

# Far-field coseismic forcing of giant rockslides in the 2017 Sarpol-Zahab Earthquake (Iran)

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## Abstract

On November 12th 2017, the largest earthquake (Mw 7.3) ever recorded in the Zagros mountains occurred near the town of Sarpol-Zahab, Iran. While this region encompasses clusters of giant ancient rockslides, this seismic event is an excellent case-study to decipher the controlling factors of earthquake-induced landslides. Here, we address this issue by deriving an original earthquake-induced landslide inventory, encompassing landslides of various velocities (from rapid rockfalls to slow-moving landslides). This inventory displays clear differences in the spatial and volumetric distributions of earthquake-induced landslides, with 360 rockfalls triggered around the epicenter, and 9 giant active and ancient rockslides coseismically accelerated at locations up to 180 km from the epicenter. This distant triggering is explained by the earthquake source properties coupled with the local geological conditions. Our study documents a rare example of slow-moving landslides accelerated by an earthquake, and opens perspectives for the study of the landslide triggering over various time-scales.

1 **Far-field coseismic forcing of giant rockslides in the 2017 Sarpol-Zahab Earthquake**

2 **(Iran)**

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15  
16 **Key Points:**

- 17
- 18 • Novel approach for generating a comprehensive earthquake-induced landslide inventory  
(by combining various satellite data and methods).
  - 19 • The forcing of several giant pre-existing rockslides in the far-field (140 to 180 km) of the  
20 Sarpol-Zahab earthquake epicenter.

21 **Abstract**

22 On November 12th 2017, the largest earthquake (Mw 7.3) ever recorded in the  
23 Zagros mountains occurred near the town of Sarpol-Zahab, Iran. While this region encompasses  
24 clusters of giant ancient rockslides, this seismic event is an excellent case-study to decipher the  
25 controlling factors of earthquake-induced landslides. Here, we address this issue by deriving an  
26 original earthquake-induced landslide inventory, encompassing landslides of various  
27 velocities (from rapid rockfalls to slow-moving landslides). This inventory displays clear  
28 differences in the spatial and volumetric distributions of earthquake-induced landslides, with 360  
29 rockfalls triggered around the epicenter, and 9 giant active and ancient rockslides  
30 coseismically accelerated at locations up to 180 km from the epicenter. This distant triggering is  
31 explained by the earthquake source properties coupled with the local geological conditions. Our  
32 study documents a rare example of slow-moving landslides accelerated by an earthquake, and  
33 opens perspectives for the study of the landslide triggering over various time-scales.

34 **Plain Language Summary**

35 Landslides are one of the main secondary effects of earthquakes, with up to several  
36 thousands of landslides triggered during large magnitude earthquakes. The spatial and size  
37 distribution of these landslides is function of the earthquake source and site specificities. The  
38 factors that control earthquake-induced landslides can be diverse and combine in complex ways.  
39 In this study, we address this issue by focusing on the landslides induced by the Mw7.3 Sarpol-  
40 Zahab earthquake that struck the Zagros mountains (Iran/Irak border) on November 12th 2017.  
41 We developed an original approach to detect and monitor landslides of different velocities, from  
42 rapid rockfalls (m/s) to slow-moving landslides (m/yr to mm/yr), by using a set of various

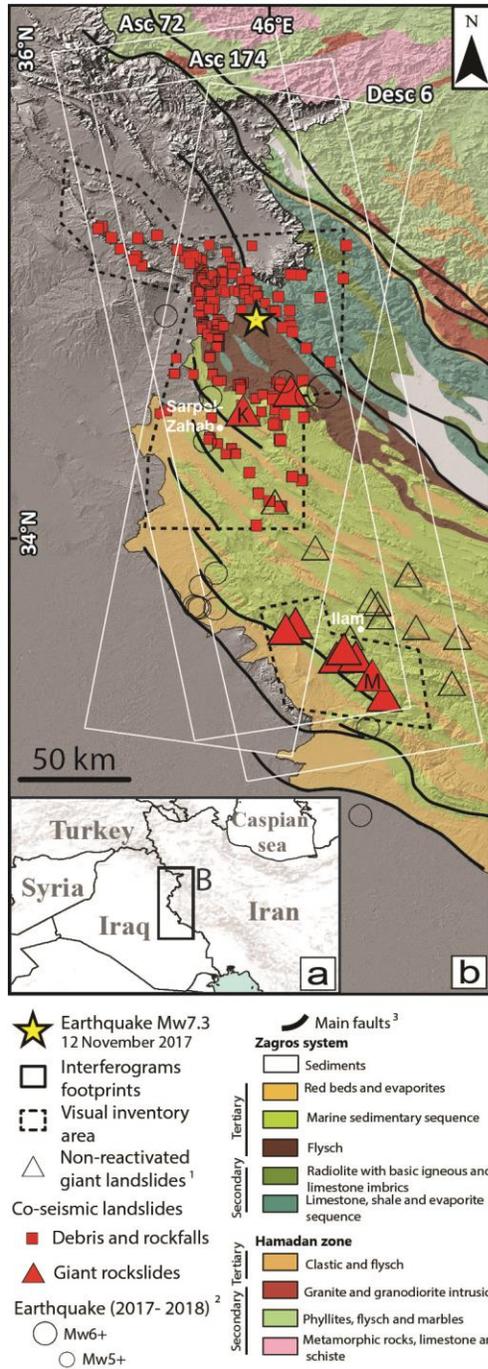
43 satellite data and techniques. The striking elements of our landslide inventory is (1) the very  
44 little number of slope movements induced by the shaking, and (2) the unprecedented detection of  
45 giant rockslides ( $\sim 10^9 \text{ m}^3$ ) in the far-field of an earthquake. This distant triggering is explained  
46 by a combination of the earthquake source properties coupled with the local geological  
47 conditions. These giant and slow-moving landslides, which have probably been active for several  
48 millennia, are unique objects for the study of earthquake forcing on landslides over time.

## 49 **1 Introduction**

50 Slope failures are one of the main secondary effects of earthquakes (Marano et al., 2010), with a  
51 large part of co- and post-seismic damage caused by landslides in mountainous areas (Fan et al.,  
52 2019; Keefer, 2002; Marc et al., 2015). Many factors can contribute to the heterogeneous spatial  
53 distribution of earthquake-induced landslides: geology (Roback et al., 2018), topography  
54 (Meunier et al., 2008), groundwater (e.g., Wang et al., 2014), and earthquake source  
55 characteristics (e.g., Gorum et al., 2011). In addition, shaking can have delayed effects on slope  
56 stability at time scales from days to years. These effects may include changes in groundwater  
57 circulation (Wang & Chia, 2008), a modification in soil permeability (Rojstaczer & Wolf, 1992),  
58 or a degradation of the mechanical properties of the slope thereby making it more susceptible to  
59 landslides in future earthquakes (Bontemps et al., 2020; Marc et al., 2015). The quantification of  
60 these controlling factors and their subsequent mechanisms is, however, based on a limited  
61 number of earthquake-induced landslide inventories (e.g. Tanyas et al., 2017), and on the  
62 instrumentation of a small number of low-velocity active landslides in seismic areas (Bontemps  
63 et al., 2020; Lacroix et al., 2014), thus highlighting the need to document and analyze more  
64 earthquake-induced landslides across a wider range of seismic and climatic settings.

65 On November 12th 2017, a Mw7.3 earthquake struck the northwestern part of the Zagros  
66 Mountains, close to the town of Sarpol-Zahab (Figure 1). This major earthquake occurred at the  
67 end of the dry season in a semi-arid area that encompasses a high density of giant paleo-  
68 landslides of volumes between 0.01-30 km<sup>3</sup> affecting mostly carbonate lithology (Ghazipour &  
69 Simpson, 2016). Following this earthquake, a few coseismic landslides of various types (debris  
70 fall, boulder/rock fall) were reported near the epicenter (Miyajima et al., 2018; Vajedian et al.,  
71 2018). The earthquake triggered the giant Mela-Kabod landslide (4-km-long, 1-km-wide)  
72 ~40 km south of the epicenter, with a coseismic displacement of ~30 m (Goorabi, 2020;  
73 Vajedian et al.,2018). No reactivation was reported for the many other giant landslides in this  
74 region.

75 The large Sarpol-Zahab earthquake provides a unique opportunity to study the forcing  
76 mechanisms of landslides under strong seismic stressing in a semi-arid region. To this end, we  
77 have used optical satellite and multi-temporal InSAR methods to establish a complete inventory  
78 of rapid (~m/s) and slow (~mm to m/yr) landslides and investigate their response to the  
79 earthquake.



80

81 **Figure 1.** (a) Study area location. (b) Coseismic landslide inventories (geology adapted from the  
 82 1:2,500,000 tectonic map of the National Iranian Oil Company, 1978). Empty triangles  
 83 correspond to giant landslides mapped by Ghazipour and Simpson (2016). The landslides most  
 84 mentioned in this study are the Mela-Kabod landslide (K) and the Mehr rockslide (M).

85 Earthquakes ( $M_w > 5$ ; period 2017-2018) are reported from <sup>2</sup>US Geological Survey. The main  
86 faults are from <sup>3</sup>Hessami et al. (2003).

## 87 **2 Geological settings**

88 The Zagros fold and thrust belt (ZFTB, Figure 1) formed in response to the collision  
89 between the Arabian and Eurasian plates, which initiated at  $\sim 35$  Ma (McQuarrie et al., 2003) and  
90 continues at the present day with a convergence rate of 8-23 mm/yr (Masson et al., 2014). This  
91 N-S convergence produced (1) a succession of asymmetrical, NW-trending, inverted folds  
92 affecting a 7-12 km thick pile of sedimentary rocks, comprising limestones, siltstones, shales and  
93 salts dating from the Cenozoic to Palaeozoic, and (2) major NW-striking active thrust faults  
94 associated with significant seismicity ( $M_w > 6$ ) at the interface between basement and  
95 sedimentary cover units at  $\sim 20$  km depth (Figure 1b; Tavani et al., 2018). The high relief of the  
96 Zagros, culminating at about 3650 m altitude, is strongly controlled by resistant calcareous  
97 anticlines, which form a succession of ridges separated by narrow valleys developed along the  
98 synclinal axes. This semi-arid zone receives about 230 mm/yr of rainfall annually, which falls  
99 mostly between November and May. The  $M_w 7.3$  Sarpol-Zahab earthquake (12/11/2017)  
100 occurred along a near-horizontal blind thrust fault, located between 14-20 km depth (Barnhart et  
101 al., 2018; Chen et al., 2018; Gombert et al., 2019; Nissen et al., 2019). Previous studies inferred:  
102 (1) a fault rupture of  $\sim 50$ -km-long and  $\sim 30$ -km-wide with a maximum coseismic slip of  $5.5 \pm 0.5$   
103 m, and (2) a high impulsive source with a robust southward rupture directivity which produced  
104 the largest ground motions ( $PGA = \sim 700$  cm/s<sup>2</sup>) in the Sarpol-Zahab town ( $\sim 40$  km south from  
105 the epicenter). Moreover, high horizontal peak ground accelerations ( $\sim 100$  cm/s<sup>2</sup>) were recorded  
106 up to 100 km south of the source (Mahani & Kazemian, 2018).

## 107 **3 Materials and Methods**

### 108 3.1 Optical Satellite Images Comparison and Correlation

109 A detailed coseismic landslide inventory was conducted by the visual comparison of PlanetScope  
110 satellite images (3 m resolution) acquired before and after the earthquake (19/10/2017 and  
111 13/11/2017) covering an area of 12,000 km<sup>2</sup> centered on the epicenter (see footprint on Figure 1  
112 and Figure S1 in the supporting information). We typically detected and mapped new rockfall  
113 scars and debris deposits induced by the earthquake. We also calculated earthquake-induced  
114 horizontal ground displacement from the correlation of pre/post-earthquake optical satellite  
115 images (Leprince et al., 2007), using both SPOT6/7 images (orthorectified at 1.5 m resolution,  
116 following Beyer et al. (2018) see Section S1 for further details), and PlanetScope satellite  
117 images-see Table S1, Figure S1, and text S1.

### 118 3.2 InSAR Processing

119 To detect and monitor smaller ground motions associated with earthquake-induced slow-moving  
120 landslides, we derived Sentinel-1 InSAR time-series (Doin, et al., 2011) for each landslide using  
121 72 Sentinel-1 SAR images spanning a 20 month time period (beginning 10 months before the  
122 earthquake). We generated differential interferograms using the NSBAS (New Small BAseline  
123 Subset) (Doin et al., 2011) processing chain based on the ROI\_PAC software (Rosen et  
124 al., 2004). We used two ascending (174 and 72, subswath iw1 for both) and one descending (6,  
125 subswath iw2) tracks of Sentinel 1A and 1B covering an area of 33,500 km<sup>2</sup> (see footprint on  
126 Figure 1b and Table S1), with a revisit time of 12 days. Initially, we re-sampled all secondary  
127 SLC (Single Look Complex) images in a single reference SLC geometry and co-registered  
128 secondary to reference using precise orbits and an ASTER digital elevation model (30 m

129 resolution), combined with empirical offsets between secondary and reference images. Then, a  
130 small baseline subset is defined using temporal and perpendicular baseline constraints (Doin, et  
131 al., 2011). After calculating differential interferograms we corrected them from atmospheric  
132 delays using ERA-5 ECMWF reanalysis (Doin et al., 2009). Finally, we made an empirical  
133 correction for topographically-correlated atmospheric-delays. The coseismic interferograms were  
134 then inspected for landslide-like patterns throughout the region. The coseismic signals were too  
135 large to be unwrapped due to phase ambiguities across landslide boundaries. For this reason, we  
136 bound the amplitude of the coseismic motion using (1) the number of fringes on the coseismic  
137 interferogram (which provides a lower limit), and (2) the optical image correlation (an upper  
138 limit). For the same reason, we also analyze the time-series independently to determine the pre-  
139 and post- seismic landslide kinematics (see text S5).

### 140 3.3 Geomorphological Analysis

141 Finally, we conducted geomorphological and geological analysis coupling the stereo-derived  
142 high resolution DEMs (text S1), Google Earth satellite imagery and geological maps to better  
143 constrain the typology and failure modes of the detected landslides.

## 144 4 Results

### 145 4.1 Rockfalls

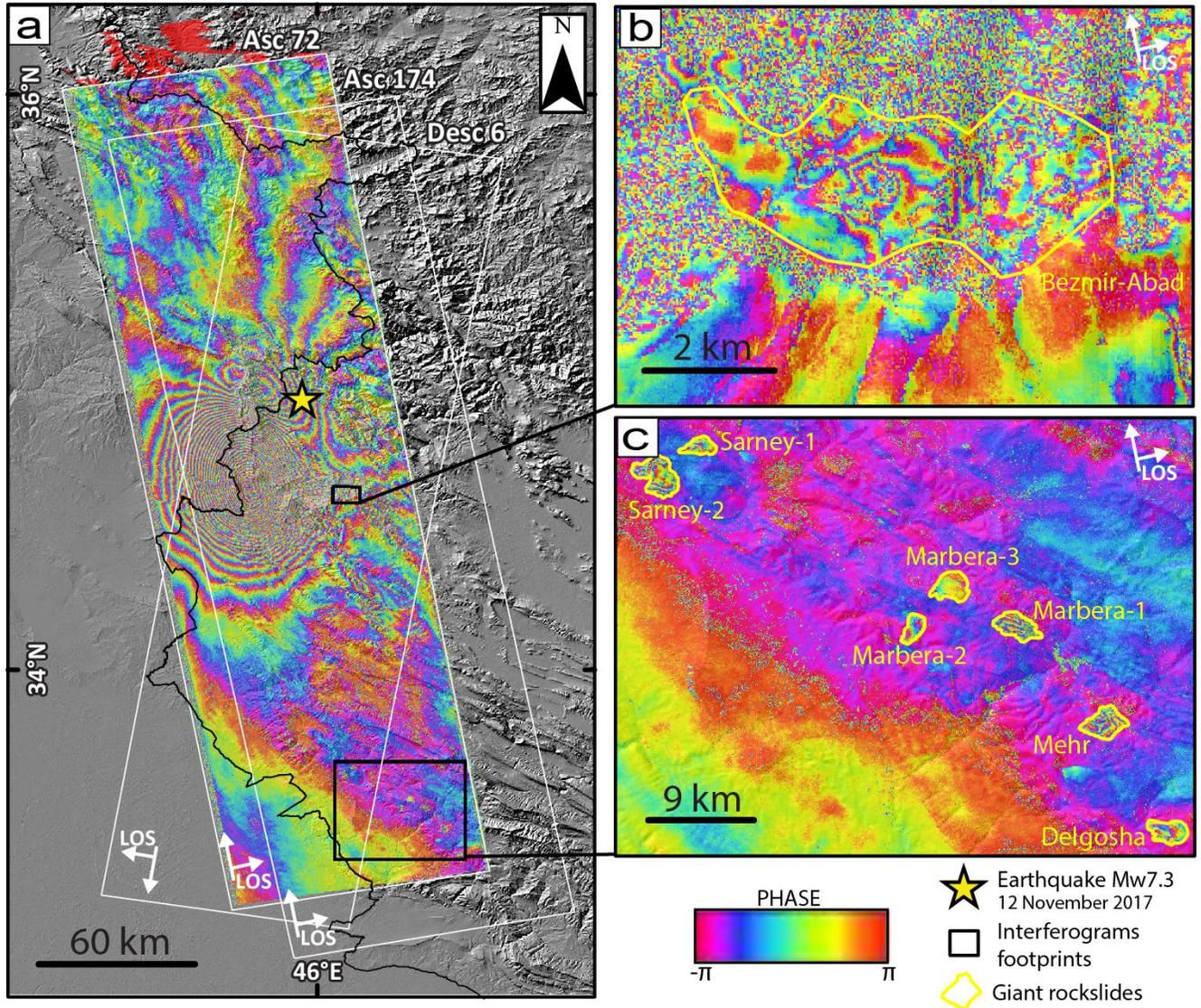
146 We map 360 coseismic rockfalls, and their associated debris cones (areas ranging between 200-  
147 and 20,000 m<sup>2</sup>), which affect mainly limestones and flyschs (Figure 1b, S2). About 85% of the  
148 detected rockfalls are concentrated within a radius of 40 km (smaller than the fault rupture  
149 length) of the epicenter (Figure S3). They mostly occurred on slopes between 40° and 80°,  
150 which are significantly steeper than the mean ~18° slope of the area (Figure S4).

## 151 4.2 Giant rockslides

## 152 4.2.1 Coseismic Detection

153 Nine instances of coseismic landslide motion were detected (Figure 1b), one from the coseismic  
154 correlation of optical images (Mela-Kabod) and eight from the coseismic interferograms analysis  
155 (areas between 2-15 km<sup>2</sup>, see Table S2). They are all located south of the epicenter, two at ~40  
156 km , and the remaining seven clustered between 140-180 km from the epicenter (Figure 1b,  
157 Table S2), in a region where no rapid slope-failures were detected in the PlanetScope imagery.  
158 The giant Mela-Kabod landslide displays a coseismic motion of about 35 m (Figure S5), a value  
159 consistent with previous estimates (Valkaniotis et al., 2018).

160 InSAR analysis reveals activity of 8 landslides, which are generally characterized by 3-4 fringes  
161 outlined by sharp phase discontinuities with the surrounding area (Figure 2) during the coseismic  
162 period. They correspond to a coseismic motion of at least 30 mm in the Line Of Sight (LOS),  
163 reaching more than 100 mm in most cases (Figure 2b, c, Table S2 and text S4). Four of the  
164 detected patterns are correlated with giant rockslides from the inventory of Ghazipour and  
165 Simpson (2016); the other four reflect newly mapped slope failures (Table S2).



166

167 Figure 2. (a) Example of a coseismic interferogram computed over the study area from the  
 168 ascending track 72 between 11/11/2017 and 17/11/2017 (the looking angle of the radar varies  
 169 between  $27.3^{\circ}$ - $32.5^{\circ}$ ). (b) and (c) zooms show the 8 landslides detected with InSAR near the  
 170 earthquake and in the southern far-field, respectively.

171

#### 4.2.2 Annual kinematics

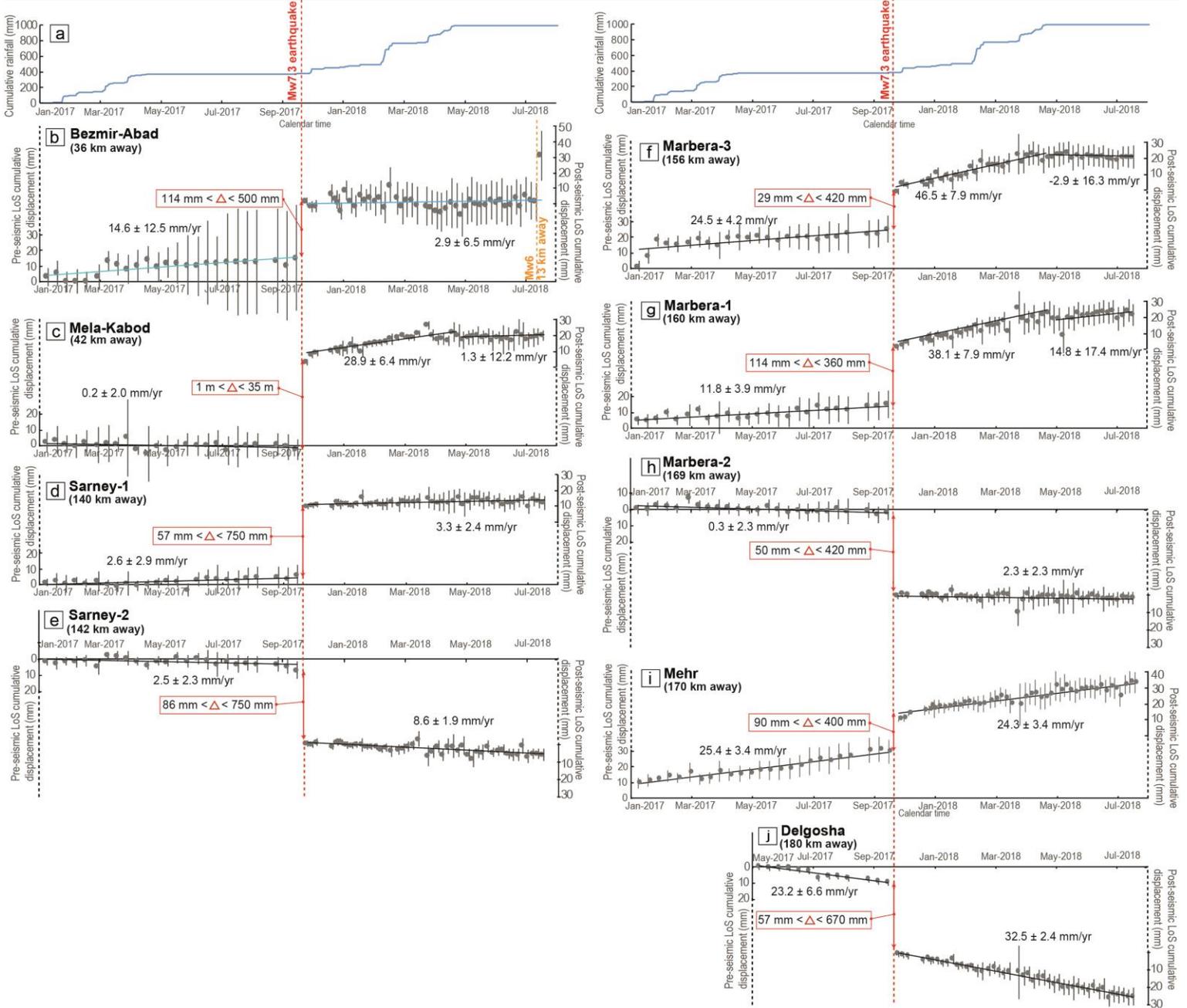
172

The cumulative LOS displacement time-series computed shows different pre- and post-seismic

173

rockslide behaviors. Rockslide velocities range between 0-25 mm/yr, and 2-46 mm/yr for the

174 pre- and post-seismic periods, respectively (Figure 3 and Table S2). The pre-seismic period  
175 shows either dormant rockslides (almost zero velocity within the error limits: Figure 3c, d, e, h)  
176 or active rockslides with constant velocities (Figure 3b, f, g, i, j). The coseismic motion is  
177 followed by a transient relaxation over 20 days clearly seen at several sites: Mela-Kabod,  
178 Marbera-3, Marbera-1 and Mehr (Figure 3c, f, g, i). Following this, three different post-seismic  
179 patterns emerge: (1) rockslides with constant post-seismic velocity equivalent to the pre-seismic  
180 one (Figure 3b, d, h, i), (2) rockslides with constant post-seismic velocity higher than the pre-  
181 seismic one (Figure 3e, j) and (3) rockslides showing a transient increase in velocity of several  
182 months before returning to their pre-seismic rates (Figure 3c, f, g). In this latter case, the  
183 succession of those two ultimate post-seismic phases coincides with the rainy and dry seasons as  
184 shown by the comparison with cumulated rainfall (Figure 3a). Finally, an offset of the last  
185 acquisition date at the end of the Bezmir Abad time-series (Figure 3b) may correspond to the co-  
186 seismic effects of a Mw6.0 earthquake that occurred 13 kilometers away (Figure 1, S7).



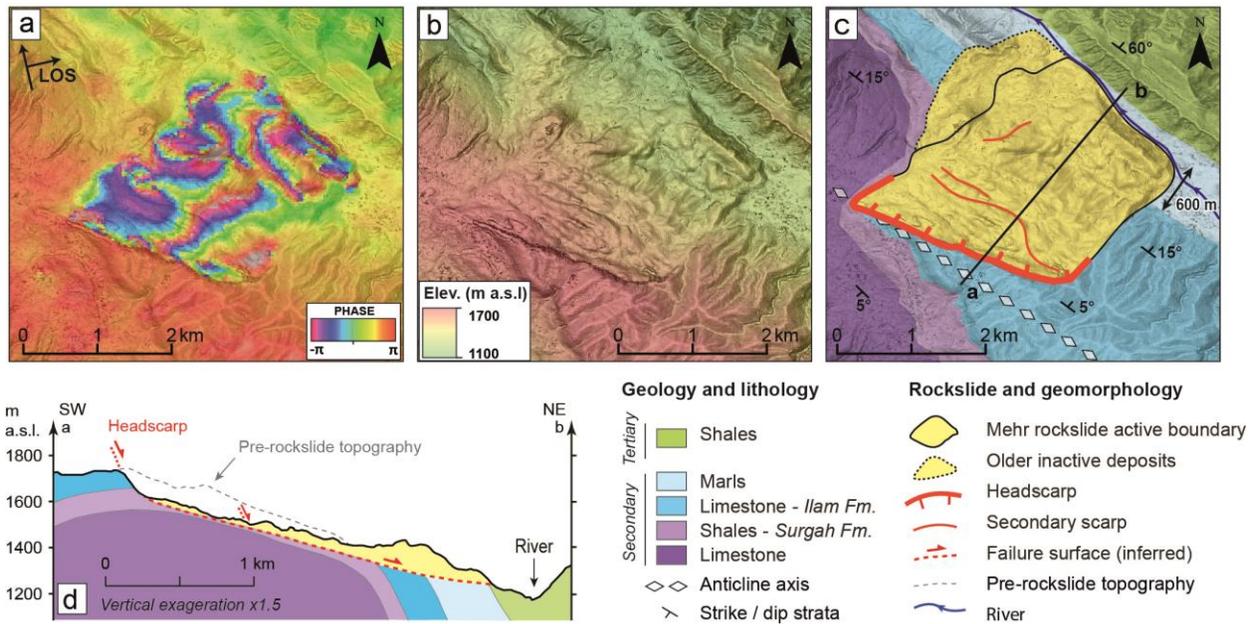
189 Figure 3. (a) Cumulative rainfall collected at the Ilam meteorological station (Figure 1b; National  
 190 Climatic Data Center). (b) to (j) show the cumulative LOS displacement time-series with the  
 191 error bars, computed for all the detected rockslides from InSAR over 18 months spanning the  
 192 Mw7.3 Sarpol-Zahab earthquake and revealing the rockslide coseismic motion ( $\Delta$ ). The pre-

193 seismic InSAR time-series for the Delgosha rockslide (j) span for only 6 months before the  
194 earthquake.

#### 195 4.2.2 Geomorphological characterization

196 Results from our geomorphological analysis of the Mehr rockslide, located 170 km south of the  
197 epicenter, are shown in Figure 4 (see Figure S6 and table S2 for detailed results of the other  
198 rockslides). The coseismic motion extent, clearly visible in the interferogram (Figure 4a),  
199 delineates a region 3 km long by 2.5 km wide, which is bounded to the SW by a ~160 m high  
200 headscarp and to the northeast by the toe of debris deposits that propagate ~600 m over the  
201 valley floor (see DEM in Figure 4b). At the SE limit, the lateral rockslide boundary is well-  
202 defined in the geomorphology. The geological map (Figure 4c) and the cross-section (Figure 4d)  
203 indicate that the rockslide occurred at the contact between Ilam limestones and overlying Surgah  
204 shales, along the northern flank of a NW-SE-striking anticline. The rockslide consists of  
205 limestone blocks sliding on the shale layer, which dips 5-15° to the NE. Comparison between a  
206 topographic profile extracted along the rockslide and another along the undisturbed slope  
207 suggests a maximum depth of 200 m for the slip surface (Figure 4d), implying a rockslide  
208 volume of ~0.5 km<sup>3</sup>.

209 Observations are similar for the other rockslides (Figure S6 and Table S2): (1) all the detected  
210 interferometric patterns match the positions of pre-existing giant rockslides with an estimated  
211 volume range from 0.16 to 2.2 km<sup>3</sup>, and (2) six of those rockslides occur at the contact between  
212 limestone and shales from the Ilam and the Surgah formations.



213

214 Figure 4. The Mehr giant rockslide (see location in Figure 1b) presented from (a) a coseismic  
 215 interferogram computed along the Sentinel-1 ascending track-72 between 11/11/2017 and  
 216 17/11/2017, (b) a pre-seismic SPOT6-7 hillshaded DEM (4 m resolution, 09/09/2014), (c) a  
 217 geological map adapted from Llewellyn (1974), (d) a cross-section built along the profile “ab”  
 218 shown in c.

219 **5 Discussion**

220 5.1 Coseismic landslide database

221 The Sarpol-Zahab earthquake induced coseismic displacements for two types of landslides: 360  
 222 small rockfalls clustered in a radius of a few tens of kilometers around the epicenter, and 9 giant  
 223 rockslides mainly located in the far-field (up to 4 times the fault length). The high concentration  
 224 of rockfalls in the epicentral area (compared to the wider zone of fault slip) can be explained by  
 225 the impulsive source (Gombert et al., 2019), which leads to stronger ground-motions close to the

226 epicenter. This spatial distribution highlights the dynamic triggering of these rockslides (e.g.,  
227 Meunier et al., 2007).

228 The number of recorded landslides is low for a Mw7.3 earthquake, which would be expected to  
229 trigger a few thousand landslides in such a mountainous area (Keefer, 2002, Tanyas et al., 2017).  
230 This low number may be explained by (1) the aridity of the region, which limits weathering and  
231 soil production (Lacroix et al., 2013, Roback et al., 2018), and (2) blind thrust faulting, which  
232 induces lower ground motions than surface rupturing earthquakes (Aki, 1987).

## 233 5.2 Far-field seismic forcing

234 The striking feature of this earthquake resides in the coseismic motion of several giant pre-  
235 existing rockslides located at epicentral distances of 140 km to 180 km (Figure 1b). Keefer  
236 (2002) documents several case studies where small landslides were triggered at large distances  
237 from the epicenters, possibly due to low seismic attenuation and/or extraordinary susceptibility  
238 of some sites. However, the huge size ( $\sim 0.1\text{-}2\text{ km}^3$ ) of the earthquake-induced rockslides here is  
239 intriguing and, to our knowledge, has never been reported so far from the epicenter.

240 Rockslide forcing in the far field south of the epicenter can first be explained by the strong  
241 directivity of the rupture toward the south (Chen et al., 2018; Gombert et al., 2019), as well as by  
242 the stronger movements perpendicular to the fault (Mahani & Kazemian, 2018), which may  
243 favor the triggering of rockslides oriented NE-SW. In the ZFTB geological context, the ground  
244 shaking can also be amplified by both local topographic and geological effects (Maufroy et al.,  
245 2015; Murphy, 2015), more specifically as the rockslides developed on flanks of anticline ridges  
246 and in mechanically heterogeneous lithologies. Calculation of a 1D resonance frequency of the  
247 destructured slump body overlying the thick rigid layer shows wave amplifications at low  
248 frequencies (text S6), around 1Hz, compatible with the frequency content of the earthquake

249 source at such large distances (Mahani & Kazemian, 2018), which can thus favor landslide  
250 triggering.

251 Interestingly, several other ancient landslides previously mapped in this region (Ghazipour &  
252 Simpson, 2016) were not reactivated during the Sarpol-Zahab earthquake (Figure 1b). However,  
253 these landslides mostly occurred in different lithologies (Oligocene and Eocene units) with a  
254 structure (thick moderately incompetent layer of calcareous shale underlying the carbonate  
255 slump body) less sensitive to sliding and seismic amplification.

256 Finally, our time-series reveal a transient increase of the post-seismic velocity compared to the  
257 pre-seismic annual velocity for at least 5 of the studied rockslides (Figure 3c, e, f, g, j). For 3 of  
258 the rockslides (Mela-Kabd, Marbera-1, Marbera-3) we document a higher velocity during the  
259 2018 rainy season (Figure 3c, f, g) followed by a subsequent decrease in velocity, eventually  
260 returning to the pre-earthquake velocity during the dry season. This observation may suggest a  
261 seasonal motion and threshold rainfall effects on the landslide kinematics (Zerathe et al., 2016).  
262 Furthermore, moderate earthquakes (Mw5-6) that occurred close to the landslides (Figure 1, S7)  
263 may have contributed to the transient velocity increase by damaging rock, thus promoting water  
264 infiltration (Bontemps et al., 2020) up to the impermeable Surgah formation. However, exploring  
265 these issues will require longer time-series as well as detailed field measurements of hydrologic  
266 and kinematical parameters (e.g. Schulz et al., 2009).

## 267 **6 Conclusion**

268 We used a new approach to generate a comprehensive inventory of earthquake-induced  
269 landslides, combining InSAR, image correlation and visual change detection. Applying it to the  
270 Mw 7.3 Sarpol-Zahab earthquake, we detected 369 coseismic landslides of different sizes and  
271 kinematics, including 360 rockfalls and 9 giant rockslides.

272 The striking element of this earthquake-induced landslide database is the coseismic motion of 9  
273 pre-existing giant rockslides (3-30 km<sup>2</sup>) in the far field, all initially generated by sliding of a  
274 thick limestone layer over a shale formation on both limbs of anticlinal structures, in directions  
275 perpendicular to the anticline axis (i.e. NE-SW). The coseismic motion of these slow-moving  
276 giant rockslides (<40 mm/yr) was at least 30 mm, reaching up to 35 m. At least three of them  
277 were also accelerated over the rainy season following the earthquake, showing either a seasonal  
278 climatic forcing or an increased rainfall infiltration enhanced by the landslide bulk damage  
279 produced by the ground shaking. We show that the coseismic motion of these rockslides may be  
280 related to a complex combination of the southward directivity of the source, the NE-SW  
281 polarization of the motions, their sensitivity to low-frequencies (~1 Hz), and the site effect due to  
282 the seismic impedance contrast on the flanks of the anticlines.

283 The detection of significant coseismic motion of several ancient giant rockslides using InSAR  
284 may open new perspectives on the understanding of large-scale gravitational deformations in arid  
285 settings. Most of the rockslides investigated here display huge cumulated headscarps (100's m);  
286 it is therefore likely that repeated large earthquakes over longer time scales constitute one of  
287 the predominant forcings for this displacement. These landslides have certainly been active over  
288 several millennia, as observed for other giant landslides of the area (Roberts & Evans,  
289 2013). Dating of landslide headscarps is therefore a key issue in understanding how earthquakes  
290 control landslide dynamics on different time scales.

## 291 **Acknowledgments, Samples, and Data**

292 This research is funded by the Lebanese National Council for Scientific Research (CNRS). The  
293 authors also thank the support of the French Spatial Agency (CNES). Supplementary figures,  
294 tables and graphs are provided in the supporting information accompanying the main text. The

295 data archiving is underway on the PerSCiDO platform. Meanwhile, the data are available at the  
296 following link <https://1drv.ms/u/s!Alv20Xzpi7Jig2gksQk9RHNuN3Ym?e=AanU5p>.

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*Geophysical Research Letters*

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Supporting Information for

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**Far-field coseismic forcing of giant rockslides in the 2017 Sarpol-Zahab Earthquake (Iran)**

5

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19

**Introduction**

20

This document provides supplementary information on the used data and methods, the

21

uncertainty analysis presented in the main text, an extended description of the results and

22

detailed explanation of the giant rockslides site effect assessment.

23

As our study area is large and partially not well documented (Iran-Iraq boundary/area of

24

conflict), the available geological maps were only: (1) a large geological map (1:2,500,000) for

25

the entire Iran and (2) a 1:250,000 geological map for the Ilam province in Iran (southern far

26

field of the earthquake). No geological maps were available for Iraq (see more details in text S1

27

and Figure S1).

## 28 **Text S1. Data Set**

29 Three kinds of DEMs were used in our study: (1) the ASTER GDEM of 30 m resolution (regional  
30 view of our study area), (2) 4 m resolution pre and post-earthquake DEMs (Table S1) that were  
31 generated in the region of the epicenter (Figure S1) from tri-stereo pairs of SPOT6-7 images  
32 using Ames-Stereo Pipeline (Beyer et al., 2018), and finally (3) a 4 m resolution pre-earthquake  
33 DEM (also generated with Ames-Stereo Pipeline using tri-stereo SPOT7 images (acquired in  
34 2014) covering the southern part of our study area, in the region of the far southern rockslides,  
35 Figure S1). See Table S1 for more details on the data.

36 SPOT data was provided via the CNES-funded ISIS program (now integrated with DINAMIS:  
37 Dispositif Institutionnel National d'Approvisionnement Mutualisé en Imagerie Spatiale).

38 The pre- and post-earthquake SPOT6-7 images (around the epicenter, Figure S1) of 1.5 m  
39 resolution were orthorectified using the high resolution DEMs generated from the same data.

40 The PlanetScope images are available as orthorectified tiles, 10 km long and 25 km wide  
41 containing four bands (blue, green, red and infrared).

42 The 72 Sentinel-1 images cover 10 months before and after the earthquake. They were acquired  
43 from ESA with Interferometric Wide Swath (IW) mode from both A and B satellites and feature  
44 a 250 km swath, a spatial resolution of 5x20 m and a repeat cycle of 12 days.

45 A 1:2,500,000 regional geological map of the Iran republic (National Iranian Oil Company,  
46 NIOC) was used in our study, alongside a more detailed 1:250,000 geological map (Llewellyn,  
47 1974), covering the Ilam region in Iran (Figure S1).

## 48 **Text S2. Methods**

49 Our working strategy aimed at detecting the maximum possible number of earthquake-induced  
50 landslides in our study area, extending 200+ km along the Iran-Iraq border. Thus, we used  
51 different methods: the scars of rapid coseismic landslides were mapped by a comparison of pre-  
52 and post-seismic Planet-scope images (Manual **visual comparison**), whereas slow-moving  
53 landslides (m/yr-mm/yr) were detected by deriving the ground deformation from optical (**Optical**  
54 **images correlation**) and radar (**Interferometric Synthetic Aperture Radar**) satellite images.

### 55 **2.1. Visual comparison**

56 To detect the rapid slope-failures, the available pre- and post-earthquake PlanetScope data were  
57 merged then compared in ArcGIS software using the "**swipe**" tool. To accomplish this inventory in the

58 best way, we used the available DEMs and the earth view base maps in order to verify that the detected  
59 landslide scars occur on topographic slopes and try to visualize them if possible.

## 60 **2.2. Optical image correlation**

61 The COSI-CORR iterative correlator was used to measure the horizontal displacements on the Earth's  
62 surface using georeferenced optical images (Leprince et al., 2007). Each correlation yields a north-  
63 south and an east-west displacement fields, as well as a signal to noise ratio map. It allows usually the  
64 detection of displacements higher than 10% of the image's pixel resolution during the time interval  
65 between the two correlated images.

66 First we correlated the mosaic of SPOT6-7 images covering the area around the epicenter (a minimum  
67 distance of 10 km and a maximum distance 75 km to the epicenter). The correlation was conducted  
68 pixel by pixel in the frequency domain using a sliding window of 64 pixels in both iterations.

69 In a second step we correlated the available PlanetScope images in the southern part of our study area  
70 (see Figure S1). The aim of this step is to see if we can detect any displacement fields on the body of  
71 the rockslides detected from the coseismic interferograms. Thus, the green bands were correlated in  
72 the frequency domain for each pixel using a sliding window of 64 pixels also for both iterations.

73 Each time, several tests were done before adopting the final sliding window sizes.

74 The obtained results were then detrended in ENVI software and corrected afterward with Matlab by  
75 eliminating the pixels of high signal to noise ratios then subtracting the median of all the displacement  
76 field from each pixel.

## 77 **2.3. Interferometric Synthetic Aperture Radar (InSAR)**

78 We generated differential interferograms using NSBAS (New Small BAseline Subset) (Doin et al., 2011)  
79 process chain based on the ROI\_PAC software (Rosen et al., 2004). See more details in the manuscript.

## 80 **Text S3. Results**

81 In total, 369 earthquake-induced landslides were detected. We divided them into two main categories:  
82 rockfalls and giant rockslides.

### 83 **3.1. Rockfalls**

84 360 scars of rockfalls were mapped around the epicenter using the visual comparison of Planet-scope  
85 images. In the following figures we will be showing an example of how we detected the scars of the  
86 debris cones (Figure S2), their density analysis (Figure S3) and their occurrence on the available slopes  
87 (Figure S4).

### 88 **3.2. Giant rockslides**

89 9 giant rockslides were detected. One of them was the Mela-kabod rockslide detected from optical  
90 images correlation that moved coseismically for about 35 m toward the south-west (Figure S5). While  
91 8 landslide-like patterns were detected from the coseismic interferograms (Figure 2) and then  
92 interpreted to be old giant rockslides. The characteristics of all the rockslides are detailed in Table S2  
93 and presented in Figure S6.  
94 However, in Ghazipur and Simpson (2016), the areas of those rockslides are systematically  
95 underestimated by up to an order of magnitude compared to the surface areas determined from our  
96 results (Table S2).

#### 97 **Text S4. Quantification of Giant Rockslides Coseismic Displacement**

98 While a coseismic movement of all the rockslides is observed in the coseismic interferograms (Figure  
99 2), its precise quantification is not possible due to (1) the sharp limits of the patterns in the coseismic  
100 interferograms, that prevent extraction of the phase ambiguity during the unwrapping process, and  
101 (2) the absence of pattern in the correlation of optical images. However, those two sources of data  
102 nevertheless provide constraints on the coseismic movement between several cm (~10 cm) in the radar  
103 LOS and 0.9 m maximum for all the rockslides (Table S2). The minimal coseismic displacement can be  
104 quantified by counting the number of fringes inside each rockslide pattern (formula: (number of  
105 fringes\*wave length)/4pi). The maximum value of the coseismic displacement can be estimated by the  
106 uncertainty of the horizontal displacement field obtained from optical PlanetScope images correlation  
107 (explained above).

#### 108 **Text S5. Time-series analysis**

109 After detecting and characterizing the slow movements triggered by the Sarpol-Zahab earthquake in  
110 our study area, we computed their radar LOS time-series for each pixel in the interferograms stack  
111 over 10 months before and after the earthquake. To do that, we divided the interferograms into pre-  
112 and post-earthquake groups, then we inverted the phase delays of the unwrapped interferograms pixel  
113 by pixel in order to solve the total phase delay, relative to the first date (Doin et al., 2011). Time-series  
114 were then constructed in Matlab using the cumulative deformation maps obtained from the inversion  
115 (one map at each date of the Sentinel-1 images). So we calculated the mean displacement over a  
116 selected window, of about 25x25 pixels on the landslide body at each date, relative to a mean  
117 displacement extracted from a surrounding stable area of hundreds of meters around it. The final  
118 displacement was computed from the differences between the two means. After that, the deviation

119 of the displacement was estimated from the mean absolute deviation of the displacement in the  
120 reference area.

121 In a next step, we calculated the pre-and post-seismic landslide mean velocities and their associated  
122 uncertainties. Each point “i” of the time-series is considered as a random variable of normal distribution  
123 ( $\mu_i$ ,  $\sigma_i$ ). 10,000 realizations of this random variable are randomly picked at each point of the  
124 time series, and the associated 10,000 pre and post velocities are calculated by a linear regression with  
125 time. The mean and standard deviation of these 10,000 velocities gives us an estimate of the mean  
126 velocity and its uncertainty.

### 127 **Text S6. Site Effect Assessment of the Giant Rockslides**

128 Six of the giant landslides affected the same 200 m thick Ilam formation (limestone) overlying a  
129 100 m thick shale layer (Surgah formation). This structure constitutes a dynamic oscillator on  
130 the thick carbonate Sarvak formation.

131 During the slide of the rockslides, the block disintegrated and dragged part of the Surgah  
132 formation, creating a highly destructive deposit with a maximum thickness  $t$  of around 150 m,  
133 consisting of a mixture of shale and limestone. The amplification of the seismic waves resulting  
134 from the earthquake is due to the seismic impedance contrast (product of the density  $\rho$  and  
135 shear wave velocity  $V_s$ ) between this deposit and the underlying, mainly calcareous,  
136 substratum. For vertically incident waves and 1D structure, the resonance frequency  $f_0$  and the  
137 corresponding amplification  $Af_0$  are given by (Kramer, 1996):

$$138 \quad f_0 = \frac{V_{SD}}{4t} \quad (1)$$

$$139 \quad Af_0 = \frac{\rho_B V_{SB}}{\rho_D V_{SD}} \quad (2)$$

140 where  $V_{SD}$  and  $V_{SB}$  are the shear wave velocities of the rockslide deposit and the bedrock,  
141 respectively, and  $\rho_D$  and  $\rho_B$  are the corresponding densities.

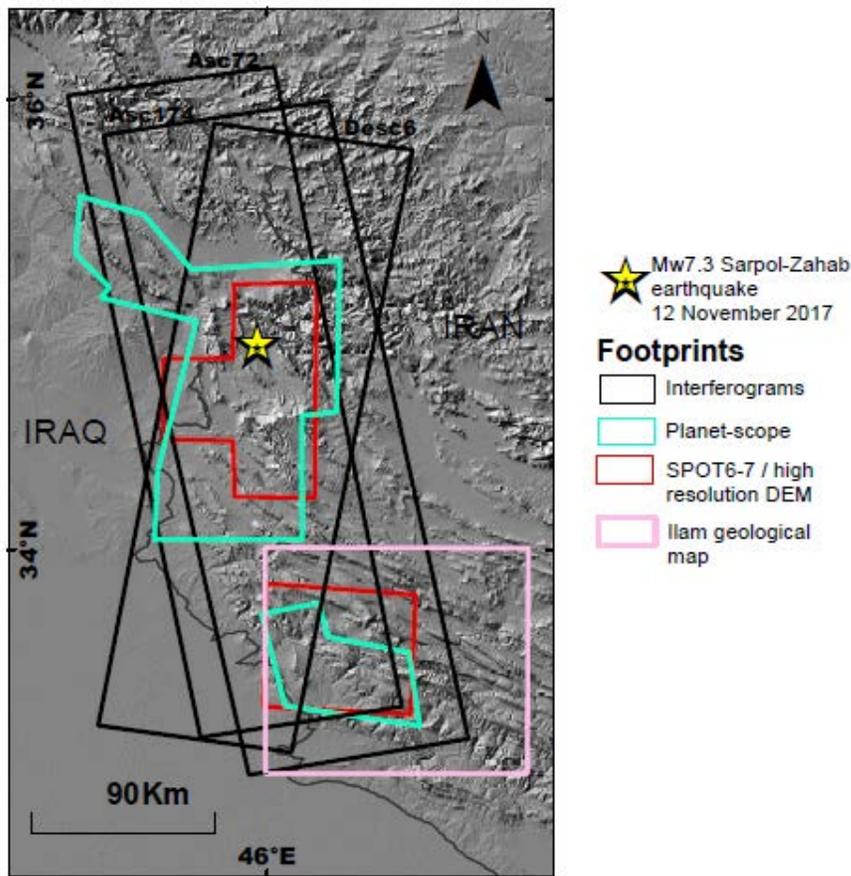
142 Rockslide deposit and bedrock  $V_s$  values at these sites are not available, but plausible values  
143 can be taken from a similar rockslide for the deposit mixing limestone and marl (Socco et al.,  
144 2010) and in the literature for bedrock (Telford et al., 1990):

$$145 \quad V_{SD} = 600 \text{ m/s}; V_{SB} = 3000 \text{ m/s}; \rho_D = 1.9; \rho_B = 2.5.$$

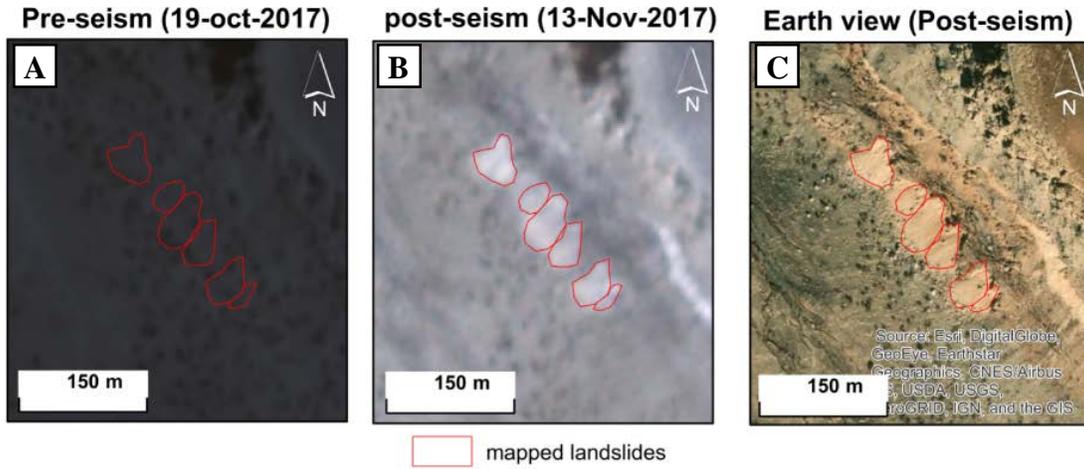
146 Considering these values, we obtain a resonance frequency  $f_0 \approx 1$  Hz associated to a 1D  
147 amplification over 6.

148 Thus the 1D resonance frequency of the carbonate Sarvak formation before the rockslide can  
149 be estimated to be around 1 Hz, taking plausible values of dynamic material moduli. After the

150 rockslide, the destructured slump body, characterized by lower rigidity and smaller thickness  
 151 (varying between 75 m and 150 m), also has a resonance frequency in the low range (1-2 Hz).  
 152 Topographic amplification is maximum for a wavelength comparable to the width of the  
 153 mountain (Geli et al., 1988), a condition that is again fulfilled in the low frequency range  
 154 (around 1 Hz) if we consider a mountain a few km wide with a velocity  $V_s$  of the order of 3 km/s.  
 155 These results suggest that ground motion parallel to the slope may have been significant  
 156 around 1 Hz at these 6 sites during the Sarpol-Zahab earthquake, due to the combined effect of  
 157 a particular directivity of the source and site amplification that can generate ground motions 5  
 158 to 10 times stronger than normal (Murphy, 2015). Interestingly, the presence of multiple ridges  
 159 can even increase the topographic effect (Geli, et al, 1988).

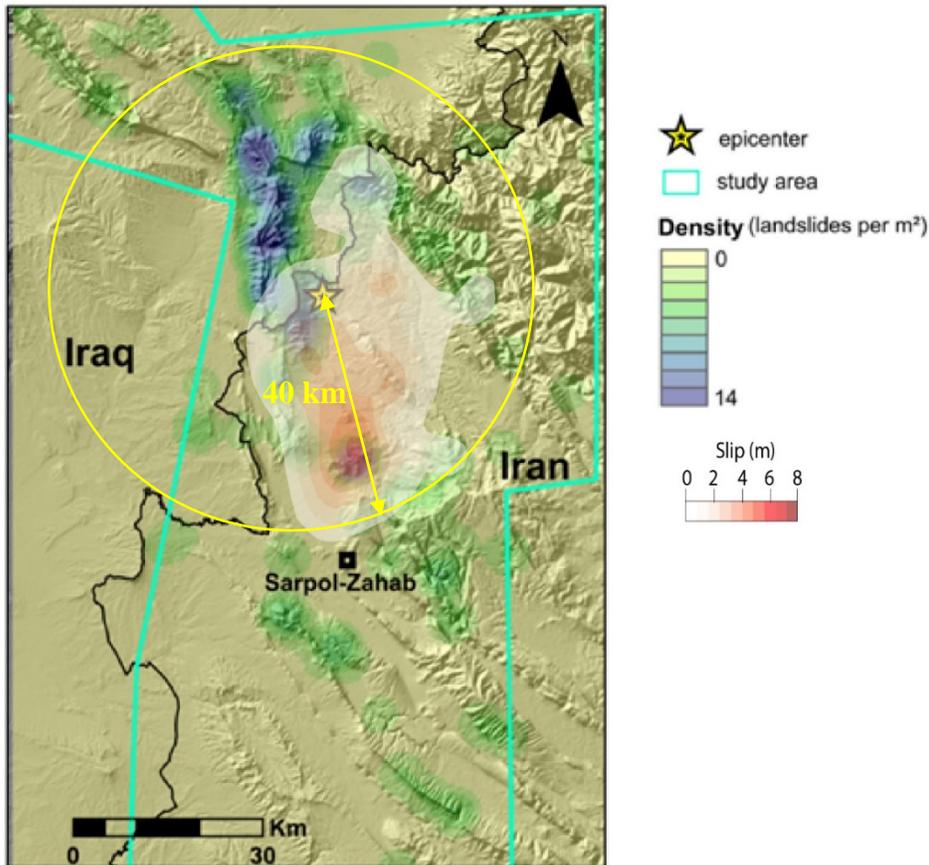


160  
 161 **Figure S1.** Footprints of the data used in our study. When pre- and post-earthquake data are  
 162 available, the common area is presented.



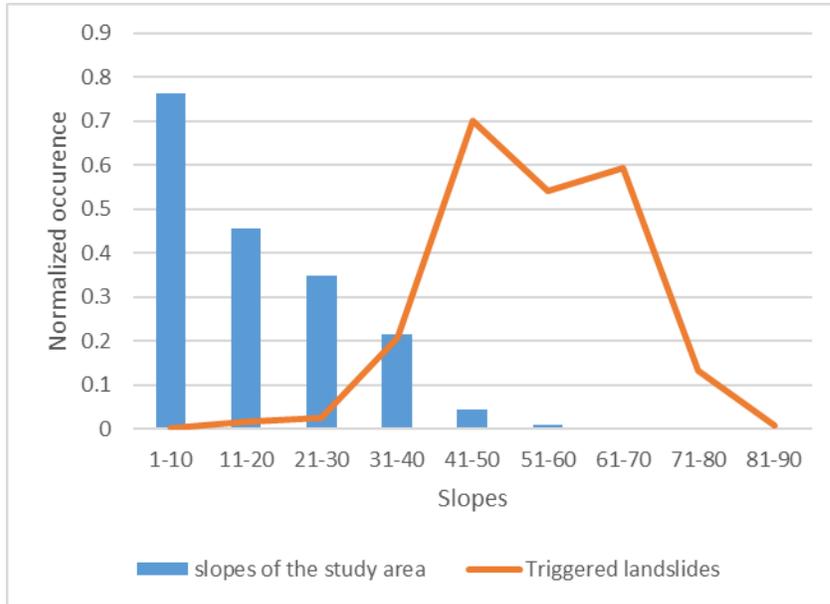
163

164 **Figure S2.** Typical example of rapid landslides mapped from PlanetScope images (3 m  
 165 resolution). (A) and (B) show the view of the same area from PlanetScope images before and  
 166 after the Sarpol-Zahab earthquake respectively. (C) shows the Google Earth view of the same  
 167 extent.



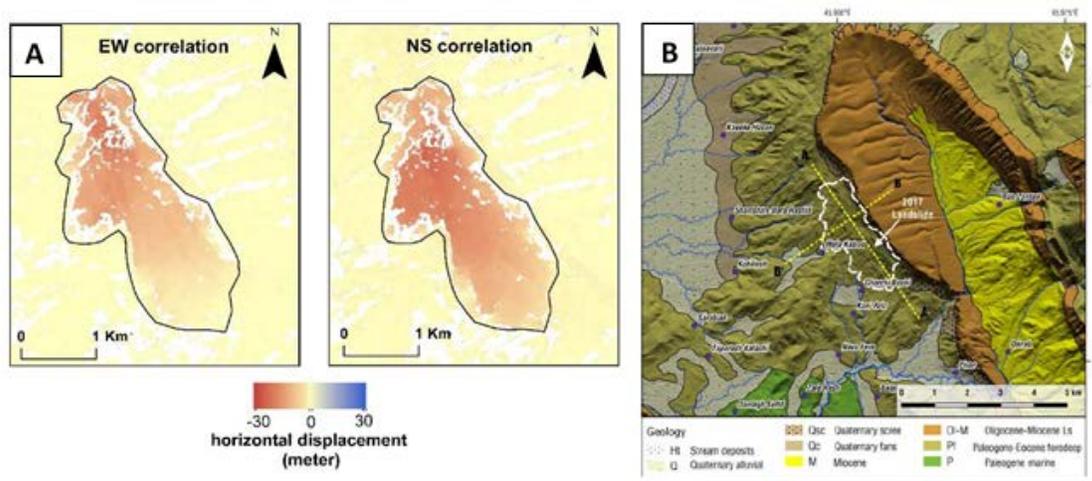
168

169 **Figure S3.** Rapid rockfalls density map. This map was calculated using the Kernel density tool in  
 170 ArcGIS software by evaluating the density of points within a 5 km radius. The cumulative slip at  
 171 12 s was added from Gombert et al (2019).



172

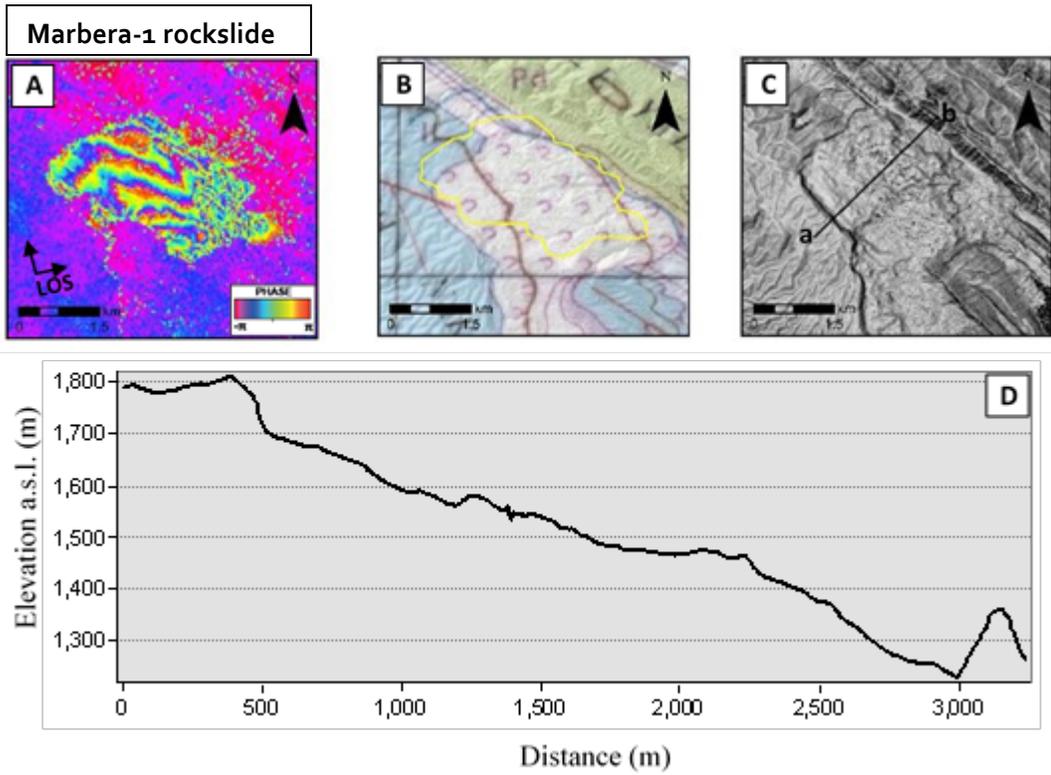
173 **Figure S4.** Plot showing the distribution of 276 detected rockfalls in respect to the available  
 174 slopes of our study area.



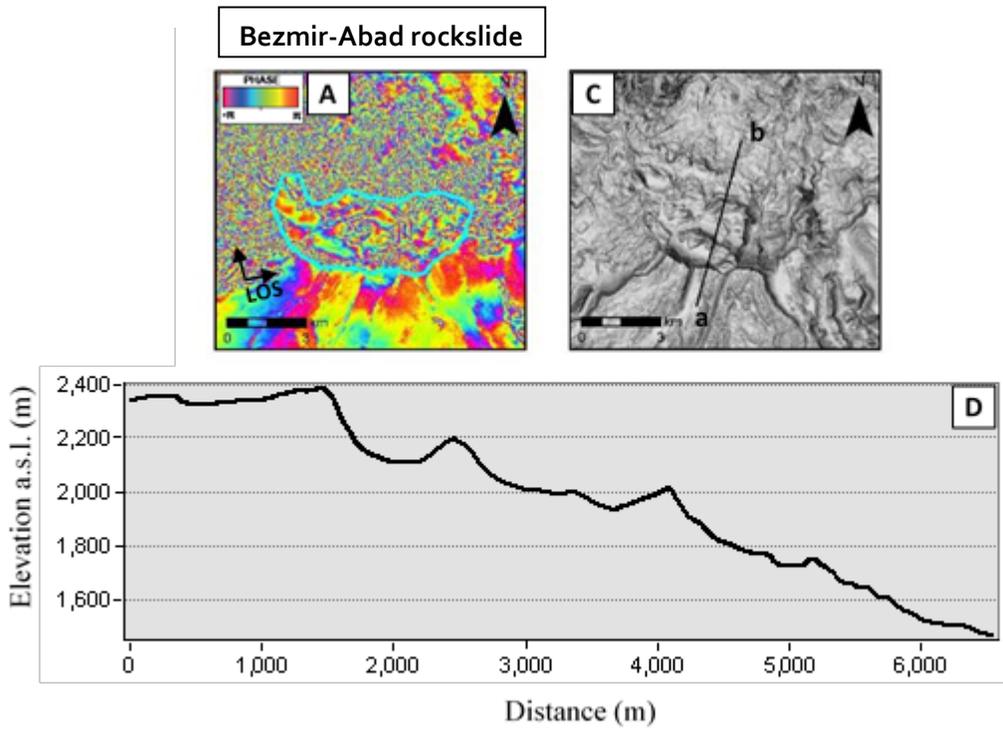
175

176 **Figure S5.** (A) Spot-6 images (Table S1) correlation results for the Melah-Kaboud landslide  
 177 obtained using the COSI-Corr tool, showing the coseismic displacement during the Sarpol-  
 178 Zahab earthquake. (B) Geological map of the Mela-Kaboud landslide region from the study of

179 Valkaniotis et al (2018). The white contour shows the limit of the displacement field detected  
180 from high resolution images in their study.

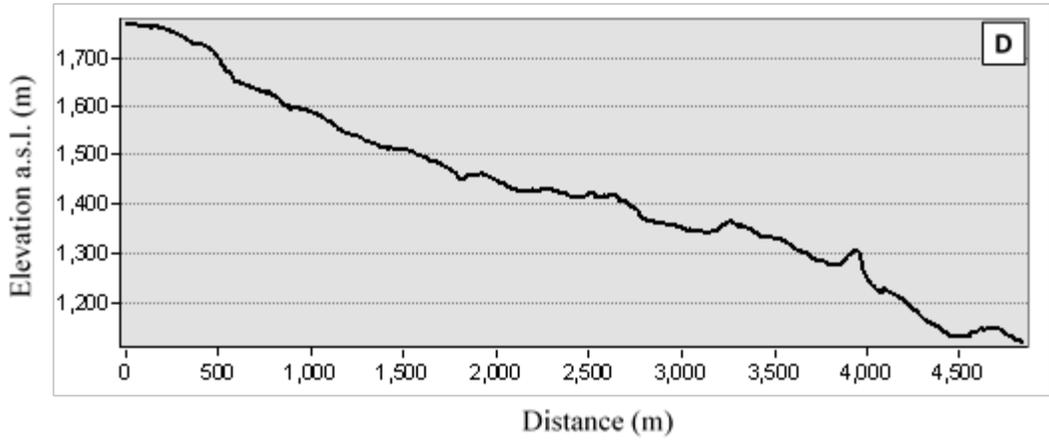
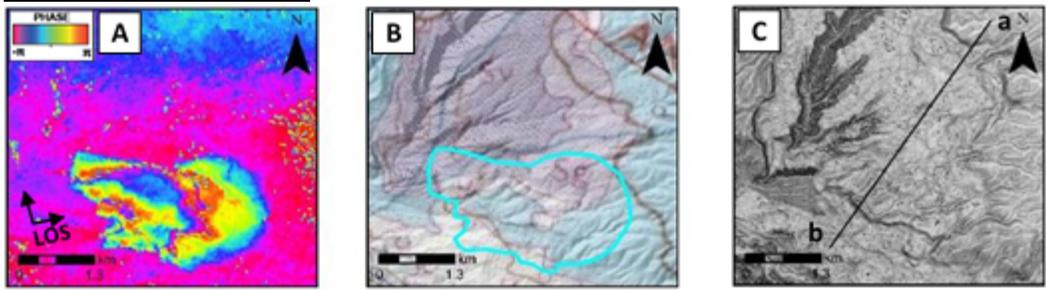


181



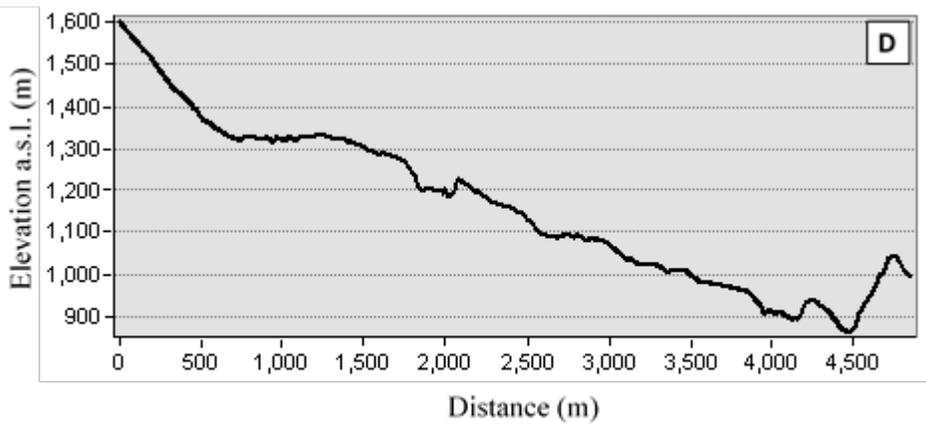
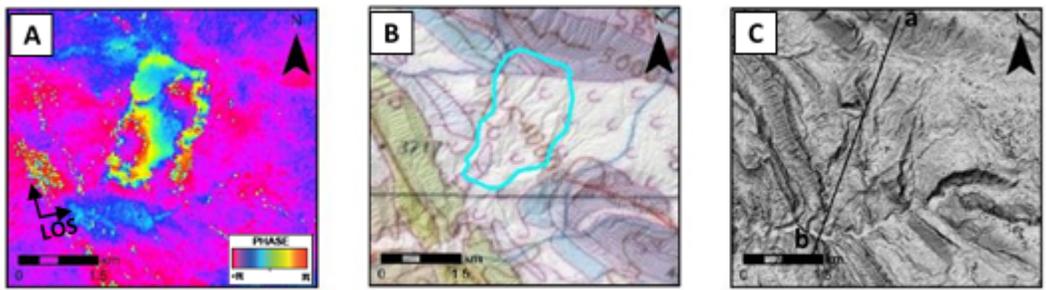
182

**Delqosha rockslide**



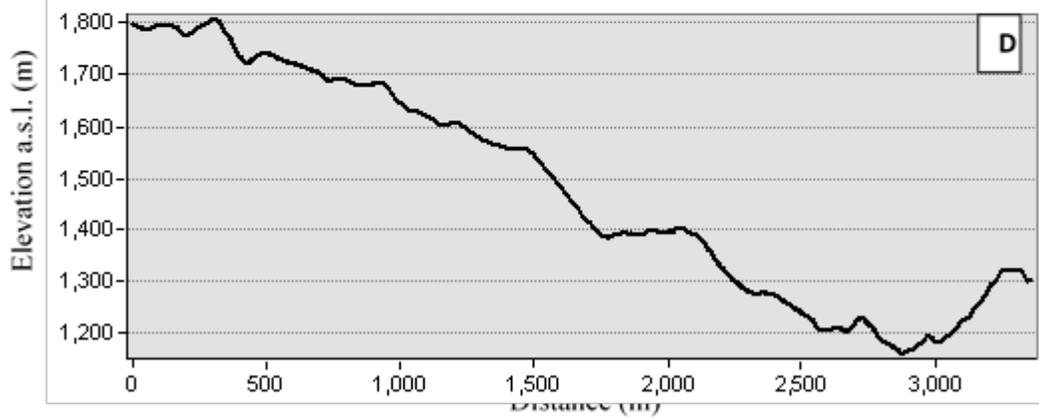
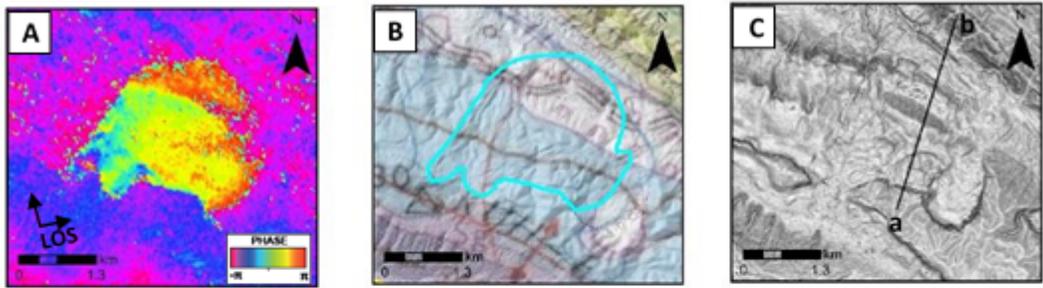
183

**Marbera-2 rockslide**



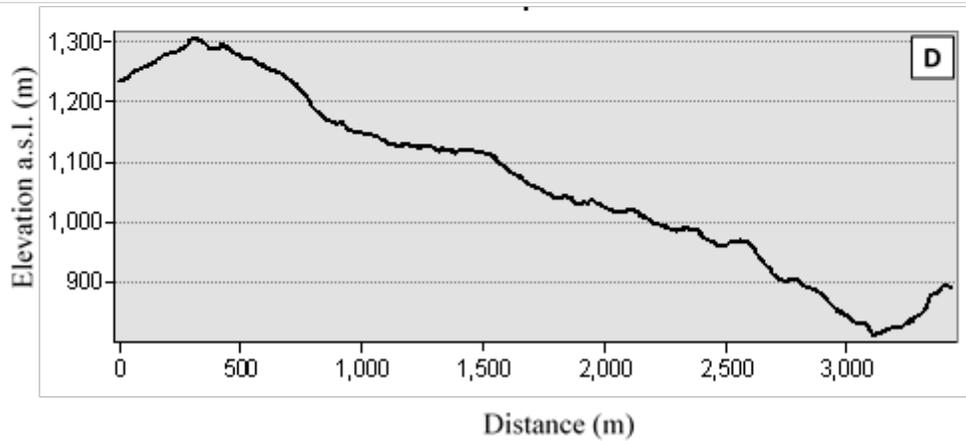
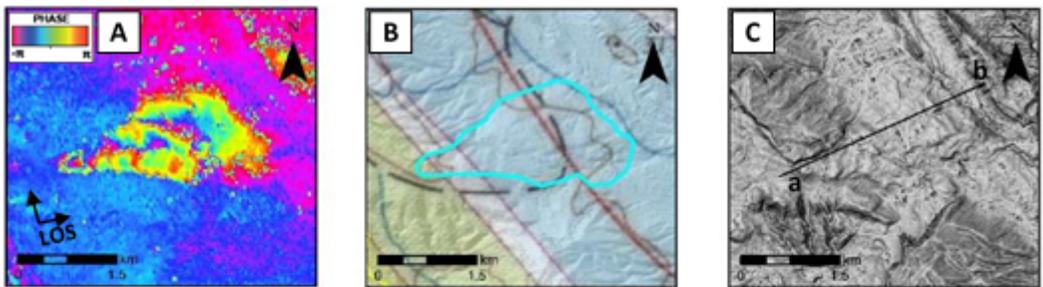
184

**Marbera-3 rockslide**

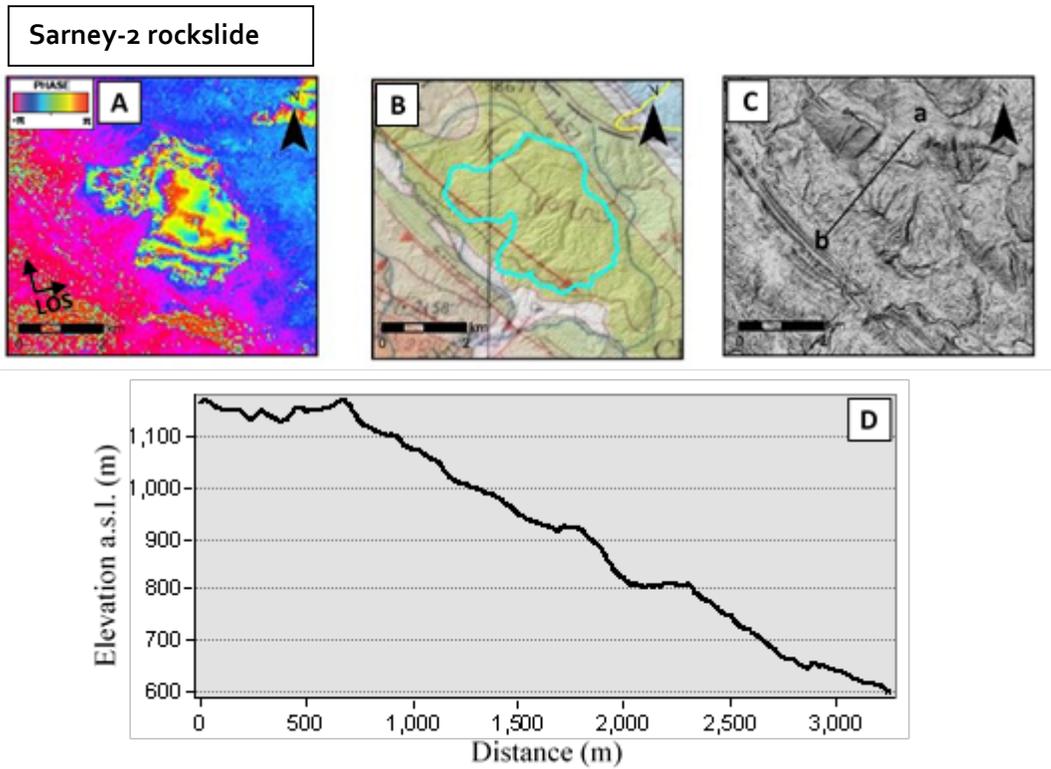


185

**Sarney-1 rockslide**



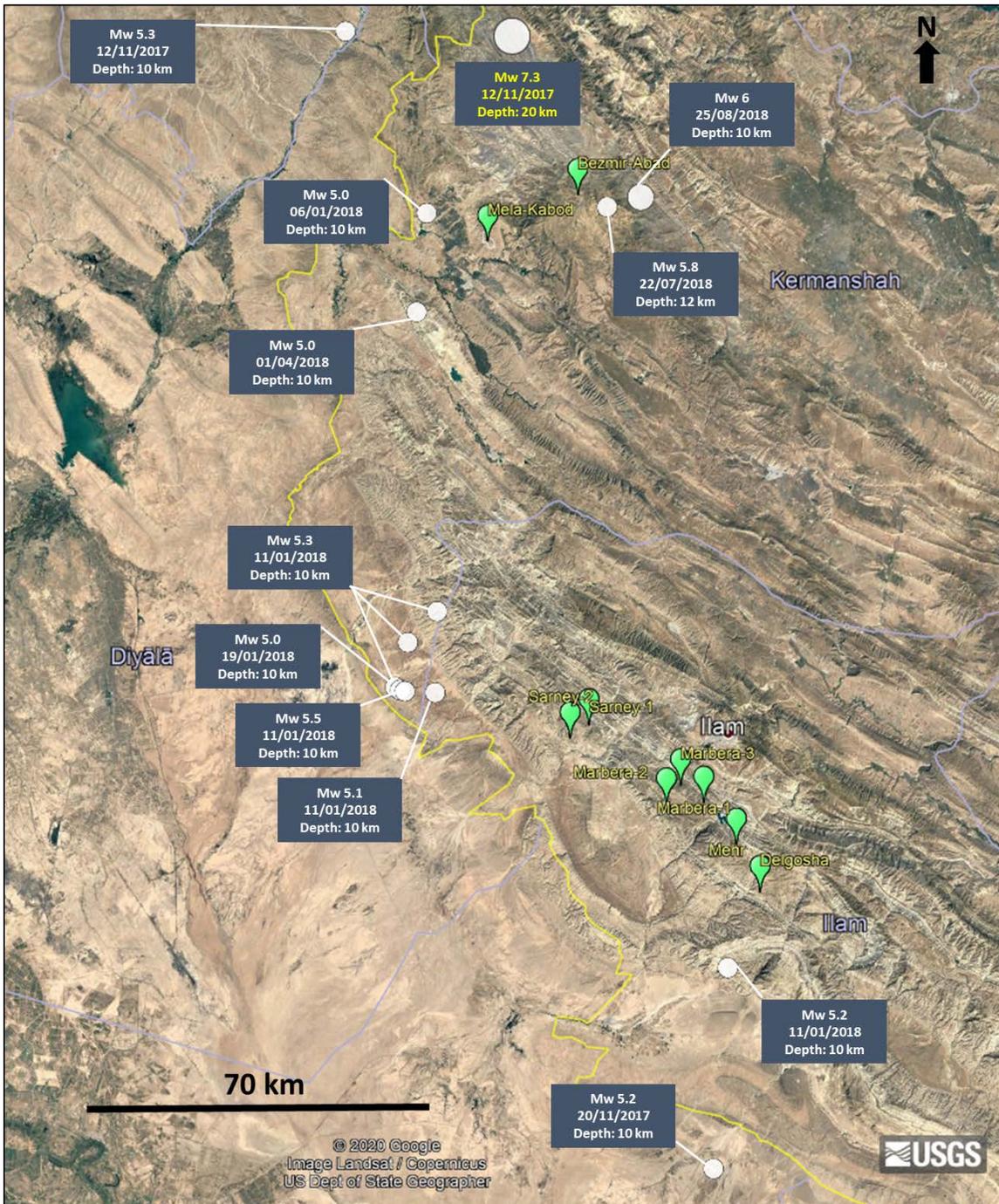
186



187

188 **Figure S6.** Figures showing (A) the interferogram pattern (the interferogram is computed  
 189 along the ascending track 72 between 11/11/2017 and 17/11/2017), (B) the geological map, (C)  
 190 the DEM and (D) a topographic profile 'ab' along the instance of the rockslides (other than the

191 Mehr rockslide). No detailed geological map is available for the region around the epicenter  
192 (the region of the Bezmir-Abad and Mela-Kabod landslide).



193

194 **Figure S7.** Details of all the seismic events that took place during the period of the  
195 accomplished time-series analysis (10-01-2017 and 27-08-2018).

Data type and origin		Date of acquisition		Resolution (meter)	Application and use
		Pre-earthquake	Post-earthquake		
Optical	PlanetScope	19-Oct-2017 (North*)	13-Nov-2017 (North*)	3 m	-Image correlation (COSI-Corr)
		07-Nov-2017 (South*)	17-Nov-2017 (South*)	3 m	-Visual comparison
	Spot-6	13-Oct-2013, 24-Apr-2014, 04-May-2014, 14-Aug-2014	29-Nov-2017, 12-Dec-2017	1 m	-Image correlation (COSI-Corr) - Dem generation
Radar	Sentinel-1 AB	10-Jan-2017 to 06-Nov-2017 (repeat cycle each 12 days)	12-Nov-2017 to 27-Aug-2018 (repeat cycle each 12 days)	5x20 m	New Small Baseline Subset (NSBAS)
Digital Elevation Models	ASTER	-----		30 m	Correction of interferograms
	Spot-6/7	13-Oct-2013, 24-Apr-2014, 04-May-2014, 14-Aug-2014	29-Nov-2017, 12-Dec-2017	2 m	-Interpretation of landslides -Generation of cross-sections
		09-Nov-2014 (south*)		4 m	

196 **North\*** correspond to the area around the epicenter

197 **South\*** correspond to the area of the rockslides detected in the far field to the south from the  
198 epicenter

199 **Table S1.** Synthesis of satellite data used in this study and their characteristics.

200 **Table S2.** Characteristics of the giant slow-moving rockslides detected from Sentinel-1  
201 interferograms and optical images correlation. The area of the rockslides already identified by  
202 Ghazipur and Simpson (2016) is given for comparison. The area deduced from this study

203 corresponds to the area of the interferogram (see Results for details). The volume is calculated  
204 based on the empirical law adopted in the study of Ghazipur and Simpson (2016) for the Zagros  
205 region.  $\Delta H$  is the elevation difference between the landslide toe and its headscarp. The average  
206 slope is calculated from the headscarp top to the landslide toe. Landslide orientation gives the  
207 direction toward which the landslide is sliding; "North-East" means the landslide orientation is  
208 from South-West to North-East. Line Of Sight (LOS) velocities correspond to linear  
209 interpolation of accumulated displacements from time-series computed over several months  
210 (see Figure 4 and text S5 for details).