# Chondrule flattening by shock metamorphism in unequilibrated chondrites

Masaaki Miyahara<sup>1</sup>, Junnosuke Edanaga<sup>1</sup>, Akira Yamaguchi<sup>2</sup>, Takamichi Kobayashi<sup>3</sup>, Toshimori Sekine<sup>4</sup>, and Ayaka Nakamura<sup>1</sup>

<sup>1</sup>Hiroshima University <sup>2</sup>National Institute of Polar Research <sup>3</sup>National Institute for Materials Science <sup>4</sup>The University of Osaka

November 24, 2022

#### Abstract

Shock recovery experiments using ALH-78084 H3 and Y-793375 L3 chondrites are conducted in the shock pressure range of 11–43 GPa to reproduce shock-induced melting and chondrule flattening. Shock experiments prove that shock-induced melting occurs beyond 11 GPa at least. The melting occurs at the boundaries between chondrules and matrices. The melts include fine-grained silicate minerals, glasses, and amoeba or spherical metallic Fe-Ni or metallic Fe-Ni–iron-sulfide with a eutectic texture, which coincides with shock-induced melts in shocked natural chondrites. Shock experiments also prove that shock-induced flattening of chondrules occurs and the flattening degree increases with increasing shock pressure. Taking account of not only the shock experiments of ordinary chondrites but also carbonaceous chondrites. Considering the shock experiments of the Allende CV3 and Murchison CM2 carbonaceous chondrites along with present shock experiments using H/L3 ordinary chondrites, the aspect ratios of chondrules in unequilibrated chondrites (R<sub>cho</sub>) can be expressed as follows: R<sub>cho</sub> = 0.011 ( $\pm$ 1) × Pressure (GPa) + 1.18 ( $\pm$ 3). The long axes of chondrules in shocked ALH-78084 H3 and Y-793375 L3 chondrites have preferred orientations and the degree increases with increasing shock pressure. Natural L/LL3 ordinary chondrites with shock-induced melts have higher aspect ratios and preferred orientations than those without shock-induced melts although it is difficult to determine quantitatively shock pressure using the empirical formula between the aspect ratios of chondrules and shock pressure.

| 1        | Chondrule flattening by shock metamorphism in unequilibrated chondrites   |
|----------|---|
| 2        |   |
| 3        | Masaaki Miyahara <sup>1*</sup> , Junnosuke Edanaga <sup>1</sup> , Akira Yamaguchi <sup>2</sup> , Takamichi Kobayashi <sup>3</sup> , |
| 4        | Toshimori Sekine <sup>4,5</sup> , and Ayaka Nakamura <sup>1</sup>   |
| 5        |   |
| 6        |   |
| 7        | <sup>1</sup> Graduate School of Advanced Science and Engineering, Hiroshima University,   |
| 8        | Higashi-Hiroshima, 739-8526, Japan  |
| 9        | <sup>2</sup> National Institute of Polar Research, Tokyo 190-8518, Japan  |
| 10       | <sup>3</sup> National Institute for Materials Science, Tsukuba 305-0047, Japan  |
| 11       | <sup>4</sup> Center for High Pressure Science and Technology Advanced Research, Pudong,   |
| 12       | Shanghai 201203, P.R. China   |
| 13       | <sup>5</sup> Graduate School of Engineering, Osaka University, Suita, Osaka 565-0871 Japan  |
| 14       |   |
| 15       |   |
| 16       | *Corresponding author: Dr. Masaaki Miyahara   |
| 17       | Graduate School of Advanced Science and Engineering, Hiroshima University,  |
| 18       | Higashi-Hiroshima, 739-8526, Japan.   |
| 19       | Postal code: 739-8526   |
| 20       | Telephone: +81-824-24-7461  |
| 21       | FAX: +81-82-424-0735  |
| 22       | E-mail address: miyahara@hiroshima-u.ac.jp  |
| 23       |   |
| 24       | Key points:   |
| 25       | • Shock experiments of H/L3 chondrites were conducted in the range of 11–43   |
| 26       | GPa.  |
| 27       | • Shock-induced melting occurs above 11 GPa and the aspect ratios of chondrules   |
| 28       | increase with increasing shock pressure.  |
| 29       | • No differences in chondrule flattening between carbonaceous and ordinary  |
| 30       | chondrites.   |
| 31       |   |
| 32       | Abstract  |
| 33       | Shock recovery experiments using ALH-78084 H3 and Y-793375 L3 chondrites are  |
| 34<br>05 | conducted in the shock pressure range of 11–43 GPa to reproduce shock-induced   |
| 35       | melting and chondrule flattening. Shock experiments prove that shock-induced  |
| 36       | melting occurs beyond 11 GPa at least. The melting occurs at the boundaries   |
| 37       | between chondrules and matrices. The melts include fine-grained silicate minerals,  |

38 glasses, and amoeba or spherical metallic Fe-Ni or metallic Fe-Ni-iron-sulfide with a 39 eutectic texture, which coincides with shock-induced melts in shocked natural 40 chondrites. Shock experiments also prove that shock-induced flattening of chondrules occurs and the flattening degree increases with increasing shock pressure. 41 42 Taking account of not only the shock experiments of ordinary chondrites but also 43 carbonaceous chondrites, the flattening degree does not depend significantly on the 44 densities, porosities, and chondrule/matrix ratios of chondrites. Considering the 45 shock experiments of the Allende CV3 and Murchison CM2 carbonaceous chondrites along with present shock experiments using H/L3 ordinary chondrites, the 46 47 aspect ratios of chondrules in unequilibrated chondrites (R<sub>cho</sub>) can be expressed as follows:  $R_{cho} = 0.011 (\pm 1) \times Pressure (GPa) + 1.18 (\pm 3)$ . The long axes of chondrules 48 49 in shocked ALH-78084 H3 and Y-793375 L3 chondrites have preferred orientations 50 and the degree increases with increasing shock pressure. Natural L/LL3 ordinary chondrites with shock-induced melts have higher aspect ratios and preferred 51 52 orientations than those without shock-induced melts although it is difficult to 53 determine quantitatively shock pressure using the empirical formula between the 54 aspect ratios of chondrules and shock pressure.

55

#### 56 Plain Language Summary

57 An early ordinary chondrite parent-body had an onion-shell structure and was 58 disrupted by an impact. Some ordinary chondrites derived from the parent-body 59 surface have melts, high-pressure polymorphs, and flattened chondrules, which may 60 be due to shock metamorphism. Hence, shock recovery experiments using ordinary 61 chondrites were conducted in the shock pressure range of 11-43 GPa to reproduce 62 shock-induced melting and chondrule flattening. Shock-induced melting occurs 63 beyond 11 GPa. The melting occurs at a boundary between chondrules and 64 surrounding fine-grained materials. Shock-induced flattening of chondrules occurs 65 and the flattening degree increases with increasing shock pressure. Taking account of 66 not only the shock experiments of ordinary chondrites but also carbonaceous 67 chondrites, the flattening degree does not depend significantly on the densities, 68 porosities, and chondrule/matrix ratios of chondrites. The aspect ratios of chondrules 69 in unequilibrated chondrites (Rcho) can be expressed as follows: Rcho =  $0.011 (\pm 1)$ 70  $\times$  Pressure (GPa) + 1.18 (±3). The long axes of chondrules in experimentally shocked 71 ordinary chondrites have preferred orientations and the degree increases with 72 increasing shock pressure.

73

#### 74 1. Introduction

One of the classic structure models for ordinary chondrite parent-bodies is the onion shell model derived from maximum metamorphic temperature and cooling rate recorded in each petrologic type ordinary chondrite (e.g., Trieloff et al., 2003). Following the onion shell model, early ordinary chondrite parent-bodies had consisted of petrologic type 3–6 ordinary chondrites toward the inside from the outside. The early ordinary chondrite parent-bodies were disrupted by impacts and the fragments escaped from the parent-bodies.

82 Shock features recorded in ordinary chondrites have been investigated to clarify the disruption histories. Most works have focused on petrologic type 5 and 6 83 84 ordinary chondrites derived from the inner portions of the parent-bodies (e.g., Chen 85 et al., 1996; Ohtani et al., 2004). In contrast, few works have worked on shock 86 features recorded in petrologic type 3 and 4 ordinary chondrites derived from the 87 outer portions of the parent-bodies (Ruzicka et al., 2015a-b), because, in general, the shock features in petrologic type 3 and 4 ordinary chondrites are not distinct 88 89 compared to those in petrologic type 5 and 6 ordinary chondrites.

90 Miyahara et al. (2021) conduct systematic investigations of shock-induced melts 91 and high-pressure polymorphs in ordinary chondrites and report that some petrologic 92 type 3 ordinary chondrites have shock-induced melts and high-pressure polymorphs. 93 Also, some petrologic type 3 ordinary chondrites with shock-induced melts include 94 flattened chondrules. Foliation in chondrites is formed during the sedimentation of 95 chondrules on the parent-bodies (Dodd, 1965; Martin & Mills, 1980). The foliation 96 in chondrites also can be formed by impacts because when a shock wave propagates 97 in a rock, uniaxial compaction occurs (Sneyd et al., 1998; Gattacceca et al., 2005).

We raise the possibility that the flattened chondrules in type 3 ordinary chondrites
are due to shock metamorphism and the aspect ratios of chondrules depend on shock
pressure. Several shock recovery experiments using equilibrated ordinary chondrites
were conducted to clarify the shock features (e.g., Stöffler et al., 1991; Kohout et al.,
2020). In contrast, shock recovery experiments using unequilibrated ordinary
chondrites have not been conducted so far.

104 The average size of chondrules in H3 chondrites is distinct from those of L/LL3 chondrites: the mean diameters of chondrules in H3 and L/LL3 are about 450 and 105 106 500-550 µm, respectively (Friedrich et al., 2015). The flattening degree of 107 chondrules may depend on the size distribution of chondrules. Hence, we conduct 108 shock recovery experiments using H3 and L3 chondrites to i) reproduce 109 shock-induced melt and chondrule flattening and ii) clarify the relationship between 110 flattening degree and shock pressure. This study also uses several Antarctic and 111 non-Antarctica petrologic type 3 ordinary chondrites with or without a

shock-induced melt to verify the shock recovery experiments.

113

#### 114 2. Materials and experimental methods

#### 115 2.1 Materials

116 Shock experiments used two petrologic-type 3 ordinary chondrites: the Allan 117 Hills (ALH)-78084 H3 and Yamato (Y)-793375 L3. Sliced ALH-78084 H3 (Area: about 20 x 40 mm<sup>2</sup>, thickness: about 2.73 mm) and Y-793375 L3 (Area: about 22 x 118 38 mm<sup>2</sup>, thickness: about 2.08 mm) samples were allocated from the National 119 Institute of Polar Research (NIPR), Japan. These two chondrites show a 120 121 well-preserved chondritic texture suitable as starting materials for the shock 122 experiments. Fifteen H3, twenty-three L3, and twenty-three LL3 ordinary chondrites 123 were selected for the comparisons and verifications of the shock recovery 124 experiments. The classification of individual samples (chemical group and petrologic 125 type) followed the Meteoritical Society Bulletin database.

#### 126 **2.2 Shock experiments**

127 The allocated ALH-78084 H3 and Y-793375 L3 chondrite samples were 128 adhered to glasses using Crystalbond<sup>TM</sup> 509. Doubly polished samples with a 129 thickness of 1.5 mm were made by a surface grinding machine. Disc samples ( $\varphi = 6$ 130 or 12 mm) were cored out from the doubly polished samples by an ultrasonic 131 machine. The cored-out disc samples were put in acetone for thirty hours to remove 132 Crystalbond<sup>TM</sup> and dried at 373 K for thirty hours. The prepared disc samples were 133 kept in a closed plastic bag with a silica gel until shock experiments.

134 The densities of individual disc samples were calculated by measuring their 135 thicknesses, diameters, and weights (Table 1). The disc samples were put into 136 stainless (SUS 304) containers (30 mm diameter x 30 mm long). The disc samples 137 were placed at a depth of 3 mm from the impact surface of the stainless container. 138 Shock experiments used a 30-mm bore propellant gun installed at the National 139 Institute for Materials Science (NIMS), Tsukuba, Japan. 2-mm thick steel (SUS 304) 140 or tungsten (W) flyer plate, attached to the head of a projectile, impacted the stainless 141 containers. Velocity just before the impact was measured with a magneto flyer 142 method (Kondo et al., 1977). The peak shock pressure was calculated by the 143 impedance match method assuming the sample pressure reached equilibrium with the 144 container pressure because the experimental procedures are similar to those 145 described by Yamaguchi & Sekine (2000). Table 1 shows the details of individual 146 experimental conditions.

#### 147 **2.3 Sample analysis**

148 Stainless containers were recovered after individual shock experiments and were

sliced open with a diamond saw and lathe to observe the cross-sections of disc samples. 1-3 slices were made from each recovered stainless container. The slices were embedded into an epoxy resin and were polished with an abrasive paper under the dry condition to reduce the exfoliation of the disc samples.

153 We took back-scattered electron (BSE) images of polished cross-sections by a 154 field-emission scanning electron microscope (FE-SEM): JEOL JSM-7100F (installed 155 at NIPR) at an accelerating voltage of 15 kV. A dedicated image processing 156 application the Oxford AZtec installed in JEOL JSM-7100F combined all individual 157 BSE images into one BSE image depicting the whole area of the cross-section. 158 Subsequently, individual melts and deformation textures in the cross-sections were 159 observed under higher magnification BSE image mode. The short and long axes of 160 chondrules in the cross-sections were measured using the combined BSE images and 161 dedicated image processing application SmartGrain (Tanabata et al., 2012). The measured axis lengths are apparent. The measured axis lengths (2D) were converted 162 163 to the actual lengths (3D) by multiplying a factor of 1.273 (Kong et al., 2005). The angles of the long axes to an arbitrary axis were measured by the image processing 164 165 application ImageJ. The total lengths of open cracks in individual cross-sections 166 were also measured by ImageJ.

167 The whole area BSE images of selected petrologic type 3 ordinary 168 chondrites were also observed by FE-SEM to measure the short and long axes of 169 chondrules and the angles of the long axes to an arbitrary axis. Individual melts and 170 deformation textures were observed under higher magnification BSE image mode. 171 The measurement procedures of the chondrules are as same as the samples recovered 172 after shock experiments. The mineralogy of melts was determined using a laser 173 micro-Raman spectrometer: Renishaw inVia (at NIPR). An optical microscope was 174 used to focus the excitation laser beam (the 532 nm line of green laser). The laser 175 power was kept at < -7 mW to reduce laser beam damage. The acquisition time was 176 10-60 s. For each phase, the Raman spectrum was acquired in the spectral region of 200 to 1500 cm<sup>-1</sup>. 177

178

### 179 **3. Results**

#### 180 **3.1 Petrologic descriptions of starting materials**

Figures S1–2 show the petrologic thin sections of unshocked ALH-78084 H3 and Y-793375 L3 chondrites. No distinct shock features such as undulatory extinction and fractures were observed in the petrologic thin sections. Figures S3a–b show the representative petrological textures of ALH-78084 H3 and Y-793375 L3 chondrites, respectively. Melts were not found in ALH-78084 H3 and Y-793375 L3 chondrites.
A small amount of weathering vein occurred in ALH-78084 H3 chondrite (Fig. S3a).

187 The long and short axes of chondrules in ALH-78084 H3 and Y-793375 L3 188 chondrites were measured using the petrologic thin sections. The distributions of 189 long and short axes did not follow the Gaussian distribution (Figs. S4a-d). The long 190 and short axes of chondrules in ALH-78084 H3 chondrite were respectively 429 µm 191 and 352  $\mu$ m (median, n = 258) (Table 2). Those in Y-793375 L3 chondrite were 773 192  $\mu$ m and 634  $\mu$ m (n = 143) (Table 2). The aspect ratios of chondrules in ALH-78084 193 H3 and Y-793375 L3 chondrites were respectively 1.19 and 1.22 (median) (Figs. 194 S4e-f, Table 2). The number densities of chondrules in ALH-78084 H3 and 195 Y-793375 L3 chondrites were 2.12 and 1.18 mm<sup>-2</sup>.

#### 196 3.2 Experimentally shocked ALH-78084 H3 chondrite

197 Shock experiments were conducted in the shock pressure range from 11 to 43 GPa (Table 1). Six samples were successfully recovered after shock experiments. 198 199 Three polished slices were prepared from each recovered sample except for Shot No. 200 H\_7. Figure S5 shows the whole area BSE images depicting the cross-sections of 201 recovered samples. Some recovered samples were concave in the shapes. With 202 increasing shock pressure, the degree of concave became bigger (Fig. S5). A part of 203 sample Shot No. H\_6 (42.9 GPa) came out into the stainless container (Fig. S5e). 204 Open cracks cutting chondrules occurred in most samples and arranged sub-parallel 205 to shockwave front (Figs. 1a, S5). Figure 2a shows the drawing of open cracks in 206 each shot. The length of open cracks per unit square in each shot is shown in Table 2.

207 Some olivine grains in shocked samples had irregular and/or planar fractures 208 (Fig. 1b). Melts were found from all the shocked samples. Melts occurred along 209 boundaries between chondrules and matrices consisting of fine-grained materials 210 (Fig. 1c). In most cases, each melt was isolated and not interconnected with each 211 other. The melts consisted of quenched melts and mineral and/or chondrule 212 fragments. The quenched melts consisted of fine-grained silicate minerals (olivine 213 and pyroxene), glasses, and metals (Fig. 1d). The mineral and chondrule fragments 214 entrained in the melts were rounded due to melting (Fig. 1d). The metals, which were 215 in the shapes of amoeba or spherule, consisted of metallic Fe-Ni or metallic Fe-Ni-216 iron-sulfide with a eutectic texture (Figs. 1d-f). Voids occurred in the melts (Figs. 217 1e-f). No high-pressure polymorphs were found from the quenched melts or mineral 218 and/or chondrule fragments.

The long and short axes of chondrules and chondritic fragments in the
shocked samples were measured. The aspect ratios are shown in Table 2 and Figure
3a. The number of measurable chondrules decreased with increasing shock pressure

because the deformation degree of sample containers increased (Fig. S5). All the
aspect ratios of chondrules in the shocked samples (1.19–1.57) were bigger than that
of starting material (1.19)(Table 2). Azimuths of long axes to an impact surface were
measured (Table S1, Fig. S6). Chondrules in Shot No. H\_7 were not measured
because its cross-section was too small for measurements. Only textural observations
by FE-SEM were conducted for Shot No. H\_7.

#### 228 **3.3 Shocked Y-793375 L3 chondrite**

229 Shock experiments were conducted in the shock pressure range from 11 to 230 43 GPa for Y-793375 L3 (Table 1). Six samples were successfully recovered after 231 shock experiments. Three polished slices were prepared from each recovered sample. 232 Figure S7 shows the whole area BSE images depicting the cross-sections of 233 recovered samples. Some recovered samples were concave in the shapes and the 234 degree became bigger with increasing shock pressure (Figs. 2b, S7). A part of sample 235 Shot No.  $L_6$  (42.5 GPa) came out into the stainless container (Fig. S7f). Open 236 cracks arranging sub-parallel to shockwave occurred in the shocked samples (Fig. 237 4a). Figure 2b shows the drawings of open cracks in each shot. The length of open 238 cracks per unit square in each shot is shown in Table 2.

239 Some olivine grains in shocked samples had irregular and/or planar fractures 240 (Fig. 4b). Melts were found from all the shocked samples. Isolated melting occurred 241 along with boundaries between chondrules and matrices (Fig. 4c). The textures and 242 constituents of melts were similar to those of ALH-78084 H3 chondrite (Figs. 4c-f). 243 Some melted chondrule glasses had a flow-like texture (Fig. 4e). One-third of the 244 cross-section of Shot No. L\_4 (26.4 GPa) was melted (Figs. S7d, S8a). The quenched 245 melts consisted mainly of fine-grained pyroxene with a dendrite texture, glasses, and 246 metallic Fe-Ni-FeS with a eutectic texture (Figs. S8b-c). Voids occurred in the 247 glasses and also in some metals (Fig. S8c). No high-pressure polymorphs were found 248 from all the shocked samples.

The long and short axes of chondrules (and chondritic fragments) and the aspect ratios in the shocked samples are shown in Table 2. All the aspect ratios of shocked samples (1.36–1.75) were bigger than that of starting material (1.22)(Table 2). The azimuths of long axes to an impact surface are shown in Table S1 and Figure S9.

#### 254 **3.4 Petrologic type 3 ordinary chondrites**

Shock-induced melts and aspect ratios of chondrules (and chondritic fragments) in selected H/L/LL3 ordinary chondrites were investigated. Typical whole area BSE images of investigated H/L/LL3 ordinary chondrites are shown in Figure S10. The modes of melting are shown in Tables S2–4 and the definition of 259 melting followed Miyahara et al. (2021): i) pocket, ii) line, and iii) network. The 260 textures and constituents of shock-induced melts were similar to those of shocked 261 ALH-78084 H3 and Y-793375 L3 chondrites (Figs. 5a-b). Melting occurred also in 262 some fragments of chondrule glasses in contact with the shock-induced melts (Figs. 263 5c-d). Void, which was ubiquitously observed in shocked ALH-78084 H3 and 264 Y-793375 L3 chondrites, did not occur in the shock-induced melts of H/L/LL3 265 ordinary chondrites. Distinct open cracks cutting chondrules, which was observed in 266 shocked ALH-78084 H3 and Y-793375 L3 chondrites, were not observed in 267 H/L/LL3 ordinary chondrites. By contrast, open cracks occurred in the 268 shock-induced melts of some L/LL3 ordinary chondrites (Fig. 5e).

269 The long and short axes of chondrules (and chondritic fragments) in H/L/ 270 LL3 chondrites are shown in Tables S2-4. In fifteen H3 chondrites, only Y-981139 271 H3 chondrite included a shock-induced melt (Table S2). No high-pressure 272 polymorphs occurred in Y-981139 H3 chondrite. The difference in the aspect ratios 273 between H3 ordinary chondrites without (1.17-1.28) or with shock-induced melts 274 (1.24) was not clear (Fig. 6, Table S2). In twenty-three L3 chondrites, sixteen L3 275 chondrites included shock-induced melts (Table S3). Y-981327 chondrite had a 276 pervasive shock-induced melt (Figs. S10b-c). Asuka (A)-880870, A-881096, 277 Y-86706, and Northwest Africa (NWA) 8664 L3 chondrites included Na-pyroxene 278 and coesite as high-pressure polymorphs (Figs. 5e-f, Table S3). The aspect ratios of 279 chondrules in L3 ordinary chondrites without or with shock-induced melts were 280 1.20-1.36 and 1.27-1.75, (Table S3). The aspect ratio of chondrules in Y-981327 281 chondrite was higher (1.75) than others (Table S3, Fig. 7). In twenty-three LL3 282 ordinary chondrites, five LL3 ordinary chondrites included shock-induced melts 283 (Table S4, Fig. 8). A-881199 and A-881981 LL3 chondrites included Na-pyroxene 284 and coesite as high-pressure polymorphs. The aspect ratios of chondrules in LL3 285 ordinary chondrites without or with shock-induced melts were 1.16-1.41 and 1.31-286 1.54 (Table S4). The angles of long axes to arbitrary axes in H/L/LL3 ordinary 287 chondrites were also measured (Tables S2-4).

288

#### 289 4. Discussion

#### 290 4.1 shock-induced melting

291 Melting occurs in all the shocked ALH-78084 H3 and Y-793375 L3 292 chondrite samples and the melts have evidence of immiscibility between silicate 293 melts and metallic melts under high-pressure (e.g., Kato & Ringwood, 1989). The 294 melts in the shocked samples consist of fine-grained silicate minerals (olivine and/or 295 pyroxene), glasses, and the spherules of metallic Fe-Ni or metallic Fe-Ni– iron-sulfide with a eutectic texture (Figs. 1c–f, 4c–d). Some spherules of fine-grained
silicate mineral assemblages are included in the spherules of metallic Fe-Ni or
metallic Fe-Ni–iron-sulfide with a eutectic texture (Figs. 1c–d, 4f). All these features
indicate that melting occurs under high-pressure.

300 The temperature in the shocked ALH-78084 H3 and Y-793375 L3 chondrite 301 samples during the impact is not homogeneous. Most chondrules in shocked samples 302 do not show clear evidence of melting (Figs. S5, S7). Melting occurs in matrices in 303 contact with chondrules (Figs. 1c, 4c). The porosity of matrices is higher compared 304 to that of chondrules because the matrices are fine-grained crystal assemblages. 305 When compression wave propagates in un-equilibrated chondrites, their matrices 306 with relatively low shock impedance decrease their volumes more drastically than 307 their chondrules with relatively high shock impedance. The large volume decrease 308 induces a temperature spike resulting in local melting in the matrices.

309 The petrologic structure of un-equilibrated chondrites, which consists of 310 coarse-grained chondrules fine-grained materials and (matrices), makes 311 shock-induced melting easier under low pressure. Hirata et al. (2009) conducted 312 shock experiments using mimic porous primitive chondritic materials (initial 313 porosity: 35±5%). Shock veins occur above 25 GPa in the shocked mimic porous 314 primitive chondritic materials. By contrast, shock-induced melting occurs in the 315 un-equilibrated H3 and L3 ordinary chondrites with low porosities even at 11 GPa. A 316 local temperature spike under low shock pressure in un-equilibrated chondrites is 317 predicted by numerical simulation (Bland et al., 2014). The petrologic structure of a 318 target becomes a more important factor in some cases than its porosity for 319 shock-induced melting.

320 If petrologic type 3 ordinary chondrites expose on asteroids, the surfaces 321 have experienced melting along with brecciation. The onion shell model proposes 322 that petrologic type 3 ordinary chondrites have made up the outmost layers of 323 ordinary chondrite parent-bodies (e.g., Trieloff et al., 2003). Shock experiments 324 prove that even if the impact of 0.57 km/sec causes shock-induced melting in 325 petrologic type 3 ordinary chondrites. Following the Rankine-Hugoniot equation, 326 shock pressure depends on impact velocity. In general, the existence of 327 shock-induced melts in ordinary chondrites implies high shock pressure. However, 328 petrologic type 3 ordinary chondrites do not apply the criterion. The impact velocity of 0.57 km/sec is much lower compared to the most probable impact velocity in the 329 330 asteroid belt (4.4 km/sec) (Bottke et al., 1994). The surfaces of S-type asteroids may 331 like lunar regolith that includes brecciated rocks and melts.

332 Heating by shock-induced melts does not significantly affect thermal 333 metamorphism in petrologic type 3 ordinary chondrites. Numerical simulations 334 indicate that the temperature of matrices during the impact becomes higher with 335 increasing chondrule/matrix ratios at the same impact speed in unequilibrated 336 chondrites (Bland et al., 2014). The maximum temperature in petrologic type 3 337 ordinary chondrites is higher compared to carbonaceous chondrites because the 338 former includes a few amounts of matrices (Weisberg et al., 2006). The 339 shock-induced melts in the matrices may work as a heat source for thermal 340 metamorphism. However, the heat sources in the petrologic type 3 ordinary 341 chondrites are small due to the limited amounts of matrices. Hence, the 342 shock-induced melts do not affect the bulk-rock temperature of petrologic type 3 343 ordinary chondrites.

344 Voids occur in the shock-induced melts of shocked ALH-78084 H3 and 345 Y-793375 L3 chondrite samples (Figs. 1f, 4f), which is due to the degassing of 346 volatile elements during the decompression stage. Parts of shocked chondrites melt, 347 and volatile elements are dissolved into the melts under high-pressure. The degassing 348 occurs before the melts are quenched. Decompression is initiated when a rarefaction 349 wave catches up the compression wave. Assuming that the longitudinal wave 350 velocities of chondrite and stainless container at ambient pressure conditions are 6 351 and 5.77 km/sec., respectively, the propagations of compression and rarefaction 352 waves in the sample containers are calculated. The rarefaction wave achieves the 353 backside of samples in  $1-2 \mu$  sec after the impact. The quenching of melts is initiated 354 by thermal conduction into the stainless container. The degassing is due to the 355 limited duration of high-pressure compared to the time required for the quenching.

Shock-induced melts in investigated H/L/LL3 chondrites do not include 356 357 voids (Fig. 5), which is due to the difference in the durations of high-pressure 358 between shock experiments and natural impact events. The duration of high-pressure 359 recorded in several ordinary chondrites is estimated using the kinetics of 360 high-pressure polymorphs or diffusion rates of trace elements (Ohtani et al., 2004; 361 Beck et al., 2005; Xie et al., 2006). The durations (several milliseconds to seconds) 362 are much longer than the shock experiments (several microseconds). Hence, the 363 melts can be quenched under high-pressure.

Some quenched melts include fractures (Figs. 5e–f), which is due to a heterogeneous compression in un-equilibrated chondrites. The quenched melts, which are quenched under high-pressure after melting, is denser than original matrices. Chondrules are less compacted due to low porosity during the compression stage. The fractures in the melts are induced by the tensile failure of the quenchedmelts during the decompression stage.

#### 370 4.2 Shock-induced flattening of chondrules

371 The aspect ratios of chondrules in shocked ALH-78084 H3 and Y-793375 372 L3 chondrites increase with increasing shock pressure (Fig. 3c), indicating that the 373 chondrules are deformed by shock metamorphism. Shock recovery experiments 374 prove that both the chondrule flattening along with melting occur in type 3 ordinary 375 chondrites by shock metamorphism. Liner fitting is conducted for plotted data by the 376 least-squares method (Fig. 3c). The aspect ratios at ambient pressure are fixed with 377 the aspect ratios of chondrules in starting materials. The calculated slopes for ALH-78084 H3 and Y-793375 L3 chondrites are 0.88 x  $10^{-2}$  and 1.09 x  $10^{-2}$ . The 378 aspect ratios of chondrules in experimentally shocked Murchison CM2 and Allende 379 380 CV3 chondrites (Tomeoka et al., 1999; Nakamura et al., 2000) are also plotted in 381 Figure 3c. The calculated slopes for Murchison CM2 and Allende CV3 chondrites are  $1.40 \times 10^{-2}$  and  $2.38 \times 10^{-2}$ . 382

383 Considering the interquartile range, the flattening degree of chondrules in 384 Y-793375 L3 chondrite is a bit higher than that in ALH-78084 H3 chondrite (Figs. 385 3a-c), which is due to the grain refining of chondrules. There are no distinct 386 differences in the densities and chondrule/matrix ratios between ALH-78084 H3 and 387 Y-793375 L3 chondrites as starting materials (Table 1). By contrast, the size of 388 chondrules in ALH-78084 H3 chondrite is finer than that of Y-793375 L3 chondrites 389 (Table 1). A petrologic-type 3 ordinary chondrite is an assemblage of chondrules, 390 which can be regarded as a polycrystal. Yield stress  $(\sigma_v)$  can be described by the 391 Hall-Petch relationship,

392

 $\sigma_{\rm y} = \sigma_0 + k_{\rm y} d^{-1/2}$ 

where  $\sigma_{0, k}$ , and *d* are the internal stress, the strengthening coefficient, and grain size, respectively. Following the Hall-Petch relationship, the yield stress of petrologic type 3 ordinary chondrites increases with decreasing the size of chondrules. The strengthening by grain refinement inhibits the flattening of chondrules in ALH-78084 H3 chondrite.

The size sorting of chondrules also affects the flattening degree of chondrules. The size of chondrules in ALH-78084 H3 chondrite is sorted, whereas that in Y-793375 L3 chondrite is scattered (Fig. S4). In petrologic type 3 ordinary chondrites consisting of chondrules with varied sizes, the yield stress is heterogeneous. The heterogeneous yield stress decreases the strength of Y-793375 L3 chondrite as a bulk-rock, which increases the flattening degree of chondrules. 404 The relative abundance of matrices (+ lithic or mineral fragments) does not 405 significantly affect the flattening degree of chondrules. The relative abundances of 406 chondrules and matrices (+ lithic or mineral fragments) in Allende CV3 chondrites 407 are 35–43 % and 41– 55 % (McSween, 1977; Ebel et al., 2016). Those in Murchison 408 CM2 are ~16 % and ~75 % (McSween, 1979). The relative abundances of 409 chondrules in ALH-78084 H3 (chondrule: 65 %, matrix: 15 %) and Y-793375 L3 410 (chondrule: 63 %, matrix: 23 %) chondrites are much higher than Murchison CM2 and Allende CV3 chondrites. However, the flattening degree of chondrules in 411 412 Murchison CM2 chondrite is similar to those of ALH-78084 H3 and Y-793375 L3 413 chondrites.

414 A bulk-rock porosity does not also affect the flattening degree of chondrules. 415 The porosities of Allende CV3 and Murchison CM2 are 16.9-27.9 % and 18.8-416 24.9 % (Macke et al., 2011). Although the porosities of ALH-78084 H3 and Y-793375 L3 chondrites are not measured, the average porosities of H and L 417 418 ordinary chondrites are 7.0 and 5.6 % (Consolmagno et al., 2008). However, there 419 are no distinct differences in the flattening degrees between H/L3 ordinary chondrites 420 and Murchison CM2 chondrite. The flattening degree of chondrules in Allende CV3 421 chondrite is much higher than Murchison CM2 chondrite although there are no 422 distinct differences in their physical properties (Fig. 3c). We could not find an 423 adequate factor to explain the difference.

424 Long axes of flattened chondrules have a weak preferred orientation because 425 of uniaxial compression during shock metamorphism. The long axes of chondrules in 426 ALH-78084 H3 and Y-793375 L3 chondrites do not arrange homogeneously (Figs. 427 S6, S9). We first draw the Lorenz curves using the azimuths of long axes in starting 428 materials and each shocked sample (Table S1, Figs. S11a-b) and calculate the Gini 429 coefficients to evaluate the preferred orientation (Table 2). The Gini coefficients of 430 long axes in ALH-78084 H3 and Y-793375 L3 chondrites as starting materials are 0.33 and 0.32, respectively (Table 2). If the long axes do not have preferred 431 432 orientations, the Gini coefficient becomes null. Although the Gini coefficients of 433 ALH-78084 H3 and Y-793375 L3 chondrites keep small, the values become bigger 434 slightly with increasing shock pressure (Fig. S11c).

Open cracks in shocked samples are induced by tensile failure when compressive stress is reflected at the free surface of a stainless container and becomes tensile stress. The lengths of open cracks in shocked ALH-78084 H3 and Y-793375 L3 chondrites increase with increasing shock pressure (Fig. 3d). However, the linear relationship between shock pressure and lengths is not clear. It is difficult to adopt the relationship for the estimation of shock pressure.

#### 441 **4.3 Estimation of shock pressure by aspect ratios of chondrules**

442 The shock pressure recorded in each H/L/LL3 ordinary chondrite is 443 estimated using by adopting the aspect ratios of chondrules (median) into the 444 empirical formulae obtained from shock experiments to estimate shock pressure 445 (Tables S2-4). The empirical formulae obtained from ALH-78084 H3 and Y-793375 446 L3 chondrites are adopted for H3 and L/LL3 ordinary chondrites, respectively. 447 Figures 6-8 show box plots showing the aspect ratios of chondrules in each H/L/LL3 448 ordinary chondrite and estimated shock pressure. The ordinary chondrites in each 449 figure are sorted according to the aspect ratios.

450 No difference is found in estimated shock pressure between H3 ordinary 451 chondrites with a shock-induced melt (Gray-colored box) or without a shock-induced 452 melt (White-colored melt) (Fig. 6). The estimated shock pressure of Y-981139 H3 453 chondrite with a shock-induced melt is 4.8 GPa (Table S2). The estimated shock 454 pressure of H3 ordinary chondrites without shock-induced melts is up to about 10 455 GPa. Several estimated shock pressure values are below zero, which is impossible.

456 By contrast, the estimated shock pressure of L/LL3 ordinary chondrites with 457 shock-induced melts is higher than those of L/LL3 ordinary chondrites without 458 shock-induced melts (Tables S3-4, Figs. 7-8). The estimated shock pressure of L3 459 ordinary chondrites without shock-induced melts is less than 7.8 GPa except for 460 A-881244 L3 chondrite. The estimated shock pressure of most L3 ordinary 461 chondrites with shock-induced melts is higher than 7.8 GPa. The chondrules of 462 Y-981327 L3 chondrite have a very high aspect ratio (1.75) and the estimated shock 463 pressure is also very high (48.2 GPa)(Table S3). Y-981327 L3 chondrite is heavily 464 shocked because it has the extensive shock-induced melts (Figs. S10b-c).

465 No obvious correlation is found between estimated shock pressure and 466 high-pressure polymorphs (Tables S3-4). No wonder that there is no correlation 467 among shock pressure estimated from the aspect ratios of chondrules, shock-induced 468 melts, and high-pressure polymorphs. The formations of high-pressure polymorphs 469 are controlled by kinetics. Even if both high-pressure and -temperature conditions 470 required for the formation of high-pressure polymorphs are achieved, high-pressure 471 polymorphs cannot form if their duration is not enough. It is also possible that 472 high-pressure polymorphs, which formed during compression and equilibrium stages 473 once, back-transforms into their low-pressure polymorphs during the subsequent 474 adiabatic decompression stage if the cooling is slow.

475 Some estimated shock pressure values are invalid, which is due to the 476 deviation of aspect ratios of intact chondrules that are free from shock 477 metamorphism. The mean aspect ratio of intact chondrules in LL3 chondrites is  $1.2 \pm$  478 0.18 (n = 719) (Nelson & Rubin, 2002), which coincides with the aspect ratios of 479 chondrules (and chondritic fragments) in ALH-78084 H3 and Y-793375 L3 480 chondrites as starting materials (Table 2). Considering the standard deviation, in the 481 case of less shocked ordinary chondrites, some shock pressure values estimated from 482 the aspect ratios of chondrules becomes invalid. In contrast, some estimated shock 483 pressure values become overestimated.

484 The preferred orientations of long axes of chondrules become distinct in 485 highly shocked L/LL3 chondrites, especially including shock-induced melts (Tables 486 S2-4, Fig. S12). The Gini coefficients increase with increasing shock pressure in 487 L/LL3 chondrites. The L/LL3 chondrites including shock-induced melts have higher 488 Gini coefficients than those without melts. The degrees of preferred orientations in 489 L/LL3 chondrites are higher compared to those of experimentally shocked chondrites. 490 The variance is due to the differences in the duration of compression between shock 491 experiments and natural impacts although all impacts in nature do not have long 492 compression time.

493 To conclude, it is difficult to use the aspect ratios of chondrules for the 494 quantitative estimation of shock pressure in petrologic type 3 ordinary chondrites. 495 Repeated low-velocity impacts may increase the aspect ratios of chondrules 496 gradually, which is proved by the shock experiments of Allende CV3 chondrite 497 (Nakamura et al., 2000). The aspect ratios of chondrules in pre-heated Allende CV3 498 chondrite become bigger than those of chondrules at room temperature. Ordinary chondrite parent-bodies had a heart source such as <sup>26</sup>Al when they were born. If 499 impacts occur on hot parent-bodies, the flattening degree of chondrules becomes 500 501 bigger than that of chondrules on cold parent-bodies. Although chondrules with a 502 high aspect ratio and preferred orientation become a screening standard for strong 503 shock metamorphism, it does not work as a quantitative shock pressure barometer.

504

#### 505 5. Conclusions

- Shock experiments using ALH-78084 H3 and Y-793375 L3 chondrites prove that shock-induced melting occurs in petrologic type 3 chondrites beyond 11
   GPa at least. The melting occurs at boundaries between chondrules and matrices.
- 509 2) Shock-induced flattening of chondrules occurs in petrologic type 3 chondrites.
  510 The flattening degree increases with increasing shock pressure.
- 511 3) There is no significant difference in the flattening degree of chondrules between
  512 carbonaceous and ordinary chondrites. The flattening degree does not depend
  513 significantly on the densities, porosities, and chondrule/matrix ratios of
  514 chondrites. Finally, by considering the relationships between the aspect ratios of

- 515 chondrules and shock pressure in carbonaceous and ordinary chondrites, the 516 aspect ratios in unequilibrated chondrites ( $R_{cho}$ ) can be expressed following the 517 linear equation:  $R_{cho} = 0.011 (\pm 1) \times Pressure (GPa) + 1.18 (\pm 3)$ .
- 518 4) The long axes of chondrules have a preferred orientation and the degree519 increases with increasing shock pressure.
- 5) Natural L/LL3 ordinary chondrites with shock-induced melts have higher aspect
   ratios and preferred orientations of chondrules than those without shock-induced
   melts.
- 523 6) It is difficult to determine quantitatively shock pressure in petrologic type 3
  524 chondrites using the empirical formula between the aspect ratios of chondrules
  525 and shock pressure.
- 526

#### 527 Acknowledgments

This study was supported by Grants-in-Aid for Scientific Research, no. 18H01269 from the Ministry of Education, Culture Sports, Science, and Technology (MEXT) to M.M. This study was also financially supported by NIPR through General Collaboration Projects nos. 26-31 and KP-307. We acknowledge Dr. M. Kayama for providing the design drawing of a sample container used in this shock experiment. The measured aspect ratios and azimuth of chondrules are available in Miyahara (2021).

535

#### 536 References

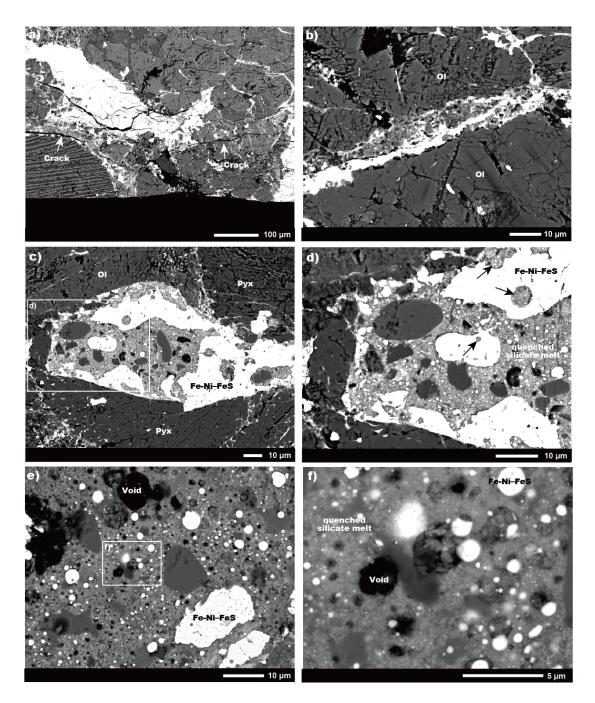
- Beck P., Gillet Ph., El Goresy A., & Mostefaoui S. (2005). Timescales of shock
  processes in chondritic and martian meteorites. *Nature* 435, 1071–1074.
- 539 Bland P.A., Collins G.S., Davison T.M., Abreu N.M., Ciesla F.J., Muxworthy A.R., &
- Moore J. (2014). Pressure-temperature evolution of primordial solar system solids
  during impact-induced compaction. *Nature Communications* 5, 5451 doi:
- 542 10.1038/ncomms6451.
- 543 Bottke Jr. W.F., Nolan M.C., Greenberg R., & Kolvoord R.A. (1994). Velocity
  544 distributions among colliding asteroids. *Icarus* 107, 255–268.
- 545 Chen M., Sharp T.G., El Goresy A., Wopenka B., & Xie X. (1996). The
  546 majorite-pyrope + magnesiowüstite assemblage: Constraints on the history of
  547 shock veins in chondrites. *Science* 271, 1570–1573.
- 548 Consolmagno G.J., Brittb D.T., & Macke R.J. (2008). The significance of meteorite
  549 density and porosity. *Chemie der Erde* 68, 1–29.
- 550 Dodd Jr. R.T. (1965). Preferred orientation of chondrules in chondrites. *Icarus* 4,
  551 308–316.

- Ebel D.S., Brunner C., Konrad K., Leftwich K., Erb I., Lu M., Rodriguez H., Ellen J.
  Crapster-Pregont E.J., Friedrich J.M., & Weisberg M.K. (2016). Abundance,
  major element composition and size of components and matrix in CV, CO and
  Acfer 094 chondrites. *Geochimica et Cosmochimica Acta* 172, 322–356.
- Friedrich J.M., Weisberg M.K., Ebel D.S., Biltz A.E., Corbett B.M., Iotzov I.V.,
  Khan W.S., & Wolman M.D. (2015). Chondrule size and related physical
  properties: A compilation and evaluation of current data across all meteorite
  groups. *Chemie der Erde* 75, 419–443.
- Gattaccecaa J., Rochette P., Denise M., Consolmagno G., & Folco L. (2005). An
  impact origin for the foliation of chondrites. *Earth and Planetary Science Letters*234, 351–368.
- Hirata N., Kurita K., & Sekine T. (2009). Simulation experiments for shocked
  primitive materials in the Solar System. *Physics of the Earth and Planetary Interiors* 174, 227–241.
- Kato T. & Ringwood A.E. (1989). Melting relationships in the system Fe-FeO at high
  pressures: Implications for the composition and formation of the earth's core. *Physics and Chemistory of Minerals* 16, 524–538.
- Kondo K.I., Sawaoka A., & Saito S. (1977). Magneto flyer method for measuring
  gas-gun projectile velocities. *Review of Scientific Instruments* 48, 1581–1582.
- Kong M., Bhattacharya R.N., James C., & Basu A. (2005). A statistical approach to
  estimate the 3D size distribution of spheres from 2D size distributions. *Geological Society of America Bulletin* 117, 244–249.
- Kohout T., Petrova E.V., Yakovlev G.A., Grokhovsky V.I., Penttilä A., Maturilli A.,
  Moreau J.G., Berzin S.V., Wasiljeff J., Danilenko I.A., Zamyatin D.A.,
  Muftakhetdinova R.F., & Heikkilä M. (2020). Experimental constraints on the
  ordinary chondrite shock darkening caused by asteroid collisions. Astronomy &
- 578 Astrophysics, doi: https://doi.org/10.1051/0004-6361/202037593.
- 579 Martin P.M. & Mills A.A. (1980). Preferred chondrule orientations in meteorites.
  580 *Earth and Planetaey Science Letters* 51, 18–25.
- Macke R.J., Consolmagno G.J., & Britt D.T. (2011). Density, porosity, and magnetic
  susceptibility of carbonaceous chondrites. *Meteoritics and Planetary Science* 46,
  1842–1862.
- McSween Jr. H.Y. (1977). Petrographic variations among the carbonaceous
  chondrites of the Vigarano type. *Geochimica et Cosmochimica Acta* 41, 1777–
  1790.
- 587 McSween Jr. H.Y. (1979). Alteration in CM carbonaceous chondrites inferred from
  588 modal and chemical variations in matrix. *Geochimica et Cosmochimica Acta* 43,

589 1761–1765.

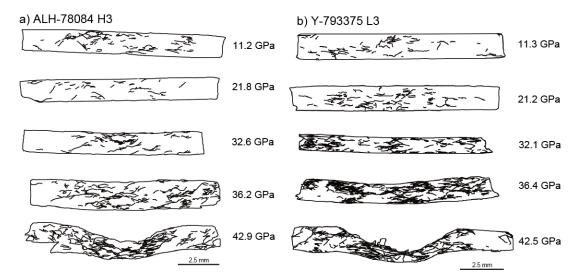
- Miyahara M., Yamaguchi A., Saitoh M., Fukimoto K., Sakai T., Ohfuji H., Tomioka
  N., Kodama Y., & Ohtani E. (2020). Systematic investigations of high-pressure
  polymorphs in shocked ordinary chondrites. *Meteoritics and Planetary Science*.
  doi: 10.1111/maps.13608.
- 594 Miyahara M. (2021) Aspect ratios and azimuth of chondrules in ordinary chondrites.
   595 doi.org/10.6084/m9.figshare.14060870
- Nakamura T., Tomeoka K., Takaoka N., Sekine T., & Takeda H. (2000).
  Impact-induced textural changes of CV carbonaceous chondrites: Experimental reproduction. *Icarus* 146, 289–300.
- Nelson V. & Rubin A.E. (2002). Size-frequency distributions of chondrules and
  chondrule fragments in LL3 chondrites: Implications for parent-body
  fragmentation of chondrules. *Meteoritics and Planetary Science* 37, 1361–1376.
- Ohtani E., Kimura Y., Kimura M., Takata T., Kondo T., & Kubo T. (2004). Formation
  of high-pressure minerals in shocked L6 chondrite Yamato 791384: Constraints on
  shock conditions and parent body size. *Earth and Planetary Science Letters* 227,
- 605 505-515.
- Ruzicka A.M., Hutson M., Friedrich J.M., Bland P.A., & Pugh R. (2015a). Northwest
  Africa 8709: A rare but revealing type 3 ordinary chondrite melt breccia. 78<sup>th</sup> *Annual Meeting of the Meteoritical Society*, 5348.pdf.
- Ruzicka A.M., Grossman J., Bouvier A., Herd C.D.K., & Agee C.B. (2015b). *The Meteoritical Bulletin*, No. 101.
- 611 Sneyd D.S., McSween H.Y., Sugiura N., Strangway D.W., & Nord G.L. (1988).
  612 Origin of petrofabrics and magnetic anisotropy in ordinary chondrites. *Meteoritics*613 23, 139–149.
- 614 Stöffler D., Keil K., & Scott E.R.D. (1991). Shock metamorphism of ordinary
  615 chondrites. *Geochimica et Cosmochimica Acta* 55, 3845–3867.
- Tanabata T., Shibaya T., Hori K., Ebana K., & Yano M. (2012). SmartGrain:
  High-throughput phenotyping software for measuring seed shape through image
  analysis. *Plant Physiology* 160, 1871-1880.
- Tomeoka K., Yamahana Y., & Sekine T. (1999). Experimental shock metamorphism
  of the Murchison CM carbonaceous chondrite. *Geochimica et Cosmochimica Acta*63, 3683–3703.
- Trieloff M., Jessberger E.K., Herrwerth I., Hopp J., Fiéni C., Ghélis M.,
  Bourot-Denise M., & Pellas P. (2003). Structure and thermal history of the
  H-chondrite parent asteroid revealed by thermochronometry. *Nature* 422, 502–
  506.

- Weisberg M.K., McCoy T.J., & Krot A.N. (2006). Systematic and evolution of
  meteorite classification. In Meteorites and the early solar system II (e.d. by
  Lauretta and McSween), The university of Arizona Press, Tucson, pp19–52.
- Xie Z., Sharp T. G., & DeCarli P.S. (2006). High-pressure phases in a shock-induced
   melt vein of the Tenham L6 chondrite: Constraints on shock pressure and duration.
- 631 *Geochimica et Cosmochimica Acta* 70, 504–515.
- 632 Yamaguchi A. & Sekine T. (2000). Monomineralic mobilization of plagioclase by
- 633 shock: an experimental study. *Earth and Planetary Science Letters*175, 289–296.
- 634



636 Figure 1. BSE images showing shocked ALH-78084 H3 chondrite. a) Open cracks 637 sub-parallel to a shockwave front (H\_1\_3). b) Planar or irregular cracks occurred in 638 some olivine grains (H\_2\_2), c) Melt between chondrules (H\_3\_3), d) 639 High-magnification image of a white-colored box in c). Mineral fragments entrained 640 into the melt were rounded. Metals include spherules of quenched silicate melt 641 (indicated by dark-colored arrows). e) Melt consisting of quenched silicate melts, 642 metals, mineral fragments, and voids (H\_4\_3), f) High-magnification image of a 643 white-colored box in e). Ol: olivine, Pyx: pyroxene, Fe-Ni: metallic Fe-Ni, FeS: iron-sulfide. 644

645



646

647 Figure 2. Representative drawings of shocked samples. a) ALH-78084 H3 and b)

648 Y-793375 L3 chondrites. The upper direction is the impact surface.

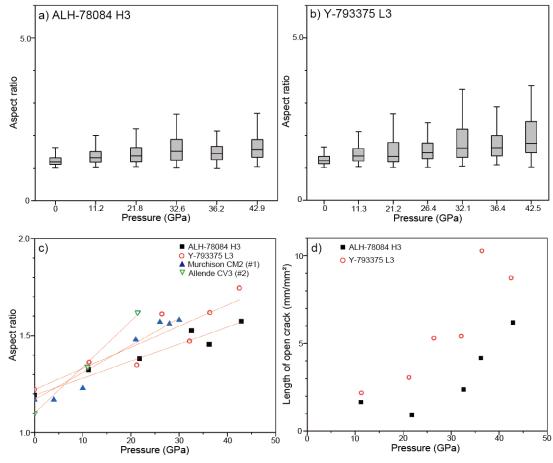
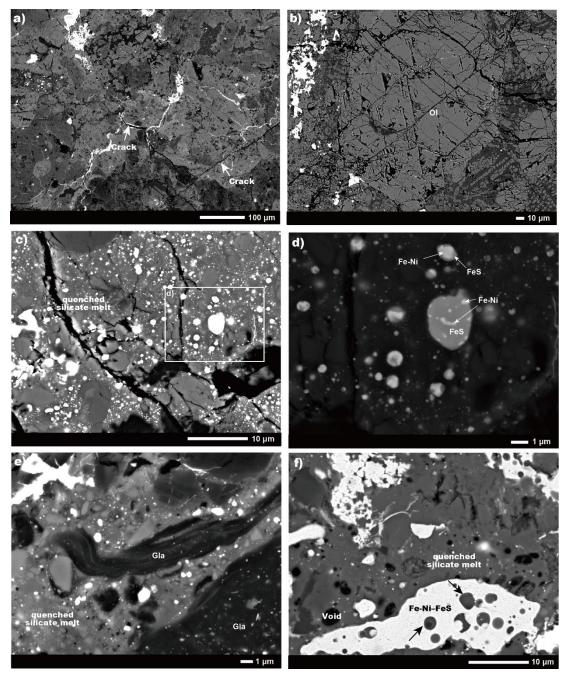


Figure 3. The morphological properties of chondrules. a) Relationship between
shock pressure and the aspect ratios of chondrules. b) Relationship between shock
pressure and the lengths of open cracks. #1 Tomeoka et al. (1999), #2 Nakamura et al.
(2000)





**Figure 4**. BSE images showing shocked Y-793375 L3 chondrite. a) Open cracks sub-parallel to a shockwave front  $(L_1_2)$ , b) Planar- and irregular-fractures in some olivine grains  $(L_2_3)$ , c) Melt between chondrules  $(L_2_3)$ , d) High-magnification image of a white-colored box in c). The spherules of metallic Fe-Ni–FeS with a eutectic texture. e) Melted chondrule glasses in the melt  $(L_6_1)$ . f) Metals including the spherules of quenched silicate melts  $(L_5_3)$ . OI: olivine, Pyx: pyroxene, Fe-Ni: metallic Fe-Ni, FeS: iron-sulfide, Gla: chondrule glass.

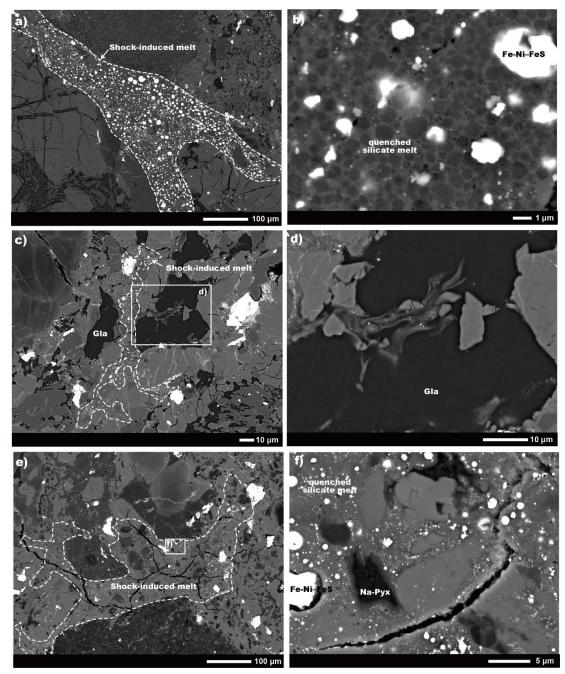
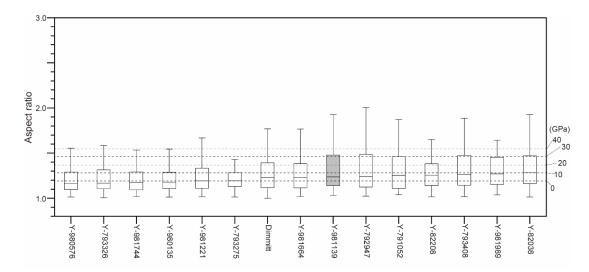




Figure 5. BSE images showing shock-induced melts in petrologic type 3 ordinary
chondrites. a) Pocket type melt in the Aba Panu L3 chondrite. b) High-magnification
image of shock-induced melt in a), c) Pocket type melt in A-87170 L3 chondrite, d)
High-magnification image of a white-colored box in c). A flow-like texture appears
in chondrule glasses, e) Pocket type melt in NWA 8664 L3 chondrite. f)
High-magnification image of a white-colored box in e). Na-pyx: Na-pyroxene.
Fe-Ni: metallic Fe-Ni, FeS: iron-sulfide.



675 Figure 6. Box plots showing the aspect ratios of individual H3 ordinary chondrites. 676 The H3 ordinary chondrites are sorted according to the aspect ratios. White-colored 677 and gray-colored boxes are H3 ordinary chondrites without melts and those with 678 melts, respectively. Dashed lines correspond to each shock pressure estimated from 679 the shock experiments of ALH-78084 H3 chondrite.

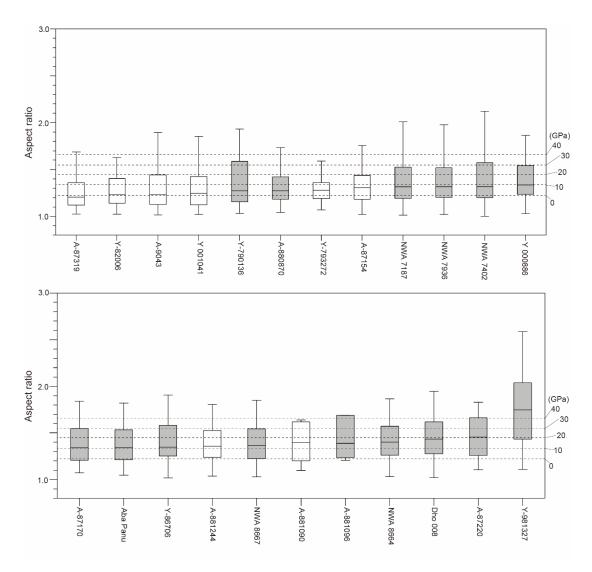
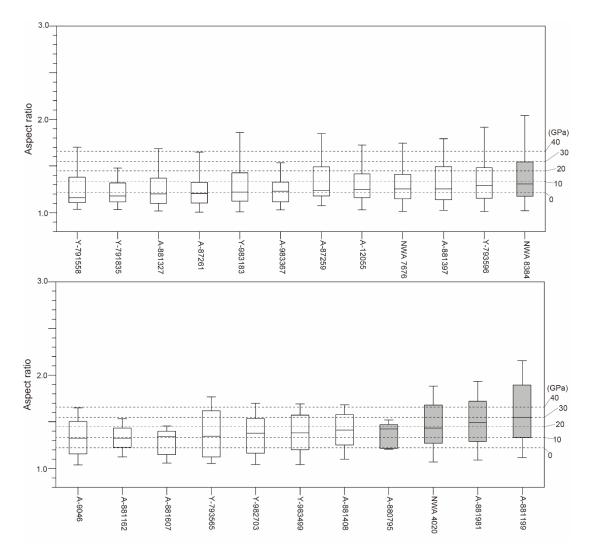




Figure 7. Box plots showing the aspect ratios of individual L3 ordinary chondrites. The L3 ordinary chondrites are sorted according to the aspect ratios. White-colored and gray-colored boxes are L3 ordinary chondrites without melts and those with melts, respectively. Dashed lines correspond to each shock pressure estimated from the shock experiments of Y-793375 L3 chondrite.



688

Figure 8. Box plots showing the aspect ratios of individual LL3 chondrites. The LL3 ordinary chondrites are sorted according to the aspect ratios. White-colored and gray-colored boxes are LL3 ordinary chondrites without melts and those with melts, respectively. Dashed lines correspond to each shock pressure estimated from the shock experiments of Y-793375 L3 chondrite.

|              |     | Target    |             |                   | F  | Flyer    |                   |                              |                                  |                           |  |
|--------------|-----|-----------|-------------|-------------------|--|----------|-------------------|------------------------------|----------------------------------|---------------------------|--|
| Shot No.     |     | φ<br>(mm) | Mass<br>(g) | Thickness<br>(mm) | Initial<br>density<br>(g/cm <sup>3</sup> ) | Material | Thickness<br>(mm) | Impact<br>velocity<br>(km/s) | Equilibrium<br>pressure<br>(GPa) | Sample recovery           |  |
| ALH-78084 H3 | H_1 | 12        | 0.554       | 1.51              | 3.24                                       | SUS      | 2                 | 0.57                         | 11.2                             | Three slices              |  |
|              | H_2 | 12        | 0.543       | 1.49              | 3.18                                       | SUS      | 2                 | 1.03                         | 21.8                             | Three slices              |  |
|              | H_3 | 12        | 0.563       | 1.52              | 3.30                                       | SUS      | 2                 | 1.46                         | 32.6                             | Three slices              |  |
|              | H_4 | 12        | 0.564       | 1.50              | 3.30                                       | SUS      | 2                 | 1.59                         | 36.2                             | Three slices              |  |
|              | H_5 | 12        | 0.544       | 1.49              | 3.19                                       | -        | -                 | -                            | -                                | Broken during preparation |  |
| A            | H_6 | 12        | 0.551       | 1.50              | 3.23                                       | W        | 2                 | 1.40                         | 42.9                             | Three slices              |  |
|              | H_7 | 6         | -           | 1.50              | -  | SUS      | 2                 | 1.17                         | 25.1                             | One disc                  |  |
|              | L_1 | 12        | 0.536       | 1.49              | 3.14                                       | SUS      | 2                 | 0.57                         | 11.3                             | Three slices              |  |
| L3           | L_2 | 12        | 0.542       | 1.49              | 3.18                                       | SUS      | 2                 | 1.01                         | 21.2                             | Three slices              |  |
| Y-793375 I   | L_3 | 12        | 0.536       | 1.52              | 3.14                                       | SUS      | 2                 | 1.44                         | 32.1                             | Three slices              |  |
|              | L_4 | 12        | 0.580       | 1.49              | 3.40                                       | SUS      | 2                 | 1.22                         | 26.4                             | Three slices              |  |
|              | L_5 | 12        | 0.577       | 1.47              | 3.38                                       | SUS      | 2                 | 1.60                         | 36.4                             | Three slices              |  |
|              | L_6 | 12        | 0.589       | 1.49              | 3.45                                       | W        | 2                 | 1.38                         | 42.5                             | Three slices              |  |

## 694 Table 1. Run table for shock experiments.

695 SUS: SUS 304, W: Tungsten

696

| S            | hot No.           | Shock pressure | Long axis <sup>#2</sup> | Short axis <sup>#2</sup><br>(μm) | Aspect<br>ratio <sup>#2</sup> | Number of analyses | Length of open cracks | Gini<br>coefficients |
|--------------|-------------------|----------------|-------------------------|----------------------------------|-------------------------------|--------------------|-----------------------|----------------------|
| 5            | 101110.           | (GPa)          | (µm)                    |                                  |                               | п                  | (mm/mm <sup>2</sup> ) | coefficients         |
|              | H_0#1             | 0              | 429                     | 352                              | 1.19                          | 258                | -                     | 0.33                 |
| g            | H_1               | 11.2           | 384                     | 290                              | 1.32                          | 197                | 1.65                  | 0.39                 |
| 84 H         | H_2               | 21.8           | 414                     | 273                              | 1.38                          | 162                | 0.93                  | 0.46                 |
| ALH-78084 H3 | H_3               | 32.6           | 427                     | 283                              | 1.53                          | 161                | 2.38                  | 0.40                 |
| LH-          | H_4               | 36.2           | 389                     | 254                              | 1.46                          | 187                | 4.17                  | 0.41                 |
| A            | H_6               | 42.9           | 445                     | 259                              | 1.57                          | 153                | 6.18                  | 0.44                 |
|              | H_7               | 25.1           | -                       | -                                | -                             | -                  | -                     | -                    |
|              | L_0 <sup>#1</sup> | 0              | 773                     | 634                              | 1.22                          | 143                | -                     | 0.32                 |
|              | L_1               | 11.3           | 707                     | 506                              | 1.36                          | 89                 | 2.19                  | 0.25                 |
| 5 L3         | L_2               | 21.2           | 634                     | 393                              | 1.35                          | 115                | 3.07                  | 0.23                 |
| 337:         | L_3               | 32.1           | 749                     | 482                              | 1.47                          | 79                 | 5.42                  | 0.27                 |
| Y-793375 L3  | L_4               | 26.4           | 1296                    | 770                              | 1.61                          | 36                 | 5.31                  | 0.31                 |
|              | L_5               | 36.4           | 821                     | 507                              | 1.62                          | 83                 | 10.28                 | 0.31                 |
|              | L_6               | 42.5           | 711                     | 426                              | 1.75                          | 77                 | 8.74                  | 0.49                 |

| 698 | Table 2. | Data | table for | recovered | samples. |
|-----|----------|------|-----------|-----------|----------|
|-----|----------|------|-----------|-----------|----------|

699 #1 Shot No. H\_0 and L\_0 are starting materials. #2 median values. H\_5 is removed because it is lost. -: No data