passive margin since  $z_{LAB,rifted}$  is already quite shallow towards the inner COB.

In the southern part, where the African passive margin is rich in volcanic material, the lithospheric thickness shows intermediate values ~100 km towards the inner COB. The across410 margin gradient is relatively high with values lower than 50 km towards the outer COB.
411 Compared to the South American counterpart, the African side shows less along-margin
412 heterogeneity.

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Figure 10. Lithospheric thickness for the deforming region of the South American and African 415 passive margins in Gondwana framework, rotated back to 83 Ma using GPlates software (Müller 416 et al., 2018). In between, the age of the oceanic lithosphere is visualized (Seton et al., 2020). 417 Thin dashed line represents the location of the Mid Atlantic Ridge at 83 Ma. AF=Ascension 418 Fracture Zone, RF=Rio de Janeiro Fracture Zone. Grey lines represent flowlines intersecting the 419 passive margins and are used to calculate margin width (see Fig. 12-14). Cratons on the 420 continent are draped with geology. Craton boundaries are taken from Celli et al. (2020). Cratonic 421 422 units are abbreviated: AM=Amazonia, SF=Sao Francisco, RDP=Rio de la Plata, WA=West Africa, CO=Congo, KA=Kalahari. Dark red polygons show location of LIPs, taken from the 423 Johansson et al. (2018) data base. 424

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#### 426 **5 Discussion**

- 428 5.1 The role of magmatic underplating for the LAB structure along the South Atlantic 429 passive margins
- 430

The two-part lithospheric structure with deeper lithosphere in the equatorial part and shallower, 431 432 more heterogeneous lithosphere in the southern part along the South American passive margin offers insights in the rifting mechanism that controlled the break-up of Pangea. The smooth 433 lithosphere at the equatorial margins, especially on the South American side, implies that far-434 field edge forces are the dominant mechanism, causing only minor thinning of the pre-break-up 435 continental lithosphere. The lithospheric thickness is in a range that can be expected for stable 436 continental platforms (e.g., Artemieva, 2012). A lower amount of lithospheric thinning indicates 437 that less magmatic material has been involved in the rifting process. This is the case for the 438 equatorial passive margin of South America, where LIPs are absent. 439

At the southern part of the passive margins the initial lithosphere has been thinned more extensively. This suggests a longer initial slow pre-rift phase with sufficient time to destruct the initial continental lithosphere. In some areas, like the Colorado Basin on the South American side, the LAB is even shallower than 50 km. Such a shallow LAB is rather unlikely and is a result of the assumed simplifications governing the initial isostatic equation.

The heterogeneous lithospheric structure of the southern South American margin points to a more variable rifting process, which is strongly connected to underplating of magmatic material. On the one hand, magmatic underplating thickens the crust by adding partial melts from a stable magma chamber to the base of the crust and cooling over a longer time period (e.g., Cox, 1993; Thybo and Artemieva, 2013). On the other hand, crustal thickening due to magmatic underplating might not necessarily be an indicator for lithospheric thickening or thinning.

The occurrence of LIPs (see polygons in Figures 6 and 10) shows that the bulk of the southern 451 margin has been affected by volcanic underplating. In the Colorado Basin, volcanic underplating 452 is associated with SDRs. McDermott et al. (2018) distinguish the SDRs in this area in two 453 different types, representing the continuum from continental rifting to full plate separation with 454 formation of new magmatic crust. While the first type was formed during stretching of the crust, 455 the second type was formed as narrow lava flows due to the Tristan plume activity (McDermott 456 et al., 2018). Even though the passive margin does not capture the entire sequence of SDRs, the 457 modeled lithospheric thickness is characterized by very low values in this area. A transition of 458 two different SDR types cannot be observed in terms of lithospheric thickness. 459

According to McDermott et al. (2018) SDRs smoothly transition from combined passive continental rifting/plume activity to wider lavas, representing magmatic activity only. This transition zone is located offshore Uruguay/Brazil in the Pelotas Basin and correlates with a distinct transition of lithospheric thickness over a relatively short distance. While increased plume activity induces reinforced lithospheric thinning, our model shows the opposite behavior as the LAB increases from ~40 km to ~100 km under the Pelotas Basin (Figure 6).

Morgan et al. (2020) showed that the formation of the volcanic rifted margins in the South 466 Atlantic is a result of an asymmetric lateral drainage of the Tristan Plume. This asymmetric flow 467 is triggered by along-strike variations in the geometry and opening of the rift, as well as lateral 468 variations in the initial continental lithosphere. In the initial model setup, cratons are assumed to 469 have thicker lithosphere than the surrounding regions. Using 3D numerical modeling, this 470 structure is roughly preserved throughout the first 28 Ma of rifting (see Figure 3 and S4 in 471 Morgan et al., 2020). The shallower lithosphere in the Morgan et al. (2020) model pulls the 472 plume material southwards and coincides with the deeper lithosphere in the Santos and Pelotas 473

474 Basins, as featured in our model.

Both the SDR distribution as observed by McDermott et al. (2018) and the model of lateral plume drainage of Morgan et al. (2020) do not match our observation of deep lithosphere in the

Santos and Pelotas Basins. This raises two general questions regarding our modeled LAB: 1. Can 477 small-scale patterns like different types of SDR be identified and distinguished? 2. Can larger-478 scale patterns that arise from plume-rift interaction be identified and distinguished? 479

SDRs are an indicator of increased magmatic activity. They are emplaced at a narrow time 480 window in the synrift phase, whereas the underlying lithosphere can be modified over a much 481 longer time frame afterwards. The distribution of SDRs points to different melting mechanisms 482 in the lithosphere. Even though SDRs represent different episodes of lithospheric thinning, their 483 size is too small to be recovered by our lithospheric model. 484

The lithospheric structure of the Morgan et al. (2020) model at the Santos and Pelotas Basins is 485 controlled by two parameters: location of the starting plume head and definition of the initial 486 lithospheric thickness. Morgan et al. (2020) showed that different plume locations and plume 487 fluxes do not impede southward migration of plume material (see their supplementary material). 488 The shallow lithosphere in their model, located in the Santos and Pelotas Basin, is a combination 489 of both parameters, but always based on a certain assumed geometry. Varying the initial 490 geometry would cause a different pattern of the initial lithosphere. 491

Qualitatively, our LAB model distinguishes patterns of both SDRs and location of volcanic 492 material during rifting. Quantitatively, we cannot correlate our LAB model with different SDR 493 types and the asymmetric plume flow, as observed by Morgan et al. (2020). Thus, the amount of 494 volcanic material emplaced during rifting and plume activity remains enigmatic. 495

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5.2 Comparison of stretching factors

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The stretching factors of Müller et al. (2019) serve as a basis to define a global deforming plate 499 motion model. The total stretching factor is defined by stacking all stretching factors over time 500 for a certain point or area. This is beneficial when multiple rifting or collisional events occurred. 501 As the stretching factors of Müller et al. (2019) are valid for plate motions since the Triassic, the 502 opening of the South Atlantic is the only event contributing to the stretching factors. Müller et al. 503 (2019) showed that globally most extensions range in stretching factors  $\beta = 1-2$ . Stretching factors 504 larger than 2 reflect highly extended areas. 505

In Figure 11a the total stretching factors  $\beta$  for the South American passive margin are plotted. 506 The stretching factor  $\beta$  varies between 1-3, indicating extensional tectonics only. Both the 507 equatorial and southern passive margins comprise a heterogeneous distribution. The Amazonas-508 Marajo Basin in the north and Bahia Basin, Espirito Santo Basin and Campos Basin in the south 509 show highest values of stretching larger than 3. 510

Rescaling the stretching factors of our approach (Figure 3c) to the same amplitude as in Figure 511 11a reveals a considerably lower amount of stretching (Figure 11b). The pattern of stretching 512 with the two-part structure of lower stretching in the equatorial segment and higher stretching in 513 the southern segment is less pronounced. However, the highest stretching factors in the Campos 514

515 and Santos Basins ( $\beta = 2$ ) correlate with the distinct area of high stretching in Figure 11a ( $\beta > = 3$ ).

Other similarities are the lower stretching factors in the Pernambuco-Parnaiba Basin and adjacent 516

Borborema Province, as well as the varying pattern in the Punta del Este and Colorado Basin 517

518 further south.

519 The high stretching factors of the Müller et al. (2019) model imply that crustal thickness of

unthinned crust is more than three times larger than thinned crust. Our model does not reflect 520

521 such high values due to averaging of the continental crust into certain segments (Figure 3a and

3b). In Müller et al. (2019), crustal thickness prior to extension is defined by seismological 522

523 measurements along a given zone of extension, causing higher variability of the initial crustal 524 thickness and consequently locally higher stretching factors.

The underlying plate deformation in the Müller et al. (2019) plate motion model is approximated 525 by a pure-shear, uniform extension model. As this is the only way to define a kinematic plate 526 model without capturing ductile flow, rift-internal variations of stretching, representing strain 527 localization or depth-dependent stretching, are not included. As a consequence, the stretching 528 factors represent "wide rifts that lack margin-orthogonal strain rate and crustal thickness 529 gradients" (Müller et al., 2019). In Figure 11a this becomes obvious as the pattern of stretching 530 varies only along strike. The stretching factors of our approach are derived independent of a plate 531 motion model by the crustal thickness model only. They comprise heterogeneities of the crustal 532 model as margin-perpendicular and margin-alongside gradients. 533

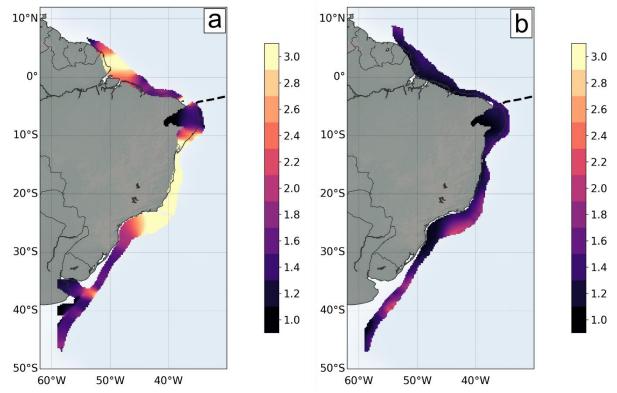
In Müller et al. (2019), the thinned crust is defined by integrating a surface dilatation rate over time. This may be a source of error because the dilatation rate is not exactly known at each time step. We approximate the recent crustal thickness in the deformable region as thinned crust. This method might be another source of error because not each crustal unit of the deformable region

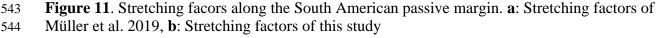
represents thinned crust due to rifting only but might be thickened by magmatic underplating as

well. Less underplating would cause lower crustal thickness in the passive margin and higher

540 stretching factors.

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5.3 Relation of lithospheric thickness and passive margin width

The South Atlantic passive margins are strongly related in age. While classical 2D-rifting models like McKenzie (1978) suggest uniform extension of both margin sides, recent studies showed that initial rift asymmetry leads to different widths of conjugate margins (Brune et al., 2014).

551 This is also reflected in the deforming regions of the Müller et al. (2019) plate motion model.

552 Comparing the lithospheric thickness of both margin sides can be an indicator of how much 553 lithospheric thinning has contributed to rift asymmetry.

554 We used the conjugate Bahia/Espirito Santo/Campos Basins on the South American side and the

555 (offshore) Congo/Kwanza Basins on the African side to study the lithospheric thickness of both

- 556 margins jointly because the asymmetry of conjugate margin width is particularly pronounced in 557 this area. Using flowlines, the difference is determined by subtracting the lithospheric 558 thicknesses at the outer COB of conjugate margin pairs. The width of a passive margin is 559 calculated by shifting flowlines, propagating from the Mid Atlantic Ridge, towards the inner 560 COB of the respective passive margin (see Figure 10).
- The conjugate margins are characterized by large differences in the lithospheric thickness 561 (Figure 12). In the northern Bahia and Congo Basins the difference is moderate with 20 km 562 deeper lithosphere in the Congo Basin. For the southern Bahia/Congo Basins this changes 563 abruptly. Here, the LAB is 80 km deeper in the South American Bahia Basin, coincident with 564 distinct narrowing of the margin. This pattern proceeds throughout the Espirito Santo/Lower 565 Congo Basins with 50-60 km thicker lithosphere at the South American margin. For these 566 particular conjugate basins, a considerable amount of magmatic underplating has been mapped 567 on the South American side (Johansson et al., 2018), whereas these structures are absent on the 568 African side. Towards the Campos/Kwanza Basins the lithospheric thickness pattern gradually 569 changes to 30 km thinner lithosphere at the South American margin. This part of the Campos 570 Basin is characterized by a very wide margin, while the Kwanza Basin narrows towards the Rio 571

572 de Janeiro Fracture Zone.

Plotting the differences of LAB depth and margin width against each other shows a linear 573 correlation of both parameters (Figure 13). Locally, the South American margin is up to 600 km 574 wider than the African conjugate. That is because the wide extension of the Campos Basin is 575 amplified by an oblique distribution of the flowlines. In this area, the South American 576 lithosphere is up to 40 km thinner than the African lithosphere. Following the linear trend shows 577 that up to 20 km shallower South American LAB occurs for margin width differences between 578  $\pm 100$  km, independent of the margin side. For the narrower South American margin, the linear 579 trend of increasing lithospheric thickness difference is more distinctive than for the wider 580 margin. A maximum difference of -200 km in margin width correlates with 80 km deeper LAB 581 for the South American margin. 582

The highest LAB differences are found in the Bahia Basin, where the passive margin is locally 583 narrower than 100 km. In this area, the across-margin gradients of lithospheric thickness are less 584 pronounced, causing a relatively deep LAB at the outer COB (Figure 6 and Figure 10). Possible 585 586 magmatic underplating could have strengthened this effect. The African conjugate shows a very shallow LAB, mainly caused by shallow crust and high stretching factors in the deforming 587 region. Besides that, no magmatic underplating has been mapped. Even for the narrow margin in 588 the southern Kwanza Basin the LAB is rather shallow, causing only moderate differences in 589 lithospheric thickness for the conjugate margin pairs. 590

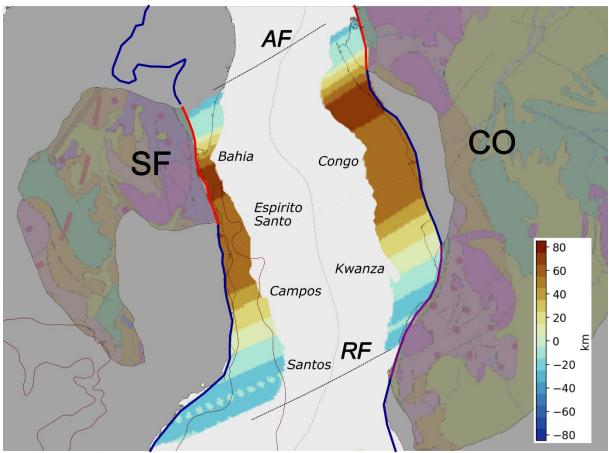
A contribution of magmatic underplating to lithospheric thinning cannot be defined by our models. But they show that magmatic underplating might not necessarily be accompanied by extensive melting at the lithospheric base. Our results suggest that the asymmetry of the rifts rather causes differences in lithospheric thickness of both margins. However, this is only valid

for a scenario where lithospheric thickness is initially constant for both margin sides. Our initial

isostatic LAB could be regarded as pre-Gondwana-breakup, but its amplitude is too close to our

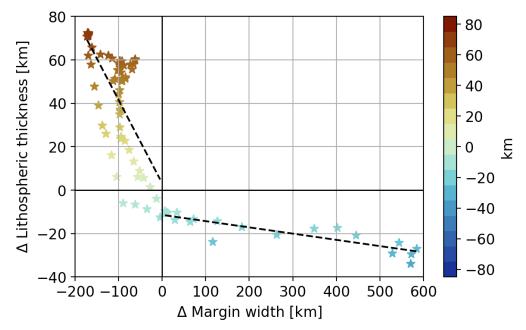
final model. However, establishing a LAB model in a Gondwana framework is beyond the scopeof this paper.

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**Figure 12.** Difference of lithospheric thickness for selected conjugate South American and African basins, rotated back to 105 Ma. Red colours indicate deeper LAB in South America, blue colours deeper LAB in Africa. Thick lines at the inner COBs indicate directly adjacent cratons (Red: Sao Francisco and northern Congo Craton, purple: Southern Congo Craton, Blue: No craton adjacent). Red polygons indicate LIPs, for abbreviations see Figure 10.



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**Figure 13.** Difference of margin width of conjugate margin pairs versus difference of lithospheric thickness at the respective points at the outer COB. Red colours indicate deeper LAB in South America, blue colours deeper LAB in Africa. Dashed lines show the linear fit for the pairs with wider African margin (negative  $\Delta$  Margin width) and for the pairs with wider South American margin (positive  $\Delta$  Margin width). Colour map is same like in Figure 12.

5.4 The role of cratons and Parana Flood Basalts for lithospheric thickness in deforming
 regions

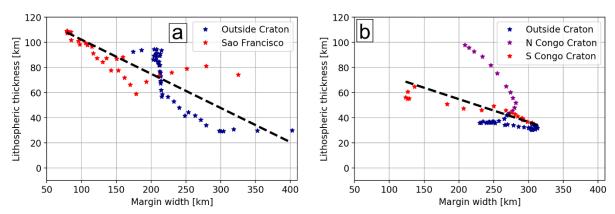
Distinguishing the individual lithospheric thickness and margin width for the margin segments per continent shows a linear trend that is surrounded by several outliers (Figure 14). To investigate how much the cratons contribute to these deviations, we marked each point by colors. For the South American margin, thicker lithosphere occurs in the vicinity of the Sao Francisco

619 Craton, which is also characterized by narrow margin widths (Figure 14a).

For the African side, we distinguish between the southern and northern Congo Craton. Together 620 621 with the area where the southern Congo Craton reaches the coastline the points outside the craton boundary form a linear trend. Stronger deviations belong to those points, where the northern 622 Congo Craton intersects the deforming region. Overlapping geometries of the craton and 623 deforming region can introduce errors in the modeled lithospheric thickness. In this case, 624 unthinned crust contains some portions of cratonic lithosphere. Consequently, stretching is 625 underestimated, which causes subsequent overestimation of the lithospheric thickness. This trend 626 can clearly be reproduced in the distribution of the outliers (Figure 14b). 627

The profiles of margin width versus lithospheric thickness for both continents evince that the location of the cratons controls the LAB depth. Deeper values of LAB are preferably observed in areas, where the inner COB reaches the craton. In the south, the Parana Flood Basalts are located between the Sao Francisco and Rio de la Plata Cratons. In this area, the deep LAB in the adjacent Santos and Pelotas Basins reflects high stretching factors and a strong contrast between unthinned and thinned crust (Figure 6). Our interpretation is that the relatively short, but intense magmatic activity, forming the Parana Flood Basalts (Thiede and Vasconcelos, 2010), caused a 635 crustal thickening in the area of the Parana Basin. This can also be seen in Figure 3b, where the 636 average crustal thickness of the Parana Basin is similar to the cratons. However, in the deforming 637 region, the volcanic activity occurred in a later stage and lead to rifting with crustal and 638 lithospheric thinning.

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Figure 14. Margin width of conjugate margin pairs versus lithospheric thickness for the
 individual passive margins. Linear trend is visualized as black dashed line. a: South America, b:
 Africa

#### 644 6 Conclusions

We have derived a new thermal lithospheric thickness model for the South Atlantic passive margins. Our model is calculated as a function of rifting time, crustal thickness, and stretching factor. The stretching factors are obtained by the amount of unthinned crust divided by thinned crust, using state-of-the-art crustal models for the South American and African continent. The new stretching factors account for across rift crustal gradients at the passive margin and are a refinement compared to the Müller et al. (2019) model.

From our model, the lithospheric structure of the conjugate passive margins can be distinguished 651 in two parts: the equatorial part with deep lithosphere along the relatively narrow margins and 652 the southern part with rather thin lithosphere. For the South American passive margin, the 653 lithospheric thickness reveals a very heterogeneous structure, which can be related to different 654 rifting mechanisms that lead to the opening of the South Atlantic. Magmatic underplating and 655 SDRs indicate more complex tectonics with a large amount of volcanic material, which is 656 reflected in the heterogeneous LAB depth. The subsequent equatorial opening of the South 657 Atlantic, driven by far-field forces, is evident in a minor destruction of the initial continental 658 lithosphere. For Africa, the lithospheric thickness is rather shallow along the entire margin. 659

Analyzing the lithospheric thickness in a Western Gondwana framework evinces its large variability in conjugate basins of the South American and African passive margins. The differences in lithospheric thickness are up to 80 km. We propose that these differences reflect highly asymmetric melting and lithospheric thinning prior to rifting. This is also in accordance with the margin widths. Wider margins have higher across gradients and a shallower LAB at the outer COB compared to conjugate basins of the other margin. This asymmetry is manifested in the different distribution of margin width and lithospheric thickness for each continental margin.

667 Our modeling approach does not capture the amount of melting and volcanic underplating. This 668 is reflected in very shallow LAB regions in the southern passive margins. The mechanism of 669 underplating control on lithospheric thickness cannot be resolved by our model. Furthermore, the

- 670 lithospheric thickness strongly depends on the initial crustal model and the governing isostatic 671 equation. Despite these simplifications, our LAB model represents the thermal state of the rifted 672 margins at present day, which generally agrees with other studies
- margins at present day, which generally agrees with other studies.

Future efforts should focus on including melting and magmatic underplating, once a comprehensive data set outlining underplated crustal thickness on both conjugate margins is available. This could be easily included in the governing isostatic equation. The modeling of the thermal structure can be extended to 2D rifting scenarios instead of the 1D approach that we are using. Given the potential of improvements, we are confident that our approach opens a new pathway for more extensive analysis of the lithospheric structure of passive margins. Our procedure can be easily extended to other passive margins on the globe. Ultimately, this would

fill the gaps of reconstructed lithospheric models for the Gondwana Supercontinent.

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- Most of the plots have been created using Matplotlib with color-blind friendly color maps
- 691 provided (Crameri, 2018; Hunter, 2007).
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# 693 **Data**

- Models of crustal thickness and density for South America are published via the GFZ Data
- 695 Services: https://doi.org/10.5880/GFZ.1.3.2020.006 (Finger et al., 2021). For Africa, the data is
- in preparation. The rifting times of the South Atlantic opening are part of the deforming plate
- model of Müller et al. 2019, which is available at:
- 698 <u>https://www.earthbyte.org/webdav/ftp/Data\_Collections/Muller\_etal\_2019\_Tectonics/</u>
- A version of the code RiftSubsidence.py that generates the presented results will be made
- available for the public. This code will be distributed, probably with GitHub, within the final
- submission of the manuscript.
- 702

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