Atmospheric Gravity Waves in Aeolus Wind Lidar Observations

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Abstract

Aeolus is the first Doppler wind lidar in space. It provides unique high-resolution measurements of horizontal wind in the sparsely-observed upper-troposphere/lower-stratosphere (UTLS), with global coverage. In this study, Aeolus' ability to resolve atmospheric gravity waves (GWs) is demonstrated. The accurate representation of these small-scale waves is vital to properly simulate dynamics in global weather and climate models. In a case study over the Andes, Aeolus GW measurements show coherent phase structure from the surface to the lower stratosphere, with wind perturbations >10 m/s, a vertical wavelength 8 km and an along-track horizontal wavelength 900 km. Good agreement is found between Aeolus and colocated satellite, ground-based lidar and reanalysis data sets for this example. Our results show that data from satellites of this type can provide unique information on GW sources and propagation in the UTLS, filling a key knowledge gap that underlies known major deficiencies in weather and climate modelling.

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Key Points:

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10	•	First satellite observations of atmospheric gravity waves using ADM-Aeolus
11	•	A case study is presented of an orographic gravity wave over the Southern Andes,

- with coherent phase structure down to the surface
- Results reproduce well in satellite observations and reanalysis data

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14 Abstract

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²⁸ Plain Language Summary

Gravity waves are an important driver of the global atmospheric circulation but 29 are difficult to observe due to their scale size and location. Existing satellite observations 30 reveal these waves in temperature perturbations, but tend to be limited in either ver-31 tical or horizontal resolution. Since they are arguably best described in a wind-based math-32 ematical framework and due to their influential behaviour in the upper-troposphere lower-33 stratosphere region, an observing platform that satisfies both of these requirements could 34 prove very significant. This study explores the capability of the first Doppler wind li-35 dar in space, Aeolus, to measure gravity waves and provide unique information about 36 their sources and propagation through the atmosphere. Significantly, Aeolus measures 37 wind speed directly and is well suited to observe the upper-troposphere lower-stratosphere 38 region. Here, a case study is presented showing observations of a strong gravity wave pro-39 duced by the enhanced orography of the Southern Andes, which are the most prominent 40 hotspot of gravity wave activity globally. Results are validated against two other obser-41 vational instruments and atmospheric reanalysis, and give confidence in Aeolus' ability 42 to measure these phenomena. 43

44 **1** Introduction

Atmospheric gravity waves (GWs) are small-scale propagating disturbances that 45 arise due to the vertical forcing of air parcels by a disturbance in the flow. They are gen-46 erated by a variety of meteorological processes, including flow over orography, atmospheric 47 deep convection, and jet stream, cyclonic and frontal instabilities. GWs play a wide range 48 of key roles in the atmospheric system, particularly in the transfer of energy and mo-49 mentum (e.g. Fritts & Alexander, 2003). They are responsible for driving the large-scale 50 circulation in the middle atmosphere, primarily through accelerations to the mean-flow 51 by the convergence of GW momentum flux (Fritts, 1984). They also modulate phenom-52 ena such as the Quasi-Biennial Oscillation (Dunkerton, 1997; Ern et al., 2014), affect stratosphere-53 mesosphere coupling during Sudden Stratospheric Warmings (Siskind et al., 2007; Wright 54 et al., 2010), and can produce turbulence when they dissipate that is dangerous to air-55 craft (Lilly, 1978; Bramberger et al., 2018). 56

Both climate and numerical weather prediction (NWP) models rely on accurately representing the propagation of GWs from sources near the surface to the upper-troposphere/lowerstratosphere (UTLS) region. This accuracy is needed to correctly simulate important features of the atmosphere such as the strength of the northern and southern hemispheric jet streams (Holton, 1983; McFarlane, 1987; M. Alexander et al., 2010; Sato et al., 2012; Ehard et al., 2017). As such, limited understanding of the drag forces GWs exert on the mean winds has historically proven to be a significant barrier to advances in NWP, and improvements in GW simulation have often led to model-wide improvements across a range of processes and scales (Palmer et al., 1986; Eichinger et al., 2020). As model resolutions continue to improve, particularly in the vertical dimension, capturing GW processes is becoming an increasingly important problem, and this trend is likely to continue in the near future (Jones et al., 1997; Kim et al., 2003; Watanabe et al., 2015).

Most GWs are currently parameterised in general circulation models (GCMs) (M. Alexan-69 der & Barnet, 2007; Geller et al., 2013). However, in practice these parameterisations 70 are not well constrained observationally, due to both the relatively small spatial scales 71 of most GWs and because the GW field varies dramatically across a large range of spa-72 tiotemporal frequencies (M. J. Alexander, 1998; M. Alexander et al., 2010). This has long 73 been identified as an important knowledge gap, and since the first empirical measure-74 ments of GWs (Hines, 1960), observations have been made using a wide variety of meth-75 ods. 76

In-situ observations such as those from radiosondes, aircraft and meteorological rock-77 ets often form an anchor point against which other measurements can be validated (B. Sun 78 et al., 2010; Krisch et al., 2017). Satellite-based observations of GWs are provided by 79 nadir-sounders such as AIRS and AMSU, as well as limb-sounders such as HIRDLS and 80 the COSMIC GPS constellation. These instruments typically measure infra-red radiances, 81 from which wind perturbations must be inferred using a set of GW dispersion relations 82 (Ern et al., 2004; Hindley et al., 2015; Wright, Hindley, & Mitchell, 2016; Wright, Hind-83 ley, Moss, & Mitchell, 2016). Finally, there are ground-based radars and lidars, which 84 provide good temporal coverage at a reasonable vertical resolution (N. Kaifler et al., 2020). 85 These however are fixed to one location and cannot provide a global climatology of GW 86 activity by themselves. 87

Radars and lidars in particular provide direct measurements of the wind pertur-88 bations induced by GWs (e.g. Larsen et al., 1982; Vincent & Reid, 1983), which is im-89 portant because GWs are arguably best-described at a theoretical level in a wind-based 90 mathematical framework. However, the existing inability to systematically measure winds 91 from space means that relatively few GW measurements have been made using wind per-92 turbations directly other than at these radar and lidar sites, and none to date on a global 93 domain. In 2018 however, the first spaceborne wind lidar instrument was launched aboard 94 the European Space Agencys Aeolus mission. This novel ability to systematically mea-95 sure winds from space offers the potential to significantly advance our understanding of 96 how GWs propagate in the real atmosphere, and in turn to advance weather and climate 97 modelling. 98

To demonstrate the benefits of using Aeolus and its proposed successors as a plat-99 form for systematic GW observations, presented here is a case study using Aeolus data 100 to examine the structure of a large GW observed in winter (July) 2019 over the Andes 101 mountains. The Andes are a fantastic natural laboratory for observing GWs due to the 102 ridge at their southern end which is transverse to the prevailing westerly winds. Glob-103 ally, the Andes are by far the most prominent hotspot of GW activity, and the strong 104 orographic forcing often present can produce waves of large magnitude which propagate 105 significant distances into the middle and upper atmosphere (Jiang et al., 2002; M. Alexan-106 der & Teitelbaum, 2011; Hoffmann et al., 2013). Additionally, Aeolus' high inclination 107 orbit is oriented approximately parallel to the southern Andean ridge line, measuring a 108 horizontal line-of-sight (HLOS) wind that is close to zonal and approximately transverse 109 to the mountains at this latitude. Detection conditions here are therefore well-suited for 110 a study of this type, which will be generalised to more complicated cases at the global 111 scale in future work. 112

Section 2 describes the data sources and methodology for this study, with an outline of both Aeolus and the other observing systems that are used to validate these measurements. Section 3 shows the first example of GWs measured in Aeolus data, and uses
 other observations and ERA5 reanalysis to validate these results. Section 4 discusses the
 limitations of the methods used and summarises the key points from this study.

¹¹⁸ 2 Data and Methods

The purpose of this case study is to determine whether Aeolus is a suitable platform for observing GWs, and to give a first suggestion of the possibilities it presents for wider GW studies. First, a large GW event on the 26th July 2019 is established as a good candidate using carefully selected along-track vertical profiles from Aeolus. The data is detrended using a band-pass filter to extract wind perturbations, and the coinciding meteorological and geographical context is assessed to determine if the observed GW structure is plausible.

Validation of this GW event in the Aeolus observations is then carried out using data from the CORAL lidar in Tierra del Fuego, the AIRS instrument onboard the Aqua satellite, and output from the ERA5 reanalysis. For the comparison between each, an Aeolus profile which provides a good demonstration of the broader GW signature is empirically selected, and then co-location profiles from CORAL, AIRS and ERA5 are found nearby in time and space. Differences between the datasets mean slightly different procedures are required to extract GW perturbations for each, as described below.

2.1 Aeolus

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Aeolus is the first satellite with a space-borne wind lidar instrument onboard (ESA, 134 1989, 2008; Chanin et al., 1989; Stoffelen et al., 2005; Reitebuch, 2012). This instrument, 135 known as the Atmospheric Laser Doppler Instrument (ALADIN), probes the lowermost 136 30 km of the atmosphere and provides high vertical resolution profiles of wind, aerosol 137 and cloud along its orbital path. The satellite has a sun-sychronous orbit with 15.6 or-138 bits per day and a repeat cycle of 7 days. The orbit's inclination is 96.97° and its mean 139 altitude is 320 km, with an ascending-node local equator-crossing time of 18:00. Both 140 the laser and telescope are directed at 35° off-nadir, perpendicular to the direction of 141 travel. 142

A single wind component v_{LOS} is measured along this line-of-sight (LOS), which is converted into the HLOS wind speed v_{HLOS} by assuming the vertical wind speed wis small. Equations (1) and (2) show how these parameters relate to the three cartesian wind components u, v and w; where θ is the elevation of the target-to-satellite pointing vector (55°) and Ψ is the bearing of the satellite track.

$$v_{\rm LOS} = v_{\rm HLOS}\cos(\theta) + w\sin(\theta) \tag{1}$$

$$v_{\rm HLOS} = -u\sin(\Psi) - v\cos(\Psi) \tag{2}$$

ALADIN measures backscattering from atmospheric molecules (Rayleigh scatter-148 ing), and aerosol and hydrometeors (Mie scattering) in the path of light from its laser, 149 which is operated at a wavelength of 355 nm. The backscattered light is received using 150 a 55 kg telescope which is 1.5 m in diameter, and the Doppler shift of this signal rela-151 tive to the laser pulse frequency is recorded. These data are processed to produce a mea-152 surement of the HLOS wind speed for both Mie and Rayleigh scattering, throughout the 153 depth of the atmospheric profile (Rennie et al., 2020). The vertical levels are split into 154 range bins which are between 250-2000m in depth and can be chosen and altered from 155 the ground, with the instrument arranging data into these bins according to the time 156 difference between the laser pulse being transmitted and the return signal being received. 157

Aeolus data used in this study are aggregated into profiles according to grouping identifiers within each file (De Kloe et al., 2020). We use the L2B product; for this prod-

uct, each observation is categorised as occurring in either clear or cloudy conditions, and 160 an associated HLOS error estimate is also provided for quality control. Only Rayleigh 161 clear observations are used in this study to maximise data coverage in combination with 162 data quality. An 8 ms^{-1} cut off is used as a filter on the random error for each data point. 163 The data used here has been reprocessed using baseline B10 of ESA's L2B processor. A 164 number of campaigns have validated the HLOS winds from Aeolus, with the latest es-165 timates for this reprocessed data showing a systematic bias of $<1 \text{ ms}^{-1}$ for the Rayleigh 166 wind product (Abdalla et al., 2020). Since this study is primarily concerned with wind 167 perturbations, any systematic biases in the data are considered to have a negligible im-168 pact on the results. 169

To plot along-track profiles, the data are linearly interpolated onto a 500m verti-170 cal grid. Wind perturbations are then calculated by running an along-track Savitzky-171 Golay bandpass filter through the data points, with lower and upper bounds of 7 and 172 25 profiles respectively. GWs with along-track horizontal wavelengths of ~ 6002000 km 173 are detectable using this method. This detrending method is chosen because it best high-174 lights the orientation of the GW for this particular case study. Unlike at higher altitudes, 175 there is no clear scale separation between GWs and other atmospheric features; such as 176 jet streaks, synoptic-scale Rossby waves and various mesoscale phenomena (Perlwitz & 177 Graf, 2001; C. Sun et al., 2014). In this example, the low cutoff point should remove smaller-178 scale perturbations which might not be related to GW activity, however caution ought 179 to be applied if using this detrending method in the general case. 180

2.2 CORAL lidar

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The Compact Rayleigh Autonomous Lidar (CORAL) instrument is an autonomous 182 ground-based lidar system designed to provide temperature and density profiles of the 183 middle atmosphere (B. Kaifler & Kaifler, 2020). Situated at Tierra del Fuego on the south-184 ern tip of South America (54° S, 68° W), it is positioned at a prime location for measur-185 ing strong orographic GW activity (N. Kaifler et al., 2020). The altitude range covered 186 by CORAL measurements extends from 15 - 90 km, and measurements have a 900 m 187 vertical resolution oversampled onto a 90 m grid, with a temporal resolution of 20 min. 188 CORAL measures backscattered photons detected in three Rayleigh channels (532 nm) 189 and one Raman channel (608 nm), in clear-sky conditions only. In this study, the Ra-190 man channel is used for altitudes below 31 km. 191

As discussed by Ehard et al. (2015), estimating temperature perturbations using 192 a lidar such as CORAL can be challenging, especially where there are sudden changes 193 in the vertical temperature gradient, such as at the stratopause. Using temporal filter-194 ing would alleviate this issue, however due to the short observational periods on this par-195 ticular day, such a method is not possible for this case study. Instead, to obtain temper-196 ature perturbations a vertical Savitzky-Golay high-pass filter is run with a cut-off wave-197 length of 20 km. This is in line with previous studies (e.g. B. Kaifler et al., 2015; N. Kai-198 fler et al., 2020) and should be sufficient to observe the important characteristics of this 199 GW. 200

2.3 AIRS

The Atmospheric Infrared Sounder (AIRS) is a 2378-channel infrared nadir sounder 202 onboard NASA's Aqua satellite (Aumann et al., 2003; Chahine et al., 2006). Part of the 203 A-Train satellite constellation, Aqua is in a sun-synchronous orbit with 14.55 orbits per 204 day, a repeat cycle of 16 days, an orbital inclination of 98.20° and an ascending-node lo-205 cal equator-crossing time of 13:30. The AIRS instrument scans across track in a $\pm 49.5^{\circ}$ 206 wide swath, measuring radiances. The horizontal resolution varies from ~ 13.5 km $\times 13.5$ 207 km at nadir to 41 km \times 21.4 km at the track edge, and measurement data are stored 208 in "granules", each corresponding to 6 minutes of data (Wright et al., 2017). 209

In order to provide enough context for the Aeolus overpass, one AIRS granule be-210 fore and one after the Aeolus overpass are chosen, each with good spatial coverage of the 211 Southern Andes. The retrieval outlined by Hoffmann and Alexander (2009) has been used 212 to estimate air temperature on a 3 km vertical grid from 18–55 km altitude, and results 213 are shown at the 30 km altitude level to avoid vertical edge-truncation effects and as a 214 representative snapshot. Temperature perturbations have been calculated by running 215 a cross-track 4th-order polynomial fit through the data as demonstrated in M. Alexan-216 der and Barnet (2007). GWs with cross-track horizontal wavelengths of \sim 50–1000 km 217 are detectable using this method (e.g. Hoffmann et al., 2014; Ern et al., 2016). 218

$2.4 \quad \text{ERA5}$

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ERA5 is a reanalysis dataset provided by the ECMWF which combines an Earth-220 system model with assimilated observations to provide a historical archive of the state 221 of the atmosphere (Hersbach et al., 2020). The dataset has a spatial resolution of $0.25 \,^{\circ} \times 0.25 \,^{\circ}$ 222 $(\sim 31 \text{ km})$, 137 vertical levels from the surface up to 0.1 hPa, and a temporal resolution 223 of one hour. Since Aeolus observations are not assimilated into ERA5, the two datasets 224 are entirely independent from each other. As the Aeolus overpass studied here occurs 225 at almost exactly 10:00, this is also the time chosen for the ERA5 vertical profile. In or-226 der to directly compare data from the Aeolus overpass with ERA5, the latter is inter-227 polated bi-linearly onto each Aeolus measurement point in space. This is done by pro-228 jecting the ERA5 u and v wind components onto the Aeolus comparable HLOS, as shown 229 in equation (2). The data are then further interpolated onto the same 500 m vertical grid 230 as done previously, so that the same horizontal Savitzky-Golay filtering approach can 231 be used. For the multi-dataset validation of the GW, the single nearest ERA5 profile to 232 the chosen Aeolus profile is selected rather than interpolating to the measurement lo-233 cation, in order to avoid any aliasing caused by averaging between data points. 234

Topographical context is provided using $0.25 \,^{\circ} \times 0.25 \,^{\circ}$ resolution elevation data from the TBASE archive provided by NCAR, and the elevation for each profile is calculated using a bi-linear interpolation of this dataset. Mean Sea-Level Pressure (MSLP) and cloud fraction layers for the meteorological context are also provided from hourly ERA5 data at a $0.25 \,^{\circ} \times 0.25 \,^{\circ}$ resolution.

240 3 Results

Figure 1 demonstrates clearly for the first time that GWs can be observed using 241 space-borne wind lidar instruments. The figure shows winds measured during an Aeo-242 lus pass over the Andes mountain range on the 26th July 2019. The meteorological con-243 text is a deep depression to the west of the Drake Passage, which drives a strong west-244 erly wind over the raised topography of the Southern Andes. This pattern causes sig-245 nificant surface wind stresses, which typically translate to strong upward-propagating 246 orographic GWs; such a source mechanism is consistent with the characteristics of the 247 wave observed here. Aeolus is on the descending node of its orbit at this time, and trav-248 els down the length of South America before intersecting the mountain range and head-249 ing out over the southern Pacific Ocean. The cloud fraction overlay from ERA5 at the same time shows clear skies along much of the satellite ground-track, providing confi-251 dence that it is appropriate to use the Rayleigh-clear HLOS wind product. 252

The wave itself can be seen clearly in the HLOS wind field between 10:00 and 10:02 in Figure 1b as absolute measured values and in Figure 1c as perturbations to the background wind field. It coincides geographically with a region of raised topography and appears to be propagating upwards into the stratosphere with a vertical wavelength of around 8 km and an amplitude of around 10 m/s. Also of note is the polar night jet, which can be seen in the lower stratosphere south of 50°S. The Savitzky-Golay filtering emfigures/Aeolus_figure_v4.pdf

Figure 1. Aeolus overpass of 2019-07-26 with a) a map showing its geographical context, overlayed with ERA5 data on a single-level for cloud fraction and mean sea-level pressure, and terrain data from the TerrainBase Global Terrain Model; b) raw winds in an along-track time-series of the L2B HLOS Rayleigh wind speed product, with colour range set by the mean and two standard deviations of the domain; c) wind perturbations calculated using a 5 - 25 data point Savitzky-Golay band pass filter. Cloudy or missing data points are obscured with a grey mask. The profile used for the validation of the GW in Aeolus is marked with a black arrow, and corresponds to 2019-07-26 10:00:34.

figures/Co-location_figure_v4.pdf

Figure 2. a) Timeseries of kinetic temperature perturbations from the CORAL lidar at Tierra del Fuego for 2019-07-26 at 15 minute resolution. b,c) Data granules 55 and 190 from the AIRS instrument for 05:32 and 19:02 respectively and at an altitude of 30 km. d) 3 time snapshots comparing data from Aeolus with CORAL and AIRS observations and ERA5 reanalysis. AIRS has been multiplied by 0.5 for ease of comparison.

phasises the GW in the wind perturbations in Figure 1c, showing diagonally oriented wave
 fronts which are most pronounced during the same time-frame.

Validation of the GW seen in the Aeolus data has been carried out using the CORAL
lidar, AIRS, and ERA5 reanalysis, and is shown in Figure 2. Distinct quasi-stationary
wave-fronts can be seen in the CORAL lidar timeseries (Figure 2a), with a vertical wavelength of around 10 km and an amplitude of around 15 K peaking near the stratopause.
AIRS overpasses from before and after the Aeolus pass also show very pronounced GW
structures emanating eastwards from the Southern Andes (Figure 2b-c).

Figure 2d shows a comparison between a selected profile from the 10am Aeolus overpass and co-located profiles from each observing system, further contextualised with the nearest ERA5 profile in both temperature and projected HLOS wind. All measurements are taken from geographically close to the Aeolus profile location, with the exception of
the CORAL lidar, which is some 600 km to its south-east. Since the plot shows a combination of temperature and wind measurements, it is important to note that according to theory there is an expected phase difference of 90° between these two fundamental variables.

Between 5-6am strong temperature perturbations can be seen in both CORAL and 275 AIRS, increasing in amplitude with height. The vertical positions of the AIRS temper-276 ature peaks match well with CORAL at 25 and 37 km, but appear to be in anti-phase 277 around 50 km. The likely reason for this is the poor vertical resolution of AIRS which 278 aliases the wave; this issue is explored further in Section 4. ERA5 at the same time also 279 shows perturbations in both temperature and wind, particularly below 20 km in the wind 280 profile. Once again there are phase differences with AIRS which are likely a result of alias-281 ing, whereas the differences with CORAL are more likely to be a consequence of the 600 282 km distance between the two profiles. 283

At 10am there is very good agreement between the Aeolus and ERA5 wind pro-284 files, particularly in wave amplitude, although with a slight vertical phase offset which 285 increases gradually with height. The large temperature perturbations in the CORAL pro-286 file confirm the expectation of strong GW activity coinciding with the Aeolus overpass. 287 The relative change in amplitude at each height between 5-6am and 10am is consistent 288 with ERA5. The reanalysis however does not quite capture the amplitude that CORAL 289 does, which could be a consequence of either the wave dissipating too quickly in the re-290 analysis model, or the stratopause interfering with the filtering process for CORAL. 291

At 7pm, only the AIRS observations are available for comparison, and this partic-292 ular profile shows the wave structure from 5-6am largely persisting through the day un-293 til the 7pm overpass. It is difficult to compare with ERA5 due to the significant deficien-294 cies in both datasets, especially at higher altitudes, however the most notable observa-295 tion is the presence of a persistent large amplitude GW in each. Throughout the mea-296 surement period, the amplitude of the wave in AIRS is significant larger than in the other 297 datasets. The reasons for this are unclear but may involve either inaccurate model physics 298 or our choice of detrending approach. 299

Figure 3 shows ERA5 data projected onto the HLOS wind points observed by Aeolus. The middle and right panels show striking morphological agreement with Figures 1b and 1c. As in the observations, we see a strong orographic GW propagating upwards into the stratosphere above the sharp topography of the Andean ridge. The vertical wavelength is of a similar magnitude to that seen in the Aeolus observations, although importantly the amplitude is not as high as in the Aeolus data, suggesting the model may have a tendency to underestimate the amplitude of such waves as well as exhibiting the phase differences described above.

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4 Discussion and Conclusions

The primary limitation for measuring GWs using Aeolus is the complex dynam-309 ics of the troposphere and the challenges this presents when detrending the observed data 310 to identify wave perturbations. Unlike in the stratosphere and above, where there is a 311 clear spectral difference between GW perturbations and planetary wave activity or other 312 non-GW related phenomena, in the troposphere a host of meteorological processes can 313 interfere with the analysis. While we have addressed this problem in this case study by 314 empirically selecting a filter length which highlights well the wave of interest, solving this 315 problem in the general case is a larger technical challenge, and will require significant 316 research before our results can be generalised to the global scale (e.g. Rapp et al., 2018). 317

A further limitation is that Aeolus only measures in one wind direction and only provides data along its flight track. Thus, no information about the orientation or real figures/ERA5_figure_v4.pdf

Figure 3. ERA5 validation of the Aeolus overpass of 2019-07-26 with a) a map showing its geographical context, overlayed with ERA5 data on a single-level for cloud fraction and mean sea-level pressure, and terrain data from the TerrainBase Global Terrain Model; b) ERA5 HLOS projected winds to match with the Aeolus along-track profile, with colour range set by the mean and two standard deviations of the domain; c) wind perturbations calculated using a 5 - 25 data point Savitzky-Golay band pass filter. Cloudy or missing data points are obscured with a grey mask. The profile used for the validation of the GW in Aeolus is marked with a black arrow, and corresponds to 2019-07-26 10:00:34.

horizontal wavelength of the measured GWs can be inferred. Instead, only the projection of the horizontal wavelength along the satellite track can be determined, which usually gives an overestimation of the real horizontal wavelength. In the case of orographic GWs over the Andes, this poses a strong restriction on the information Aeolus can provide about their horizontal structure; especially when the satellite bearing is near-parallel to the mountain range as is the case in this study.

Additionally, the HLOS wind calculation of the Aeolus L2B processor assumes a 326 vertical wind of 0 ms^{-1} . This is unproblematic under normal conditions, however where 327 there is strong GW activity this assumption is less justified due to the strong vertical 328 motions that can be present. Since for all structures observed by Aeolus the horizontal 329 domain dominates, regions where the real vertical wind w is large tend to be too small 330 to be resolved. For the GW presented in this paper, the error in the HLOS wind is de-331 termined to be below 2%; a figure that can be calculated using the gravity wave polar-332 isation relations (Fritts & Alexander, 2003). Further to this, the orientation of the HLOS 333 winds is not near-zonal close to the pole, leading to a strong bias of the measured waves 334 with respect to latitude. This again may add additional technical difficulties for any global 335 studies of GW activity using data from Aeolus. 336

Finally, there remain the general problems of filtering for and spectrally analysing 337 GWs in observational data, a problem extensively discussed by e.g. Preusse et al. (2008); 338 Wright, Hindley, Moss, and Mitchell (2016); Strube et al. (2020); Krisch et al. (2020), 339 among others. The data here have been interpolated onto a regular grid in order to carry 340 out the Savitzky-Golay filtering, a process which itself will tend to smooth the peaks of 341 each wave and reduce their amplitude. Furthermore, the transmission of the filter used 342 is inherently imperfect within the wavelength window analysed; this is a general prob-343 lem for spectral analysis, but one which adds to the uncertainties in such work. Here, 344 as in Hindley et al. (2015), the Savitzky-Golay filter is selected as a trade-off between 345 a desirably sharp transition at each end and Gibbs ringing at the discontinuity, but it 346 will be important to assess how well alternative filters perform in more general future 347 work. 348

Nonetheless, the strong morphological and quantitative agreement in wave prop-349 erties between the HLOS wind profiles from Aeolus and the temperature based profiles 350 from CORAL and AIRS leads to a high level of confidence in Aeolus' ability to observe 351 GWs. A clear phase structure is visible from near the surface up to the stratosphere, with 352 the CORAL lidar supplementing this higher up and AIRS providing information about 353 its geographical orientation. Good qualitative agreement is found between these wind and temperature measurements, suggesting that good phase relationship is observed, even 355 if there is sometimes a phase offset from one location to the next. Our results demon-356 strate the benefit of these spaceborne wind lidar measurements for GW studies, which 357 can be used to better constrain GW parameterisations in models and improve our un-358 derstanding of small-scale GW processes. 359

360 Data Availability Statement

Aeolus data were provided by the European Space Agency, and can be accessed via https://aeolus-ds.eo.esa.int/oads/access/. The AIRS data were provided by NASA; L1 radiance data can be acquired via https://disc.gsfc.nasa.gov/, and were retrieved to L2 temperatures using the method described by Hoffmann and Alexander (2009). ERA5 data can be accessed from the Copernicus Climate Data Store, https://cds.climate.copernicus.eu/. CORAL Lidar data are not routinely archived publically, but the data used for this study have been archived at https://halo-db.pa.op.dlr.de/dataset/7620.

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Figure 1.



Figure 2.



Figure 3.

