How frequent are Antarctic sudden stratospheric warmings in present and future climate?

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Abstract

Southern Hemisphere (SH) Stratospheric Sudden Warmings (SSWs) result in smaller Antarctic ozone holes and are linked to extreme midlatitude weather on subseasonal to seasonal timescales. Therefore, it is of interest how often such events occur and whether we should expect more events in the future. Here, we use a pair of novel multi-millennial simulations with a stratosphere-resolving coupled ocean-atmosphere climate model to show that the frequency of SSWs, such as observed 2002 and 2019, is about one in 22 years for 1990 conditions. In addition, we show that we should expect the frequency of SSWs - and that of more moderate vortex weakening events - to strongly decrease by the end of this century.

How frequent are Antarctic sudden stratospheric warmings in present and future climate?

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Key Points:

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10	•	Antarctic sudden stratospheric warmings occur once every 22 years in present-day
11		(1990) climate conditions.
12	•	The warmings will become much rarer under future climate change, irrespective
13		of their exact definition.
14	•	The future decrease in frequency is linked to a strengthening of the Antarctic po-
15		lar vortex.

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Abstract 16

Southern Hemisphere (SH) Stratospheric Sudden Warmings (SSWs) result in smaller Antarc-17 tic ozone holes and are linked to extreme midlatitude weather on subseasonal to seasonal 18 timescales. Therefore, it is of interest how often such events occur and whether we should 19 expect more events in the future. Here, we use a pair of novel multi-millennial simula-20 tions with a stratosphere-resolving coupled ocean-atmosphere climate model to show that 21 the frequency of SSWs, such as observed 2002 and 2019, is about one in 22 years for 1990 22 conditions. In addition, we show that we should expect the frequency of SSWs-and that 23 of more moderate vortex weakening events-to strongly decrease by the end of this cen-24 turv.

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Plain Language Summary 26

The stratosphere at 10-50 km height can influence surface weather for several months. 27 In 2002 and 2019, the stratosphere warmed over Antarctica within a few days to weeks. 28 This caused dry and hot summers in Australia and South America. And it reduced the 29 size of the ozone hole. Since these warming events are rare, it is difficult to say how of-30 ten they occur. We therefore use long computer simulations to answer that question. We 31 find that without climate change, warming events occur about every 22 years. But with 32 climate change, the warming events will happen only once every 300 years. From this, 33 we believe that the quick succession of two events in 2002 and 2019 will remain special 34 35 in history.

1 Introduction 36

37 The stratospheric polar vortex forms in the winter hemisphere due to the lack of solar heating at high latitudes and the resulting strong equator-to-pole temperature gra-38 dient. In the Northern Hemisphere (NH), strong and planetary scale waves originating 30 in the troposphere from orographic forcing and land-sea contrast periodically propagate 40 upward into the stratosphere and perturb the polar vortex via momentum deposition when 41 the waves break (Eliassen & Palm, 1960; Charney & Drazin, 1961; Matsuno, 1971). In 42 extreme cases, this disruption of the polar vortex leads to a rapid warming and rever-43 sal of wind directions in the polar stratosphere, a so-called (major) Sudden Stratospheric 44 Warming (SSW) (Butler et al., 2015). These SSWs occur around every other winter in 45 the NH. 46

However, over the six decades that we have station records (and later satellite ob-47 servations) of the Southern Hemisphere (SH) polar vortex, only one such wind reversal 48 has been recorded in 2002 (Roscoe et al., 2005; Esler et al., 2006). This event substan-49 tially decreased the size of the ozone hole thanks to higher than usual stratospheric po-50 lar temperatures and transport of ozone-rich air from lower latitudes into the polar re-51 gions (Fig. S2a) (Stolarski et al., 2005). There was also a dynamical effect of the 2002 52 SSW at the surface, as an extreme negative polarity of the Southern Annular Mode (SAM) 53 was recorded at the surface for the 10-90 day period following the event (Thompson et 54 al., 2005). Even though no wind reversal at 60° S and 10 hPa was registered in 2019, the 55 polar vortex in this more recent event weakened dramatically and also lead to a smaller 56 ozone hole (Fig. S2b) with almost 30% higher total column ozone values compared to 57 the previous decade (Safieddine et al., 2020). The event has also been linked to the se-58 vere bushfire season in South Eastern Australia the following spring and summer (Lim 59 et al., 2021). 60

Due to the impacts on stratospheric ozone and surface weather on the subseasonal 61 to seasonal timescale, it is important to determine how rare SSWs are in the SH, and 62 whether we should expect more or less frequent SSWs under future climate change. How-63 ever, given the shortness of the observational record it is impossible to get an observa-64

tional estimate of how often SSWs do occur on average. Recently, Wang et al. (2020) 65 analyzed hindcasts of a seasonal forecasting system and found an average Antarctic SSW 66 frequency of one every 25 years. However, the underlying model of this study had a strong 67 mean westerly wind bias, raising some doubts on the validity of their results. Here, we 68 revisit the question of how frequent Antarctic SSWs are in present climate and also ad-69 dress possible changes under future climate change. This is accomplished by investigat-70 ing two nearly 10,000-year long simulations with a well-performing stratosphere-resolving 71 coupled ocean-atmosphere model based on present-day (1990) and future (increased CO2) 72 conditions and by considering integrations from the sixth Climate Model Intercompar-73

⁷⁴ ison Project (CMIP6).

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⁷⁵ 2 Model data and SSW definitions

2.1 Multi-millennial coupled GCM simulations

We use a set of two 9,900-year long simulations with the stratosphere-resolving ver-77 sion of the Geophysical Fluid Dynamics Laboratory's CM2.1 atmosphere-ocean cou-78 pled climate model (Delworth et al., 2006; Horan & Reichler, 2017), which has been used 79 in particular for studies of stratosphere-troposphere coupling in the past (Horan & Re-80 ichler, 2017; Jucker & Reichler, 2018). The model has 48 vertical levels with approxi-81 mately half of the levels situated in the stratosphere and a model top at 0.002 hPa. The 82 horizontal resolution is approximately 2° in latitude and 2.5° in longitude. The bound-83 ary conditions are set to perpetual 1990 conditions. More specifically, ozone in the year 84 1990 is comparable to both 2002 and the 2010s (Newman & Nash, 2019). The two sim-85 ulations differ in their greenhouse gas forcing; CO₂ is set to 353 ppm in the 'present-day' 86 and 1120 ppm in the 'future' simulation, which is a quadrupling relative to pre-industrial 87 CO_2 concentration (and 3.2 times present-day concentration). This is the only difference 88 between the two simulations. Atmospheric variables are stored on a daily frequency to 89 allow for detailed dynamical analysis, including Eliassen-Palm fluxes. 90

In agreement with Horan and Reichler (2017), who have shown that this model com-91 pares well to reanalysis in the troposphere and northern hemisphere stratosphere, both 92 the southern hemisphere stratospheric zonal mean zonal wind and vertical component 93 of the Eliassen-Palm flux from our present-day simulation show excellent agreement with 94 those from ERA5 reanalysis (1979-2019) (Hersbach et al., 2020), for both mean and stan-95 dard deviation (Figs. 1a,c and S1). We also note that the model intercomparison work 96 by Reichler and Kim (2008) showed that CM2.1 had the best performance index among 97 CMIP3 models, even though that version had only half the number of vertical levels com-98 pared to the version used here. Besides its performance in the atmosphere, which is of 99 particular relevance here, the oceanic component has been validated extensively and also 100 found to have a good representation of tropical (including ENSO, Wittenberg et al., 2006) 101 as well as extratropical southern hemisphere ocean dynamics (Gnanadesikan et al., 2006). 102

Having multi-millennial simulations with a model showing such small bias will allow us to robustly estimate SSW frequencies. In addition, having future projections will make it possible to address the question of whether or not we should expect another SSW to occur in the future, and we will show that increased greenhouse gas concentrations have a strong impact on SSW frequency.

2.2 SSW definitions

We follow the most common definition of Sudden Stratospheric Warmings as the reversal of u₁₀₆₀, the zonal mean zonal wind at 60°S and 10 hPa ('SSW-reversal', Charlton & Polvani, 2007). However in observations, only the September 2002 event is an SSWreversal event, while the 2019 event is widely considered an SSW but did not show wind reversal at 60°S and 10 hPa. Therefore, we have performed our analysis with an addi-

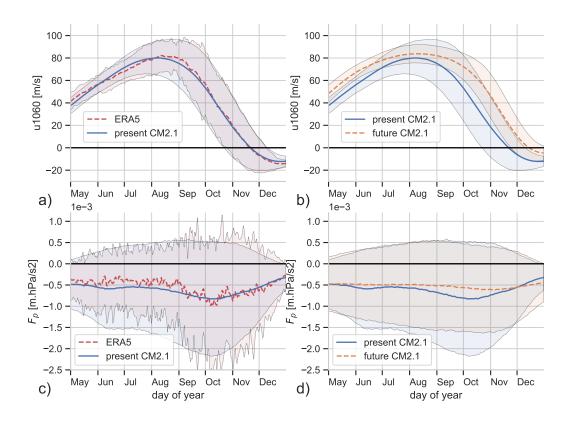


Figure 1. (top) Climatological mean (solid) and two interannual standard deviations (shaded) of zonal mean zonal wind at 60° S and $10 \text{ hPa} (u_{1060})$ for (a) present-day CM2.1 and ERA5 and (b) present-day and future CM2.1. (bottom) same but for vertical EP flux. The present-day simulation (blue, solid) reproduces both mean and variability of the ERA5 reanalysis (1979-2019; red, dashed) in both u_{1060} (a) and vertical EP flux (c). The future simulation (orange, dashed) shows a clear strengthening of the polar vortex throughout the year (b) and a weakening of the vertical EP flux (d), in particular during the spring.

tional definition, allowing for a more general determination of SSW frequency and fu-ture change.

We found that the simplest method to define SSWs in the SH which detects both 2002 and 2019 as the only events during the satellite era is that the zonal mean zonal wind anomaly with respect to the day of the year at 60°S and 10 hPa passes below -40 m/s. The onset date is then defined as the day when the zonal mean zonal wind anomaly crosses -20 m/s for the last time before crossing -40 m/s. These 'SSW-weak' events follow the common features of stratosphere-troposphere coupling in the SH in their significant surface impact on monthly timescales (Fig. S3).

For both definitions, two events have to be separated by at least 20 days, and the onset date has to be at least 20 days before the vortex breakdown, which is defined as the last day of the year when u_{1060} becomes negative.

Finally, we follow Lim et al. (2018) who showed that weaker events can also have 126 an impact at the surface, and we will also report results from their detection method based 127 on the yearly timeseries of the first Principal Component of de-seasonalized monthly mean 128 zonal mean zonal wind between 55 and 65°S. The corresponding Empirical Orthogonal 129 Function is two-dimensional but in month of the year-pressure space (instead of the con-130 ventional longitude-latitude space) and is centered around the vortex breakdown in spring 131 (the 'L18' method). This method does not provide onset dates, as there is only one value 132 per year, and L18 is closely related to variations in the date of the vortex breakdown (pos-133 itive for earlier breakdown; the correlation coefficient between the first Principal Com-134 ponent and the vortex breakdown date is r = 0.79 in ERA5 data, not shown). Follow-135 ing Lim et al. (2019), we apply a threshold of 0.8 standard deviations, which detects many 136 more events than the other two definitions. 137

¹³⁸ 3 Occurrence of SSWs in the Southern Hemisphere

The present-day 9,900-year simulation produces 458 SSW-weak and 159 SSW-reversal 139 events, corresponding to an average frequency of about one SSW-weak every 22 years and 140 one SSW-reversal every 59 years. This compares well with the single SSW-reversal and 141 only two SSW-weak events in the 42-year long satellite observation record (and the 63-142 year long non-satellite observational record since 1957 (Roscoe et al., 2005; Naujokat & 143 Roscoe, 2005)), as well as Wang et al. (2020). In addition to yearly occurrence, we also 144 analyze the seasonal occurrence of SSWs and find that the SSW-weak criterion detects 145 events during the entire winter, with a peak occurrence in late August to September (Fig. 2d) 146 and a mean occurrence of 27 August (note that early events in June and July have a sim-147 ilar impact to later events, not shown). The 2002 SSW occurred in late September, a 148 time of the year when we estimate the mean return time of SSW-weak events to be 113 years, 149 and the 2019 SSW occurred in early September, when the mean return time is estimated 150 to be 102 years (Fig. 2a). Irrespective of time of the year, our present-day simulations 151 indicate that we should expect between 0 and 6 SSW-reversals and between 0 and 12 152 SSW-weak events per century, with most likely numbers of 0-2 SSW-reversal and 3-6 SSW-153 weak events per century (25th and 75th percentiles, Figs. 2b,e). As indicated before, L18 154 events are much more abundant, with an occurrence of 7-36 events per century and a 155 mean return time of one in 5 years (Fig. 2h). 156

To get an estimate of when the next SSW might occur, we perform a return time analysis where we produce a histogram of the number of SSWs which occur within a given time interval (Fig. 2c,f,i). If SSWs are independent and random events, we can compare the observed return time distribution to a theoretical distribution (Text S6). The return time histogram follows closely the theoretical distribution for all methods, suggesting that in the SH, SSWs are independent and random, with a mean return time of about 59 years for SSW-reversal and 22 years for SSW-weak, or an annual probability of occurrence of

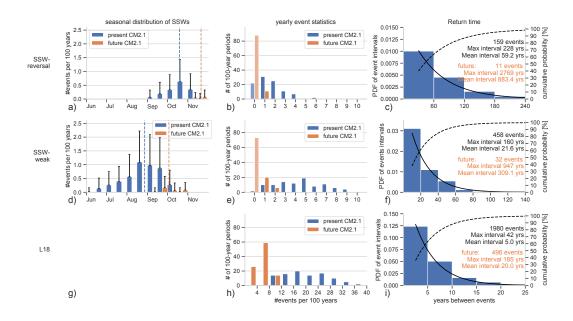


Figure 2. Event statistics: (left) Seasonal distribution, (middle) histogram of number of events per century, and (right) return time distribution histograms (bars) and theoretical distribution (black lines) for probability (solid) and cumulative distribution functions (dashed). Statistics are shown for (top) SSW-reversal, (middle) SSW-weak and (bottom) L18. For all plots, the present-day simulation is in blue and increased CO_2 ('future') in orange. On the left panels, statistics are shown for half-monthly intervals, the black whiskers show the standard deviation, and the vertical dashed lines indicate the mean date of occurrence. Panel (g) is empty as there is no seasonal information for L18. Note the differences in scales between rows. In panels (b) and (e), bars are drawn for each year, whereas in panel (h), the bars are drawn within intervals designated by the tick marks. Bars showing the number of centuries without event are pale.

1.6% for SSW-reversal and 4.6% for SSW-weak. Using the theoretical survival function, 164 we can then compute the probabilities of various scenarios (reported in Table 1). All of 165 these probabilities are consistent with the observational record of one SSW-reversal and 166 two SSW-weak events during the satellite era. Finally, neglecting any changes in climate 167 from further greenhouse gas forcing since 1990, we estimate from the present-day sim-168 ulation that the probability of at least one SSW by the end of the century (next 80 years) 169 would be 74% for SSW-reversals and 98% for SSW-weak events. Of course, this is only 170 hypothetical as greenhouse gas concentrations have already risen since 1990 and are pro-171 jected to further increase in the future. 172

¹⁷³ 4 Enhanced greenhouse gas forcing

To estimate the impact of enhanced greenhouse gas forcing on the occurrence of 174 SSWs in the SH, we conducted a second 9,900 year long simulation using increased CO_2 175 corresponding to the end of the century (1120 ppm instead of 353 ppm, henceforth called 176 'future'). The occurrence of SSWs in this simulation decreases drastically. The number 177 of SSW-reversals reduces from 159 SSWs for present-day to only 11 in the future sim-178 ulation, while SSW-weak events decrease from 458 to only 32 (Fig. 2). This translates 179 into a return time of one SSW-reversal every 883 and one SSW-weak every 309 years, 180 and a maximum of 1 SSW-reversal and 2 SSW-weak events per century. Note how the 181 most probable outcome by far for any given 100-year period is zero SSWs (median is zero 182 for both SSW-reversal and SSW-weak; Fig. 2b and e, orange). From the theoretical fit, 183 the probability of occurrence of at least one SSW-weak event in 80 years is now about 184 23% (2.8% for at least two SSWs; Table 1). The analysis also suggests that SSW-reversals 185 become very rare (probability of 8.7% within 80 years). SSWs not only become much rarer, 186 but are also occurring later in the year, with a mean date of 3 October for SSW-weak, 187 i.e. more than one month later than in the present-day simulation. For all definitions, 188 there is a strong tendency for fewer SSWs in the future-including L18, which reduce to 189 0-11 events per century. Thus, while the 2019 event is consistent with the occurrence rate 190 in our present-day simulation, it is inconsistent with the rate seen in our future simu-191 lation. Given the trend in SSW frequency, and that we are already one-third of the way 192 towards the year 2080 (when the greenhouse gas concentrations are projected to reach 193 the levels of our future simulation), we conclude that this latest event should not be at-194 tributed to increased CO_2 forcing, and might indeed be the last observed event this cen-195 tury. 196

The decrease in SSW frequency in the future is accompanied by a strengthening 197 of the SH polar vortex (Fig. 1b), which can be linked to stronger radiative cooling un-198 der increased greenhouse gas concentrations (Thompson et al., 2012; Santer et al., 2013). 199 In addition, our simulations suggest a decrease in wave forcing, more so during spring 200 than other times of the year (Fig. 1d). Together with an earlier study, which found a di-201 rect link between the SSW-reversal frequency and polar vortex strength (Jucker et al., 202 2014), our results suggest that the projected strengthening of the polar vortex along with 203 a decrease in wave forcing are responsible for a substantial decrease in the probability 204 of occurrence of SSWs. 205

²⁰⁶ 5 Comparison to NH

The occurrence of SSWs in the NH is very different from the SH, not just because of the much higher SSW frequency at present, but also in terms of future projections of both polar vortex strength and SSW frequency. As discussed in detail by Horan and Reichler (2017), our model climatology and variability in the NH compares well to reanalysis products (Fig. 3), and it produces about five SSWs per decade in the NH, in accordance with observations (Jucker & Reichler, 2018). Therefore, we perform the same analysis for the NH and briefly report our findings here.

ak SSW-reversal
% 1.6%
% 52%
% 35%
% 15%
rs 41 years
% 0.1%
rs 612 years
% 8.7%
% 0.4%
0 ⁰ an 3 ⁰ an 3 ⁰

Table 1. Results from the theoretical fitting of the return times (Figs. 2c and f). Yearly probability is the probability of an event occurring during any given year (1/mean return time), probability of exact observation is computed for 2 SSW-weak and 1 SSW-reversal in 41 years. Time periods give the interval after which an SSW is more probable than not (probability of one or more events > 50%). The labels 'present' and 'future' refer to the relevant CM2.1 simulations, and we use an 80-year period to compare to the time span 2021-2100 in the future simulation, but noting that this has CO_2 concentrations that are more representative of the end of the 21. century. Note that the observation percentages in the present simulation add to 101 instead of 100 due to rounding errors.

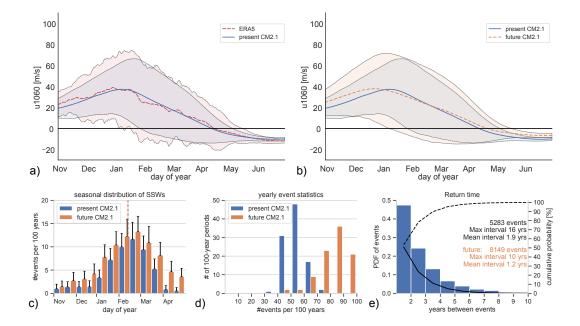


Figure 3. (top) u_{1060} for the Northern Hemisphere (NH) for (a) present-day and (b) increased CO₂ ('future'), similar to Fig. 1. (bottom) NH SSW-reversal statistics for (c) seasonal distribution, (d) number of events per century and (e) return time, similar to Fig. 2. Note the differences in scale of the bottom row compared to Fig. 2, which is a result of the higher occurrence rate for the NH.

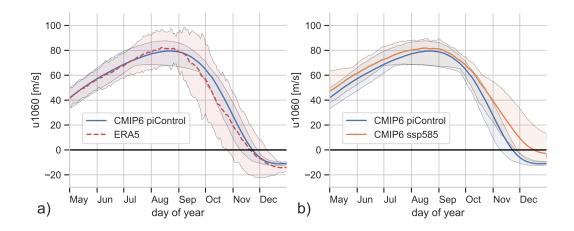


Figure 4. Analysis similar to Fig. 1, but using CMIP6 data. SSP585 data represents the climatology over 2080-2100. Shading corresponds to the range of model means (min to max) and the thick lines the multi model means. piControl is shown in blue, and SSP585 in orange, similar to Fig. 1.

The return time distribution shows that at intervals shorter than four years, NH 214 SSWs are not independent and random (Fig. 3e), probably reflecting the influence of slowly 215 evolving large scale climate modes, such as the El Niño Southern Oscillation or the Quasi-216 Biennial Oscillation, on the occurrence of SSWs (Holton & Tan, 1980; Taguchi & Hart-217 mann, 2006; Anstey & Shepherd, 2014). The NH polar vortex is also weaker and more 218 influenced by upward propagating planetary waves from the troposphere, resulting in a 219 more variable polar vortex than in the SH (Fig. 3, top). Our simulations suggest a slightly 220 weaker polar vortex and more SSWs in the future NH (Fig. 3, bottom; SSW-reversal only). 221 However, we have less confidence in this result because strong dynamical coupling be-222 tween the troposphere and the stratosphere in the NH complicates future projections, 223 and also because several past studies were unable to reach a consensus on possible fu-224 ture changes of SSW occurrence rates over the NH (Manzini et al., 2014; Ayarzagüena 225 et al., 2018; Wu et al., 2019; Ayarzagüena et al., 2020). There is also no consensus about 226 the future strength of the polar vortex (Simpson et al., 2018), which is in agreement with 227 our conclusion that the polar vortex strength is important for the frequency of SSWs. 228

229 6 CMIP6

To check the robustness of our single model simulations, we repeat our analysis with 230 CMIP6 data (see supplementary text S4 for details). We find that these models show 231 a positive polar vortex strength bias (Fig. 4) and generally struggle to produce the ob-232 served frequency of SSWs, with a range of 0.3-2.4 SSW-weak events on average in 80 years 233 for piControl (Table S1). The low SSW frequency in CMIP6 was also briefly noted in 234 recent work (Ayarzagüena et al., 2020). However, the statistical analysis again suggests 235 a decrease in SSWs in the future, with three models producing one single and two mod-236 els producing no SSW-weak event in SSP585 between 2021 and 2100 (Table S1b). Sim-237 ilar to our CM2.1 simulations, the CMIP6 models consistently project a strengthening 238 of the SH polar vortex (Fig. 4), suggesting that our main conclusion that SSWs will be-239 come much rarer in the future is robust. 240

Our enhanced CO₂ CM2.1 simulation only considers future increases in CO₂. Changes in other radiatively active gases, in particular the expected recovery of the ozone hole by 2080 (Dhomse et al., 2018), are not included. However, our 1120 ppm CO₂ concen-

tration is equal to the CO_2 concentration at the end of the century following the SSP585 244 scenario (which in addition to CO_2 also increases other greenhouse gases such as methane 245 and nitrous oxide (O'Neill et al., 2016; Meinshausen & Nicholls, 2020)). Consequently, 246 u_{1060} of our future simulation compares well to the end of the 21^{st} century in CMIP6 247 SSP585 model data (Fig. 4b). This is consistent with previous findings that over the long 248 term, the greenhouse effect from increasing CO_2 concentrations dominates the effect of 249 the ozone hole recovery (Barnes & Polvani, 2013). The similarities in u_{1060} and CO₂ con-250 centrations between our CM2.1 simulations and CMIP6 models gives us confidence that 251 our enhanced CO_2 simulation is relevant for end-of-century projections. 252

253 7 Conclusions

The 2002 and 2019 SSWs both resulted in exceptionally small ozone holes as have 254 not been observed since the 1980s. They were also followed by extended periods of neg-255 ative Southern Annular Mode at the surface, and 2019 in particular was linked to the 256 catastrophic fire season in South Eastern Australia. While possibly predictable on the 257 seasonal time scale, it has been difficult to determine how often SSWs should be expected 258 in the southern hemisphere, due to a relatively short observational record on one hand 259 and large model biases in the southern hemisphere stratosphere in most comprehensive 260 climate models on the other hand. Using a pair of exceptionally long and low bias cli-261 mate model runs, we found that while SSWs in the SH have significant impacts on strato-262 spheric ozone and surface weather, such events are rare and will become even rarer as 263 CO_2 concentrations increase. In our simulation based on 1990 conditions, the mean re-264 turn time for events similar to the 2002 and 2019 SSWs is about 22 years, with a 57%265 chance of at least two and a 30% chance of three or more SSW-weak events happening 266 within the time period spanned by the satellite record. Thus, it is no surprise that two 267 events have been observed, and there would be a fair chance of another SSW (of either 268 flavor) in the near future if CO_2 levels were kept constant. However, we show that one 269 should not make predictions of future occurrence from past data; given that the world 270 follows a high emissions pathway, our projections suggest that events similar to 2002 and 271 2019 will become extremely rare, with a mean return time of one in 309 years (or 0.3%272 each year) by the end of the century. 273

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Supporting Information for "How frequent are Antarctic sudden stratospheric warmings in present and future climate?"

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Contents of this file

- 1. Text S1 to S6
- 2. Figures S1 to S3
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Text S1. CM2.1 validation In addition to the discussion in the main text, Fig. S1 shows the latitude-pressure zonal mean zonal wind seasonal climatologies of the mean and interannual standard deviation. The model shows good agreement at all levels and

seasons, although there is some overestimation of southern upper stratospheric variability in September-October-November and the model does not simulate the Quasi-Biennial Oscillation, resulting in underestimates in tropical interannual variability.

Text S2. The 2002 and 2019 events Fig. S2 shows the evolution of u_{1060} and polar cap stratospheric ozone during the springs of 2002 and 2019 from ERA5.

Text S3. Surface impacts Just as for the observed SSWs, in our present-day simulation SSWs are followed by a negative phase of the SAM on a monthly to seasonal timescale (Fig. S3a; only composites for SSW-weak are shown) (Thompson et al., 2005), accompanied by colder and wetter conditions over New Zealand and South America as well as warmer and drier conditions over Eastern Australia (Figs. S3b and S3c). These surface impacts agree well with previous work (Gillett et al., 2006; Lewis, 2019; Lim et al., 2019) and the reanalysis data from the 2002 and 2019 events, confirming that our model reproduces the dynamical evolution of SSWs well and that our definition based on anomalous u_{1060} does indeed capture events with considerable surface impact. We note that the surface impact of early SSW-weak events (e.g. those occurring in June and July) is similar to the impact of later events (not shown).

Text S4. CMIP6 model selection We consider pre-industrial control (piControl) and Shared Socioeconomic Pathway 585 (SSP585) (O'Neill et al., 2014) simulations (which include e.g. ozone hole recovery, Fig. 4 and Table S1). The models from the CMIP6 archive were chosen based on the availability of daily data for both piControl and SSP585 scenarios, and given the lack of stratospheric variability in low top models (Charlton-Perez et al., 2013), we require a well resolved stratosphere with at least 30 vertical levels and

a model top at or above 1 hPa. For piControl we required at least 100 years of data for sufficient statistics. The five models that fulfill all these conditions are CESM2-WACCM, CanESM5, GFDL-CM4, INM-CM5-0, MIROC6, and the data used comprise a total of 3,341 years of piControl as well as 5x80 years of SSP585 (from 2021 to 2100). One ensemble member (r1i1p1f) for each model was considered.

Text S5. Uncertainty estimates For the two CM2.1 simulations, frequency uncertainties in Fig. 2 are computed by splitting the 9,900 years (after 90 years spinup) into 99 century-long non-overlapping segments, and computing the mean and standard deviation from this ensemble. For CMIP6 there are not enough events for similar statistical calculations, and the raw results are reported in Table S1.

Text S6. Return time If SSWs are random and independent, we should be able to model them as a Poisson process. For such a process, the return or waiting time can be computed using an exponential distribution with an expectation value equal to the mean occurrence frequency: PDF = $\lambda \exp(-\lambda x)$, where λ is the average frequency (e.g. 1/21.6 years for present-day SH SSW-weak events) and x is the waiting time in years (Gumbel, 1941). This is an approximation to a binomial distribution assuming large sample size and low probability. Since in our case we do not always have large sample size, we compute the return time using the binomial distribution. Then, the return time distribution is determined by the probability of zero events during a given time period (k = 0, n =number of years, p = 1/mean return time to be fitted). This has the advantage of being able to compute the probabilities for an arbitrary number of events, while still being able to check the validity of randomness and independence. The cumulative distribution function of the

exponential distribution is an approximation (again large sample size and low probability) for the survival function of a binomial distribution for zero events. Therefore, we use the latter to compute the probability of one or more events within a given time period, provided the events are independent and random.

Explicitly, the return time PDF of a random and independent process follows a binomial distribution of zero events, as the return time corresponds to the probability of no event happening within a given time interval:

$$P(y) = \left[\sum_{k=0}^{k=N} \binom{y}{k} \left(\frac{1}{\tau}\right)^k \left(1 - \frac{1}{\tau}\right)^{y-k}\right]_{N=0} = \left(1 - \frac{1}{\tau}\right)^y,\tag{1}$$

where y is the time interval in years, N(=0) is the number of events, and τ is the mean time interval between two SSWs (e.g. 21.6 years for present-day SSW-weak events). This is what is shown as solid black line in the return time plots of Fig. 2.

The probability of one or more events within a given time interval is simply 1 - P(y), which is shown as dashed black line in the return time plots. This so-called 'survival function' is used along with the cumulative probability function (as shown above but not setting N = 0) and the mass probability function (without the summation) to compute the various probabilities reported in the text and Table **??**. For instance, the probability of at least one SSW-reversal in 80 years is $1 - (1 - 1/883)^{80} \approx 8.7\%$, and the probability of exactly two SSW-weak events in 41 years of present-day conditions is

$$\binom{41}{2} \left(\frac{1}{21.6}\right)^2 \left(1 - \frac{1}{21.6}\right)^{41-2} \approx 28\%,\tag{2}$$

as reported in Table 1.

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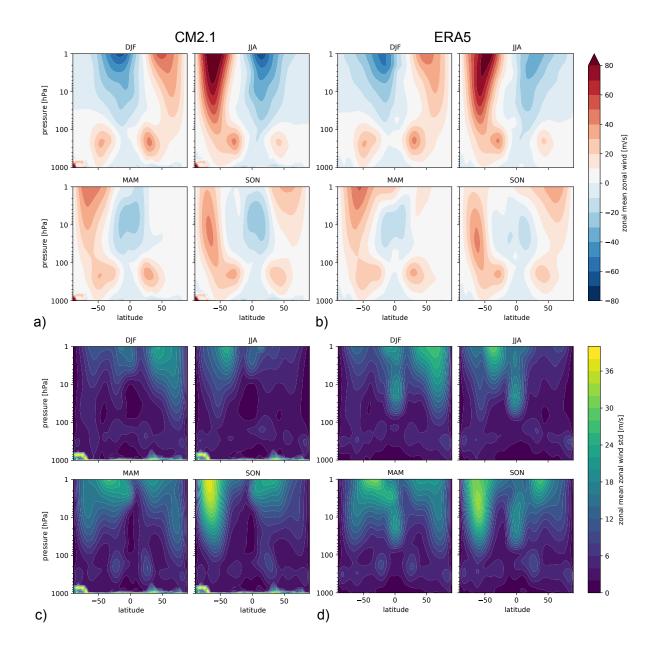
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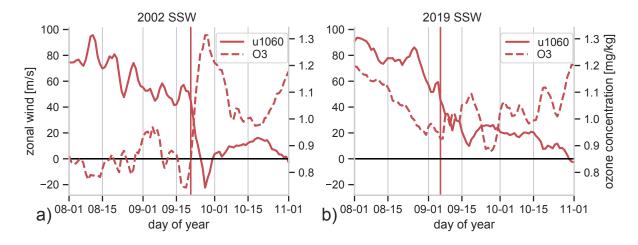
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Figure S1. Comparison of (left) CM2.1 and (right) ERA5 zonal mean zonal wind climatology (a,b) and interannual standard deviation (c,d).





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Figure S2. u_{1060} (solid), and polar cap (60-90°S) averaged ozone mass mixing ratio at 50 hPa (dashed) for the springs of (a) 2002 and (b) 2019 from ERA5 reanalysis. The solid vertical lines denote the onset date based on the SSW-weak definition.

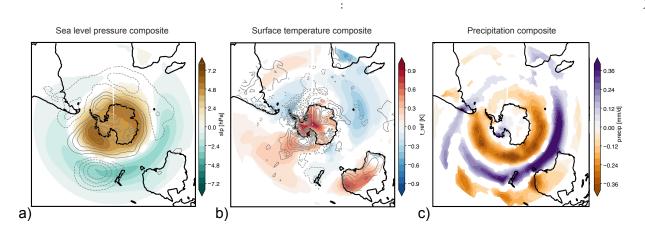


Figure S3. Composited surface anomalies averaged 0-60 days after the onset day for (a) surface pressure, (b) surface temperature and (c) precipitation for present-day CM2.1 SSW-weak events. The composites of the 2002 and 2019 events from ERA5 are added in gray contours for direct comparison except for precipitation which is too noisy in ERA5. Anomalies are relative to daily climatology, and only values which are statistically significant at the 5% level (two-sided *t*-test) are plotted.

Model	# years	# SSW-	mean return	# SSW-weak	# SSW-weak
		weak	time [years]	per 80 years	2021-2100
CESM2-WACCM	500	15	33.3	2.40	0
CanESM5	1000	6	166.7	0.48	1
GFDL-CM4	140	1	140.0	0.57	1
INM-CM5-0	1201	5	240.2	0.33	0
MIROC6	500	5	100.0	0.80	1

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Table S1. Statistical information for CMIP6 data. All columns except the last refer to piControl simulations, whereas the last reports results from the SSP585 simulations. The second last column normalizes the number of SSW-weak events in piControl to a 80-year period for direct comparison to 2021-2100. All models except CESM2-WACCM strongly underestimate the number of SSWs, and no model produces more than one single event between 2021 and 2100.