# Shock remanent magnetization intensity and stability structures of single-domain titanomagnetite-bearing basalt sample

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### Abstract

Knowledge of the shock remanent magnetization (SRM) structure is crucial to interpret the spatial changes in magnetic anomalies observed over the impact crater. This study reports the SRM intensity and stability structures of single-domain titanomagnetite-bearing basalt based on the SRM acquisition experiments, remanence measurements for divided subsamples, and impact simulations. The SRM properties systematically change with increasing pressure, and three distinctive aspects are recognized at different pressure ranges: (1) constant intensity below 0.1 GPa, (2) linear trend as intensity is proportional to pressure up to 1.1 GPa, and (3) constant intensity and increasing stability above 1.9 GPa. The SRM intensity and stability structures suggest that the crustal rocks containing the single-domain titanomagnetite originally had an SRM intensity structure according to the distance from the impact point, which changed depending on the remanence stability after the impact.

1	Shock remanent magnetization intensity and stability structures of single-domain
2	titanomagnetite-bearing basalt sample
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## 17 Abstract

18 Knowledge of the shock remanent magnetization (SRM) structure is crucial to interpret 19 the spatial changes in magnetic anomalies observed over the impact crater. This study 20reports SRM intensity stability single-domain the and structures of 21titanomagnetite-bearing basalt based on the SRM acquisition experiments, remanence 22measurements for divided subsamples, and impact simulations. The SRM properties 23systematically change with increasing pressure, and three distinctive aspects are 24recognized at different pressure ranges: (1) constant intensity below 0.1 GPa, (2) linear 25trend as intensity is proportional to pressure up to 1.1 GPa, and (3) constant intensity 26 and increasing stability above 1.9 GPa. The SRM intensity and stability structures 27suggest that the crustal rocks containing the single-domain titanomagnetite originally 28had an SRM intensity structure according to the distance from the impact point, which 29changed depending on the remanence stability after the impact.

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## 31 **1. Introduction**

Shock remanent magnetization (SRM) is acquired as a result of the shock

33	wave propagation in a magnetic field (Nagata, 1971). A clear relationship between the
34	formation ages and magnetic anomaly intensities observed over the impact craters on
35	terrestrial planets (e.g., Acuña et al., 1999; Lillis et al., 2008; Mitchell et al., 2008)
36	indicates the SRM acquisition and/or the impact-induced demagnetization of crustal
37	rocks at the time of impact events. The SRM records of impact craters are vital in
38	reconstructing the evolution of the planetary field. Knowledge of a three-dimensional
39	distribution of the SRM intensity is crucial for interpreting the spatial change in
40	magnetic anomalies observed over the crater and reconstructing the paleo-planetary
41	field based on the anomaly data. However, the intensity distribution is an unexplained
42	phenomena concerning SRM properties owing to the lack of subsample magnetization
43	measurements for the experimental SRM-imparted samples.
44	Investigations of the SRM acquisition and measurement of the whole samples
45	showed that the SRM intensities of the natural basalt, Apollo 12 crystalline rocks
46	(Nagata, 1971), and basalt samples containing both the single-domain (SD) and
47	multidomain (MD) titanomagnetite (Pohl et al., 1975) were proportional to the applied
48	pressure. This suggests that the SRM intensity depends on the location in the shocked

49	samples owing to the variation in the pressure and temperature during the shock
50	propagation. Srnka et al. (1979) conducted an impact-induced SRM acquisition
51	experiment with natural remanent magnetization- (NRM) bearing basalt plate samples
52	containing MD titanomagnetite, and measured the remanence intensities of core
53	samples drilled from the shocked basalt. They qualitatively demonstrated that the SRM
54	intensities decreased with increasing distance from the impact point. Gattacceca et al.
55	(2008) conducted laser-induced SRM acquisition experiments using pseudo-SD (PSD)
56	titanomagnetite-bearing basalt and MD magnetite-bearing microdiorite samples. They
57	cut cylindrical samples that were 10 mm high and 9.5 mm in diameter into
58	parallelepipedic subsamples with a thickness of 1 mm and measured the SRM
59	intensities. The SRM acquisitions were homogeneous in the cylindrical samples, and
60	this was further supported by the superconducting quantum interference device
61	(SQUID) microscope measurement for the SRM-bearing basalt sample (Gattacceca et
62	al., 2010). Although the SRM structure should depend on the composition and magnetic
63	domain state of magnetic minerals, there is no consensus on the SRM structure owing to
64	limited papers, and further investigation is clearly required.

65	To investigate the SRM intensity and stability structures using a magnetically
66	well-characterized basalt sample bearing fine-grained SD titanomagnetite, we
67	conducted the newly designed SRM acquisition experiments and remanence
68	measurements for cube-shaped subsamples cut from the SRM-imparted samples. The
69	pressure and temperature changes during the shock wave propagation were estimated
70	from the impact simulations. Based on the SRM experiments, remanence measurements,
71	and impact simulations, this study reports the relationships between SRM properties and
72	pressure/temperature changes during the shock wave propagation.
73	
74	2. Method
75	2.1 Experimental sample
76	A natural basalt sample (Linxi, Inner Mongolia) was used for the SRM
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76 77 78	A natural basalt sample (Linxi, Inner Mongolia) was used for the SRM experiments. The basalt consists of olivine phenocrysts approximately 0.6 mm in diameter and fine-grained plagioclase, clinopyroxene, olivine, glass, and opaque
76 77 78 79	A natural basalt sample (Linxi, Inner Mongolia) was used for the SRM experiments. The basalt consists of olivine phenocrysts approximately 0.6 mm in diameter and fine-grained plagioclase, clinopyroxene, olivine, glass, and opaque minerals in the groundmass (Figure S1). The NRM intensity of the basalt sample was

81	field demagnetization (AFD) treatment of 80 mT (Figure S2). A strong-field
82	thermomagnetic curve shows a Curie temperature ( $T_{\rm C}$ ) of 237 °C (Figure S3a) and
83	zero-field cooling and field cooling remanence curves of low-temperature remanence
84	measurements show a remanence loss at approximately 60 K (Figure S3b), indicating
85	that Ti-rich titanomagnetite is the main remanence carrier of the basalt sample (Hunt et
86	al., 1995; Moskowitz et al., 1998). The ulvöspinel content $x$ (Fe <sub>3-x</sub> Ti <sub>x</sub> O <sub>4</sub> ) estimated from
87	the $T_{\rm C}$ value was $x = 0.51$ (Hunt et al., 1995). Magnetic hysteresis parameters and
88	first-order reversal curves (FORC) indicate the presence of SD grains with slight
89	magnetostatic interactions (Day et al., 1977; Roberts et al., 2000). The titanomagnetite
90	concentration in the basalt samples is estimated as 0.7 wt% using the saturation
91	magnetization values of the magnetic hysteresis loop.
92	
93	2.2 Shock remanence acquisition experiment
94	A two-stage light gas gun at the Institute of Space and Astronautical Science
95	(ISAS) of the Japan Aerospace and Exploration Agency (JAXA) was used for the SRM

96 acquisition experiments. A schematic diagram of the experimental system for SRM

97	acquisition is shown in Figure S4. A three-layered magnetic shield with a cylindrical
98	form was set in a vacuum experimental chamber connected to a two-stage light gas gun.
99	The diameter and length of the shield were 32 and 100 cm, respectively. The residual
100	field in the shield was less than 0.3 $\mu T\!.$ A solenoid coil with a diameter of 26 cm was
101	placed in the magnetically shielded cylinder.
102	The basalt samples were shaped into cylinders that were 10 cm both in
103	diameter and length, and used as targets in the SRM acquisition experiments. A
104	cylindrical coordinate system was used to describe the shock remanence. The sample's
105	cylindrical axis and radial directions were defined as the Z- and $R$ -axes, respectively
106	(Figure S5). The basalt sample was placed at the center of the solenoid coil, and the
107	basalt cylinder, solenoid coil, and magnetic shield were coaxially placed in the
108	experimental chamber (Figure S4). The Z-axis was set in the direction parallel to the
109	projectile trajectory. Before the shock remanence experiments, the basalt samples were
110	subjected to a one-axial (Z-axis) AFD of 80 mT using a DEM-8601C AF demagnetizer
111	(Natsuhara-Giken) comprising a large solenoid coil with a diameter of 12 cm.
112	Two SRM acquisition experiments were conducted under different applied

113	field conditions (Table S1). An aluminum sphere with a diameter of 2 mm was used as
114	the projectile, and a nylon slit sabot was used to accelerate the projectile (Kawai et al.,
115	2010). The impact velocity was set to approximately 7 km/s, and the impact angle was
116	fixed at 90°, measured from the top flat surface of the basalt cylinder. The magnetic
117	fields of 0 and 100 $\mu$ T were applied during the shock experiments, and the applied field
118	direction under the 100 $\mu$ T condition was a positive direction in the Z-axis. The spatial
119	change in the magnetic field intensity around the basalt cylinder was below 4% (Figure
120	S5).
121	

#### 1222.3 Remanence measurement

After the SRM acquisition experiments, the target samples were cut into 123124cubes using rock cutters. Slabs that were 3 mm thick and 24 mm wide were cut from the 125basalt cylinder. These were subsequently pasted on glass slides using a cyanoacrylate 126adhesive and divided into the oriented cube-shaped samples approximately 3 mm in length (Figure S6). The cube-shaped basalt and glass slide underneath were separated 127using acetone. Hereinafter, the cubic samples are denoted as  $RiZ_j$ , where the indexes i 128

129	and $j$ are sample numbers from the impact point in the $R$ and $Z$ directions, respectively
130	(Figure S6b). The magnetic field distribution around the rock cutters was measured
131	using a Model 4048 Gauss Meter (F. W. BELL), and the field intensity was below 2 mT.
132	Fifteen cube samples were cut from the unshocked basalt and the intensities of
133	anhysteretic remanent magnetization (ARM) for these cube samples were imparted with
134	DC and AC fields of 50 $\mu T$ and 80 mT, respectively, to evaluate the inhomogeneity. The
135	average and two standard deviations of the ARM intensity were 5.84 $\times$ $10^{-4}$ and 1.37 $\times$
136	$10^{-4}$ Am <sup>2</sup> /kg, respectively; the inhomogeneity among the 3 mm cube samples was
137	estimated to be 23%.
138	The remanence measurements were conducted using a Model 755 SQUID
139	magnetometer (2G Enterprise) at the Center for Advanced Marine Core Research,

140 Kochi University. An acrylonitrile butadiene styrene (ABS) sample holder and 141 measurement methods specially designed for single crystal remanence measurements 142 (Sato et al. 2015) were employed for the cube sample measurements. Stepwise AFD 143 treatments up to 80 mT were conducted using a DEM-95C alternating field 144 demagnetizer (Natsuhara-Giken) in the axial direction (*Z*-axis). After the SRM

145	measurements, the ARM with DC and AC fields of 50 $\mu T$ and 80 mT, respectively, was
146	measured for the 12 cube samples to evaluate the change in magnetic properties. The
147	cube samples selected for the magnetic measurements are shown in Figure S6b.
148	Considering that the laboratory IRM is below 2 mT and the residual of the original
149	NRM coercivity is higher than 80 mT, the remanence components of the coercivity
150	ranging between 2-80 mT are characterized as the SRM components in this study.
151	
152	2.4 Impact simulation
153	A series of impact simulations using the two-dimensional version of the
154	iSALE shock physics code (Amsden et al., 1980; Ivanov et al., 1997; Wünnemann et al.,
155	2006) was conducted to estimate the peak pressure $P_{\text{peak}}$ and temperature $T_{\text{peak}}$ values in
156	the SRM acquisition experiments. The impact velocity and shapes of the projectile and
157	target in the simulation were set to the same value achieved in the SRM acquisition
158	experiments. The Tillotson EOS with parameters pertaining to aluminum (Tillotson,
159	1962) and the analytical equations of state (ANEOS, Thompson and Lauson, 1972) with
160	parameters set for basalt were employed for the projectile and target respectively. A

161 parameter set of ANEOS for basalt was constructed by fitting the experimental data of 162 Sekine et al. (2008) and listed in Table S4. The ROCK model in the iSALE package 163 (Collins et al., 2004), with parameters set for basalt (Bowling et al., 2020), was 164 employed to treat the elasto-plastic behavior in the basalt target. The end time of the 165simulation was set to the time taken for a generated compressional wave to sweep the 166 entire target. Lagrangian tracer particles were inserted into each computational cell such 167 that the pressure and temperature values in the simulation were stored on the tracers. The mass-weighted averaged values of the  $P_{\text{peak}}$  and  $T_{\text{peak}}$  in each 3 mm cube region 168 169 were calculated to compare the calculated pressure and temperature changes with the 170 experimentally measured SRM properties. Further details of the impact simulation are 171provided in the Supporting Information. 172

## 173 **3. Results**

A representative result for the stepwise AFD treatment of the SRM component is shown in Figure S7 as an orthogonal vector plot. The SRM component is confirmed to be a single component in one direction as they linearly decreased in all the 177 cube samples. The average SRM directions deviated by  $3^{\circ}$  from the direction of the 178 applied field and the 95% confidence limit ( $\alpha_{95}$ ) was estimated to be  $3^{\circ}$ . The SRM was 179 likely aligned to the applied field direction, although the preliminary orienting method 180 for the cube samples had a large orientation uncertainty.

The SRM intensity was calculated as  $|J_{2 \text{ mT}} - J_{80 \text{ mT}}|$ , where  $J_{X \text{ mT}}$  is the 181 182remanence vector at the X mT AFD step, and is plotted as a function of the distance 183 from the impact point in Figure 1a. The SRM intensity in the case with a zero field shows an almost constant value (ranging from  $0.06-0.89 \times 10^{-4}$  Am<sup>2</sup>/kg and an average 184 of  $0.40 \times 10^{-4}$  Am<sup>2</sup>/kg) in a random direction. In contrast, the SRM intensity in the case 185186 with an applied field of 100 µT systematically changes with the distance; the maximum 187 is at approximately 10-15 mm from the impact point, which subsequently decreases 188monotonically with increasing distance. The systematic change in the SRM intensity is 189 also clear in the two-dimensional intensity map (Figure 1c). The SRM intensity in the 190 case with an applied field of 100  $\mu$ T is larger than that of zero, except for the two cube 191 samples. Considering the inhomogeneities in the magnetic minerals among the 3 mm 192cube samples (23%) and applied field intensities during the SRM acquisition experiment (4%), it was confirmed that the basalt sample acquired the remanent
magnetization as a result of shock wave propagation in an applied magnetic field of 100
µT.

196 A representative AFD curve for the SRM component of the cube sample in the case with an applied field of 100  $\mu$ T is shown in Figure 2 with the AFD curves of 50 197 198µT thermoremanent magnetization (TRM) and 500 mT IRM imparted for a 1-inch core 199 basalt sample. The stability of the SRM component varies in the cube samples; the AFD 200 curve of the SRM component was as stable as the TRM in one sample, while it was 201magnetically softer than the IRM in another sample. The stability of the SRM 202component was evaluated as  $|J_{14 \text{ mT}} / J_{2 \text{ mT}}|$  and plotted as a function of the distance from 203 the impact point in Figure 1b. A two-dimensional stability map of the SRM component 204is shown in Figure 1d. The stability monotonously decreased with increasing distance 205from the impact point (within approximately 15 mm), and likely converged 0.3–0.4 mm 206 from the 15 mm.

207 The average and two standard deviations of the ARM intensity values for the 208 selected cube samples were  $5.73 \times 10^{-4}$  and  $0.88 \times 10^{-4}$  Am<sup>2</sup>/kg, respectively. There

209	was no significant difference between the ARM intensity values for the shocked and
210	original samples, indicating that the magnetic properties of the basalt sample used in
211	this study did not undergo alteration owing to the shock wave propagation.
212	The results of the impact simulations are illustrated as a two-dimensional map
213	for the $P_{\text{peak}}$ and $T_{\text{peak}}$ values in Figures S11a and 11b. The target basalt sample
214	experienced a $P_{\text{peak}}$ ranging from 10 GPa near the impact point to below 0.1 GPa at the
215	bottom of the basalt cylinder in the SRM acquisition experiment. A significant
216	temperature rise was restricted to the region within 10 mm from the impact point, and
217	the target basalt sample experienced a $T_{\text{peak}}$ of up to 600 K in the region. The averaged
218	$P_{\text{peak}}$ of the cube sample monotonously increased with increasing average $T_{\text{peak}}$ (Figure
219	S12a).
220	
221	4. Discussion
222	The SRM intensity and stability are plotted as functions of the average $P_{\text{peak}}$
223	(Figure 3). When the $P_{\text{peak}}$ ranges below 1.1 GPa, the SRM intensity linearly changes

224 with the  $P_{\text{peak}}$ , while the samples with a  $P_{\text{peak}}$  higher than 1.9 GPa deviate from the

225 linear trend. The cube samples showing the linear trend did not experience a 226 temperature rise during the shock wave propagation ( $T_{peak}$  less than 315 K), while the 227 deviating samples experienced a significant temperature change ( $T_{peak}$  values between 228 340–590 K). The  $P_{peak}$  dependence of the SRM intensity was calculated for the sample 229 with a  $P_{peak}$  below 1.1 GPa that is given as

$$\frac{J_{SRM}}{Am^2kg^{-1}} = 3.50 \times 10^{-4} \times \frac{P_{peak}}{GPa} + 1.18 \times 10^{-4} \ (1),$$

where  $J_{\text{SRM}}$  is the SRM intensity. Assuming the proportionality in the applied field 230231intensity, the efficiencies for the SRM acquisition are estimated to be 3.50 and  $5.0 \times 10^2$  $Am^{2}kg^{-1}T^{-1}GPa^{-1}$  for the basalt and titanomagnetite contained in the basalt sample, 232respectively. The efficiency for the TRM acquisition is estimated to be  $46.0 \text{ Am}^2 \text{kg}^{-1} \text{T}^{-1}$ , 233and the SRM acquisition efficiency when the  $P_{\text{peak}}$  is 1 GPa is 7.61% of that of the TRM. 234The intercept coefficient of  $1.18 \times 10^{-4}$  Am<sup>2</sup>/kg in the linear regression is larger than the 235236 SRM intensity of the zero field. Although the SRM structure depends on the nature of 237the magnetic mineral (composition and domain state), the decreasing trend in the SRM intensity is consistent with the mixture of SD and MD titanomagnetite (Pohl et al., 2382391975) and MD titanomagnetite (Srnka et al., 1979), while the SRM properties were

241	Regarding the origin of SRM observed in this study, multiple dominant
242	factors can be described for the three different aspects: the constant $J_{\text{SRM}}$ below 0.1 GPa
243	$J_{\text{SRM}}$ proportional to the $P_{\text{peak}}$ up to 1.1 GPa, and the $J_{\text{SRM}}$ deviating from the linear trend
244	above 1.9 GPa. The basalt sample used in this study did not experience the alteration of
245	magnetic properties due to the shock wave propagation. This is consistent with previous
246	studies showing that changes in magnetic properties after the shock experiment were
247	distinct above 10 GPa (Gattacceca et al., 2007; Bezaeva et al., 2016). Thus, the above
248	characteristics arose from the magnetically reversible changes during the shock wave
249	propagation.
250	The linear trend up to 1.1 GPa likely arose from pressure effects. In the case
251	of grains exhibiting uniaxial magnetic anisotropy under a uniaxial stress $\sigma$ applied

252 parallel to the easy direction of anisotropy, the uniaxial stress effect on the

253 microcoercivity  $H_{\rm K}$  is expressed as

$$H_{K}' = H_{K} - \frac{3\lambda_{s}\sigma}{\mu_{0}M_{s}} = \frac{2K_{u}}{\mu_{0}M_{s}} - \frac{3\lambda_{s}\sigma}{\mu_{0}M_{s}}$$
(2),

254 where  $H_{\rm K}$ ' is the modified microcoercivity,  $\lambda_{\rm s}$  is the averaged magnetostriction for

255	randomly oriented crystals, $\mu_0$ is the permeability of free space, $M_s$ is the spontaneous
256	magnetization, and $2K_u/\mu_0 M_s$ expresses the microcoercivity due to uniaxial anisotropy
257	(Dunlop and Özdemir, 1997). The uniaxial stress reduces the microcoercivity, although
258	the magnitude of reduction varies with the relative orientations of the stress, magnetic
259	field, and easy direction. The IRM-like AFD curves for the SRM of this region support
260	this interpretation. The SRM is probably acquired as a result of the microcoercivity
261	decrease/increase cycle due to the pressure increase/decrease cycle during the shock
262	wave propagation. In-situ measurements of the magnetic properties of titanomagnetite,
263	such as the magnetostriction constants (Nagata and Kinoshita, 1967) and coercivity (e.g.,
264	Gilder et al., 2004; Sato et al., 2015) are prospective future studies to confirm the
265	acquisition mechanism.
266	The constant $J_{\text{SRM}}$ value can be interpreted as the remanence component
267	saturated at less than 0.1 GPa. The magnetoelastic anisotropy significantly contributes
268	to the $H_{\rm K}$ in titanomagnetite with $x = 0.6$ (Dunlop and Özdemir, 1997) and is probably
269	dominant for the titanomagnetite used in this study ( $x = 0.51$ ). However, there might be

a certain number of grains with dominating shape and magnetocrystalline anisotropy,

and the microcoercivity of these grains reduce to almost zero during the uniaxialcompression below 0.1 GPa, resulting in the saturation of their remanence.

273The deviation from the linear trend apparently considers multiple factors. The 274sample of this region experienced a significant change in temperature (340-590 K). The 275SRM stability increases with increasing  $T_{\text{peak}}$  up to the TRM-like AFD curve, while the 276SRM intensity is almost unchanged or decreases slightly with increasing  $T_{\text{peak}}$ . This 277suggests that these remanences are distinctively different from the simple TRM. 278Considering the pressure effect on the  $T_{\rm C}$  of titanomagnetite as ~15 K/GPa (Schult, 1970),  $M_s$  decreases with increasing temperature, while the elevating  $T_c$  due to 279280 increasing pressure reduces the temperature effect. Time series for the shape of the 281three-dimensional energy surface considering the relative orientations of the stress (both original and shock wave), magnetic field, and titanomagnetite grain should be 282283calculated for temperature and pressure changes during the shock wave propagation in 284future studies to understand the origin of SRM for high pressure regions.

The SRM structure observed in this study has implications for the source of the magnetic anomaly observed over the Martian impact craters. There is a clear

287	relationship between the formation ages of the Martian impact craters and the intensities
288	of magnetic anomaly over the crater (Lillis et al., 2008). Additionally, the SD
289	titanomagnetite could be a possible source of the Martian magnetic anomaly (Dunlop
290	and Arkani-Hamed, 2005). Assuming that the SD titanomagnetite is the main remanence
291	carrier of the Martian crust, the crustal rock acquired the SRM with varying intensities
292	and stabilities at the time of impact. Subsequently, depending on the SRM stability, the
293	SRM intensity relaxed after the impact, and its structure changed from the original
294	distribution. The remanence relaxation tends to emphasize the magnetization around the
295	crater center because of its high stability. However, a detailed relaxation calculation
296	based on the magnetic properties of titanomagnetite (e.g., Sato et al., 2018) should be
297	conducted in a future study. The three-dimensional distribution of the SRM intensity in
298	the crust probably creates a unique spatial pattern of magnetic anomalies over the
299	impact craters. Deciphering the magnetization distribution based on the experimentally
300	constructed SRM distribution model can provide information on the paleo-planetary
301	field evolution.

## **5. Conclusion**

304	This study conducted SRM acquisition experiments and remanence
305	measurements for cube-shaped subsamples using SD titanomagnetite-bearing basalt
306	samples to understand the SRM intensity and stability structures. The SRM intensity
307	and stability systematically change with distance from the impact point. In addition to
308	the SRM experiments, impact simulations were conducted to estimate the pressure and
309	temperature changes during the shock wave propagation and compare the calculated
310	pressure and temperature changes with the observed SRM properties. Three distinctive
311	aspects of SRM properties are recognized at different pressure ranges: (1) the SRM
312	intensity is almost constant below 0.1 GPa, (2) the SRM intensity linearly increases
313	with increasing pressure up to 1.1 GPa, and (3) the SRM intensity is almost constant,
314	while the SRM stability increases with increasing pressure above 1.9 GPa. Regarding
315	the SRM acquisition mechanisms, the pressure effect was likely dominant below 1.1
316	GPa, while multiple factors can be considered in the high-pressure range. The
317	systematic changes in the SRM intensity and stability suggest that the crustal rocks
318	containing the SD titanomagnetite had an SRM intensity structure at the time of impact,

319 and this structure changed subsequently.

320

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335	pacl	kage has been available from Thompson SL, Lauson HS, Melosh HJ, Collins GS
336	and	Stewart, ST, M-ANEOS: A Semi-Analytical Equation of State Code, Zenodo,
337	http	://doi.org/10.5281/zenodo.3525030.
338		
339	Ref	erences
340	1.	Acuña, M. H., Connerney, J. E. P., Lin, R. P., Mitchell, D., Carlson, C. W.,
341		McFadden, J., et al. (1999). Global distribution of crustal magnetization discovered
342		by the Mars Global Surveyor MAG/ER experiment. Science, 284(5415), 790–793.
343	2.	Amsden, A., Ruppel, H., and Hirt, C. (1980). SALE: A simplified ALE computer
344		program for fluid flow at all speeds, Los Alamos National Laboratories Report,
345		LA-8095:101p.
346	3.	Bezaeva, N. S., Swanson-Hysell, N. L., Tikoo, S. M., Badyukov, D. D., Kars, M.,
347		Egli, R., et al. (2016). The effects of 10 to >160 GPa shock on the magnetic
348		properties of basalt and diabase. Geochemistry, Geophysics, Geosystems, 17,
349		4753–4771.

350 4. Bowling, T. J., Johnson, B. C., Wiggins, S. E., Walton, E. L., Melosh, H. J., and

- Sharp, T. G. (2020). Dwell time at high pressure of meteorites during impact
  ejection from Mars. Icarus, 343, 113689.
- 353 5. Collins, G. S., Melosh, H. J., and Ivanov, B. A. (2004). Modeling damage and
- deformation in impact simulations. Meteoritics and Planetary Science, 39, 217–
  231.
- 356 6. Day, R., M. Fuller, and V. A. Schmidt (1977), Hysteresis properties of
- 357 titanomagnetites: Grain-size and compositional dependence, Phys. Earth Planet.
  358 Inter., 13, 260–267.
- 359 7. Dunlop, D. J., and Arkani-Hamed, J. (2005). Magnetic minerals in the Martian
- 360 crust. Journal of Geophysical Research, 110, E12S04.
- 361 8. Dunlop, D. J., and Ö. Özdemir (1997), Rock Magnetism: Fundamentals and
- 362 Frontiers, Cambridge University Press, 573 pp.
- 363 9. Gattacceca, J., A. Lamali, P. Rochette, M. Boustie, and L. Berthe (2007), The
- 364 effects of explosive-driven shocks on the natural remanent magnetization and the
- 365 magnetic properties of rocks, Phys. Earth Planet. Inter., 162, 85–98.
- 366 10. Gattacceca, J., L. Berthe, M. Boustie, F. Vadeboin, P. Rochette, and T. De

367	Resseguier (2008), On the efficiency of shock magnetization processes, P	'hys.
368	Earth Planet, Inter., 166, 1–10.	

- 369 11. Gattacceca, J., M. Boustie, E. Lima, B. P. Weiss, T. de Resseguier, and J.-P.
- 370 Cuq-Lelandais (2010), Unraveling the simultaneous shock magnetization and
- demagnetization of rocks, Phys. Earth Planet. Inter., 182, 42–49.
- 372 12. Gilder, S., M. LeGoff, J.-C. Chervin, and J. Peyronneau (2004), Magnetic
- properties of single and multi-domain magnetite under pressures from 0 to 6 GPa,
- Geophys. Res. Lett., 31, L10612.
- 13. Hunt, C.P., Moskowitz, B.M., and Banerjee, S.K. (1995), Magnetic properties of
- 376 rocks and minerals, in Rock Physics and Phase Relations: A Handbook of Physical
- 377 Constants, Vol. 3, pp. 189–204, ed. Ahrens, T.J., American Geophysical Union.
- 14. Ivanov, B. A., Deniem, D., and Neukum, G. (1997). Implementation of dynamic
- 379 strength models into 2-D hydrocodes: Applications for atmospheric breakup and
- impact cratering. International Journal of Impact Engineering, 20(1–5), 411–430.
- 381 15. Kawai, N., K. Tsurui, S. Hasegawa, and E. Sato (2010), Single microparticle
- 382 launching method using two-stage light-gas gun for simulating hypervelocity

383		impacts of micrometeoroids and space debris, Rev. Sci. Instrum., 81, 115105.
384	16.	Lillis, R. J., Frey, H. V., and Manga, M. (2008). Rapid decrease in Martian crustal
385		magnetization in the Noachian era: Implications for the dynamo and climate of
386		early Mars. Geophysical Research Letters, 35, L14203.
387	17.	Mitchell, D. L., J. S. Halekas, R. P. Lin, S. Frey, L. L. Hood, M. H. Acuña, and A.
388		Binder (2008), Global mapping of lunar crustal magnetic fields by Lunar
389		Prospector, Icarus, 194, 401–409.
390	18.	Moskowitz, B.M., Jackson, M., and Kissel, C. (1998), Low-temperature mag- netic
391		behavior of titanomagnetites, Earth planet. Sci. Lett., 157, 141–149.
392	19.	Nagata, T. (1971), Introductory notes on shock remanent magnetization and shock
393		demagnetization of igneous rocks, Pure Appl. Geophys., 89, 159–177.
394	20.	Nagata, T., and H. Kinoshita (1967), Effect of hydrostatic pressure on
395		magnetostriction and magnetocrystalline anisotropy of magnetite, Phys. Earth
396		Planet. Inter., 1, 44–48.
397	21.	Pohl, J., U. Bleil, and U. Hornemann (1975), Shock magnetization and
398		demagnetization of basalt by transient stress up to 10 kbar, J. Geophys., 41, 23–41.

399	22.	Roberts, A. P., C. R. Pike, and K. L. Verosub (2000), First-order reversal curve
400		diagrams: A new tool for characterizing the magnetic properties of natural samples,
401		J. Geophys. Res., 105, 28,461–28, 475.
402	23.	Sato, M., S. Yamamoto, Y. Yamamoto, Y. Okada, M. Ohno, H. Tsunakawa, and S.
403		Maruyama (2015), Rock-magnetic properties of single zircon crystals sampled
404		from the Tanzawa tonalitic pluton, central Japan, Earth Planets Space, 67, 150.
405	24.	Sato, M., Yamamoto, Y., Nishioka, T., Kodama, K., Mochizuki, N., Usui, Y., and
406		Tsunakawa, H. (2015). Pressure effect on magnetic hysteresis parameters of
407		single-domain magnetite contained in natural plagioclase crystal. Geophysical
408		Journal International, 202(1), 394–401.
409	25.	Sato, M., Yamamoto, Y., Nishioka, T., Kodama, K., Mochizuki, N., Ushioda, M.,
410		Nakada, R., and Tsunakawa, H. (2018) Constraints on the source of the Martian
411		magnetic anomalies inferred from relaxation time of remanent magnetization.
412		Geophysical Research Letters, 45, 6417–6427.
413	26.	Schult, A. (1970). Effect of pressure on the Curie temperature of titanomagnetites
414		[(1-x)·Fe3O4-x·TiFe2O4]. Earth and Planetary Science Letters, 10(1), 81–86.

415	27.	Sekine, T., T. Kobayashi, M. Nishio, and E. Takahashi (2008), Shock equation of
416		state of basalt, Earth Planets Space, 60, 999–1003.
417	28.	Srnka, L., G. Martelli, G. Newton, S. Cisowski, M. Fuller, and R. Schaal (1979),
418		Magnetic field and shock effects and remanent magnetization in a hypervelocity
419		impact experiment, Earth Planet. Sci. Lett., 42, 127-137.
420	29.	Thompson, S. L., and H. S. Lauson (1972), Improvements in the Chart-D radiation
421		hydrodynamic code III: Revised analytical equation of state, pp. SC-RR-71 0714
422		119 pp., Sandia Laboratories, Albuquerque, NM.
423	30.	Tillotson, J. H. (1962). Metallic equations of state for hypervelocity impact Rep.,
424		141 pp, DTIC Document.
425	31.	Wünnemann, K., Collins, G. S., and Melosh, H. J. (2006). A strain-based porosity
426		model for use in hydrocode simulations of impacts and implications for transient
427		crater growth in porous targets. Icarus, 180(2), 514–527.
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429	Fig	ıres

430 Figure 1. Shock remanence (a) intensity and (b) stability plotted as a function of

distance from the impact point. Two-dimensional structure map for the shockremanence (c) intensity and (d) stability.

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434 Figure 2. Stepwise alternating field demagnetization (AFD) curves for shock 435 remanences. Normalized remanence intensity is plotted as a function of applied AF field 436 peak. AFD curves of the thermoremanent magnetization (TRM) and isothermal 437 remanent magnetization (IRM) are shown in grey lines.

438

Figure 3. Shock remanence (a) intensity and (b) stability plotted as a function of peak pressure during the shock wave propagation. Grey circles indicate that the sample is subjected to peak pressure higher than 1.9 GPa. Dashed line in (a) shows a linear regression line for the sample subjected to peak pressure below 1.1 GPa. Figure 1.



Figure 2.



Figure 3.

