Seismic and aseismic fault slip during the initiation phase of the 2017 Mw=6.9 Valparaíso earthquake.

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Abstract

Transient deformations associated with foreshocks activity has been observed before large earthquakes, suggesting the occurrence of a detectable pre-seismic slow slip during the initiation phase. In this respect, a critical issue consists in discriminating the relative contributions from seismic and aseismic fault slip during the preparation phase of large earthquakes. We focus on the April-May 2017 Valparaíso earthquake sequence, which involved a Mw=6.9 earthquake preceded by an intense foreshock activity. To assess the relative contribution of seismic and aseismic slip, we compare surface displacements predicted from foreshock source models to the transient motion measured prior to the mainshock. The comparison between observed and predicted displacements shows that only half of the total displacement can be explained by the contribution of foreshocks. This result suggests the presence of aseismic preslip during an initiation phase preceding the mainshock.

Seismic and aseismic fault slip during the initiation phase of the 2017 $M_W = 6.9$ Valparaíso earthquake

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¹⁰ Key Points:

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11	•	The 2017 Valparaíso $M_W = 6.9$ earthquake presents a pre-seismic transient dis-
12		placement.
13	•	We evaluate the contribution of foreshock-induced displacement to the pre-seismic
14		GPS observations.
15	•	Results suggest that $50\pm11\%$ of the pre-seismic displacement is caused by aseis-
16		mic slip.

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17 Abstract

Transient deformation associated with foreshocks activity has been observed before large 18 earthquakes, suggesting the occurrence of a detectable pre-seismic slow slip during the 19 initiation phase. A critical issue consists in discriminating the relative contributions from 20 seismic and aseismic fault slip during the preparation phase of large earthquakes. We 21 focus on the April-May 2017 Valparaíso earthquake sequence, which involved a $M_W =$ 22 6.9 earthquake preceded by intense foreshock activity. To assess the relative contribu-23 tion of seismic and aseismic slip, we compare surface displacement predicted from fore-24 shocks source models with transient motion measured prior to the mainshock. The com-25

26 parison between observed and predicted displacements shows that only half of the to-

tal displacement can be explained by the contribution of foreshocks. This result suggests

the presence of aseismic preslip during an initiation phase preceding the mainshock.

²⁹ Plain Language Summary

Several studies suggest that some large earthquakes are preceded by aseismic fault 30 slip. Such preslip could explain foreshock activity and transient displacements observed 31 before some large earthquakes. However, a large portion of observed pre-seismic defor-32 mations could be associated with the displacement field caused by each individual fore-33 shock earthquakes. This study focuses on the 2017 $M_W = 6.9$ Valparaíso (Chile) earth-34 quake that was preceded by a noticeable GPS displacement and numerous foreshocks. 35 By combining geodetic and seismic observations, our results show that only half of pre-36 seismic displacement can actually be explained by the contribution of foreshocks. This 37 confirms that the Valparaíso earthquake was preceded by detectable aseismic fault slip 38 accelerating into the main dynamic rupture. 39

40 **1** Introduction

Experimental and theoretical studies suggest that earthquakes begin with aseis-41 mic slow slip accelerating into a dynamic, catastrophic rupture (Das & Scholz, 1981; Kaneko 42 et al., 2016; Latour et al., 2013; Ohnaka, 2000). Laboratory-derived rate-and-state mod-43 els depict different evolution of preslip within nucleation zones of various sizes (Ampuero 44 & Rubin, 2008; Kaneko & Ampuero, 2011). With technological advances such as high-45 speed photoelastic techniques, the progressive acceleration from slow stable slip to fast 46 dynamic slip can be accurately monitored in laboratory conditions (e.g., Latour et al., 47 2013). Despite these advances, the detectability of such nucleation phases on natural faults 48 is still an open question. In addition to the nucleation itself, observations of the precur-49 sory phase leading to an earthquake indicate that earthquakes are often preceded by fore-50 shocks that could potentially be triggered by aseismic preslip (Bouchon et al., 2011, 2013; 51 Kato et al., 2012). Nonetheless, the role of foreshocks during this precursory phase re-52 mains unclear. At present, two end-member conceptual models compete in explaining 53 the occurrence of foreshocks. In the first model, foreshock stress changes contribute to 54 a slow cascade of random failures, leading eventually to the mainshock (Ellsworth & Bu-55 lut, 2018; Helmstetter & Sornette, 2003; Marsan & Enescu, 2012). The second model 56 proposes that foreshocks are triggered by aseismic slip corresponding to the nucleation 57 process of the mainshock (Bouchon et al., 2011; Dodge et al., 1996). 58

The continued development of geophysical networks in active tectonic regions pro-59 vides new opportunities to better capture the genesis of earthquakes. Geodetic obser-60 vations provide strong evidences of pre-seismic transient deformations at various time-61 scales (Ito et al., 2013; Mavrommatis et al., 2014; Ozawa et al., 2012; Socquet et al., 2017; 62 Yokota & Koketsu, 2015). However, the interpretation of such observations is often dif-63 ficult. This is particularly evident for the 2014 $M_W = 8.4$ Iquique (Chile) earthquake, 64 which was preceded by an active foreshock sequence that started 8 months before the 65 mainshock (Kato & Nakagawa, 2014). This foreshock sequence was accompanied by clear 66

GPS transient displacements, corresponding at least to some extent to aseismic fault slip preceding the mainshock (S. Ruiz et al., 2014; Socquet et al., 2017). The aseismic behavior of the observed pre-seismic transient is however debated as it might largely correspond to the cumulative co-seismic displacement of the foreshocks and associated afterslip (Bedford et al., 2015; Schurr et al., 2014). A reliable estimate of the relative contribution of seismic and aseismic deformations during nucleation is essential to better capture fault processes at the onset of earthquakes (Herman et al., 2016).

On 24 April 2017, a $M_W = 6.9$ earthquake occurred offshore Valparaíso in the cen-74 tral segment of the Chilean megathrust (33.089°S, 72.116°W, 21:38:28 UTC; Centro Sismólogico 75 National, CSN). This event is relatively moderate given that this region of the Chilean 76 subduction experienced earthquakes of magnitudes $M_W > 8$ (Comte et al., 1986; Dura 77 et al., 2015). This earthquake, however, caught the attention of seismologists because 78 it was preceded by a vigorous foreshock activity in the ~ 2 days preceding the mainshock. 79 This precursory activity has also been captured by GPS stations indicating a pre-seismic 80 trenchward motion over a similar time-scale (S. Ruiz et al., 2017; J. A. Ruiz et al., 2018). 81 A preliminary analysis of seismological and geodetic observations suggests that 80% of 82 pre-seismic GPS displacement is due to aseismic fault slip preceding the mainshock (S. Ruiz 83 et al., 2017). This first order estimate is obtained by comparing inverted preslip with 84 the seismic moment of foreshocks assuming they are all located on the subduction in-85 terface. This assumption is questionable as seismicity catalogs depict a significant dis-86 persion of earthquake locations around the plate interface (S. Ruiz et al., 2017; J. A. Ruiz 87 et al., 2018), most events being located at depths larger than the slab 1.0 model (Hayes 88 et al., 2012). Such dispersion, probably related to depth uncertainty, implies a signifi-89 cant non-random bias in seismic moment for dip-slip earthquakes. For example, if an earth-90 quake at $20 \,\mathrm{km}$ depth is mislocated at $25 \,\mathrm{km}$, the moment is underestimated by nearly 91 20% using long-period teleseismic records (Tsai et al., 2011). Such mis-estimation of seis-92 mic moment may lead to non-negligible errors in the contribution of foreshocks to ob-93 served pre-seismic deformations. 94

The primary goal of this study is to assess the relative contribution of seismic and 95 aseismic slip during the few days preceding the 2017 Valparaíso earthquake. Estimat-96 ing the seismic contribution to observed geodetic displacement is difficult as we deal with 97 moderate-sized foreshocks ($M_W < 6$) for which a co-seismic offset is not clearly visi-98 ble on GPS time-series. The seismic contribution to the observed displacement can be 99 estimated by modeling the source of foreshocks from seismic data. However, this pro-100 cess should be done carefully as source models and the corresponding predictions can 101 be affected by significant uncertainties. In this work, we obtain a moment-tensor cat-102 alog and predict the corresponding co-seismic offsets at GPS stations accounting for ob-103 servational and modeling uncertainties. In particular, we account for prediction uncer-104 tainties associated with inaccuracies in the Earth model. We find that about half of the 105 observed GPS pre-seismic displacement is aseismic and is caused by preslip in the vicin-106 ity of the impending mainshock hypocenter. Such pre-seismic deformation is unlikely to 107 be explained by afterslip induced by preceding foreshocks. This suggests that aseismic 108 preslip played an important role in the 2017 Valparaíso sequence. 109

¹¹⁰ 2 Pre-seismic Transient Displacements captured by GPS

We process GPS data of 68 stations in South America from several networks (CSN, 111 LIA Montessus de Ballore, Ministerio de Bienes Nacionales, RAMSAC, RBMC-IP, IGS, 112 IGM Bolivia, see supplementary information S1 for references). Processing is done us-113 ing a differential approach (Herring et al., 2018) including tropospheric delays and hor-114 izontal gradients. The results are computed in the ITRF 2014 reference frame (Altamimi 115 et al., 2016) and converted in a fixed South-America frame (Nocquet et al., 2014). We 116 use daily solutions except for the last position before the mainshock, which is obtained 117 from data up to one hour before the event. We remove a trend corresponding to inter-118



Figure 1. The 2017 Valparaíso earthquake sequence. (a) Earthquake locations including foreshocks (blue circles), mainshock (green star), and aftershocks (white circles). The red colormap indicates the preslip distribution resulting from the inversion of GPS data (see section 5). The black arrows show the cumulative observed GPS surface displacements (up to one hour before the mainshock). Orange dots indicate the seismicity distribution from 2017/01/01 until 2017/10/05 according to the microseismicity catalog obtained by S. Ruiz et al. (2017). (b) GPS Time-series in the vicinity of Valparaíso. The vertical red dashed line indicates approximate onset of the transient displacement visible on the time-series. The cumulative number of earthquakes from S. Ruiz et al. (2017) is shown at the bottom of the figure. The purple star represents the largest $M_W = 6.0$ foreshock.

seismic motion from the time-series by fitting a linear regression in a 4 months time-window
before the mainshock. Finally, we subtract the first sample of the time-series (i.e., which
we consider as displacement zero) and obtain the corresponding offsets.

Figure 1-b and S3 show the resulting horizontal displacements for stations in the 122 vicinity of the study area. There is a clear westward motion starting about 3 days be-123 fore the main shock and reaching ${\sim}8\,\mathrm{mm}$ close to the coast. Figure 1-b compares GPS 124 time-series with the cumulative number of earthquakes in the micro-seismicity catalog 125 obtained by S. Ruiz et al. (2017). Interestingly, the pre-seismic GPS transient starts be-126 fore a noticeable increase in seismicity. In Figure 1-b, we can see that the slope of cu-127 mulative seismicity rate does not change significantly at the beginning of the transient. 128 The increase in seismicity rate is delayed by about 24 hours and only starts with a $M_W =$ 129 6.0 foreshock on April 23 (purple star in Figure 1-b). This suggests that aseismic pres-130 lip initiated on the fault before the increase in foreshock activity. 131

¹³² 3 Centroid Moment Tensor catalog

To constrain the contribution of foreshocks to the observed GPS displacement, we 133 estimate Centroid Moment Tensor (CMT) parameters for moderate to large earthquakes 134 during the Valparaíso earthquake sequence (from 2017/04/05 up to 2017/05/30). We use 135 records from broadband seismic stations located within 12° from the mainshock hypocen-136 ter. These stations are mostly included in the C and C1 regional networks maintained 137 by the Centro Sismológico Nacional (CSN) of the Universidad de Chile (Universidad de 138 Chile, 2013). We also use stations operated by GEOSCOPE, and IRIS/USGS network 139 (Institut de Physique du Globe de Paris and Ecole et Observatoire des Sciences de la Terre 140 de Strasbourg (EOST), 1982; Albuquerque Seismological Laboratory (ASL)/USGS, 1993, 141 1988). 142

We use a modified version of the W-phase algorithm adapted to regional distances 143 and the magnitude range of the Valparaíso sequence (Kanamori & Rivera, 2008; Zhao 144 et al., 2017). Estimated parameters are the deviatoric moment tensor, the centroid lo-145 cation, the centroid time, and the half-duration of an isosceles triangular moment rate 146 function. The inversion is performed by fitting full waveforms in a 180s time-window start-147 ing at the P-wave. We filter data between 12s and 100s using different pass-bands for 148 different magnitude events (see Table S1 in the online supplementary). We compute Green's 149 functions for the source inversion in a 1D layered structure extracted from the 3D Earth 150 model of S. Ruiz et al. (2017) in the area of Valparaíso (Figure S4). 151

The resulting CMT catalog is shown in Figure 2 and in table S2. Most earthquakes (more than 90% of the total catalog) have thrust mechanisms. Interestingly, foreshocks are mostly concentrated close to the mainshock hypocenter (see Figure 1 and Figure 2a). On the other hand, aftershocks show a different behavior, surrounding the region where foreshocks have previously occurred.

The cumulative scalar seismic moment released by foreshocks before the mainshock 157 is largely dominated by two events with $M_W \geq 5.5$ (cf., Figure 2-b). These foreshocks 158 of magnitude $M_W = 6.0$ and $M_W = 5.5$ occurred respectively 43 hours and 26 hours 159 before the mainshock. As our CMT catalog only consists of $M_W \geq 3.8$ earthquakes, 160 the contribution of microseismicity is not included in our estimates of cumulative seis-161 mic moment before the mainshock. Even though the individual contribution of these small 162 earthquakes to the observed displacement is negligible, their large number may contribute 163 to surface displacement. To assess the contribution of small earthquakes, we consider the 164 frequency-magnitude distribution of our CMT catalog assuming a completeness magni-165 tude of $M_c = 3.9$ (Figure S5). We compare our catalog with previous moment tensor 166 catalogs of the same sequence (S. Ruiz et al., 2017; J. A. Ruiz et al., 2018), which are 167 qualitatively consistent with our estimates (Figure S5). We then compute the Gutenberg-168 Richter (GR) law using the methodology proposed by Aki (1965) for the whole sequence, 169 and the foreshocks sequence. Even though the GR laws show some discrepancies, they 170 are in good agreement considering the uncertainties on our estimates (Figure S5). The 171 foreshocks GR law is then extrapolated to lower magnitudes, and the cumulative mo-172 ment of magnitudes below the magnitude of completude is included to correct for the 173 influence of small, hence not detected earthquakes. Our CMT catalog suggests a cumu-174 lative moment $M_0 = 1.474 \times 10^{18}$ N·m. The cumulative seismic moment of foreshocks 175 with magnitudes below completeness is $M_0 = 4.966 \times 10^{15}$ N·m (i.e., $M_w = 4.4$). The 176 contribution of microearthquakes is therefore negligible compared to seismic events. 177

To evaluate the contribution of foreshocks to observed surface displacements, we calculate synthetic static displacements using our CMT catalog and the same 1D velocity model employed to obtain our CMT solutions. Synthetics are computed using the CSI package (http://www.geologie.ens.fr/ jolivet/csi) incorporating the approach of Zhu and Rivera (2002) to compute static displacement in a layered model. Results on Figure S6 indicate that the largest foreshock ($M_W = 6.0$) largely dominates the co-seismic



Figure 2. CMT solutions of the 2017 Valparaíso earthquake sequence and cumulative moment (a) CMT solutions of the 2017 Valparaíso earthquake sequence. Focal mechanisms are contoured in blue and black for foreshocks and aftershocks respectively. The size of beach balls scales with the moment magnitude. Color of the compressive quadrants represents the event depth. (b) Cumulative scalar seismic moment of the 2017 Valparaíso sequence. The mainshock scalar moment is not included in this figure. The red dashed line outlines the approximate onset of transient displacements visible on GPS time-series. The green line indicates the origin time.

contribution to the observed GPS transient while $M_W < 6.0$ events in our catalog generate relatively small surface displacement. Assuming that microearthquakes are located in the vicinity of $M_W \ge 3.8$ foreshocks, they should also have a negligible contribution to the observed surface displacement (given their small cumulative scalar moment). As the $M_W = 6.0$ foreshock plays a important role in the sequence, we assess uncertain-

ties associated with the corresponding CMT parameters.

¹⁹⁰ 4 Uncertainty on predicted co-seismic displacements

Synthetic co-seismic surface displacements are sensitive to uncertain earthquake 191 source parameters. For large magnitude foreshocks, uncertainties on centroid location 192 and moment tensor affect our estimates of the co-seismic contribution to the transient 193 displacement observed before the mainshock. Source parameters uncertainties can ei-194 ther result from observational errors, or from errors in the forward model (prediction/theoretical 195 errors). For example, there might be innacuracies in the velocity model, which is known 196 to induce non-negligible errors in CMT solutions (Duputel et al., 2012, 2014; Morales-197 Yañez et al., 2020). The point source assumption is another source of uncertainty in the 198 forward model. As for the observations, temporally and spatially variable noise level at 199 seismic stations is a major source of uncertainty. 200

In order to assess uncertainties associated with the CMT solution of the largest $M_W =$ 201 6.0 foreshock, we perform a new CMT inversion within a Bayesian framework, follow-202 ing Duputel et al. (2012, 2014). Each source of uncertainty considered here is integrated 203 in the problem as a covariance matrix. The covariance matrix C_d , associated with ob-204 servational errors, is derived after a first CMT inversion. From this inversion, an aver-205 age correlation function is derived from residuals between synthetic and observed wave-206 forms at each station. This allows us to estimate the correlation between neighbor data 207 samples, and include it into C_d . The standard deviation for each channel is fixed to 4 208 times the corresponding average absolute residuals. This empirical procedure provides 209 a conservative estimate of observational uncertainty associated with each waveform. 210

Forward modeling uncertainties are represented by the matrix C_p , which assesses 211 the influence of inaccuracies in the Earth model. We use the same velocity model as in 212 section 3 assuming log-normal uncertainties on elastic parameters as shown in Figure S4. 213 Uncertainty in each layer is estimated by assessing the spatial variability of the 3D Earth 214 model of S. Ruiz et al. (2017) in the epicentral region and by comparison with other re-215 gional models (e.g., J. A. Ruiz et al., 2018). To evaluate the corresponding variability 216 in the predictions, we employ the first-order perturbation approach described in Duputel 217 et al. (2014), assuming that prediction error is linearly related with uncertainty on the 218 elastic parameters. A test is described in supplementary information S2 and Figures S7-219 S8 to assess the validity of this approach. 220

The posterior ensemble of plausible source locations and moment tensors is appraised using a strategy similar to Sambridge (1999). At a fixed point-source location in time and space, the posterior distribution of moment tensor parameters is Gaussian and can be written as (Tarantola et al., 1982):

$$p(\mathbf{m}|\mathbf{d}_{obs}, \mathbf{x}) = N(\widetilde{\mathbf{m}}, \widetilde{\mathbf{C}}_m) \tag{1}$$

where **m** are the moment tensor parameters, \mathbf{d}_{obs} is the data vector containing the concatenated observed waveforms and **x** is the point source location. The right-hand member of this equation is a Gaussian distribution of mean $\tilde{\mathbf{m}}$ and covariance $\tilde{\mathbf{C}}_m$. The posterior mean $\tilde{\mathbf{m}}$ is the maximum *a posteriori* moment tensor given by:

$$\widetilde{\mathbf{m}} = \left(\mathbf{G}^{t}\mathbf{C}_{\chi}^{-1}\mathbf{G}\right)^{-1}\mathbf{G}^{t}\mathbf{C}_{\chi}^{-1}\mathbf{d}_{obs},\tag{2}$$

where **G** is the Green's function matrix while $\mathbf{C}_{\chi} = \mathbf{C}_d + \mathbf{C}_p$ is the covariance matrix reflecting observational (\mathbf{C}_d) and prediction uncertainties (\mathbf{C}_p). The posterior covariance matrix is given by:

$$\widetilde{\mathbf{C}}_m = \left(\mathbf{G}^t \mathbf{C}_{\chi}^{-1} \mathbf{G}\right)^{-1} \tag{3}$$

To get the joint posterior distribution on moment tensor \mathbf{m} and source location \mathbf{x} , we first calculate $\tilde{\mathbf{m}}$ and $\tilde{\mathbf{C}}_m$ on a 3D grid of possible point-source locations around the hypocenter. Starting from the initial location \mathbf{x}_c determined in section 3 (corresponding a moment tensor \mathbf{m}_c), we then employ an hybrid metropolis algorithm by repeating the following iterations until a sufficiently large number of model samples is generated:

- 1. Randomly generate a candidate point-source location $\mathbf{x}^* = \mathbf{x}_c + \delta \mathbf{x}$ where $\delta \mathbf{x}$ is a small perturbation randomly generated from a Gaussian distribution with a standard deviation of 0.1° in latitude/longitude and $\sigma=0.1$ km in depth.
 - 2. Extract $\widetilde{\mathbf{m}}$ and \mathbf{C}_m from the grid point closest to \mathbf{x}^* and generate a random model \mathbf{m}^* from $p(\mathbf{m}|\mathbf{d}_{obs}, \mathbf{x}^*)$ in eq. (1).
 - 3. Accept or reject \mathbf{m}^* and \mathbf{x}^* using a standard Metropolis approach:
 - Draw a random number $\alpha \sim U(0,1)$
 - Accept \mathbf{m}^* and \mathbf{x}^* if $\alpha < \min\left(1, \frac{p(\mathbf{m}|\mathbf{d}_{obs}, \mathbf{x}^*)}{p(\mathbf{m}_c |\mathbf{d}_{obs}, \mathbf{x}_c)}\right)$
 - Otherwise duplicate \mathbf{m}_c and \mathbf{x}_c

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Figure 3. Bayesian point-source model for the M_W =6.0 foreshock on 2017-04-23. Blue circles and lines in the figure represent model samples randomly drawn from the posterior distribution. a) Samples from the posterior PDF depecting uncertainties in the point source location. The red and orange stars are the initial solution (i.e. starting model) and the posterior mean model respectively. b) Focal mechanism uncertainty. c) Marginal posterior PDF of the scalar seismic moment. The red and orange lines are the initial and the posterior mean model.

Figure 3 shows 4500 model samples generated using the approach described above. 246 The posterior distribution shows a location uncertainty of about 10 km. We observe a 247 good fit between observed and synthetic seismograms (Figure S9). However, we also no-248 tice a trade-off between longitude and depth, which probably results from the distribu-249 tion of stations used for inversion (Figure S10). To evaluate the uncertainty on the pre-250 dicted co-seismic displacement, we simulate static displacement for each model samples 251 shown in Figure 3. The resulting stochastic co-seismic displacements are shown in Gray 252 in Figure 4a for GPS stations that are closest to the mainshock epicenter. This shows 253 prediction uncertainties ranging from 0.25 to 0.4 mm on the east component of displace-254 ment. Despite these uncertainties, the predicted cumulative co-seismic offsets are still 255

significantly smaller than the observed pre-seismic displacements (~ 6 to 8 mm of the east component for the closest stations).

5 Partitioning between seismic and aseismic fault slip

In Figure 4, we compare the total cumulative foreshock co-seismic offset with the 259 observed pre-seismic GPS displacement. Predicted co-seismic displacements include the 260 contribution of microearthquakes below the magnitude of completeness, assuming a to-261 tal scalar moment derived from our GR analysis with a location and mechanism simi-262 lar to the $M_W = 6.0$ foreshock. As discussed earlier, only the largest foreshock $M_W =$ 263 6.0 is significantly contributing to co-seismic displacements (see Figure 4a and S6). The 264 contribution of earthquakes smaller than $M_W = 6.0$ has a minimal impact on the fi-265 nal result. 266

To get a total budget of seismic and aseismic displacement before the mainshock, 267 Figure 4b compares GPS data 1 hour before the mainshock with the corresponding cu-268 mulative foreshock displacement. Observed displacement are on average between 4 and 269 6 mm larger than co-seismic offsets. Such differences cannot be explained by uncertain-270 ties on the observations and the predictions. These results clearly suggest that a signif-271 icant portion of the observed pre-seismic deformation is actually aseismic and cannot be 272 caused by foreshocks. We estimate that about $51\pm11\%$ of the displacement measured 273 at the GPS stations originates from aseismic slip on the megathrust. As shown in Figure, 4c, 274 the portion of aseismic deformation is quite consistent between stations suggesting that 275 a common source located in the vicinity of the foreshocks could explain those results. 276

To further explore this hypothesis, we then conduct two inversions: a first slip in-277 version of the total GPS pre-seismic displacement and another inversion after removing 278 the contribution of foreshocks (i.e., aseismic displacement only). To build a fault geom-279 etry, we use the CSI package to mesh the Slab 2.0 model with triangles of variable sizes 280 as shown in Figure 4e-f. We invert for slip values at the triangular nodes using AlTar, 281 a Markov chain Monte Carlo sampler based on the algorithm described by Minson et al. 282 (2013). Continuous fault slip distribution is represented as a linear interpolation of the 283 slip values at the triangular nodes. Green's functions are computed in the same stratified elastic model used for our CMT catalog (Figure S4). Given the limited amount of 285 available observations, we enforce a positive Laplacian prior distribution with a scale pa-286 rameter of 1 m. Such sparsity-inducing prior will favor "simple" models with slip only 287 where it is requested by the data. Results in Figure 4e-f shows that GPS observations 288 can be explained by slip in the vicinity of the mainshock hypocenter. Aseismic slip dis-289 tribution appears to be somewhat more spread out, which may be an effect of the larger 290 uncertainty associated with GPS data after removing the contribution of foreshocks (as 291 the co-seismic prediction uncertainty propagates in the corrected GPS data). 292

²⁹³ 6 Discussion and conclusion

We investigate the seismic and aseismic motions during the preparation phase of 294 the 2017 $M_w = 6.9$ Valparaíso earthquake. We first evaluate the contribution of foreshock-295 induced displacement to pre-seismic GPS observations. Co-seismic offsets are largely dom-296 inated by a $M_W = 6.0$ foreshock that occurred ~43 hours before the mainshock. As 297 pointed out in section 2, the transient GPS signal starts before the increase in seismic-298 ity rate. More specifically, we can see in Figure 4a that the observed displacement on April 299 22 mainly corresponds to aseismic slip as no significant foreshock occurs on that day. On 300 the other hand, the position on April 23 results from a combination of seismic and aseis-301 mic fault slip. The detailed evolution of the partitioning between seismic and aseismic 302 slip is difficult to interpret using daily GPS time-series in which each position corresponds 303 to an average over 24 hours. This analysis is also subject to large observational and pre-304



Figure 4. Slip during the Valparaíso foreshock sequence. a) Time series of GNSS data (blue) and stochastic foreshock-induced co-seismic displacement (gray). Red dots represent the average of stochastic co-seismic offsets. Green cross corresponds to the total foreshock displacement, including the contribution of earthquakes below the magnitude of completeness. b) Distributions of observed pre-seismic displacement and predicted cumulative co-seismic offsets caused by foreshocks. Blue histograms represent observations assuming Gaussian uncertainties from standard errors estimated at each station. Red histograms correspond to the posterior distribution of cumulative foreshock-induced co-seismic displacement. c) Percentage of aseismic displacement for each station. d) Average postseismic signal measured on stations TRPD, VALN, BN05 and QTAY (see Figure S11). e) Slip inversion of pre-seismic GPS data. f) Slip inversion of GPS data after removing foreshock-induced displacement. Black and blue arrows are observed and predicted horizontal GPS displacements along with their 1- σ ellipses (representing observational and prediction uncertainties, respectively). Colored circles are observed (outer circles) and predicted (inner circles) vertical displacements from GPS and tide gauges, respectively.

diction uncertainties. For these reasons, we focus on the overall partitioning between seismic and aseismic slip during the preparation phase of the Valparaíso earthquake.

Our analysis shows that a significant part of pre-seismic GPS observations are not 307 explained by foreshock-induced displacement even when accounting for prediction and 308 observation uncertainties. We estimate that $\sim 50 \pm 11\%$ of GPS displacements is likely 309 caused by aseismic slip, a ratio that is fairly consistent for different stations in the vicin-310 ity of the Valparaíso sequence (Figure 4c). To check weather such pre-seismic motion could 311 be explained by slip on the plate interface, we conduct a slip inversion after correcting 312 313 GPS data from foreshock-induced displacement (cf., Figure 4f). The distribution of aseismic preslip spreads toward the west of Valparaíso city with an extension of about 50×90 km 314 and a scalar moment of $M_0 = 3.08 \times 10^{18}$ N.m (i.e., $M_w = 6.26$). This assisting mo-315 tion represents about 50% of the moment calculated for the slip model derived from un-316 corrected GPS data ($M_0 = 5.67 \times 10^{18}$ N.m, Figure 4e). Given the cumulative moment 317 of foreshocks ($M_0 = 1.48 \times 10^{18}$ N.m), we estimate that nearly 70% of the scalar mo-318 ment released during the preparation phase of the Valparaíso mainshock is aseismic, which 319 is roughly in agreement with estimates from S. Ruiz et al. (2017). The smaller portion 320 of aseismic moment derived from the comparison of slip models in Figure 4e-f likely re-321 sults from the simplistic assumption in Figure 4e that all foreshocks are located on the 322 plate interface. 323

Even if our analysis demonstrates the existence of aseismic slip prior to the Val-324 paraíso mainshock, such aseismic motion may include afterslip from preceding bursts of 325 seismicity. This has been suggested for pre-seismic displacement observed before the 2014 326 $M_W = 8.1$ Iquique earthquake, which could potentially be explained by afterslip induced 327 by foreshock seismicity (Bedford et al., 2015). Testing such possibility for the 2017 Val-328 paraíso sequence is difficult as we cannot easily isolate the afterslip signal from GPS time-329 series, which likely incorporate other contributions including preslip of the impeding main-330 shock. To assess the contribution of afterslip, we employ two approaches. In a first ap-331 proach, we use the mainshock post-seismic GPS signals as a proxy for the afterslip in-332 duced by foreshocks. The mainshock post-seismic time-series are normalized by the co-333 seismic offset of each station to evaluate the relative proportion of post-seismic displace-334 ment as a function of time. This suggests that about 10% of the co-seismic moment af-335 ter 43 hours corresponds to post-seismic deformations (see Figure 4d and Figure S11). 336 This result is consistent with values reported for earthquakes with similar or larger mag-337 nitudes (Chlieh et al., 2007; D'agostino et al., 2012; Lin et al., 2013). If we assume a sim-338 ilar behavior for the foreshocks, the post-seismic signal caused by foreshocks is below mea-339 surement uncertainties (approximately 0.7 mm for an uncertainty of 1.1 mm in GPS sig-340 nals) and can therefore be neglected. In a second approach, we make the more conser-341 vative assumption that afterslip caused by foreshocks is totally released before the main-342 shock. Following the empirical scaling relationship $M_{0(postseismic)}/M_{0(coseismic)} = 0.36 +$ 343 /-0.2 proposed by Alwahedi and Hawthorne (2019), the aseismic displacement not re-344 lated to foreshocks is reduced to about 37% + / -13% of the total pre-seismic GPS ob-345 servations (Figure S12). The total observed displacement is therefore unlikely to be ex-346 plained by the contribution of foreshocks even when adding the associated afterslip. Such 347 evaluation should be taken with caution due to the non-linear nature of the relationship 348 between slip rate and co-seismic stress change for afterslip (e.g., Perfettini & Avouac, 349 2004; Perfettini et al., 2010). 350

Diverse numerical and experimental studies bring up the potential importance of aseismic preslip in the triggering of foreshocks (e.g., Kaneko et al., 2016; McLaskey & Kilgore, 2013). If such observations apply on natural faults, foreshock locations could potentially inform us about the overall spatial extent of the nucleation zone prior to an earthquake. This idea is in fairly good agreement with our results suggesting a first-order correlation between preslip distribution and the location of foreshocks (Figure 1 and Figure 4). Even if preslip appears to be an important mechanism in the triggering of fore-

shocks, part of the foreshock activity likely results from cascading phenomena due to stress 358 changes of neighboring events. In addition, we still need to understand why most earth-359 quakes are not preceded by foreshock activity and even less with observable pre-seismic 360 motion. This lack of systematic precursory activity might partly be due to an observa-361 tional gap due to the incompleteness of current seismicity catalog (as suggested by Mignan, 362 2014) or the lack of near fault geodetic observations prior to large earthquakes. The anal-363 ysis of an highly complete earthquake catalog in Southern California showed that 72%364 of $M_W \geq 4$ earthquakes in the region are preceded by an elevated seismic activity com-365 pared with the background seismicity rate (Trugman & Ross, 2019), suggesting that fore-366 shock activity is more ubiquitous than previously thought. However, a recent reanaly-367 sis of the same catalog suggested that a much smaller portion of these foreshock sequences 368 were really anomalous and could not be attributed to temporal fluctuations in background 369 seismicity rate (van den Ende & Ampuero, 2020). Although anomalous foreshock sequences 370 currently appears to be the exceptional, the improvement of near-fault geodetic and seis-371 mological observational capabilities are essential to bridge the gap between natural fault 372 observations and laboratory experiments, where foreshocks are commonly observed. 373

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Supporting Information for Seismic and aseismic fault slip during the initiation phase of the 2017 $M_W = 6.9$ Valparaíso earthquake

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Introduction

S1. GPS proceesing

68 continuous GPS (cGPS) were processed in South America (66 stations) and Nazca (2 stations) Plates (Figures S1 and S2), from different networks that are listed below:

:

• 13 cGPS from the International GNSS service (www.igs.org): ANTC, AREQ, BRAZ, BRFT, CHPI, GLPS, ISPA, KOUR, LPGS, RIO2, SANT, UFPR, UNSA.

• 3 cGPS from the Instituto Geográfico Militar of Bolivia (www.igmbolivia.gob.bo): SCRZ, URUS, YCBA.

• 11 cGPS from the Brazilian Network (RBMC-IP, www.ibge.gov.br): CUIB, MABA, MSCG, NAUS, POAL, POVE, PRCV, ROCD, RSAL, SAVO, TOPL.

• 15 cGPS from Argentian National Network (RAMSAC, www.ign.gob.ar (Piñón et al., 2018)) AZUL, BCAR, CATA, DINO, EBYP, ESQU, MA01, NESA, PEJO, RWSN, SL01, TUCU, UNRO, UNSJ, VBCA

• 5 cGPS from the Chilean - French cooperation through LIA "Montessus de Ballore" (www.lia-mb.net): CONS, JRGN, OVLL, UAPE, UDAT.

• 2 cGPS from the Ministerio de Bienes Nacionales of Chile (www.bienesnacionales .cl): BN05, BN13

• 18 cGPS from the Centro Sismológico Nacional de Chile (CSN, www.csn.uchile.cl (Báez et al., 2018)): CHDA, CTPC, CUVI, DGF1, LVIL, MPLA, NAVI, PORT, QTAY, RCSD, ROB1, QTAY, SLMC, TLGT, TRPD, UAIB, VALN, ZAPA.

All these data were processed in double differences using GAMIT 10.7 software to obtain daily, 12 and 6 hours estimates of station positions, choosing ionosphere-free combination and fixing the ambiguities to integer values. The precise orbits from the International GNSS Service for Geodynamics, precise EOPs from the IERS bulletin B, IGS tables to describe the phase centers of the antennas, FES2004 ocean-tidal loading corrections, as well as atmospheric loading corrections (tidal and non-tidal). We used precise orbits from the International GNSS Service for Geodynamics, precise EOPs from the IERS bulletin B, IGS tables to describe the phase centers of the antennas, FES2004 oceantidal loading corrections, as well as atmospheric loading corrections (tidal and non-tidal). One tropospheric vertical delay parameter and two horizontal gradients per stations are estimated every 2 hours. Daily solutions and position time series are combined using the PYACS software (Nocquet, 2017) in a regional stabilization process. The results are mapped into ITRF 2014 reference frame (Altamimi et al., 2016) and then put in the South-American frame using the Euler pole at -83.4° E, 15.2° N, and angular velocity $0.287^{\circ}my^{-1}$ (Nocquet et al., 2014).

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X - 4

S2. Prediction error covariance matrix

We focus on prediction uncertainties due to inaccuracies in the Earth model. These uncertainties are represented by the matrix $\mathbf{C}_{\mathbf{p}}$. We note the forward model $\mathbf{g}(\boldsymbol{\Psi}, \mathbf{m})$ for a source model \mathbf{m} , and Earth model parameters $\boldsymbol{\Psi}$ (i.e., P and S wave velocities, density). We can estimate $\mathbf{C}_{\mathbf{p}}$ empirically from an ensemble of random models $\boldsymbol{\Psi}_{\mathbf{i}}$, (i = 1, ..., n)as:

$$\mathbf{C}_{\mathbf{p}} = \frac{1}{n-1} \sum_{i=1}^{n} (\mathbf{g}(\boldsymbol{\Psi}_{\mathbf{i}}, \mathbf{m}) - \bar{\mathbf{g}}(\boldsymbol{\Psi}, \mathbf{m})) (\mathbf{g}(\boldsymbol{\Psi}_{\mathbf{i}}, \mathbf{m}) - \bar{\mathbf{g}}(\boldsymbol{\Psi}, \mathbf{m}))^{T},$$
(1)

where $\bar{\mathbf{g}}$ is the mean of the ensemble of predictions $\mathbf{g}(\Psi_{\mathbf{i}}, \mathbf{m})$. In the following, we refer to $\mathbf{C}_{\mathbf{p}}$ estimated in equation (1) as the empirical prediction error covariance matrix. Alternatively, we can compute $\mathbf{C}_{\mathbf{p}}$ following a linearized perturbation approach. We assume that our forward model $\mathbf{g}(\Psi, \mathbf{m})$ is well approximated by linearized perturbations of our predictions. For an a priori Earth model $\tilde{\Psi}$ we write:

$$\mathbf{g}(\mathbf{\Psi}, \mathbf{m}) \approx \mathbf{g}(\tilde{\mathbf{\Psi}}, \mathbf{m}) + \mathbf{K}_{\mathbf{\Psi}}(\tilde{\mathbf{\Psi}}, \mathbf{m}) \cdot (\mathbf{\Psi} - \tilde{\mathbf{\Psi}}),$$
 (2)

where $\mathbf{K}_{\Psi}(\tilde{\Psi}, \mathbf{m})$ is the sensitivity kernels of the predictions with respect to elastic parameters used to compute forward predictions:

$$\mathbf{K}_{\Psi}(\tilde{\Psi}, \mathbf{m}) = \frac{\partial g_i}{\partial \Psi_j} (\tilde{\Psi}, \mathbf{m}).$$
(3)

In this first order approximation, we use the sensitivity kernel $\mathbf{K}_{\Psi}(\tilde{\Psi}, \mathbf{m})$ to estimate the covariance matrix $\mathbf{C}_{\mathbf{p}}$ (Duputel et al., 2014):

$$\mathbf{C}_{\mathbf{p}} = \mathbf{K}_{\boldsymbol{\Psi}} \cdot \mathbf{C}_{\boldsymbol{\Psi}} \cdot \mathbf{K}_{\boldsymbol{\Psi}}^{\mathbf{T}},\tag{4}$$

where \mathbf{C}_{Ψ} is the covariance matrix describing uncertainty in the Earth model. To analyze both approaches, we consider a simple test case limited to an uncertain in S-wave velocity

in a single layer (at 30 km depth) using the source parameters of the $M_W = 6.0$ foreshock on 2017-04-23 (see section 3 of the main text). For comparison, we calculate prediction error covariance matrices C_p using equation (1) and equation (4). We plot in Figure S7 the diagonal components of both matrices for a representative station. We observe that there is an overall good agreement between our first order C_p and the empirical C_p matrix. We notice some discrepancies in the variance amplitudes and a time-shift in the late part of the waveforms (after 75s in Figure S7). To explore the origin of these effects, we compare synthetic waveforms predicted from the stochastic models and the waveforms calculated with the first order approach. The results shown in Figure S8 indicate that the time-shift and amplitude difference in Figure S7 are related to the fact that the first order approach is unable to perfectly reproduce large perturbations in the Earth model.

To correct these differences, we can also estimate a covariance matrix using a second order approximation of the forward model as:

$$\mathbf{g}(\boldsymbol{\Psi},\mathbf{m}) \approx \mathbf{g}(\boldsymbol{\tilde{\Psi}},\mathbf{m}) + \mathbf{K}_{\boldsymbol{\Psi}}(\boldsymbol{\tilde{\Psi}},\mathbf{m}) \cdot (\boldsymbol{\Psi}-\boldsymbol{\tilde{\Psi}}) + \frac{1}{2!} (\boldsymbol{\Psi}-\boldsymbol{\tilde{\Psi}}) \cdot \mathbf{H}_{\boldsymbol{\Psi}}(\boldsymbol{\tilde{\Psi}},\mathbf{m}) \cdot (\boldsymbol{\Psi}-\boldsymbol{\tilde{\Psi}}), (5)$$

where \mathbf{H}_{Ψ} is the second order derivative with respect to the elastic parameters:

$$\mathbf{H}_{\Psi}(\tilde{\Psi}, \mathbf{m}) = \frac{\partial^2 g_i}{\partial \Psi_k \partial \Psi_j} (\tilde{\Psi}, \mathbf{m}).$$
(6)

The computation of H involves evaluating n^2 derivatives, where n is the number of elastic parameters (e.g., 3 parameters per layer for a 1D Earth model). However, assuming that cross-terms are negligible, we can reduce the number of 2nd order derivatives to be evaluated to n.

As shown in Figure S7 and S8, some of the imperfections obtained with the first order approach can be corrected by employing a second order approach neglecting cross-terms.

In practice, these discrepancies are more significant when we apply larger perturbations to the velocity model. Despite the fact that the inaccuracies of the first order approach have been corrected, we notice in Figure S8 that the differences between the first and second order approach are relatively small given the 1 Hz sampling frequency used in our moment tensor inversions. Our tests show that the differences are more visible when inverting waveforms with a higher sampling rate.

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Figure S1. Map of the GPS stations processed in South America and Nazca Plates. The red stations are those ones used to define the Reference Frame, while the blue ones are just used on the processing. The pink box denotes the study area (see Figure S2 to look at the stations processed in this region).



Figure S2. Map of the GPS stations processed in the study area.



Figure S3. GPS time series for the Valparaíso region network for north and east component. The images show the time series before and after the mainshock (green line) of the sequence.





Figure S4. Model variability of the P-wave, S-wave, and density as a function of depth in Valparaíso region. Black line represents the velocity layered model used for Green's Function (GF) calculation. Grey histograms are the probability density function for each parameter as a function of depth as described in Cp.



Figure S5. Gutenberg-Richter law for the 2017 Valparaíso earthquake sequence. Three different catalogs of the sequence are shown: Our CMT catalog, S. Ruiz et al. (2017) catalog, and J. A. Ruiz et al. (2018) catalog. For each catalog, both the whole sequence (foreshocks and aftershocks), and the foreshocks sequence are represented.



Figure S6. Synthetic surface displacement for different ranges of magnitude, foreshocks with $M_w \ge 5.5$ (largest foreshock $M_w = 6.0$) and foreshocks with $M_w \le 5.5$. The $M_w = 6.0$ contribution appears to dominate the signal, with respect to the cumulative contribution of smaller foreshocks.



Figure S7. Diagonal of the C_p matrix for the vertical component of the station G005. The matrix is calculated for the $M_W = 6.0$ foreshock of the Valparaíso sequence (see section 3 of the main text). The red line represents the diagonal matrix for the empirical covariance matrix (i.e., the matrix created from an ensemble of models). The blue line represents the first (top) and second-order (bottom) approaches used to compute C_p .





Figure S8. Comparison between synthetic waveforms predicted from stochastic models calculated with a log-normal distribution, and synthetic waveforms calculated using the first and second order C_p matrix. The waveforms are generated using the source model of the $M_W = 6.0$ foreshock presented in section 3 of the main text. The X-axis represents time shifts between waveforms generated with the average velocity model of the region (figure S4) and waveform predicted for randomly perturbed velocity models. The Y-axis represents time shifts between waveforms generated with the average velocity model and waveforms generated either with the first or the second order approximation (see equations (2) and (5) of text S2). The color represents the correlation coefficient of each pair of waveforms. If the comparison follows the y = xline, it means that the perturbation approximation properly estimates the empirical covariance matrix. We can observe that the second order approach better approximates actual synthetics (especially when there is a significant time-delay between waveforms).



OFF_COAST_CENTRAL_CHILE, filter = (0.01, 0.04, 4, 1), p 1/4



time, sec

time, sec

December 4, 2020, 4:06pm

time, sec



0.00

0.05

200

400

time, sec

600

0.02

-0.02

200

400

time, sec

E 0.00

OFF_COAST_CENTRAL_CHILE, filter = (0.01, 0.04, 4, 1), p 3/4

Figure S9. Waveforms fit for the $M_W = 6.0$ foreshock using CMT solution from our catalog. Observed (black) and synthetic (red) waveforms for a given station (orange). The fit (inversion) is made between red dots. The blue December exercise the CMT Procession. Yellow dots correspond to the ensemble of stations used in the inversion.

0.02

0.00

200

600

400

time, sec

600



Figure S10. Stations used for the $M_W = 6$ foreshock CMT inversion. The CMT location is shown in purple.



Figure S11. Mainshock postseismic surface displacement normalized by the coseismic displacement at each GPS station. This ratio approximates the moment ratio between postseismic and coseismic terms.





Figure S12. Same as Figure 4 of main text but with the quick postseismic contribution produced by the largest foreshock.

 Table S1.
 Bandpass filter corner frequencies used for CMT inversion

Magnitude	Low Corner Freq (Hz)	High Corner Freq (Hz)
< 4.5	0.02	0.08
> 4.5	0.015	0.06
6.0	0.01	0.04

X - 22	$Mtp N \cdot m$	2.927e+21	3.324e + 21	9.225e + 20	9.259e + 20	-5.981e + 22	-4.994e + 22	-1.311e + 22	4.470e + 21	-2.778e + 21	-6.063e + 20	-9.098e + 20	8.445e + 22	-3.027e+20	-1.181e + 21	4:655e+21	-8.387e + 21	4.638e + 20	-1.127e+21	-1.953e + 22	5.282e + 21	6.894e + 21	5.773e+21	-6.659e + 21	-3.851e+20	3.680e + 21	4.963e + 21	4.637e+19	-6.557e+20	-1.002e+20	-3.549e+21	-2.089e+20	$n \ next \ page$
	Mrp N·m	-1.83e + 21	-1.41e+23	-1.70e+22	-6.24e + 22	-9.06e + 24	-1.44e + 23	-4.66e + 22	-3.43e+22	-4.39e + 22	-4.30e+21	-1.63e + 22	-1.31e+24	-5.86e + 21	-1.12e+22	-3.11e+22	-5.55e+22	-6.13e + 21	-4.09e+21	-1.16e + 23	-4.85e + 22	-9.06e + 22	-4.56e + 23	-1.22e+22	-7.10e + 21	-5.93e+22	-1.27e+22	-3.50e+21	-7.48e + 21	-8.67e+21	-3.20e+22	-1.76e+22	Continued o
	Mrt N·m	-1.82e+20	-2.78e+22	-8.45e+20	-5.42e+21	4.10e + 23	-1.86e+22	1.29e+21	-9.10e + 21	6.45e+21	4.33e+20	-7.36e+20	5.75e+22	2.95e+21	4.68e + 20	3.65e+21	8.92e + 21	1.17e+21	-7.60e+18	2.77e+22	-5.64e+21	1.73e+22	2.10e+22	8.85e+20	1.11e+21	8.15e + 21	-6.61e+19	-6.10e+20	1.41e+20	3.61e+20	5.88e + 21	-1.85e+21	
	Mpp N·m	1.01e+21	-1.60e+23	-1.17e+22	-5.83e+22	-7.05e+24	-1.03e+23	-6.36e + 22	-3.19e+22	-6.97e+22	-3.34e+21	-9.07e+21	-1.59e+24	-1.62e+22	-7.14e+21	-3.05e+22	-5.30e+22	-4.07e+21	-3.23e+21	-9.04e + 22	-5.51e+22	-1.23e+23	-4.82e+23	-4.97e+21	-5.36e + 21	-6.58e + 22	-1.70e+22	-3.10e+21	-5.43e+21	-8.78e+21	-1.87e+22	-8.04e+21	
our catalog.	Mtt $N \cdot m$	-8.60e+20	-1.58e+22	-1.07e+21	2.57e+21	-3.96e + 23	2.41e + 22	-5.19e + 20	-7.83e+21	-5.13e + 21	2.78e + 20	-2.14e+21	-1.06e+23	1.06e + 22	9.72e + 20	-3.71e+21	-2.75e+21	-7.75e+20	1.04e + 21	3.38e + 22	-5.41e+21	-3.88e+21	-3.77e+22	2.37e+21	-2.88e+21	-1.13e+22	-5.64e+21	-5.02e+20	-5.54e+20	7.54e + 20	-3.06e+21	-2.08e+20	
olutions of	$Mrr N \cdot m$	-1.49e + 20	1.76e + 23	1.28e + 22	5.58e + 22	7.45e + 24	7.84e + 22	6.41e + 22	3.97e + 22	7.48e + 22	3.07e+21	1.12e + 22	1.69e + 24	5.58e+21	6.16e + 21	3.42e + 22	5.58e + 22	4.85e + 21	2.19e + 21	5.65e + 22	6.05e + 22	1.27e + 23	5.20e + 23	2.60e+21	8.24e + 21	7.71e+22	2.26e + 22	3.60e+21	5.98e + 21	8.03e + 21	2.18e + 22	8.25e+21	
MT s	Mw	3.63	4.83	4.15	4.55	6.0	4.76	4.54	4.40	4.55	3.76	4.12	5.5	4.05	4.01	4.37	4.53	3.86	3.73	4.70	4.52	4.73	5.15	4.04	3.93	4.58	4.19	3.73	3.92	3.99	4.32	4.13	
able S2: C	$M0 N \cdot m$	3.56e + 14	2.21e+16	2.10e+15	8.47e + 15	1.16e + 18	1.76e + 16	8.02e+15	5.04e+15	8.48e + 15	5.42e + 14	1.93e+15	2.10e+17	1.50e+15	1.31e+15	4.53e+15	7.87e+15	7.69e+14	5.03e+14	1.43e+16	7.58e+15	1.56e + 16	6.78e+16	1.45e+15	9.88e + 14	$9.33e{+}15$	2.43e+15	4.88e + 14	9.43e + 14	1.21e+15	3.85e+15	$1.95e{+}15$	
L	Depth km	35.5	17.5	21.5	22.5	19.5	21.5	25.5	14.5	26.5	26.5	25.5	21.5	30.5	22.5	24.5	21.5	23.5	28.5	35.5	20.5	22.5	22.5	22.5	32.5	23.5	17.5	25.5	24.5	19.5	27.5	21.5	
	Lat $^{\circ}$	-31.93	-33.14	-33.05	-33.03	-33.03	-33.05	-33.05	-32.86	-33.07	-33.07	-32.99	-33.05	-33.06	-33.11	-33.09	-33.10	-33.11	-33.00	-33.29	-33.17	-33.16	-33.16	-33.10	-33.00	-33.16	-33.03	-32.97	-32.97	-33.15	-33.10	-32.92	
	Lon°	-70.85	-71.96	-72.03	-72.06	-72.10	-71.89	-72.00	-71.97	-72.02	-71.99	-71.96	-72.16	-72.10	-72.04	-72.14	-72.05	-72.06	-71.93	-71.93	-72.04	-72.04	-72.09	-72.11	-71.89	-72.08	-72.28	-71.93	-71.98	-72.09	-72.00	-72.11	
	Time	01:50:23	22:46:44	23:57:13	01:49:12	02:36:06	02:43:18	02:52:38	03:00:12	03:02:17	12:52:15	16:12:54	19:40:10	20:30:50	01:19:42	03:50:50	03:54:11	06:54:36	13:17:02	23:54:45	00:17:36	01:33:15	01:43:03	01:54:30	02:33:05	03:02:23	05:56:26	06:34:15	08:15:17	08:29:06	09:33:31	10:20:23	
	Date	2017-04-15	2017-04-22	2017-04-22	2017-04-23	2017-04-23	2017-04-23	2017-04-23	2017-04-23	2017-04-23	2017-04-23	2017-04-23	2017-04-23	2017-04-23	2017-04-24	2017-04-24	2017-04-24	2017-04-24	2017-04-24	2017-04-24	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	

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Table

	Mtp N·m	-2.237e+21	1.057e+21	2.462e + 20	2.311e+21	1.217e+21	2.934e + 21	2.174e + 21	-5.471e+20	1.966e + 21	-3.074e + 19	6.534e + 20	1.225e+21	-5.287e+19	1.765e + 20	1.143e + 21	2:401e+21	1.043e + 21	7.488e+22	$\cdot 1.304e + 20$	3.411e + 21	4.253e+20	-4.765e+20	5.323e + 23	-1.243e + 22	-2.064e + 22	1.129e + 22	2.300e + 23	5.465e + 23	-3.073e+21	-6.374e+19	-2.420e + 22	3*052e+22	n gent page
	Mrp N·m	-1.43e+22 -	-6.50e+21	-9.24e + 21	-6.92e+22	-1.02e+22	-2.37e+22	-4.93e+21	-1.06e+22	-1.58e+22	-6.90e+21	-6.05e+21	-8.10e + 21	-4.19e+21	-2.64e+22	-3.00e+22	-2.42e+22	-3.61e+21	-2.13e+23	-1.42e+22	-7.88e+22	-1.47e+22	-6.69e+21	-5.03e+24	-1.58e+23	-4.91e+22	-2.18e+23	-8.39e+23	-3.88e+24	-6.67e+22	-2.64e+22	-8.34e+21	-7.79e+23	Jontinued or
	Mrt N·m	-1.29e+21	$3.36e{+}19$	3.35e+21	1.81e+22	1.08e+20	2.13e+21	-5.21e+20	4.63e+20	7.19e+21	1.11e+21	1.96e + 21	1.59e+20	-4.96e+20	6.22e+21	3.36e + 21	9.17e+21	9.98e + 20	8.57e+21	1.08e+21	6.07e+21	1.44e+21	2.02e+20	1.60e+24	4.19e + 22	4.84e+22	3.56e+22	3.88e+23	-6.71e+23	2.52e+22	9.80e+21	2.62e+22	2.17e+21	
	Mpp N·m	-1.62e+22	-7.24e+21	-1.28e + 22	-9.21e+22	-7.94e+21	-1.41e+22	-9.41e + 21	-9.49e+21	-9.84e+21	-8.83e+21	-4.69e+21	-8.51e+21	-2.03e+21	-2.82e+22	-3.31e+22	-3.88e+22	-9.59e+21	-3.43e+23	-1.16e+22	-1.12e+23	-8.01e+21	-6.50e+21	-5.10e + 24	-1.57e+23	-5.65e+22	-1.98e+23	-8.98e+23	-2.93e+24	-3.91e+22	-1.72e+22	2.38e + 22	-7.70e+23	
verious page	Mtt $N \cdot m$	-2.81e+21	-1.21e+21	-2.19e+21	-1.01e+22	-3.74e+20	7.15e+20	-2.05e+21	4.71e+20	6.87e + 20	-9.29e+20	-2.24e+21	-1.85e+21	2.60e + 20	-3.83e+21	-2.38e+20	7.82e + 20	-3.26e+20	1.23e + 22	-7.08e+20	7.86e + 21	-5.83e+20	1.21e + 21	-2.17e+23	-1.56e+22	1.25e + 22	2.36e + 22	-2.51e+22	-4.38e+23	1.64e + 22	2.05e+21	-8.40e+21	-2.82e+22	
$ued from p_1$	$Mrr N \cdot m$	1.90e + 22	8.45e + 21	1.50e + 22	1.02e + 23	8.31e + 21	1.34e + 22	1.15e + 22	9.02e + 21	9.15e + 21	9.76e + 21	6.93e + 21	1.04e + 22	1.77e+21	3.21e + 22	3.33e+22	3.80e + 22	9.92e + 21	3.31e+23	1.23e + 22	1.04e + 23	8.60e + 21	5.29e + 21	5.32e + 24	1.72e+23	4.40e + 22	1.74e + 23	9.24e + 23	3.37e+24	2.26e + 22	1.51e+22	-1.54e+22	7.98e+23	
Jontin	Mw	4.17	3.94	4.09	4.65	4.01	4.23	3.98	4.03	4.13	3.98	3.89	4.00	3.71	4.34	4.37	4.38	3.95	5.01	4.11	4.68	4.09	3.90	5.85	4.84	4.56	4.91	5.35	5.74	4.53	4.27	4.35	5.30	
where $S2 - C$	$M0 N \cdot m$	2.28e + 15	1.03e+15	1.70e+15	1.21e+16	1.31e+15	2.76e + 15	1.18e + 15	1.41e+15	1.99e+15	1.16e + 15	8.67e + 14	$1.25e{+}15$	4.63e + 14	4.05e+15	$4.49e{+}15$	4.64e + 15	1.05e+15	$4.06e{+}16$	1.86e + 15	$1.34e{+}16$	$1.70e{+}15$	8.93e + 14	7.43e+17	2.32e+16	8.66e + 15	$2.89e{+}16$	$1.32e{+}17$	5.06e + 17	7.87e+15	$3.25e{+}15$	$4.21e{+}15$	1.11e + 17	
T_{2}	Depth km	22.5	19.5	19.5	19.5	21.5	21.5	19.5	19.5	24.5	19.5	17.5	26.5	26.5	21.5	27.5	31.5	32.5	27.5	25.5	28.5	23.5	29.5	23.5	23.5	28.5	26.5	27.5	29.5	20.5	24.5	29.5	25.5	
	Lat°	-32.90	-33.03	-33.12	-33.12	-33.05	-33.17	-33.10	-33.07	-33.31	-33.13	-33.12	-32.95	-32.94	-33.14	-33.29	-33.30	-33.13	-33.31	-33.29	-33.28	-33.10	-33.30	-33.26	-33.32	-33.26	-33.31	-33.26	-33.17	-33.25	-33.23	-33.35	-33.30	
	Lon°	-72.14	-72.20	-72.27	-72.21	-72.16	-72.10	-72.21	-72.18	-72.01	-72.15	-72.06	-71.98	-71.96	-72.21	-71.85	-71.99	-71.81	-71.90	-71.88	-71.89	-72.06	-71.92	-72.02	-71.96	-71.91	-71.91	-72.05	-71.66	-71.93	-72.06	-71.93	-71.98	
	Time	10:24:35	11:22:02	11:24:09	12:13:23	12:37:37	14:26:35	15:32:07	16:38:53	16:48:36	20:57:54	21:03:13	23:58:11	00:43:00	10:05:34	14:45:55	15:14:01	01:55:05	05:09:22	06:55:45	08:24:41	08:46:34	21:17:33	15:30:05	15:33:30	15:40:24	15:49:44	15:58:34	16:05:57	17:09:40	17:21:48	17:38:09	17:41:50	
	Date	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-25	2017-04-26	2017-04-26	2017-04-26	2017-04-26	2017-04-27	2017-04-27	2017-04-27	2017-04-27	2017-04-27	2017-04-27	2017-04-28	2017-04-28	2017-04-28	2017-04-28	2017-04-28	2017-04-28	2017-04-28	2017-04-28	2017-04-28	2017-04-28	

	îp N∙m	22	~1			<u>г</u> -	÷	1.4	6.7	-1.90	1.113	-4.335	-3.070	-1.5966	1.035e	3.640e	$9:744\epsilon$	8.527€	-1.131	9.3646	4.4986	-6.968	-1.707	3.070	1.786	1.183	1.094	7.9126
	Μı	-1.88e+	-2.28e+22	-3.40e+21	-4.03e+22	-3.01e+22	-2.43e+23	2.43e + 21	-5.62e+21	-2.17e+22	-2.20e+22	-5.07e+21	-4.01e+21	-1.07e+22	-2.16e + 22	-2.99e + 21	-7.45e + 21	-4.22e+21	-3.11e+21	-4.54e+21	-2.55e+21	-3.78e + 23	-1.01e+22	-4.18e + 22	-2.60e+21	1.18e + 22	-5.11e+21	-4.68e + 21
	Mrt N \cdot m	-3.88e+21	3.90e+21	2.44e+21	-3.80e+21	2.14e+22	8.74e+22	-7.30e+20	3.07e+21	1.38e+22	8.88e+21	-6.18e + 20	4.23e+20	1.76e + 21	2.43e+21	6.57e+20	8.32e + 20	9.84e+20	1.07e+21	-4.60e+19	1.48e+19	-5.74e+22	3.93e+21	2.57e+20	7.27e+20	9.26e + 21	-1.05e+21	1.74e+21
	Mpp N·m	-1.79e+22	-1.39e+22	-2.63e+21	-3.49e+22	-3.74e+22	-4.54e+23	-5.71e+21	-4.11e+21	-1.60e+22	-1.07e+22	-5.67e+21	-8.73e+21	-1.35e+22	-4.63e+22	-4.57e+21	-7.48e + 21	-2.23e+21	-5.37e+21	-5.98e+21	-3.15e+21	-5.04e+23	1.03e+21	-2.59e+21	-1.98e+22	-6.02e+21	-6.71e+19	-6.34e + 21
revious page	Mtt $N \cdot m$	1.97e + 21	9.66e + 20	1.13e + 21	2.74e + 21	4.71e+21	-4.40e+22	-1.76e+21	-9.67e+19	7.25e+21	4.24e + 21	-2.22e+20	-1.74e+21	-1.38e+21	-1.11e+22	-2.63e+20	5.61e+20	8.61e+19	-1.53e+20	1.88e + 20	-9.09e+20	-3.86e+22	1.90e + 21	9.87e + 20	-3.86e+21	-7.61e+20	2.71e+21	-1.20e+21
ued from pr	$Mrr N \cdot m$	1.60e + 22	1.30e+22	1.49e + 21	3.22e + 22	3.27e+22	4.98e + 23	7.48e + 21	4.21e+21	8.79e+21	6.44e + 21	5.89e + 21	1.05e+22	1.49e + 22	5.74e+22	4.83e + 21	6.92e + 21	2.14e + 21	5.53e+21	5.79e+21	4.06e + 21	5.42e + 23	-2.93e+21	1.60e+21	2.37e+22	6.78e + 21	-2.65e+21	7.53e+21
Jontin	Mw	4.21	4.22	3.73	4.42	4.41	5.09	3.84	3.86	4.24	4.20	3.89	3.95	4.10	4.43	3.77	3.95	3.73	3.80	3.85	3.70	5.14	4.13	4.35	4.16	4.07	3.77	3.89
able $S2 - ($	$M0 N \cdot m$	2.57e+15	2.68e+15	4.87e + 14	$5.34e{+}15$	5.16e + 15	5.41e+16	7.30e+14	7.66e + 14	2.92e+15	$2.54e{+}15$	8.74e + 14	1.04e+15	1.79e+15	5.62e+15	5.62e + 14	1.04e+15	4.92e + 14	6.36e + 14	7.49e+14	4.44e + 14	$6.48e{+}16$	1.99e+15	$4.20e{+}15$	2.20e+15	1.61e+15	5.78e + 14	8.59e + 14
Ë	Depth km	30.5	29.5	20.5	27.5	25.5	24.5	32.5	22.5	24.5	22.5	26.5	22.5	18.5	21.5	22.5	18.5	22.5	16.5	21.5	19.5	26.5	16.5	27.5	17.5	19.5	26.5	18.5
-	Lat°	-33.36	-33.28	-33.25	-33.37	-33.23	-33.22	-33.12	-33.24	-33.25	-33.24	-33.37	-32.97	-33.01	-33.08	-33.09	-32.90	-32.82	-32.94	-33.75	-33.05	-32.94	-32.96	-32.00	-33.03	-32.94	-32.16	-32.98
-	Lon°	-71.94	-71.96	-72.02	-71.99	-72.07	-72.02	-71.73	-72.07	-72.05	-72.03	-72.02	-72.26	-72.21	-72.26	-72.15	-72.24	-71.99	-72.18	-72.27	-72.25	-72.06	-72.21	-71.65	-72.32	-72.16	-71.86	-72.19
	Time	17:57:07	18:28:23	01:06:23	$01{:}08{:}35$	01:37:16	01:46:00	02:36:24	$04{:}50{:}34$	08:30:43	08:54:02	17:55:34	21:49:02	23:38:45	16:50:22	14:31:43	01:34:46	04:42:01	10:48:21	09:22:31	11:28:32	16:54:46	02:16:29	04:36:15	00:44:56	$01{:}05{:}12$	20:39:36	06:45:58
	Date	2017-04-28	2017-04-28	2017-04-29	2017-04-29	2017-04-29	2017-04-29	2017-04-29	2017-04-29	2017-04-29	2017-04-29	2017-04-30	2017-04-30	2017-05-01	2017-05-03	2017-05-04	2017-05-05	2017-05-05	2017-05-05	2017-05-09	2017-05-09	2017-05-13	2017-05-16	2017-05-16	2017-05-18	2017-05-23	2017-05-29	2017-05-30