# Online calculation of $O_2$ clumped-isotope variations in an atmospheric chemistry model reveals an important contribution from ozone isotopologue chemistry

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#### Abstract

Tropospheric <sup>18</sup>O<sup>18</sup>O is an emerging proxy for past tropospheric ozone and free-tropospheric temperatures. The basis of these applications is the idea that isotope-exchange reactions in the atmosphere drive <sup>18</sup>O<sup>18</sup>O abundances toward isotopic equilibrium. However, previous work used an offline box-model framework to explain the <sup>18</sup>O<sup>18</sup>O budget, approximating the interplay of atmospheric chemistry and transport. This approach, while convenient, has poorly characterized uncertainties. To investigate these uncertainties, and to broaden the applicability of the <sup>18</sup>O<sup>18</sup>O proxy, we developed a scheme to simulate atmospheric <sup>18</sup>O<sup>18</sup>O abundances (quantified as [?]<sub>36</sub> values) online within the GEOS-Chem chemical transport model. These results are compared to both new and previously published atmospheric observations from the surface to 33 km. Simulations using a simplified O<sub>2</sub> isotopic equilibration scheme within GEOS-Chem show quantitative agreement with measurements only in the middle stratosphere; modeled [?]<sub>36</sub> values are too high elsewhere. Investigations using a comprehensive model of the O-O<sub>2</sub>-O<sub>3</sub> isotopic photochemical system and proof-of-principle experiments suggest that the simple equilibration scheme omits an important pressure dependence to [?]<sub>36</sub> values: the anomalously efficient titration of <sup>18</sup>O<sup>18</sup>O to form ozone. Incorporating these effects into the online [?]<sub>36</sub> calculation scheme in GEOS-Chem yields quantitative agreement for all available observations. While this previously unidentified bias affects the atmospheric budget of <sup>18</sup>O<sup>18</sup>O in O<sub>2</sub>, the modeled change in the mean tropospheric [?]<sub>36</sub> value since 1850 C.E. is only slightly altered; it is still quantitatively consistent with the ice-core [?]<sub>36</sub> record, implying that the tropospheric ozone burden likely increased less than <sup>~40</sup>% over the twentieth century.

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#### Effects of ozone isotopologue formation on the clumped-isotope composition 1 2 of atmospheric O<sub>2</sub>

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- 14 Abstract

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- 16 temperatures. The basis of these applications is the idea that isotope-exchange reactions in the atmosphere
- 17 drive <sup>18</sup>O<sup>18</sup>O abundances toward isotopic equilibrium. However, previous work used an offline box-model
- framework to explain the <sup>18</sup>O<sup>18</sup>O budget, approximating the interplay of atmospheric chemistry and 18 transport. This approach, while convenient, has poorly characterized uncertainties. To investigate these 19
- uncertainties, and to broaden the applicability of the <sup>18</sup>O<sup>18</sup>O proxy, we developed a scheme to simulate 20
- atmospheric <sup>18</sup>O<sup>18</sup>O abundances (quantified as  $\Delta_{36}$  values) online within the GEOS-Chem chemical 21
- transport model. These results are compared to both new and previously published atmospheric 22
- 23 observations from the surface to 33 km. Simulations using a simplified O<sub>2</sub> isotopic equilibration scheme
- 24 within GEOS-Chem show quantitative agreement with measurements only in the middle stratosphere;
- 25 modeled  $\Delta_{36}$  values are too high elsewhere. Investigations using a comprehensive model of the O-O<sub>2</sub>-O<sub>3</sub>
- 26 isotopic photochemical system and proof-of-principle experiments suggest that the simple equilibration
- scheme omits an important pressure dependence to  $\Delta_{36}$  values: the anomalously efficient titration of 27
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- affects the atmospheric budget of <sup>18</sup>O<sup>18</sup>O in O<sub>2</sub>, the modeled change in the mean tropospheric  $\Delta_{36}$  value
- 30 31 since 1850 C.E. is only slightly altered; it is still quantitatively consistent with the ice-core  $\Delta_{36}$  record,
- 32 implying that the tropospheric ozone burden increased less than ~40% over the twentieth century.
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- 34 9937 Words (main text)
- 35 591 Words (figure captions)
- 36 **10 Figures**
- 37 1 Table
- 38

# 39 Plain Language Summary

- 40 Oxygen in the air is constantly being broken apart and remade. Its constituent atoms are shuffled
- 41 around by light-induced chemical reactions, which cause changes in the number of heavy oxygen atoms
- 42 that are bound together. The number of these heavy-atom "clumps" is sensitive to air temperatures and
- 43 the presence of air pollution; hence, their variations are being used to understand past high-altitude
- temperatures and atmospheric chemistry. This study incorporates oxygen clumping into an atmospheric
- 45 chemistry model and compares the results to measurements of oxygen clumping in the atmosphere. We
- 46 find that the model can explain all the modern-day measurements (from the surface to 33 km altitude), but
- 47 only if the broader fates of oxygen atoms—i.e., their incorporation into other molecules beyond  $O_2$ —are
- 48 considered. Simulations of the preindustrial atmosphere are also generally consistent with snapshots of
- 49 the ancient atmosphere obtained from O<sub>2</sub> trapped in ice cores. The developments described herein will 50 thus enable models to simulate heavy oxygen-atom clumping in past cold and warm climates and enable
- 50 thus enable models to simulate heavy oxygen-atom clumping in past cold and warm climates and enable 51 simulated high-altitude atmospheric changes to be evaluated directly against ice-core snapshots of the
- 52 ancient atmosphere.

# 54 1. Introduction

55 Odd oxygen  $[O(^{3}P), O(^{1}D), \text{ and } O_{3}]$  is a key component of the atmosphere's oxidizing capacity. 56 As such, tracing its evolution over time may provide better constraints on greenhouse-gas lifetimes, stratosphere-troposphere coupling, biosphere-atmosphere interactions, and radiative forcing in the past. 57 58 Moreover, elevated odd-oxygen concentrations in the upper troposphere and stratosphere mean that a 59 globally integrated record of odd-oxygen chemistry would be a unique window on the high-altitude 60 atmosphere of the past, not just in terms of chemistry but also climate [Thompson et al., 1998; Tripati et al., 2014; Yeung et al., 2016; Loomis et al., 2017]. Observational constraints on temperatures of the 61 62 ancient high-altitude atmosphere, for example, would bring new insights on Earth's past energy balance. 63 The short lifetime of odd-oxygen species, however, precludes their direct preservation in 64 traditional atmospheric archives. Interrogating the past variability of atmospheric odd oxygen thus requires the use of proxies and models [Alexander et al., 2003; Murray et al., 2014; Alexander and 65 Mickley, 2015; Geng et al., 2017; Yeung et al., 2019]. For example, hydrogen peroxide is a stable product 66 67 of odd-oxygen chemistry (via HO<sub>2</sub> self-reaction) that is preserved in ice cores, but its sensitivity to post-68 depositional alteration limits the scenarios under which its variations reflect atmospheric odd-oxygen changes [Frey et al., 2006]. Another proxy is the <sup>15</sup>N/<sup>14</sup>N ratio of snow-bound nitrate, which is sensitive to 69 70 the local ultraviolet flux. It can offer insight on the local atmospheric column density of  $O_3$  if the local 71 accumulation rate and optical depth in the snow are known [Frey et al., 2009; Ming et al., 2020; Winton 72 et al., 2020]; however, anthropogenic changes to nitrogen cycling can overprint the these signals 73 [Hastings et al., 2009; Geng et al., 2014]. A third proxy is based on the transfer of unique O<sub>3</sub>-derived <sup>17</sup>O 74 enrichments, primarily to nitrate and sulfate species. It can provide constraints on the relative importance 75 of O<sub>3</sub>, OH, and HO<sub>2</sub> oxidation pathways, but interpretations are often model-dependent [Alexander et al., 76 2009; Kunasek et al., 2010; Sofen et al., 2011; Sofen et al., 2014; Geng et al., 2017]. Furthermore, all 77 these approaches are only sensitive to local odd-oxygen variability; global trends are not accessible. 78 Recently, Yeung et al. [2019] applied a new proxy for odd-oxygen chemistry preserved in the icecore record that offers a broader view. It is based on the proportional abundance of <sup>18</sup>O<sup>18</sup>O in tropospheric 79  $O_2$  (quantified as  $\Delta_{36}$  values; see Section 2), which is linked to odd-oxygen chemistry through  $O(^{3}P) + O_2$ 80 81 isotope-exchange reactions. The subtle decrease in tropospheric  $\Delta_{36}$  value occurring over the twentieth 82 century was interpreted to reflect a change in odd-oxygen chemistry, namely, an increase in the 83 tropospheric  $O_3$  burden. The inferred magnitude of the increase (<40%) is in broad agreement with 84 predictions made by state-of-the-art atmospheric chemistry models [Yeung et al., 2019]. 85 Temporal variations in tropospheric  $\Delta_{36}$  values can be understood within a two-box framework,

with separate tropospheric and stratospheric boxes [*Yeung et al.*, 2016]. Isotope-exchange reactions occur within each box, and mass is exchanged between them. Tropospheric chemistry leads to lower  $\Delta_{36}$  values

- than stratospheric chemistry because of the difference in characteristic temperatures:  $O(^{3}P) + O_{2}$  isotope
- 89 exchange at warmer temperatures generally yields smaller <sup>18</sup>O<sup>18</sup>O enrichments, and thus lower  $\Delta_{36}$  values,
- 90 than isotope exchange at cold temperatures [*Yeung et al.*, 2012]. The tropospheric signal is well mixed on
- 91 annual timescales, incorporating changes in stratospheric inputs on sub-decadal timescales [Yeung et al.,
- 2016]. Tropospheric  $\Delta_{36}$  values therefore reflect the cumulative effects of isotope-exchange chemistry
- 93 occurring in the troposphere ( $\sim$ 3/4 of the signal) modulated by the input of high- $\Delta_{36}$  stratospheric air ( $\sim$ 1/4
- of the signal). An increase in the rate of  $O(^{3}P) + O_{2}$  reactions in the troposphere, e.g., caused by increase
- 95 in mean O(<sup>3</sup>*P*) concentration, leads to a decrease in tropospheric  $\Delta_{36}$  values at steady state. Because the
- 96 mean tropospheric  $O({}^{3}P)$  concentration is proportional to the  $O_{3}$  burden (and has a negligible dependence
- 97 on other atomic oxygen sources), tropospheric  $\Delta_{36}$  values have been used as a proxy for past changes in 98 the tropospheric  $O_3$ . A warming of the troposphere would also lower the mean tropospheric  $\Delta_{36}$  value by
- 99 lowering the characteristic tropospheric  $\Delta_{36}$  values toward which isotope-exchange reactions drive O<sub>2</sub>.
- $\Delta_{35}$  values toward which isotope exchange reactions drive o
- 100 Yet, important details of the global  $\Delta_{36}$  budget remain uncertain. Previous quantitative estimates 101 of atmospheric  $\Delta_{36}$  variations were facilitated by global 3-D chemical transport model simulations of 102 O(<sup>3</sup>*P*) concentrations, but the  $\Delta_{36}$  calculations were still performed offline, i.e., without direct coupling of 103 atmospheric chemistry, mixing, and transport [*Yeung et al.*, 2016; *Yeung et al.*, 2019]. The expense of 104 isotopologue-specific chemistry—e.g., the O-O<sub>2</sub>-O<sub>3</sub> photochemical system itself requires >300 additional, 105 highly stiff reactions to represent—renders explicit online calculation of <sup>18</sup>O<sup>18</sup>O variations impractical.
- 106 A simplified treatment of the relevant  $O_2$  isotopologue photochemistry is possible, however, 107 because laboratory studies indicate that the approach of  $\Delta_{36}$  values toward photochemical steady state can 108 be described by pseudo-first-order kinetics [Yeung et al., 2014]. Isotope-exchange reactions between 109  $O(^{3}P)$  and  $O_{2}$  at low pressures (e.g., 1-2 mbar) drive isotopologue abundances toward isotopic 110 equilibrium, but the uncertainty in their kinetics is large, particularly at the low temperatures relevant for the upper troposphere and stratosphere (Fig. 1, left). Moreover, isotope effects in  $O_3$  formation could be 111 important—particularly the anomalous enrichments in <sup>18</sup>O<sup>18</sup>O-containing O<sub>3</sub> isotopologues, which also 112 vary with temperature and pressure (Fig. 1, right) [Mauersberger, 1981; Heidenreich and Thiemens, 113 114 1983; Thiemens and Heidenreich, 1983; Guenther et al., 1999; Mauersberger et al., 1999; Gao and Marcus, 2001; Janssen et al., 2003]. While imperfect, parameterization of these effects may offer a 115 116 practical balance between accuracy and computational cost that allows  $\Delta_{36}$  variations to be simulated online. 117
- 118 In this manuscript, we use new high-precision  $\Delta_{36}$  measurements from the modern atmosphere 119 spanning a range of latitudes in the Northern Hemisphere to evaluate two different online  $\Delta_{36}$  calculation 120 schemes within the GEOS-Chem chemical transport model. The first scheme considers only O<sub>2</sub> isotopic 121 equilibration at *in situ* temperatures, using rates determined by local O(<sup>3</sup>*P*) concentrations and O(<sup>3</sup>*P*) + O<sub>2</sub>
  - 4

122 isotope-exchange rate coefficients, while the second scheme includes a temperature- and pressure-

123 dependent parameterization for the effects of O<sub>3</sub> formation. Furthermore, we compare the results of online

124 and offline  $\Delta_{36}$  calculation schemes to improve our understanding of the  $\Delta_{36}$  budget and interpretations of

- 125  $\Delta_{36}$  changes in the recent atmospheric record.
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# 127 **2. Methods**

# 128 2.1 Measurements of new upper-tropospheric and lower-stratospheric samples

129 Twenty-three whole-air samples were collected from the NASA ER-2 aircraft during the Studies 130 of Emissions, Atmospheric Composition, Clouds and Climate Coupling by Regional Surveys (SEAC<sup>4</sup>RS) 131 campaign [*Lueb et al.*, 1975; *Toon et al.*, 2016] and selected for O<sub>2</sub> isotopologue composition 132 measurements in duplicate. Flights associated with the campaign took place over the continental United 133 States in August and September of 2013. Samples were selected from five individual flights between 134 23.7°N and 42.6°N latitude and between 87.3°E and 99.5°E longitude, with potential temperatures 135 ranging from  $\theta$  = 328 K to  $\theta$  = 498 K.

Isotopologue ratios of O<sub>2</sub> were measured on a high-resolution Nu Instruments *Perspective IS*isotope ratio mass spectrometer after gas-chromatographic separation from Ar, N<sub>2</sub>, and other trace
components of air [*Yeung et al.*, 2016; *Yeung et al.*, 2018; *Ash et al.*, 2020]. These ratios are converted to
isotopic "delta" notation according to the following definitions:

140

141 
$$\delta^{18} O = \left(\frac{{}^{18}_{\Box} R_{sample}}{{}^{18}_{\Box} R_{air}} - 1\right) \delta^{17} O = \left(\frac{{}^{17}_{\Box} R_{sample}}{{}^{17}_{\Box} R_{air}} - 1\right)$$
(1)

142

143 
$$\Delta_{35} = \left(\frac{\stackrel{35}{\square}R_{sample}}{\stackrel{35}{\square}R_{stochastic}} - 1\right) \Delta_{36} = \left(\frac{\stackrel{36}{\square}R_{sample}}{\stackrel{36}{\square}R_{stochastic}} - 1\right)$$
(2)

144

145 Here,  ${}^{18}R$  and  ${}^{17}R$  are the bulk  ${}^{18}O/{}^{16}O$  and  ${}^{17}O/{}^{16}O$  ratios in O<sub>2</sub>, whereas  ${}^{36}R$  is the  ${}^{18}O/{}^{16}O/{}^{16}O$ 

146 isotopologue ratio. The "stochastic" subscript refers to the random distribution of isotopes within O<sub>2</sub>, i.e., 147  ${}^{35}R_{stochastic} = 2{}^{17}R_{sample} {}^{18}R_{sample}$  and  ${}^{36}R_{stochastic} = ({}^{18}R_{sample})^2$ . Air sampled from an external staircase on the third 148 floor of the Keith-Wiess Geology building at Rice University several times a week during analytical 149 sessions was used as the bulk-isotope standard against which  $\delta^{18}$ O and  $\delta^{17}$ O values are reported (i.e.,  ${}^{18}R_{air}$ 150 and  ${}^{17}R_{air}$ ). Reported  $\Delta_{36}$  values are calibrated against photochemical and heated-gas standards generated 151 at known temperatures according to established techniques [*Yeung et al.*, 2014]. The photochemical 152 standards are generated by submerging a 1 L bulb filled with 1 – 2 mbar of pure O<sub>2</sub> into a constant153 temperature bath held at 22°C, -25°C, or -78°C, allowing the bulb to reach thermal equilibrium, and then 154 photolyzing it with a mercury lamp through an N<sub>2</sub>-flushed internal quartz finger for 40 minutes. The 155 effective photolysis temperature reflects the radiative equilibrium between the bath and the finger, which is warmed slightly due to the presence of the lamp and  $N_2$  gas flush (i.e., yielding effective temperatures 156 157 of 25°C, –19°C, or –65°C for each calibration standard, respectively). High-temperature standards are 158 generated by heating barium peroxide in an evacuated quartz breakseal at 800°C for 1.5 – 3h, after which 159 the breakseal is quenched by submerging it in room-temperature water. 160 Samples were analyzed between 2017 and 2019, but only one sample showed a bulk isotopic

- composition more than 0.1‰ different from air O<sub>2</sub>, indicating good isotopic fidelity despite several years' storage. The mean  $\delta^{18}$ O deviations from air across the whole dataset were 0.019 ± 0.063‰ (1 $\sigma$ , *n* = 46 analyses). While omitting the high- $\delta^{18}$ O sample improves the precision across the whole dataset (i.e.,  $\delta^{18}$ O mean of 0.007 ± 0.036‰), such small positive  $\delta^{18}$ O deviations—likely arising from diffusive gas loss rather than a contaminant—are unlikely to affect  $\Delta_{36}$  values significantly. Therefore, all the SEAC<sup>4</sup>RS measurements were used. The pooled standard deviation of replicates was ±0.026‰, ±0.071‰ and ±0.045‰ in  $\delta^{18}$ O,  $\Delta_{35}$ , and  $\Delta_{36}$  values, respectively.
- 168

# 169 2.2 Atmospheric chemical transport modeling

170 Model simulations using GEOS-Chem (version 12.9.2; http://www.geos-chem.org) and the 171 unified (UCX) chemical mechanism [*Eastham et al.*, 2014] were performed using default global 172 emissions for present-day (PD) simulations of June 1998 – June 2000 and calendar year 2015 [McDuffie 173 et al., 2020]. The UCX mechanism treats stratospheric and tropospheric chemistry under a common scheme and includes explicit accounting of atomic oxygen species, which is necessary for online 174 175 calculation of  $\Delta_{36}$  values. The model used meteorology from the Modern Era Retrospective Reanalysis-2 176 (MERRA2) [Gelaro et al., 2017] with 72 vertical levels (~38 in the troposphere), degraded to a spatial 177 resolution of  $4^{\circ} \times 5^{\circ}$  (latitude × longitude). Preindustrial era (PI) simulations were also performed using 178 2015 meteorology with natural (non-anthropogenic) emissions set to those for 1850 C.E. In those 179 simulations, methane, nitrous oxide, and carbonyl sulfide were prescribed using surface mixing ratio 180 boundary conditions that yielded atmospheric concentrations consistent with ice-core constraints [Murray 181 et al., 2014]. Anthropogenic halocarbon concentrations were also prescribed (set to zero) using surface 182 boundary conditions. Biomass burning emissions for the PI are uncertain, with estimates of PI-to-PD 183 changes varying in both sign and magnitude (e.g., PI emissions of 75 - 400% those of the PD) [McConnell et al., 2007; Marlon et al., 2008; Lamarque et al., 2010; Wang et al., 2010; Bisiaux et al., 184 2012; Daniau et al., 2012; Pfeiffer et al., 2013; Nicewonger et al., 2018; Dyonisius et al., 2020]; recent 185 186 constraints on PI methane, ethane, and acetylene emissions from biomass burning suggest that they were

187 comparable to those of the PD, or perhaps slightly higher [*Pfeiffer et al.*, 2013; *Nicewonger et al.*, 2018;

188 *Dyonisius et al.*, 2020; *Nicewonger et al.*, 2020]. Consequently, biomass burning emissions were held 189 constant for PI simulations.

190 Two online calculation schemes for  $\Delta_{36}$  values were implemented in GEOS-Chem. Both were 191 simplified schemes designed to minimize the computational burden needed to represent  $O(^{3}P) + O_{2}$ isotope-exchange chemistry. In the first scheme (" $O_2$  only"),  $\Delta_{36}$  values decay toward isotopic equilibrium 192 193 according to simple first-order kinetics, which are dependent on temperature but independent of  $O_3$ 194 isotopologue chemistry. In the second scheme ("pressure-dependent"),  $\Delta_{36}$  values decay toward an 195 isotopic steady state rather than isotopic equilibrium. The steady state is determined by local temperature 196 and pressure due to changes in the rates of both isotope-exchange and  $O_3$  isotopologue formation 197 reactions. In both schemes, the temperature, atomic oxygen concentration ( $[O(^{3}P)]$ ), and O<sub>2</sub> concentration 198 in each grid box determine the local instantaneous rates of isotope exchange. The nominal first-order rate coefficient for the  $\Delta_{36}$  system was approximated by the temperature-dependent  ${}^{16}O({}^{3}P) + {}^{18}O{}^{18}O$  isotope-199 200 exchange rate coefficient reported by *Fleurat-Lessard et al.* [2003], i.e.,  $k_{exch}(T)$ , which is more precise 201 than that originally reported in *Wiegell et al.* [1997]. At each time-step, the concentration of the  $\Delta_{36}$  tracer 202 is relaxed toward its local equilibrium or steady-state value according to the isotope-exchange lifetime 203 and the length of the time-step. After the chemistry operation, the  $\Delta_{36}$  tracer is allowed to advect between 204 grid boxes and mix.

In the O<sub>2</sub>-only scheme, the local  $\Delta_{36}$  value relaxes toward its isotopic equilibrium value,  $\Delta_{36,equil}$ (see Text S1 for the formula used). The evolution of local  $\Delta_{36}$  value from time  $t_0$  to  $t_0 + \Delta t$  proceeds according to the equation:

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- 209

 $\Delta_{36,t_0+\Delta t} = \Delta_{36,t_0} + \left( \Delta_{36,equil} - \Delta_{36,t_0} \right) \times \left( 1 - e^{-\Delta t/\tau_{exch}} \right)$ (3)

210

Here,  $\tau_{exch} = 1/(k_{exch}(T)[O(^{3}P)])$  represents the local  $\Delta_{36}$  lifetime with respect to isotope exchange. In the pressure-dependent scheme, the local  $\Delta_{36}$  value relaxes toward the steady-state value  $\Delta_{36,ss}$  instead of  $\Delta_{36,equil}$ , with the other quantities remaining the same. The particular value of  $\Delta_{36,ss}$  was determined by polynomial fits to the results of an isotope-enabled photochemical model (i.e., a photochemical kinetics model that includes all isotopologue-specific rate coefficients and species; see Section 2.3).

The online  $\Delta_{36}$  outputs included full O<sub>2</sub>-only and pressure-dependent outputs (herein  $\Delta_{36T}$  and  $\Delta_{36P}$ , respectively), as well as diagnostic  $\Delta_{36}$  values for isotope exchange occurring only within the troposphere ( $\Delta_{36Tt}$  and  $\Delta_{36Pt}$ ) or only within the stratosphere ( $\Delta_{36Ts}$  and  $\Delta_{36Ps}$ ). These latter diagnostics were calculated by turning off the operation in eq. 3 if a grid box is above or below the local tropopause, respectively, during

- 220 a time-step. The tropopause is determined during each time-step using the World Meteorological
- 221 Organization lapse-rate definition (i.e.,  $\Gamma$  < 2 K/km).
- 222

# 223 **2.3 Isotope-enabled photochemical model to determine** $\Delta_{36,ss}$ **values**

224 The 321-reaction KINTECUS photochemical model for the O-O<sub>2</sub>-O<sub>3</sub> system from Yeung et al. 225 [2014] was updated to include temperature- and pressure-dependent isotope effects for O<sub>3</sub> formation in 226 the atmosphere [*Ianni*, 2003]. Nitrogen ( $N_2$ ) was added as a separate species to yield near-atmospheric 227 mixing ratios of  $N_2$  and  $O_2$  (i.e., 79%  $N_2$  and 21%  $O_2$ ); argon was not included for simplicity and because 228 isotope effects in  $O_3$  formation in argon appear to be similar to those of  $N_2$  and  $O_2$  [Morton et al., 1990; 229 Thiemens and Jackson, 1990; Feilberg et al., 2013]. The chemistry of nitrogen oxides was omitted because the goal was to simulate the effects of pressure on the O-O<sub>2</sub>-O<sub>3</sub> isotopic system only. 230 231 The temperature dependencies of  $^{18}$ O-containing O<sub>3</sub> formation rate coefficients were obtained 232 from Janssen et al. [2003], while the pressure dependencies for the relative rate coefficients of the  $^{16}O +$  ${}^{16}O^{18}O + M \rightarrow {}^{16}O^{16}O^{18}O + M$  and  ${}^{16}O + {}^{18}O^{18}O + M \rightarrow {}^{16}O^{18}O^{18}O + M$  reactions were obtained from 233

*Guenther et al.* [1999]. The temperature and pressure dependencies of these relative rate coefficients were combined using the following expression from *Guenther et al.* [1999] for  $O_3$  isotopologue enrichments:

237

$$E_{O_3}(T, P) = \frac{E_{O_3}(T)}{1 + P/P_{1/2}}$$
(4)

238

where  $E_{03}(T)$  is the temperature-dependent enrichment of an O<sub>3</sub> isotopologue relative to <sup>16</sup>O<sub>3</sub> at 300K/200 239 Torr obtained from Janssen et al. [2003] [e.g.,  $E_{O3}(T) = 0.47 + (T - 300 \text{ K}) \times 0.00015 \text{ for } {}^{16}\text{O} + {}^{18}\text{O}{}^{18}\text{O}$ ], 240 241 and  $P_{1/2}$  is the pressure at which the enrichment is halved, obtained from *Guenther et al.* [1999] (e.g.,  $P_{1/2}$ = 4100 Torr or 5466 mbar for  ${}^{16}\text{O} + {}^{18}\text{O}{}^{18}\text{O}$ ). The ratio of rate coefficients for the formation of  ${}^{16}\text{O}{}^{16}\text{O}{}^{18}\text{O}$ 242 and  ${}^{16}O^{18}O^{18}O$ , relative to  ${}^{16}O_3$ , is thus 1 +  $E_{O3}(T,P)$ . While the 300K/200 Torr anchor point yields zero-243 pressure  $E_{03}$  values that differ slightly from those used in *Guenther et al.* [1999], they agree within 244 245 several percent; the method used here yields slightly better agreement with the measurements reported in Janssen et al. [2003]. The ratio of rate coefficients for the  ${}^{18}O + {}^{16}O^{16}O + M \rightarrow {}^{16}O^{16}O^{18}O + M$  and  ${}^{18}O + M$ 246  ${}^{18}O^{18}O + M \rightarrow {}^{18}O^{18}O + M$  reactions has a negligible dependence on pressure [*Guenther et al.*, 1999], 247 248 so the 300K/200 Torr relative rate coefficients were used and kept constant [Mauersberger et al., 1999]. 249 The 300K/200 Torr isotopologue-specific relative rate coefficients measured by Mauersberger et al. 250 [1999] were also used and held constant wherever temperature- and pressure-dependent data were not 251 available (e.g., for <sup>17</sup>O-containing O<sub>3</sub> isotopologues). Because of these missing inputs,  $\Delta_{35}$  values 252 simulated by this model at elevated pressures are unlikely to be accurate; those results are not reported.

253 Where possible, the relative rate coefficients for  $O(^{3}P) + O_{2}$  isotope exchange reactions were 254 determined assuming microscopic reversibility at isotopic equilibrium for each temperature [Hathorn and 255 *Marcus*, 2000; *Wang et al.*, 2004; *Yeung et al.*, 2014], with their absolute rates ultimately tied to the  ${}^{16}O({}^{3}P) + {}^{18}O{}^{18}O$  isotope-exchange rate constant determined by *Fleurat-Lessard et al.* [2003], i.e.,  $k_{exch}(T)$ 256 257 =  $2.7 \pm 0.4 \times 10^{-12} (T/300)^{-0.9 \pm 0.5} \text{ cm}^3 \text{ s}^{-1} (2\sigma)$ . Seven <sup>17</sup>O-containing reactions of the 18 total isotopeexchange reactions were assigned nominal 300 K rate coefficients (i.e.,  $k_{exch} = 1.5 \times 10^{-12} \text{ cm}^3 \text{ s}^{-1}$  per 258 259 product channel) because their rate coefficients are not known. Previous work has shown that factor-of-260 two variations in these rate coefficients have a negligible effect on steady-state  $\Delta_{36}$  values [Yeung et al., 261 2014].

Isotope effects in other reactions (e.g., during  $O_3$  photolysis or chemical destruction) may also be important, but their cumulative effects on <sup>16</sup>O<sup>18</sup>O and <sup>18</sup>O<sup>18</sup>O-containing species in the atmosphere are poorly known. Owing to the uncertainties in these quantities [*Liang et al.*, 2006; *Früchtl et al.*, 2015a; *Früchtl et al.*, 2015b; *Huang et al.*, 2019], this model only explores the relationship between the isotope effects for O(<sup>3</sup>*P*) + O<sub>2</sub> isotope exchange and O<sub>3</sub> formation. We acknowledge that the results obtained are necessarily incomplete, although they are still informative and relevant to the global atmosphere.

268 Photochemical simulations were performed to generate an array of steady-state solutions between 269 175 K and 325 K and 0.3 mbar and 4000 mbar total pressure. The relative rates of  $O_2$  and  $O_3$  photolysis 270 were adjusted to yield ~50 ppb  $O_3$  at tropospheric pressures. Each model was run using the mean and  $2\sigma$ 271 uncertainty bounds for  $k_{exch}(T)$ . At low temperatures, the uncertainty in  $k_{exch}(T)$  is nearly a factor of two 272 (Fig. 1, left), which dominates the overall uncertainty in  $\Delta_{36,ss}$  values.

273 A parameterization for  $\Delta_{36.ss}$  as a function of temperature and pressure was made using two 274 polynomial fits. First, third-order polynomial fits for  $\Delta_{36,ss}$  vs. pressure were obtained at each temperature 275 (ten different pressures P at each of seven temperatures). The coefficients *a*, *b*, *c*, and *d* in  $\Delta_{36.ss} = aP^3 + aP^3$  $bP^2 + cP + d$  were then plotted as a function of temperature, and fifth-order polynomial fits of those trends 276 277 were obtained. The results are intended for interpolation only under atmospheric conditions (175 – 325 K, 278 0-1100 mbar) given the arbitrary functional form of the fits. The  $\Delta_{36.ss}$  value calculated from this two-279 component scheme was then substituted for  $\Delta_{36,equil}$  in eq. 3 to represent the local pressure-dependent  $\Delta_{36}$ 280 endmember for isotope exchange. This scheme, while simplified, reproduced the  $\Delta_{36,ss}$  values in the 281 photochemical model to within 0.01‰. It offers a great reduction in cost compared to representing 282 isotopologue-specific reactions in GEOS-Chem explicitly.

283

284 **3. Results** 

# 285 **3.1 Atmospheric observations of Δ**<sub>36</sub> of **O**<sub>2</sub>

286 Figure 2 shows the new  $\Delta_{36}$  measurements on the SEAC<sup>4</sup>RS samples, which were collected by the 287 NASA ER-2 aircraft, along with previous high-precision measurements made at Rice University on 288 samples from the SOLVE mission (NASA ER-2), the DC3 mission (NASA DC-8), and high-altitude 289 balloon flights over Fort Sumner, NM [Yeung et al., 2016]. The  $\Delta_{36}$  values from SEAC<sup>4</sup>RS samples 290 decrease from ~2.7‰ at  $\theta$  = 500K toward the tropospheric value of 1.99‰, which is reached at  $\theta$  = 360K 291 and below (Fig. 2A). A maximum value of 2.77‰ was observed in the lower stratosphere at 24°N ( $\theta$  = 292 444K). The new data show remarkable consistency with previously reported observations from the 293 midlatitude upper troposphere and stratosphere when plotted against potential temperature [Yeung et al., 294 2016]. All the SEAC<sup>4</sup>RS samples below  $\theta$  = 475 K were analytically more than 2 $\sigma$  different from isotopic 295 equilibrium, with deviations as large as -1.06% relative to  $\Delta_{36,equil}$  values implied by *in situ* sampling 296 temperatures (Figs. 2B and S1). The largest disequilibria were observed in the tropopause region (e.g.,  $\theta =$ 297 360K), and were larger at 24°N than at 40°N. Notably, in the midlatitudes between  $\theta$  = 400K and 500K, 298 measured  $\Delta_{36}$  values decrease just 0.1‰ with decreasing  $\theta$ , whereas  $\Delta_{36,equil}$  values increase by as much as 299 0.4‰. The  $\Delta_{36}$  disequilbrium there thus primarily reflects an increase in  $\Delta_{36,eauil}$  values. By contrast, the  $\Delta_{36}$ 300 disequilbrium in the lowermost stratosphere largely reflects the gradient in measured  $\Delta_{36}$  values rather 301 than the more weakly varying  $\Delta_{36,equil}$  values. In the troposphere, the  $\Delta_{36}$  disequilbrium reflects the 302 homogeneity of  $\Delta_{36}$  values despite strongly varying temperatures (and therefore also  $\Delta_{36,eauil}$  values).

303 Plots of  $\Delta_{36}$  values versus trace-gas mixing ratios in the same samples show correlations in certain 304 mixing ratio ranges but a limited dependence on latitude in the lower stratosphere (Fig. 2D-F). In 305 particular,  $\Delta_{36}$  values and O<sub>3</sub> concentrations are positively correlated consistently across multiple field 306 campaigns below 500 ppb O<sub>3</sub>.  $\Delta_{36}$  values and CO concentrations are inversely correlated, whereas  $\Delta_{36}$ 307 values and CH<sub>4</sub> concentrations are positively correlated below 1700 ppb CH<sub>4</sub> and negatively correlated 308 above. These observations are consistent with high- $\Delta_{36}$  stratospheric air mixing with low- $\Delta_{36}$  tropospheric 309 air in the lower stratosphere [Yeung et al., 2016]. The change in the relationship between  $\Delta_{36}$  values and 310 trace gases where  $O_3 > 500$  ppb and  $CH_4 < 1700$  ppb reflects a weak or negligible mixing signal for  $\Delta_{36}$ 311 values in that region of the stratosphere arising from more rapid photochemical isotope reordering of  $O_2$ 312 [Yeung et al., 2014].

Near 150 ppb O<sub>3</sub>—the mixing-ratio-defined tropopause used in some atmospheric chemistry modeling studies (e.g., *Young et al.* [2013])—the SEAC<sup>4</sup>RS data imply a  $\Delta_{36}$  value of 2.21 ± 0.04‰ (1 $\sigma$ ), which is slightly lower than, but consistent with the previously reported value of 2.32 ± 0.08‰ (1 $\sigma$ ) from samples taken from the DC3 campaign [*Barth et al.*, 2015; *Yeung et al.*, 2016]. However, the SEAC<sup>4</sup>RS data show higher precision and cover both subtropical and extratropical latitudes in the Northern hemisphere. At  $\theta$  = 380K—the lower bound of the stratospheric "overworld", where isentropic transport does not cross the tropopause, and thus a common isentropic reference surface for computing

- 320 stratosphere-to-troposphere mass fluxes [Holton et al., 1995; Appenzeller et al., 1996; Schoeberl,
- 321 2004]—the SEAC<sup>4</sup>RS data imply a mean  $\Delta_{36}$  value of 2.29 ± 0.04‰ (1 $\sigma$ ) based on a linear regression of
- 322 the data between  $\theta$  = 358 and  $\theta$  = 406 K ( $R^2$  = 0.96).
- 323

# 324 **3.2** $O_2$ -only online $\Delta_{36}$ modeling in GEOS-Chem

325 Modeled  $\Delta_{36T}$  values from the O<sub>2</sub>-only online calculation agree with measurements for  $\theta > 500$  K, 326 but are 0.1% - 0.4% higher than the measurements below that potential temperature (Fig. 2A). The 327 differences are not due to local meteorological disparities, as agreement between *in situ* and MERRA2 328 temperatures and pressures is excellent (Fig. S1). At the surface, the simulations yield a global-mean  $\Delta_{36T}$ 329 value of 2.37‰, whereas measurements consistently yield values that are 0.4‰ lower [Yeung et al., 2014; 330 Yeung et al., 2016; Li et al., 2019; Yeung et al., 2019]. These modeling results are similar to those from 331 the ECHAM/MESSy atmospheric chemistry model [*Gromov et al.*, 2019]. Yet, the qualitative trends seen 332 in the observational dataset are reproduced in the model, with latitude-dependent isotopic disequilibria 333 that are greatest at low latitudes.

The measurement-model disagreement persisted upon variation of many model parameters. For example, altering the rate coefficient for  $O({}^{3}P) + O_{2}$  isotope exchange within the  $2\sigma$  uncertainty range of laboratory measurements yielded perturbations to the mean surface  $\Delta_{36T}$  value of <0.01‰. Turning off advective and convective transport changed the surface  $\Delta_{36}$  value by less than 0.14‰. Resetting  $\Delta_{36T}$ values of all air parcels that cross the tropopause to a fixed value (values between 1.1‰ and 2.3‰ were evaluated) yielded tropospheric  $\Delta_{36T}$  values that are within 0.13‰ of that prescribed value.

The diagnostic  $\Delta_{36Tt}$  and  $\Delta_{36Ts}$  values—calculated  $\Delta_{36T}$  values for isotope exchange occurring in the troposphere and stratosphere only—are 2.27‰ and 2.71‰, respectively. The  $\Delta_{36Tt}$  value corresponds to an effective equilibration temperature of –33°C characteristic of the upper free troposphere. The  $\Delta_{36Ts}$ value is within 0.1‰ of the simulated  $\Delta_{36}$  values near the  $\theta$  = 380K isentrope. Absolute isotope-exchange rates increased by four orders of magnitude from the surface to  $\theta$  = 900 K, yielding O<sub>2</sub> chemical lifetimes of 0.1 – 1000 days with respect to isotope exchange (Fig. 2C).

346

# 347 3.3. Isotope-enabled photochemical modeling

Incorporation of isotope-specific rate coefficients for  $O_3$  formation into a photochemical model of the O-O<sub>2</sub>-O<sub>3</sub>-N<sub>2</sub> system, including their pressure- and temperature dependence, revealed deviations from isotopic equilibrium when the system is at photochemical steady state. As pressure increases from 0 to 1500 mbar between 175 K and 325 K, steady-state  $\Delta_{36}$  values (i.e.,  $\Delta_{36,ss}$ ) show deficits relative to  $\Delta_{36,equil}$ values (Fig. 3; fit coefficients for  $\Delta_{36,ss}$  values are reported in Table S1). The deficits are largest at high pressures, with low-temperature deficits exceeding those at higher temperatures. While the equilibrium

354  $\Delta_{36}$ -temperature relationship is approached in the low-pressure limit, modeled  $\Delta_{36,ss}$  values remain 0.01 –

355 0.04‰ below equilibrium values even at 0.1 mbar. Varying the isotope-exchange rate coefficient  $k_{\text{exch}}(T)$ 

about its  $2\sigma$  uncertainty range changes the magnitude of the  $\Delta_{36,ss}$  pressure dependence, with the upper end

357 of the range (i.e., faster rates) yielding smaller steady-state <sup>18</sup>O<sup>18</sup>O deficits and vice versa. The resulting

358 uncertainty bounds for  $\Delta_{36,ss}$  values at atmospherically relevant conditions are largest at low temperatures

and high pressures. For example, at 175 K, the uncertainty in  $k_{exch}(T)$  led to an effective 2 $\sigma$  range of 2.3%

360 at 1000 mbar (+0.9‰ and -1.4% about the mean) but a smaller 2 $\sigma$  range of 0.26‰ at 100 mbar (+0.10‰

361 and -0.16% about the mean). The assumed relative rate of  $O_3$  and  $O_2$  photolysis also affects the

362 magnitude  ${}^{18}O{}^{18}O$  deficit, but the effect is O(0.01%).

363 Nulling out the relative rate coefficient enhancements in the  $O + O_2 + M \rightarrow O_3 + M$  reactions that form heavy  $O_3$  isotopologues—which are as large as 50% compared to the rate of the  ${}^{16}O_3$ -forming 364 reaction—eliminates the pressure-dependent deficits in <sup>18</sup>O<sup>18</sup>O relative to isotopic equilibrium when the 365 366 system is at isotopic steady state. This deviation of  $\Delta_{36,ss}$  from  $\Delta_{36,equil}$  values is most sensitive to the enhancements in the relative rate coefficients for the  ${}^{16}O + {}^{18}O{}^{18}O$  and  ${}^{16}O + {}^{16}O{}^{18}O$  reactions that form O<sub>3</sub>. 367 For example, at 1000 mbar and 200 K, the isotope-exchange pathway for <sup>16</sup>O + <sup>18</sup>O<sup>18</sup>O collisions outpaces 368 the O<sub>3</sub> production pathway by a factor of ~50. Moreover, a  ${}^{16}O + {}^{18}O{}^{18}O$  collision produces O<sub>3</sub> ~20% 369 faster than a <sup>16</sup>O + <sup>16</sup>O<sup>18</sup>O collision under these conditions [*Guenther et al.*, 1999; Janssen et al., 2003]. 370 371 This unusually fast transfer of  ${}^{18}O{}^{18}O$  into O<sub>3</sub>, especially compared to  ${}^{16}O{}^{18}O$ , lowers the  $\Delta_{36}$  value of the remaining O<sub>2</sub> pool. Ozone formation via  ${}^{16}\text{O} + {}^{18}\text{O}{}^{18}\text{O}$  collisions therefore causes a roughly  $1/50 \times 0.2 \times \frac{1}{2}$ 372 373 = 2‰ decrease in  $\Delta_{36}$  value, comparable to the 1.8‰ difference between  $\Delta_{36,equil}$  and  $\Delta_{36,ss}$  values obtained 374 from the full photochemical model (note that the factor of <sup>1</sup>/<sub>2</sub> arises from the presence of both <sup>16</sup>O<sup>18</sup>O and 375  $^{18}O^{16}O$  isotopomers, which halves the effect of the  $^{16}O + ^{18}O^{18}O$  rate enhancement for O<sub>3</sub> formation). A 376 similar calculation at 1000 mbar and 300 K yields an estimated 0.7‰ deficit in  $\Delta_{36.ss}$  compared to  $\Delta_{36.eauil}$ , close to the 0.5‰ deficit calculated by the full model. At low pressures, the sequestration of <sup>18</sup>O<sup>18</sup>O 377 378 isotopologues into  $O_3$  is small:  ${}^{16}O + {}^{18}O{}^{18}O$  isotope exchange is favored by a factor of  $10^5$  over  $O_3$ 379 formation at 1 mbar and 300 K, resulting in deficits of O(0.01‰). These results suggest that the steady-380 state deficits in  $\Delta_{36,ss}$  compared to  $\Delta_{36,equil}$  values arise largely from an efficient titration of <sup>18</sup>O<sup>18</sup>O by <sup>16</sup>O to make excess <sup>16</sup>O<sup>18</sup>O<sup>18</sup>O. 381

We note that the  $\Delta_{36,ss}$  values at pressures > 1100 mbar may be most sensitive to uncertainties in the relative temperature- and pressure-dependence of O<sub>3</sub> isotopologue formation rate coefficients; because these pressures are higher than atmospheric pressures, the corresponding  $\Delta_{36,ss}$  values are not critical for modeling atmospheric  $\Delta_{36}$  values, and we do not evaluate them further.

386 Our preliminary mercury-lamp photolysis experiments, in which air is irradiated for 40 minutes at 387 elevated pressures (500 – 1000 mbar), show large  $\Delta_{36}$  deficits relative to isotopic equilibrium. An

- 388 experiment at 25°C and 887 mbar pressure showed a 0.16‰ deficit, while an experiment at –65°C and
- 389 540 mbar showed a 0.9‰ deficit relative to identical experiments performed at < 2 mbar. These
- 390 experiments confirm the presence of a pressure-dependent photochemical steady state for  $\Delta_{36}$  values, but
- 391 the precise  $\Delta_{36}$  deficits are not yet in quantitative agreement with those predicted by the photochemical
- 392 model (0.55‰ and 0.95‰, respectively). This disparity may be related to photolytic isotope
- 393 fractionations for  $O_3$  and  $O_2$  at 254 nm and 185 nm, respectively (i.e., the emission wavelengths of the
- 394 mercury lamp), which are uncertain for  $\delta^{18}$ O and  $\delta^{17}$ O and unknown for  $\Delta_{36}$ . Nevertheless, these
- 395 experiments are a proof of principle of the <sup>18</sup>O<sup>18</sup>O "titration effect," with excellent measurement-model
- agreement at low temperatures. A thorough laboratory investigation is appropriate for future work.
- 397

# 398 **3.4 Parameterizing the pressure dependence of** $\Delta_{36}$ **values at photochemical steady state**

- 399 The combined effects of atmospheric pressure and temperature covariations on  $\Delta_{36.ss}$  values are shown in Fig. 4 (left panel), along with a comparison with  $\Delta_{36,eauil}$  values (right panel). Near the surface, 400 401 the  $\Delta_{36,ss}$  value is 0.6‰ to 0.8‰ lower than the  $\Delta_{36,eauil}$  value and somewhat insensitive to temperature; the 402 pressure dependence largely determines  $\Delta_{36,ss}$  values. In the mid- to upper stratosphere, however,  $\Delta_{36,ss}$ 403 values are close to  $\Delta_{36,equil}$  values, with temperature exerting the dominant control. In between (i.e., the 404 free troposphere up through the lower stratosphere),  $\Delta_{36,ss}$  values are consistently lower than  $\Delta_{36,eauil}$  values, 405 although temperature and pressure are of comparable importance for determining local  $\Delta_{36.ss}$  values. These 406 results imply that using  $\Delta_{36,ss}$  instead of  $\Delta_{36,equil}$  values in atmospheric chemistry simulations will lead to a 407 more pronounced contrast between tropospheric and stratospheric air, and detectable  $\Delta_{36}$  offsets from the 408 surface to the lower stratosphere.
- 409 The results of GEOS-Chem simulations utilizing local  $\Delta_{36,ss}$  values are compared to the  $\Delta_{36}$
- 410 observations in Fig. 5. The model for year 2015 shows excellent agreement with the SEAC<sup>4</sup>RS, DC3, and
- 411 Ft. Sumner measurements at all altitudes in midlatitudes, with discrepancies smaller than those incurred
- 412 by the  $2\sigma$  uncertainty range in the rate coefficient for  $O(^{3}P) + O_{2}$  isotope exchange. The high-latitude
- 413 SOLVE samples were obtained in January 2000 from in the polar vortex [*Yeung et al.*, 2009; *Wiegel et*
- 414 *al.*, 2013], which can be heterogeneous in space and time; consequently, the 2015 meteorology may not
- 415 be applicable. The simulations for January 2000, by contrast, explain the SOLVE samples well (Fig. 5,
- 416 dashed lines). At the surface, the pressure-dependent online  $\Delta_{36}$  calculation scheme yields a global-mean
- 417  $\Delta_{36P}$  value of 1.97‰ in 2015 with diagnostic  $\Delta_{36Pt}$  and  $\Delta_{36Ps}$  values of 1.83‰ and 2.43‰ for troposphere-
- 418 only and stratosphere-only isotope exchange, respectively (Table 1). The calculated surface  $\Delta_{36P}$  value is
- 419 in quantitative agreement with the long-term mean of surface  $\Delta_{36}$  measurements made at Rice University
- 420 (1.99 ± 0.02‰; 95% confidence interval), while the  $\Delta_{36Ps}$  value is ~0.14‰ higher than the  $\Delta_{36}$  value at the
- 421  $\theta$  = 380 K isentrope implied by the current set of atmospheric observations.

422

## 423 **3.5 Modeled spatial and temporal variability in** $\Delta_{36}$ values

424 Figure 6 shows modeled seasonal variations in modern-day surface  $\Delta_{36P}$  values at various sites in 425 the Northern and Southern Hemispheres. Mean  $\Delta_{36P}$  values are 0.02‰ lower in the Northern Hemisphere, 426 while the seasonal ranges are 0.03‰ and 0.02‰ in the Northern and Southern Hemispheres, respectively.

427 In the O<sub>2</sub>-only calculation scheme, the seasonal ranges are slightly smaller (0.02‰ and 0.01‰,

428 respectively). None of these subtle hemispheric differences in  $\Delta_{36}$  values are currently resolvable using 429 available datasets.

430 The preindustrial (PI) simulations yielded a global-mean surface  $\Delta_{36P}$  value that is 0.06% higher

than the present day (PD) in the pressure-dependent model scheme, whereas it is 0.05‰ higher than the

432 PD in the O<sub>2</sub>-only model scheme (Table 1). Moreover, the modeled PI-to-PD  $\Delta_{36}$  change is insensitive to

433 the 2 $\sigma$  uncertainty range in <u>k</u><sub>exch</sub>(T): the spread in the PI-to-PD  $\Delta_{36}$  change is  $\leq 0.004\%$  in both online  $\Delta_{36}$ 

434 calculation schemes. Both model schemes therefore yield results that are generally consistent with the PI-

435 to-PD tropospheric  $\Delta_{36}$  change measured in ice cores (0.03 ± 0.02 decrease; 95% confidence interval)

436 [*Yeung et al.*, 2019], despite measurement-model disagreements in absolute  $\Delta_{36}$  values. We note that these

437  $\Delta_{36}$  differences correspond to the same modeled change in the tropospheric O<sub>3</sub> burden, which is 34%

438 higher in the PD compared to the PI (i.e., 296 Tg  $O_3$  vs. 221 Tg  $O_3$ ).

439

# 440 **4. Discussion**

# 441 **4.1** Assessing the accuracy of the $\Delta_{36,ss}$ pressure dependence

442 A pressure dependence to  $\Delta_{36,ss}$  could have significant implications for interpretating  $\Delta_{36}$ 443 variations in nature. First, however, one must evaluate whether the pressure-dependent online calculation 444 scheme is accurate by evaluating other potential sources of error. The evidence for an important pressure-

445 dependent effect on atmospheric  $\Delta_{36}$  values includes: (*i*) observation-model agreement in atmospheric  $\Delta_{36}$ 

446 values above  $\theta$  = 500 K, but  $\Delta_{36}$  disagreements below in the O<sub>2</sub>-only calculation scheme, (*ii*) demonstrated

447 isotopic disequilibrium caused by nonstatistical removal of O<sub>2</sub> isotopologues during O<sub>3</sub> formation, and

448 (*iii*) excellent observation-model agreement when  $\Delta_{36,ss}$  values are used in place of  $\Delta_{36,equil}$  values in

- 449 GEOS-Chem/MERRA2.
- 450 While the modern-day atmospheric observations can be explained within the pressure-dependent
- 451 framework, other factors are still relevant to consider. In particular, measurement errors, errors in
- 452 stratosphere-to-troposphere transport (STT) in the model, and errors in the rates of oxygen-isotope
- 453 exchange could be important. We explore each of these possibilities below.

454 **4.1.1 Potential errors in measurements** 

455 The agreement between measurements and the GEOS-Chem/MERRA2 model in the middle 456 stratosphere (i.e., 500 K <  $\theta$  < 950 K), argues against significant measurement errors. The high O(<sup>3</sup>P) 457 concentrations in the middle stratosphere—reflected in both the model and prior *in situ* measurements (e.g.,  $\sim 10^8$  cm<sup>-3</sup> at 30 km in boreal autumn; [Anderson, 1975])—yields rapid isotopic reordering in O<sub>2</sub>, 458 459 with O<sub>2</sub> isotope-exchange lifetimes of a day or less [Yeung et al., 2012; Yeung et al., 2014]. Thus,  $\Delta_{36}$ 460 values are expected to be near local isotopic equilibrium in the middle stratosphere, where  $\Delta_{36.ss}$  and 461  $\Delta_{36,equil}$  values are within 0.03‰ of each other. The measurements made at Rice University are consistent 462 with this expectation. Mass-spectrometric scale distortions in  $\Delta_{36}$  measurements are accounted for by low-463 pressure (<2 mbar) photochemical and high-temperature calibration experiments performed during every 464 analytical session [Yeung et al., 2014; Yeung et al., 2018]. The associated uncertainties are an order of 465 magnitude smaller than the observation-model disagreements in the lower stratosphere and troposphere. 466 Measured differences in  $\Delta_{36}$  values between photochemical calibration standards are typically within 467 0.03‰ of those predicted by theory; these deviations are quantitatively attributable to a combination of 468 isotopic reordering during sample handling [Yeung et al., 2012] and possibly a subtle pressure-dependent effect that is not currently resolvable. Moreover, recent work demonstrating reasonable agreement 469 470 between first-principles theory and respiratory fractionation of  $\Delta_{36}$  values [Ash et al., 2020] offers 471 ancillary evidence arguing against large analytical errors (i.e., >0.05%) in the measurements made at

472 Rice University.

473 Finally, all  $\Delta_{36}$  and  $\Delta_{35}$  values measured at Rice to date covary along a mass-dependent 474 fractionation trend (Fig. 7, which shows the slope of 1.92 for isotopic equilibrium [Wang et al., 2004]). While the exact pressure-dependent relationship between  $\Delta_{36,ss}$  and  $\Delta_{35,ss}$  values is not independently 475 476 known, the O<sub>3</sub> formation reactions most relevant to  $\Delta_{35,ss}$  values (e.g.,  ${}^{16}O + {}^{17}O{}^{18}O$  and  ${}^{16}O + {}^{16}O{}^{17}O$ ) are 477 predicted to have rate coefficients similar to those most relevant to  $\Delta_{36.ss}$  values [*Hathorn and Marcus*, 478 2000]. However, the relative rate coefficients for  $O(^{3}P) + O_{2}$  isotope exchange are mass-dependent, so the 479 net titration of <sup>17</sup>O<sup>18</sup>O isotopologues is expected to occur in proportion to their steady-state abundances; 480 ozone formation should not strongly alter the relationship between  $\Delta_{36}$  and  $\Delta_{35}$  values. Consequently, the 481 measured mass-dependent covariation in  $\Delta_{36}$  and  $\Delta_{35}$  values argues against analytical biases in  $\Delta_{36}$ 482 measurements, but this inference remains to be verified by measurements of isotopologue-specific rate 483 coefficients.

484 One other laboratory has reported  $\Delta_{36}$  measurements on air samples from the midlatitude upper 485 troposphere and lower stratosphere. The  $\Delta_{36}$  values reported in *Laskar et al.* [2019] are all 0.3 – 0.4‰ 486 higher than the Ft. Sumner and SEAC<sup>4</sup>RS data (samples from September 2004 and 2013, respectively), 487 despite the common latitude range and season (34°N – 42°N, 10 – 20 km, sampled September 2016). 488 Reported  $\Delta_{36}$  values for surface air from Utrecht, The Netherlands are also higher than the values reported

489 for surface air in earlier papers [Yeung et al., 2012; Yeung et al., 2014; Yeung, 2016; Li et al., 2019]. At 490 face value, the uniformly higher tropospheric and lower-stratospheric  $\Delta_{36}$  values in *Laskar et al.* [2019] 491 might favor the  $O_2$ -only photochemical scheme; however, such a result would imply a profound 492 inaccuracy in the Rice data. Importantly, a large offset in the Rice data would imply that the midlatitude 493 middle stratosphere is far from isotopic equilibrium, contrary to expectations of fast isotopic reordering 494 based on high local  $O({}^{3}P)$  concentrations ([Anderson, 1975]; see also Fig. 5C). Rapid advection of high-495  $\Delta_{36}$  air from the tropical stratosphere to the midlatitudes could cause a  $\Delta_{36}$  excess of 0.3 - 0.4% in the midlatitude stratosphere (i.e., reflecting air that is  $12 - 16^{\circ}$ C cooler at equilibrium); however, that air 496 497 would need to travel at an average speed of  $\sim$ 30 m s<sup>-1</sup> during the traverse—ten times faster than meridional wind speeds in the MERRA2 reanalysis in the middle stratosphere—without mixing with the 498 surrounding air. We note that the average meridional wind speeds in MERRA2  $[O(1 \text{ m s}^{-1})]$  are consistent 499 500 with transport timescales from the tropics to the midlatitudes implied by chemical tracer studies [Boering 501 *et al.*, 1996; *Andrews et al.*, 1999]. Furthermore, the narrow spread in measured  $\Delta_{36}$  values along surfaces 502 of constant potential temperature suggests that dispersive mixing is pervasive and mutes any variability 503 arising from meridional temperature gradients. Consequently, the  $\Delta_{36}$  values reported in *Laskar et al.* 504 [2019] are incompatible with their atmospheric context. The origin of the interlaboratory disagreement 505 remains unknown, but could plausibly be related to differences in calibration protocols such as the 506 pressures of photolysis of photochemical standards.

507

## 4.1.2 Potential errors in stratosphere-troposphere exchange fluxes

508 Inaccurate STT fluxes could have varied effects on lower-stratospheric  $\Delta_{36}$  values because  $O_2$  is 509 not at photochemical steady state with respect to its isotopologues. Weak STT fluxes in the model would 510 cause excessive lower-stratospheric residence times, driving  $\Delta_{36}$  values in the model closer to local  $\Delta_{36,eauil}$ 511 or  $\Delta_{36.ss}$  values than in the observations. STT fluxes that are too strong would have a variety of effects 512 depending on whether the transport is diabatic (e.g., from excessive vertical velocities in the residual 513 circulation) or isentropic (e.g., an excessive diffusive flux). Stronger downward diabatic flow might increase lower-stratospheric  $\Delta_{36}$  values due to the influx of high- $\Delta_{36}$  air from above. This particular case is 514 515 difficult to distinguish from the weak-STT case, but both scenarios would underestimate the fraction of 516 tropospheric air in the lower stratosphere (i.e., air that has not been photochemically reordered since 517 crossing the tropopause). By contrast, excessive isentropic mixing would increase the fraction of 518 tropospheric air in the lowermost stratosphere, rendering modeled  $\Delta_{36}$  values there too low. 519 We therefore used the fraction of tropospheric air as a metric to evaluate the accuracy of STT in 520 the model. A two-component mass balance between tropospheric  $O_2$  and stratospherically reordered  $O_2$  in

- 521 air for each sample can be constructed to approximate the lower-stratospheric budget:
- 522

$$\Delta_{36,\text{observed}} = f_{\text{trop}} \Delta_{36,\text{trop}} + (1 - f_{\text{trop}}) \Delta_{36,\text{ss or equil}}$$
(5)

523 524

525

estimate of the fraction of tropospheric air,  $f_{trop}$ , can thus be derived from an observed  $\Delta_{36}$  value and the calculated local  $\Delta_{36,ss}$  or  $\Delta_{36,equil}$  endmember of the sampling location. While the assumption that only local reordering and mixing alters  $\Delta_{36}$  values leads to some uncertainties in this estimate for  $f_{trop}$ , the results should be sufficient to test for gross errors in STT because the lapse rate is small in the lower

Here, stratospheric O<sub>2</sub> is defined as O<sub>2</sub> that was recently reset to its local  $\Delta_{36,ss}$  or  $\Delta_{36,equil}$  value. An

530 stratosphere, resulting in smaller variations in  $\Delta_{36,ss}$  or  $\Delta_{36,equil}$ . We used  $\Delta_{36,trop} = 1.99\%$  for the

531 measurements [*Yeung et al.*, 2019] and  $\Delta_{36,trop} = 1.97\%$  (pressure-dependent scheme) or 2.37% (O<sub>2</sub>-only

scheme) for the model to maintain internal consistency.

533 To facilitate comparison with other studies of STT, the modeled isotope-exchange timescale from 534 GEOS-Chem,  $\tau_{\text{exch}}$ , was used to derive an effective mean residence time ( $\tau_{\text{residence}}$ ) from the equation 535

536

$$f_{trop} = e^{-\tau_{residence}/\tau_{exch}} \tag{6}$$

537

This effective mean residence time can be directly compared to mean residence times for the lowermost stratosphere reported in observational studies. A disagreement between the  $\tau_{\text{residence}}$  values obtained via  $\Delta_{36}$ measurements, the model, and other observational studies would indicate shortcomings in the calculation of the  $\Delta_{36,ss}$  endmembers,  $\tau_{\text{exch}}$ , or the two-endmember approach.

542 Figure 8 shows the values of  $f_{\text{trop}}$  and  $\tau_{\text{residence}}$  derived from this analysis. In both schemes, the tropospheric fraction is less than 5% above  $\theta$  = 450 K and it increases below, with excellent observation-543 544 model agreement. The data also support a subtle meridional trend in modeled  $f_{trop}$  values, in which 545 subtropical  $f_{\text{trop}}$  values are 10 – 15% higher than midlatitude values at the same potential temperature. 546 Calculated residence times range between 20 and 60 days in the lowermost stratosphere (i.e.,  $\theta < 380$  K), 547 consistent with observationally constrained estimates of 20 - 80 days in the tropical tropopause layer 548 documented during boreal winter [Andrews et al., 1999; Krüger et al., 2009]. STT mass fluxes in the 549 MERRA2 reanalysis therefore appear reasonably accurate, at least for the purposes of this measurement-550 model comparison.

551

# 4.1.3 Potential errors in O(<sup>3</sup>P) + O<sub>2</sub> isotope-exchange rates

552 The measurement-model agreement in the O<sub>2</sub>-only calculation scheme could potentially be 553 improved if lower-stratospheric  $O({}^{3}P) + O_{2}$  isotope-exchange rates are overestimated and/or tropospheric 554 isotope-exchange rates are underestimated. Because the uncertainty in laboratory measurements of  $k_{exch}(T)$ 555 has only a weak effect on the O<sub>2</sub>-only online  $\Delta_{36}$  results (cf. Fig. 2), the changes in isotope-exchange rates 556 would need to be driven by changes in  $O({}^{3}P)$  concentrations. GEOS-Chem's chemical mechanism does omit  $O_2(^{1}\Sigma)$  chemistry, which can lead to an underestimate of  $O(^{3}P)$  concentrations particularly in the

- stratosphere [*Yeung et al.*, 2014], but the accuracy of stratospheric O<sub>3</sub> concentrations calculated by the
- 559 UCX mechanism [*Eastham et al.*, 2014] disfavors this scenario: an additional odd-oxygen source in the
- 560 stratosphere would increase  $O_3$  concentrations and drive lower-stratospheric  $\Delta_{36}$  values closer to isotopic
- 561 equilibrium, exacerbating existing disagreements. Similarly, the high accuracy of tropospheric O<sub>3</sub>
- 562 concentrations in GEOS-Chem argues against a major (e.g., factor-of-two or more) deficit in tropospheric
- 563  $O(^{3}P)$  concentrations at the global scale [*Hu et al.*, 2017].
- 564 A preponderance of evidence therefore supports a pervasive titration of <sup>18</sup>O<sup>18</sup>O into O<sub>3</sub> in the 565 atmosphere, which results in lower atmospheric  $\Delta_{36}$  values than isotopic equilibration would predict.
- 566 Thus, an error in the Rice laboratory measurements does not need to be invoked; the

567 GEOS-Chem/MERRA2 model represents STT and photochemical isotope exchange reasonably well; and

uncertainties in the isotope-exchange rates cannot account for the magnitude of the  $\Delta_{36}$  offsets between

569 measurements and the O<sub>2</sub>-only photochemical scheme. Given this evidence, we can now explore the

570 implications this new pressure-dependent isotope chemistry has for our understanding of the global  $\Delta_{36}$ 571 budget.

572

# 4.2 Revisiting the global Δ<sub>36</sub> budget: reconciling online and offline calculation approaches 4.2.1 Comparing the online and offline (two-box) models

575 Tropospheric  $\Delta_{36}$  values are determined largely by the balance of two processes:  $O({}^{3}P) + O_{2}$ 576 reactions and stratosphere-troposphere mass exchange. Model predictions using pressure-dependent  $\Delta_{36,ss}$ 577 values affect the implied importance of each process, but not the overall form of the budget. Therefore, in 578 steady state, and omitting biological oxygen cycling (which has a negligible contribution to the global  $\Delta_{36}$ 579 budget [*Yeung et al.*, 2015]), one can describe the tropospheric budget as a two-endmember mixture: 580

581 
$$\Delta_{36,trop} \approx \left(\frac{E_{trop}}{F_{ST} + E_{trop}}\right) \Delta_{36,Pequil} + \left(\frac{F_{ST}}{F_{ST} + E_{trop}}\right) \Delta_{36,strat}$$
(7)

582

Here,  $E_{trop}$  (mol yr<sup>-1</sup>) is the effective integrated isotope-exchange flux in the troposphere,  $F_{ST}$  (mol yr<sup>-1</sup>) is the STT flux of O<sub>2</sub>,  $\Delta_{36,Pequil}$  is the global effective-mean  $\Delta_{36,ss}$  value for the troposphere, and  $\Delta_{36,strat}$ is the global effective-mean  $\Delta_{36}$  value of air descending from the stratosphere into the troposphere. Note that the  $\Delta_{36Pt}$  and  $\Delta_{36Ps}$  values calculated online in the simulations correspond to  $\Delta_{36,Pequil}$  and  $\Delta_{36,strat}$  values, respectively.

588 Previously, the quantities in eq. 7 were determined "offline" using a chemical tropopause (150 589 ppb O<sub>3</sub>), an empirical  $\Delta_{36,strat}$  value, and model-derived  $E_{trop}$  and  $\Delta_{36,Pequil}$  values because online calculations were not yet available [*Yeung et al.*, 2016; *Yeung et al.*, 2019]. The resulting balance yielded  $\Delta_{36,trop}$  values similar to observations. However, the online modeling results reported here suggest that the previous approach was not internally consistent, and thus may be conceptually flawed. For example, using the offline scheme described in *Yeung et al.* [2016], one would derive a  $\Delta_{36,Pequil}$  value of 1.45‰ and a  $\Delta_{36,trop}$ value of 1.64‰ for the present day, both of which are much lower than the corresponding quantities obtained from the online calculation ( $\Delta_{36Pt} = 1.83\%$  and  $\Delta_{36,trop} = 1.97\%$ ). Thus, it is useful to examine how the online and offline calculations might be reconciled, both to correct any conceptual disparities that

597 may exist and to improve our understanding of the online model results.

598 First, we will examine how the offline  $\Delta_{36,Pequil}$  and online  $\Delta_{36Pt}$  values can be reconciled. By 599 comparing the  $\Delta_{36,Pequil}$  and  $\Delta_{36Pt}$  values obtained for present-day and preindustrial simulations, we find that 600 online  $\Delta_{36Pt}$  values are most closely reproduced offline by the weighting scheme below, which is similar 601 to, but slightly different from that used in *Yeung et al.* [2016]:

602

603 
$$E_{trop} = \frac{\sum_{trop \, grid \, boxes} k_{exch}(T) \left[ O\left( {}_{\Box}^{3} P \right) \right] \left[ O_{2} \right] m_{box} \rho_{box}^{n}}{\sum_{trop \, grid \, boxes} m_{box} \rho_{box}^{n}} \times V_{trop}$$
(8)

604

605 
$$\Delta_{36,Pequil} = \frac{\sum_{trop \, grid \, boxes} k_{exch}(T) \left[ O \left( \begin{smallmatrix} 3 \\ \Box P \end{smallmatrix} \right) \right] \left[ O_2 \right] m_{box} \rho_{box}^n \Delta_{36,Pequil,box}}{\sum_{trop \, grid \, boxes} k_{exch}(T) \left[ O \left( \begin{smallmatrix} 3 \\ \Box P \end{smallmatrix} \right) \right] \left[ O_2 \right] m_{box} \rho_{box}^n}$$
(9)

606

Each equation represents the annual tropospheric mean obtained by weighting contributions from different processes within each tropospheric grid box. The first term in the numerators of both equations,  $k_{\text{exch}}(T)[O(^{3}P)][O_{2}]$ , is the local rate of isotope exchange in units of concentration per unit time;  $m_{\text{box}}$  is the mass of air used to normalize for the variable size of grid boxes, and  $V_{\text{trop}}$  is the volume of the troposphere.

612 The generalized  $\rho_{\text{box}}^{n}$  term differs from that used in Yeung et al. [2016], where n was assumed to 613 be unity. Here, the exponent *n* is introduced to acknowledge that two factors influence the progression of 614 the  $\Delta_{36}$  value toward  $\Delta_{36.ss}$  in each box on longer (e.g., monthly) timescales. The first factor, the 615 cumulative residence time in a grid box, was previously argued to vary proportionately with the number 616 density  $\rho_{\text{box}}$  (i.e., n = 1) because a well-stirred troposphere should be ergodic: the ensemble mean is equal 617 to the temporal mean. Therefore, higher mean number densities imply longer cumulative residence times for air molecules in those boxes. The second factor affecting the approach of the  $\Delta_{36}$  value toward  $\Delta_{36,ss}$  in 618 619 each box is the isotope-reordering efficiency. It scales inversely with the number density  $\rho_{\text{box}}$  (i.e., n = -1) because isotopic steady state is reached more quickly when fewer  $O_2$  molecules are present—fewer exchange events are required. This second factor was not considered in *Yeung et al.* [2016].

622 The value of *n* that reproduces the online  $\Delta_{36Pt}$  value for 2015 is -0.8, indicating that the isotope-623 exchange efficiency in each grid box is much more important than the cumulative air parcel residence 624 time for determining tropospheric  $\Delta_{36}$  values. For example, in the upper troposphere, GEOS-Chem 625 predicts molar isotope-exchange rates that are roughly ten times faster than at the surface ( $\tau_{exch} \sim 10^2$  vs. 626  $10^3$  days; see Fig. 5C). Number densities are also about tenfold lower in the upper troposphere. Together, 627 these two properties make isotope exchange  $\sim 100$  times faster in the upper troposphere than at the 628 surface. Atmospheric mixing counteracts the effects of residence-time differences between the lower and 629 upper troposphere, further amplifying the importance of isotope exchange in the upper troposphere. 630 Isotope exchange in the upper troposphere thus appears to be the dominant contribution to the  $\Delta_{36,Pequil}$ 631 value. The optimal value of *n*, however, depends on the balance of chemistry and transport in principle, 632 and thus may be model- and climate-dependent: we find that faster rates of isotope exchange decrease the 633 importance of local residence-time differences (e.g., n = -1 was optimal for the  $+2\sigma$  simulation) while 634 slower rates of isotope exchange increase their importance (e.g., n = -0.3 for the  $-2\sigma$  simulation). Yet, 635 calculated differences in  $\Delta_{36,Pequil}$  between preindustrial and present-day scenarios changed <0.002‰ upon 636 varying the value of *n* between –1 and –0.3.

637 Using eq. 8 above, one can also evaluate the relative importance of different regions to the 638 present-day  $\Delta_{36,Pequil}$  value. The results show that the major contribution to the annual-mean  $\Delta_{36,Pequil}$  value 639 comes from the low- and mid-latitude free troposphere in GEOS-Chem/MERRA2 (Fig. 9). Peaks in 640 boreal spring and summer in the extratropics were identified, indicating that both STT of  $O_3$  and anthropogenic emissions influence the seasonal  $\Delta_{36,Pequil}$  value by accelerating  $O(^{3}P) + O_{2}$  isotope 641 642 exchange reactions locally. Integrated annually and globally, the mean effective altitude for tropospheric 643 photochemistry is 10.1 km, with a broad  $1\sigma$ -equivalent width of approximately  $\pm 4.5$  km about the mean. 644 This effective mean altitude is higher than that previously reported (~5 km; [Yeung et al., 2016]) and 645 reflects the greater importance of upper-tropospheric photochemical reordering uncovered in this study. 646 The zonal distribution is similar to that of Earth's surface area, with a maximum in isotope-exchange flux in the relatively cloud-free subtropics (Fig. 9). The tropospheric contribution to the  $\Delta_{36}$  budget is therefore 647 648 sensitive to photolysis, chemistry, and temperature of the free troposphere on the global scale.

649

# 4.2.2 Revising the terms in the two-box model

650 We will now compare the online-modeled  $\Delta_{36Ps}$  value with the empirically determined  $\Delta_{36,strat}$ 651 value used in eq. 7. The latter depends strongly on how the tropopause is defined, and its accuracy is 652 limited by the number and spatiotemporal coverage of available observations. Our Northern-Hemisphere 653 observations currently suggest that the  $\Delta_{36,strat}$  value is between 2.2 and 2.3‰ when using  $\theta$  = 380 K or

654 150 ppb O<sub>3</sub> as the effective  $\Delta_{36}$  tropopause (Fig. 5D). The  $\Delta_{36PS}$  value of 2.43‰ therefore appears to be 0.1 655 -0.2% higher than the observed  $\Delta_{36,\text{strat}}$  value (Table 1). Seasonality may be responsible for this apparent 656 disagreement. Within the online model, the annual-mean  $\Delta_{36P_5}$  value is closest to  $\Delta_{36}$  values at  $\theta = 380$  K 657 during boreal spring, when net STT fluxes are strongest [Holton et al., 1995; Appenzeller et al., 1996]. 658 However, the majority of lower-stratospheric measurements to date were sampled during boreal autumn, 659 when net mass transport moves mass in the opposite direction, i.e., from the troposphere into the 660 stratosphere. One therefore expects the boreal autumn-biased  $\Delta_{36,\text{strat}}$  values to be lower than the annualmean  $\Delta_{36Ps}$  value because the fraction of unaltered tropospheric air (e.g.,  $f_{trop}$ ) is larger in boreal autumn 661 662 than in spring. The current set of observations therefore provides an incomplete constraint on  $\Delta_{36Ps}$ . 663 Nevertheless, the observed seasonal bias is consistent with that expected from the seasonality of

664 stratosphere-troposphere exchange; no inconsistencies are evident.

665 The preceding discussion thus provides a way to obtain the values of the tropospheric terms in the 666 two-box description of the  $\Delta_{36}$  budget (eq. 7): when  $\Delta_{36Pt} = \Delta_{36,Pequil}$ , the online and offline descriptions are equivalent, and yields  $E_{\text{trop}} = 1.4 \times 10^{19} \text{ mol } O_2 \text{ yr}^{-1}$ . The  $\Delta_{36\text{Ps}}$  value (2.43‰) and online  $\Delta_{36,\text{trop}}$  value ( $\Delta_{36\text{Ps}}$ 667 = 1.97‰) then allow one to solve for  $F_{ST}$ , which is inferred to be 4.3 × 10<sup>18</sup> mol O<sub>2</sub> yr<sup>-1</sup>. This  $F_{ST}$  value is 668 669 similar to the net STT flux derived from seasonal diabatic transport across the  $\theta$  = 380 K surface in the extratropics used in previous studies (i.e.,  $4.6 \times 10^{18}$  mol O<sub>2</sub> yr<sup>-1</sup>; [Appenzeller et al., 1996; Schoeberl, 670 2004]) and is consistent with the 1 - 2-month mean residence time of air in the extratropical lowermost 671 672 stratosphere (Fig. 8). The present-day  $\Delta_{36}$  value at the surface is thus comprised of 77% tropospherically 673 reordered air and 23% stratospherically reordered air in steady state.

674 We note that the MERRA2 STT flux reported in Boothe and Homeyer [2017], i.e.,  $F_{\text{ST,BH}} = 1.2 \times$ 675  $10^{19}$  mol O<sub>2</sub> yr<sup>-1</sup>, is threefold larger than the estimate above and would result in an inconsistency of ~0.1‰ 676 in  $\Delta_{36,trop}$  values, with the box model yielding  $\Delta_{36,trop}$  = 2.11‰. This inconsistency arises because the 677 residence-time threshold used for determining "irreversible" STT trajectories in Boothe and Homeyer 678 [2017] is much shorter than typical isotope-exchange lifetimes in the lowermost stratosphere (e.g., 4 days 679 vs. 1-2 months). Consequently, the STT mass flux from the analysis of Boothe and Homeyer [2017] 680 includes many tropopause-crossing transients that leave  $\Delta_{36}$  values largely unaltered in an air parcel; the value of  $F_{ST}$  relevant for the global  $\Delta_{36}$  budget should be smaller. *Škerlak et al.* [2014] found that STT 681 fluxes based on trajectory studies scale inversely with the square root of the STT residence-time 682 683 threshold. That scaling relationship predicts that high- $\Delta_{36}$  stratospheric air becomes irreversibly mixed into the troposphere about a month after descending below  $\theta$  = 380K, i.e.,  $F_{\rm ST}/F_{\rm ST,BH}$  ~ (4 days / 30 684 685 days)<sup>1/2</sup>, consistent with the expected isotope-exchange lifetime at those altitudes. 686 4.2.3 Summary of revisions to the  $\Delta_{36}$  budget

687 In summary, the two-box offline framework for the atmospheric  $\Delta_{36}$  budget can be reconciled 688 with the online simulations after a more accurate accounting of photochemical reordering in the upper 689 troposphere and the STT flux relevant for the  $\Delta_{36}$  tracer. This new analysis reveals that the effective mean 690 altitude interrogated by the tropospheric  $\Delta_{36}$  tracer is 10 km rather than ~5 km, which was previously 691 suggested based on an incomplete accounting of chemistry and transport [Yeung et al., 2016]. The 692 previous approach had underweighted isotope exchange in the upper troposphere because it did not 693 consider the faster approach to isotopic steady state at low  $O_2$  number densities. Importantly, the present 694 consistency between the online and offline approaches allows one to rapidly evaluate the importance of 695 perturbations (e.g., temperature changes) on past tropospheric  $\Delta_{36}$  values without requiring additional 696 online simulations.

697

# 698 **4.3 Changes in the** $\Delta_{36}$ **budget and** $O_3$ **since the preindustrial era**

699 Previous work documented a tropospheric  $\Delta_{36}$  decrease over the twentieth century associated with 700 an increased tropospheric O<sub>3</sub> burden [*Yeung et al.*, 2019]. Online  $\Delta_{36}$  calculations and diagnostics allow us 701 to analyze the GEOS-Chem/MERRA2 results here in more detail and estimate the magnitude of known 702 systematic uncertainties, facilitating a more quantitative comparison with the atmospheric record.

The PI and PD simulation results can be summarized as follows: the increase in tropospheric O<sub>3</sub> since 1850 CE yields a predicted  $\Delta_{36,trop}$  decrease of 0.06‰. The magnitude of the  $\Delta_{36,trop}$  decrease is caused by (*i*) the locus of O(<sup>3</sup>*P*) + O<sub>2</sub> isotope exchange shifting toward the Northern Hemisphere and lower altitudes (Fig. 10), resulting in the  $\Delta_{36,Pequil}$  value decreasing by 0.04‰ (Table 1), and (*ii*) the effective integrated isotope-exchange flux in the troposphere ( $E_{trop}$ ) increasing by ~30%. These changes also lead to a 0.01‰ increase in the interhemispheric gradient. We note that the uncertainty bounds derived from online simulations run at the 2 $\sigma$  limits for  $k_{exch}(T)$  are <0.005‰ for the PI-to-PD shift in  $\Delta_{36P}$ .

We will now estimate the known systematic uncertainties. First, the modeled changes reflect
 present-day meteorology and thus omit the ~1°C global warming since 1850 C.E. This warming has two

- 712 potentially opposing effects on the mean tropospheric  $\Delta_{36}$  value: to decrease  $\Delta_{36,Pequil}$  further and to
- 713 increase  $F_{st}$ . The two-box framework suggests that a 1°C tropospheric warming would cause an
- additional 0.01‰ decrease in  $\Delta_{36,Pequil}$ , if all else is kept unchanged. However, anthropogenic warming is
- believed to have also accelerated the lower branch of the Brewer-Dobson circulation by ~10%, which
- 716 would increase  $F_{ST}$  [*Lin and Fu*, 2013]. Using the effective present-day  $F_{ST}$  of  $4.3 \times 10^{18}$  mol O<sub>2</sub> yr<sup>-1</sup>
- across the tropopause and the model-derived quantities  $E_{\text{trop}} = 1.1 \times 10^{19} \text{ mol } O_2 \text{ yr}^{-1}$ ,  $\Delta_{36,\text{strat}} = 2.43\%$ , and
- 718  $\Delta_{36,Pequil} = 1.87\%$  for the PI, we calculate that a 10% smaller STT flux during the PI results in a 0.01‰
- smaller PI-PD shift in global-mean tropospheric  $\Delta_{36}$  value. Consequently, the two opposing effects of the
- 720 1°C warming since the preindustrial cancel within the tropospheric  $\Delta_{36}$  budget.

721Second, interannual variability in meteorology can affect both the magnitude of  $F_{ST}$  and the722effective chemistry/transport balance (i.e., the exponent *n* in eqs. 8 and 9). Our explorations of these723effects suggest that interannual variability about long-term means contributes ~0.01‰ uncertainty to the724global-mean  $\Delta_{36,trop}$  value for a given emissions scenario and climate state. For example, the mean  $\Delta_{36P}$ ,725 $\Delta_{36Pt}$ , and  $\Delta_{36Ps}$  values for GEOS-Chem/MERRA2 simulations from June 1998 – June 2000 were all

726 0.01‰ higher than in 2015 despite comparable global pollutant emissions [*McDuffie et al.*, 2020].

727 Third, the biomass burning emissions of nitrogen oxides  $(NO_x)$  and carbon monoxide (CO) are 728 still uncertain in the preindustrial atmosphere. Recent studies suggesting that preindustrial biomass 729 burning emissions were higher than they are in the present argue that emissions were as much as 40% 730 higher for CO and 20% higher for NO<sub>x</sub> [*Rowlinson et al.*, 2020], perhaps due to land-use change [*Pfeiffer*] 731 et al., 2013; Andela et al., 2017; Hamilton et al., 2018]. Such "high-fire" NO<sub>x</sub> and CO emissions from 732 biomass during the PI would reduce the PI-to-PD increase in tropospheric  $O_3$  burden from 34% to 15% 733 [*Murray et al.*, 2014]) and thus also mute the change in tropospheric  $\Delta_{36}$  value. Taking the change in 734 tropospheric  $O_3$  burden as a proxy for  $E_{trop}$ , the "high-fire" scenario would reduce the PI-to-PD change in 735 tropospheric  $\Delta_{36}$  value by 0.01‰.

Finally, the modeled tropospheric  $\Delta_{36}$  change over the GISP2 (Summit, Greenland) ice coring site is 0.01‰ larger than over the WAIS-D (West Antarctica) ice coring site in both calculation schemes, reflecting a larger O<sub>3</sub> reduction in the Northern hemisphere. However, this small difference would not be detectable by current analytical methods. Nevertheless, averaging measurements from the Northern and Southern hemispheres provides a more robust estimate of the global tropospheric  $\Delta_{36}$  value and its changes through time.

742 This explicit quantification of potential systematic uncertainties in the online  $\Delta_{36}$  calculation 743 suggests that the modeled decrease in tropospheric  $\Delta_{36}$  since 1850 C.E. is 0.06 ± 0.02‰ (2 $\sigma$  based on 744 uncertainties added in quadrature), consistent with the mean  $0.03 \pm 0.02\%$  decrease (95% confidence 745 interval) observed in the ice-core record [Yeung et al., 2019]. Despite some uncertainty in the modeled absolute  $\Delta_{36}$  values arising from the isotopic photochemistry of O<sub>3</sub>, the associated systematic uncertainties 746 747 likely do not contribute more than 0.02‰ to modeled changes in tropospheric  $\Delta_{36}$  value. Preindustrial 748 scenarios invoking higher biomass-burning emissions of NO<sub>x</sub> and CO [Rowlinson et al., 2020] would 749 likely yield an estimated  $\Delta_{36,trop}$  change closer to the centroid of the observed probability density 750 distribution.

751

# 752 5. Conclusions

The observational, modeling, and experimental evidence presented in this study all suggest that the titration of <sup>18</sup>O<sup>18</sup>O into heavy ozone molecules—an atmospheric sink for <sup>18</sup>O<sup>18</sup>O that was previously

755 thought to be negligible—drives  $\Delta_{36}$  values toward a photochemical steady state that differs from isotopic 756 equilibrium. An online parameterization of this effect within the GEOS-Chem chemical transport model, 757 based on pressure- and temperature-dependent outputs of a detailed isotope-enabled kinetics model, allows GEOS-Chem to reproduce high-precision laboratory measurements on atmospheric samples 758 759 collected over a range of latitudes, altitudes, seasons, and years. This improved understanding of the 760 factors that affect atmospheric  $\Delta_{36}$  values, however, results in only minor changes to the predicted 761 evolution in tropospheric  $\Delta_{36}$  values over the twentieth century in response to the increase in tropospheric  $O_3$ ; thus, the  $\Delta_{36}$  tracer remains a constraint on the global preindustrial  $O_3$  burden. Analytical precision in 762 763  $\Delta_{36}$  measurements and uncertainties surrounding preindustrial biomass burning emissions currently 764 dominate the uncertainty in the measurement-model comparison of PI and PD  $\Delta_{36}$  values, with the PI-PD 765 change in  $\Delta_{36}$  values consistent with no more than a 30-40% increase in the tropospheric O<sub>3</sub> burden since

766 1850 C.E.

Furthermore, the online  $\Delta_{36}$  calculation scheme described herein allows variations in odd-oxygen chemistry and free-tropospheric temperatures—manifesting as changes in tropospheric  $\Delta_{36}$  values—to be investigated for past climates within chemical-transport and chemistry-climate models. While O<sub>3</sub> photochemistry and climate are linked through biogeochemistry and atmospheric dynamics [*Alexander et al.*, 2003; *Rind et al.*, 2009; *Murray et al.*, 2014; *Geng et al.*, 2017; *Wang et al.*, 2020],  $\Delta_{36}$  values may nevertheless offer a valuable constraint on the long-term coevolution of atmospheric chemistry and climate.

774

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784

# 785 Data Availability

The measured data and model outputs used to support the conclusions can be accessed at
 <u>https://doi.org/10.5061/dryad.c866t1g6t</u>.



788Temperature (K)Temperature (K)789Figure 1. (Left) Temperature dependence of the bimolecular rate coefficient for the <sup>16</sup>O + <sup>18</sup>O<sup>18</sup>O

790 isotope-exchange reaction obtained from the combined laboratory/theory study of *Fleurat-Lessard* 

791 *et al.* [2003]. Shaded area represents experimental 2σ uncertainty range. (Right) The ratios of O<sub>3</sub>-

formation and isotope-exchange rates for  ${}^{16}O + {}^{16}O{}^{16}O$  and  ${}^{16}O + {}^{18}O{}^{18}O$  collisions at different pressures,

793 which shows relative rates of  $O_3$  formation that are markedly faster for  ${}^{16}O + {}^{18}O{}^{18}O$  collisions at high

794 pressures and low temperatures.



795

**Figure 2. Measurement-model comparison for the O**<sub>2</sub>**-only online**  $\Delta_{36}$  calculation scheme in GEOS-

797 **Chem.** The diamond in panels A and B reflect the mean surface  $\Delta_{36}$  value measured in Houston, TX.

798Error bars on the data points represent the 1σ pooled standard deviation of duplicates. Modeled monthly

means from January 2000 and September 2015 are plotted for comparison. Shaded areas represent

800 pseudo- $2\sigma$  uncertainty bounds obtained by running identical simulations at the experimental  $\pm 2\sigma$  bounds

801 for  $k_{\text{exch}}(T)$ . Note that the isotope-exchange lifetime for the 78°N, January 2000 profile in panel C is

802 scaled by  $10^{-6}$  for clarity.



804Figure 3.  $\Delta_{36,ss}$  vs. pressure from the isotope-enabled photochemical kinetics model for a range of805temperatures (filled circles) compared to polynomial fits of the trends (curves). For reference,  $\Delta_{36}$ 806values for isotopic equilibrium (i.e.,  $\Delta_{36,equil}$ ) are shown as stars on the left axis. Deviations of  $\Delta_{36,ss}$  from

 $\Delta_{36,equil}$  values at 0.1 mbar range from -0.01‰ at 175K to -0.04‰ at 300K. The lower plot shows the

808 misfit between the modeled  $\Delta_{36,ss}$  compositions and polynomial fits to those values.



811

Figure 4. Patterns in  $\Delta_{36,ss}$  (left) and  $\Delta_{36,equil}$  (right) along atmospheric covariations of temperatures

- 813 **and pressures.** Modern annual-mean atmospheric temperature profiles for tropical (30°S 30°N),
- 814 midlatitude ( $30^{\circ}S 60^{\circ}S$  and  $30^{\circ}N 60^{\circ}N$ ), and polar ( $60^{\circ}S 90^{\circ}S$  and  $60^{\circ}N 90^{\circ}N$ ) latitudes from the 815 MERRA2 reanalysis are overlain.
- 816



817



819 scheme in GEOS-Chem. The diamond in panels A and B reflect the mean surface  $\Delta_{36}$  value measured in

820 Houston, TX. Error bars on the data points represent the  $1\sigma$  pooled standard deviation of duplicates.

821 Modeled monthly means from January 2000 and September 2015 are plotted for comparison. Shaded

areas represent pseudo-2σ uncertainty bounds obtained by running identical simulations at the

823 experimental  $\pm 2\sigma$  bounds for  $k_{\text{exch}}(T)$ . Note that the isotope-exchange lifetime for the 78°N, January 2000

824 profile in panel C is scaled by  $10^{-6}$  for clarity.



826 **Figure 6. Modeled seasonal cycle of surface** Δ<sub>36P</sub> **values at various sites in 2015.** The modeled global-

827 mean surface  $\Delta_{36P}$  value is also shown. Ice-coring site locations are Summit, Greenland (GISP2; 72.6°N,

828 38.5 °W), the West Antarctic Ice Sheet Divide (WAISD; 79.5°S, 112.1°W) and East Antarctica near

829 Dome C (LOCK-IN; 74.1°S, 126.2°E).

830



832 Figure 7. Comparison of  $\Delta_{36}$  and  $\Delta_{35}$  values measured on the same samples, which show a mass-

833 dependent trend.





836 Figure 8. Measurement-model comparison of stratosphere-troposphere mixing fractions (left) and

- 837 implied lower-stratospheric residence times (right).
- 838
- 839





Figure 9. Normalized contour plot depicting where O(<sup>3</sup>*P*) + O<sub>2</sub> isotope-exchange chemistry records

842 tropospheric climate (i.e.,  $\Delta_{36Pt}$  or  $\Delta_{36,Pequil}$ ; cf. eq. 7) in GEOS-Chem/MERRA2 model year 2015.

Also shown are the integrated normalized contributions as a function of altitude and latitude (black
curves). The mean effective altitude (red line in right panel) and Earth surface area (red curve) are shown

- 845 for comparison.
- 846
- 847
- 848







**GEOS-Chem/MERRA2.** Changes are computed as the difference in normalized contribution to the  $\Delta_{36Pt}$ or  $\Delta_{36,Pequil}$  value (e.g., Fig. 9) in each scenario. Positive values indicate increased importance in the

853 present day (PD), while negative values indicate decreased importance, relative to the preindustrial (PI).
854

Table 1. Comparison of the simulated preindustrial-to-present change in  $\Delta_{36}$  values and diagnostics using the O<sub>2</sub>-only and pressure-dependent calculation schemes. 

	O <sub>2</sub> -only			Pressure-dependent		
	$\Delta_{ m 36T}$ / ‰ $^{*}$	$\Delta_{ m 36Tt}$ / ‰ $^{\dagger}$	$\Delta_{ m 36Ts}$ / ‰ $^{\ddagger}$	$\Delta_{ m 36P}$ / ‰*	$\Delta_{ m 36Pt}$ / ‰ $^{\dagger}$	$\Delta_{ m 36Ps}$ / $ m \%o^{\ddagger}$
Using the mean value for O( <sup>3</sup> P) + O <sub>2</sub> isotope-exchange rate coefficients, $k_{\text{exch}}(T)^{\$}$						
2015	2.372	2.271	2.712	1.973	1.832	2.430
1850	2.418	2.302	2.710	2.033	1.871	2.428
Change	-0.046	-0.031	0.002	-0.060	-0.039	0.002
Using the lower bound for isotope-exchange rate coefficients, $k_{\text{exch}}(\mathbf{T}) - 2\sigma^8$						
2015	2.379	2.253	2.712	1.874	1.691	2.346
1850	2.427	2.283	2.710	1.939	1.730	2.344
Change	-0.048	-0.030	0.002	-0.065	-0.039	0.002
Using the upper bound for isotope-exchange rate coefficients, $k_{\text{exch}}(\mathbf{T})$ + $2\sigma^{\$}$						
2015	2.367	2.286	2.712	2.042	1.934	2.489
1850	2.411	2.318	2.710	2.100	1.973	2.488
Change	-0.044	-0.032	0.002	-0.058	-0.039	0.001
imulated global	-mean value	at the surface				

\*Diagnostic signature for isotope exchange occurring in the troposphere 

<sup>‡</sup>Diagnostic signature for isotope exchange occurring in the stratosphere

<sup>§</sup>Mean and uncertainty bounds as reported in *Fleurat-Lessard et al.* (2003).

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