Constraining Crustal Properties with Bayesian Joint Inversion of Vertical and Radial Teleseismic P-wave Coda Autocorrelations

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Abstract

The sensitivity of seismic compressional and shear waves and their velocity ratios to rock lithology, pore fluids, and hightemperature materials makes these parameters very useful for constraining the physical state of the crust. In this study, we develop a joint inversion approach utilizing both radial and vertical components' autocorrelations of teleseismic P-wave coda for imaging the crust by simultaneously characterizing the crustal Vp, Vs and Vp/Vs ratio. Autocorrelations of the radial and vertical components contain P and S waves that are reflected back from the subsurface. Therefore, joint inversion of them can account for the variations of both Vp and Vs, and consequently, the Vp/Vs ratio. Synthetic inversions show significant improvement in the estimation of these parameters compared to those from the inversion of either, receiver functions or the autocorrelation of the vertical component. The velocity models inferred from the application of the approach on teleseismic data recorded along a north-south passive seismic profile (BILBY experiment) in central Australia reveal crustal-scale structures that clearly separate crustal domains. We found a thick crust (between 43 km and 57 km) along the BILBY profile. The general trend of the Moho structure imaged from the joint inversion corresponds well with the change of the reflectivity that is normally seen at the base of the crust and also with the Moho as interpreted using the deep seismic reflection profiling method.

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Key Points:

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9	• We introduce a new joint inversion approach to estimate V_p , V_s simultaneously,
10	and V_p/V_s ratio structures below a seismic station
11	• We jointly invert both the vertical and radial component autocorrelations of the
12	teleseismic P-wave coda to image and characterize the crust
13	• The joint inversion provides better constraints than the inversion of either P re-
14	ceiver functions or the vertical component autocorrelation

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15 Abstract

The sensitivity of seismic compressional and shear waves and their velocity ratios to rock 16 lithology, pore fluids, and high-temperature materials makes these parameters very use-17 ful for constraining the physical state of the crust. In this study, we develop a joint in-18 version approach utilizing both radial and vertical components' autocorrelations of tele-19 seismic P-wave coda for imaging the crust by simultaneously characterizing the crustal 20 V_p , V_s and V_p/V_s ratio. Autocorrelations of the radial and vertical components contain 21 P and S waves that are reflected from the subsurface. Therefore, joint inversion of them 22 can account for the variations of both V_p and V_s , and consequently, the V_p/V_s ratio. Syn-23 thetic inversions show significant improvement in the estimation of these parameters com-24 pared to those from the inversion of either, receiver functions or the autocorrelation of 25 the vertical component. The velocity models inferred from the application of the approach 26 to teleseismic data recorded along a north-south passive seismic profile (BILBY exper-27 iment) in central Australia reveal a distinct pattern of the Moho and the V_p/V_s varia-28 tions across the crustal blocks/domain. The general trend of the Moho structure cor-29 responds well with the change of the reflectivity that can normally be seen at the base 30 of the crust and also with the Moho estimated from the previous studies including the 31 deep seismic reflection profiling method. The V_p/V_s structure at depths greater than 10 32 km shows dominant high values beneath locations where the crustal domains interact 33 (e.g., at transition from one domain to another). 34

³⁵ Plain Language Summary

Understanding variations of the crustal P and S waves $(V_p \text{ and } V_s)$ can indicate rock 36 lithology. Here, we introduce a new approach to estimate these wavespeeds simultane-37 ously using vibrations from earthquakes recorded at seismic stations. Autocorrelations 38 of the distant P-wave coda recorded on the vertical and radial components of a seismo-39 gram contain P- and S-wave reflections. Thus, they are sensitive to both the crustal wavespeeds 40 and the crustal discontinuities. We test the feasibility of the new method with synthetic 41 data, which reveals a substantial improvement in recovering the input V_p/V_s models. We 42 then apply the approach to teleseismic waveforms recorded along a passive seismic ex-43 periment in central Australia to infer the crustal structure. The results reveal a distinct 44 variation of the thickness of the crust and the V_p/V_s across geological units along the pas-45 sive seismic line. The overall pattern of the crustal thickness variations also matches the 46 change of reflectivity at the base of the crust. The inferred V_p/V_s model exhibits high 47 V_p/V_s values at depths greater than 10 km beneath locations where the crustal domains 48 interact. 49

50 1 Introduction

P-wave energy from a direct teleseismic source converts to S_v energy and vice versa 51 across a sharp seismic boundary beneath a seismic station. This incoming energy con-52 tains multiple reflections from the underlying structure as well as peg-leg multiples and 53 scattered and mode converted waves (coda). Teleseismic waves recorded on the verti-54 cal and radial components of a seismogram have been widely used to retrieve the impulse-55 response of the receiver-side structure by deconvolving the vertical component (P) from 56 the horizontal or radial component (S_v) . The resulting waveform is primarily sensitive 57 to sharp V_s changes with depth and is commonly referred to as the "P receiver function" 58 (hereafter RF), which represents the converted P- to S-wave energy directly beneath a 59 seismic station (Ammon, 1991; Langston, 1979; Vinnik, 1977). 60

In addition to the RFs, single station autocorrelation and cross-correlation of teleseismic and ambient noise records can provide complementary constraints on the crustal properties beneath a seismic station, and have been successfully applied to teleseismic and ambient noise data for imaging crustal and lithospheric discontinuities (Ruigrok &

Wapenaar, 2012; Tibuleac & von Seggern, 2012; Gorbatov et al., 2013; Kennett., 2015; 65 Kennett et al., 2015; Nishitsuji et al., 2016; Taylor et al., 2016; Sun & Kennett, 2016; 66 Saygin et al., 2017; Oren & Nowack, 2017; Kim et al., 2017; Sun et al., 2018; Heath et 67 al., 2018; Romero & Schimmel, 2018; Pham & Tkalčić, 2018; Becker & Knapmeyer-Endrun, 68 2018, 2019; Tork Qashqai et al., 2019; Kim et al., 2019; Wang et al., 2020; Andrés et al., 69 2020; Casas et al., 2020). Claerbout (1968) showed that, by autocorrelating the records 70 from a surface seismic station, the transmission response can be converted to a reflec-71 tion response. Vinnik (1977) introduced a correlation-based approach to detect P-to-S 72 converted waves (RFs) in the long period P-wave coda. Galetti and Curtis (2012) and 73 Sun and Kennett (2016) showed that the cross-correlation of the vertical and radial com-74 ponents of a single-station could be considered an alternative implementation of the clas-75 sical RFs (non-classical RFs). 76

Compared to the RFs, the autocorrelations of teleseismic P-wave coda recorded on 77 the radial and vertical components of a seismogram contain additional information. The 78 autocorrelations include both P- and P-to-S converted phases, whereas the RFs contains 79 P-to-S phases primarily because the P-waves are canceled by the deconvolution opera-80 tion in the RFs (Figure 1). Thus, both V_p and V_s can simultaneously be estimated by 81 jointly inverting the radial and vertical components' autocorrelations. Recently, the au-82 to correlation of teleseismic P-wave coda recorded on the vertical component has been 83 used within a Bayesian inversion framework to image the crust beneath Australia (Tork Qashqai 84 et al., 2019). They confirmed the utility of this approach for estimating large-scale crustal 85 properties such as V_p and Moho depths. However, to the best of our knowledge, the joint 86 inversion of both radial and vertical components' autocorrelations has not yet been demon-87 strated. Such a joint inversion approach offers a framework to further reduce the uncer-88 tainties associated with the crustal seismic properties $(V_p, V_S \text{ and } V_p/V_s)$, and thus can 89 provide an improved estimate of both V_p and V_s , and consequently the V_p/V_s ratio which 90 provides an indication of the rock lithology beneath a seismic station. 91

In this study, we demonstrate the feasibility of the probabilistic joint inversion of 92 radial and vertical components' autocorrelations for recovering the true 1-D structure 93 $(V_p, V_s \text{ and } V_p/V_s)$ beneath a seismic station. We compare the results of the joint in-94 version with the single inversions of RFs and autocorrelation of the vertical component 95 through a series of synthetic inversion tests. We then use this approach to image the crust 96 beneath central Australia along a north-south seismic profile (BILBY seismic array) and 97 compare our results with the results obtained from RF analysis (Sippl, 2016), Australian 98 Moho Model-AusMoho (Kennett. et al., 2011), the deep seismic reflection profiling (Korsch 99 & Doublier, 2016) and inversion of the vertical component autocorrelation (Tork Qashqai 100 et al., 2019). 101

In the following sections, we briefly describe the autocorrelation and the RFs methods, each exploiting the arrival time delays of different phases.

104 2 Method

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2.1 Autocorrelation of teleseismic P-wave coda

In this section we briefly describe the concept of autocorrelations of the vertical and radial components and the reader is referred to Frasier (1970), Langston (1979), Ammon (1991), Galetti and Curtis (2012), Sun and Kennett (2016), and Tauzin et al. (2019) for further details about the RF, cross-correlation and autocorrelation of components and their relationships.

In a plane-layered medium, when a steep incident plane P-wave from a distant earthquake (epicentral distances $\sim 30^{\circ} - 90^{\circ}$) impinges on an interface in the subsurface (e.g., Moho), the seismogram on the surface records direct P-wave energy, its direct P-to-S conversion phase, its peg-leg multiples and scattered waves including free-surface related multiples (Figure 1). The recorded teleseismic waveforms on the horizontal and verti cal components of a seismogram can be processed to obtain the impulse-response of the
 subsurface structure in the forms of RFs and autocorrelations.

Claerbout (1968) showed that for plane waves impinging at normal incidence to 118 a horizontally stratified acoustic medium with a free surface, the positive or negative part 119 of the autocorrelation of the acoustic transmission response corresponds to the reflec-120 tion response of the Earth beneath that station. We further illustrate this in Figure 1, 121 where for simplicity, we consider a simple synthetic 1-D Earth model including one crustal 122 123 layer over a half-space. The autocorrelations of the vertical or radial components of a seismometer contain seismic waves reflected back from the free-surface, such as Pppmp 124 and Ppsms (Figure 1). These reflections have the same delay times as PmP and SmS125 that are reflected back from the bottom of the model (e.g., the Moho) and can be thought 126 of as the Earth's response to a virtual source co-located with the receiver (zero-offset) 127 at the surface of the Earth (Figure 1). RFs are generated through the deconvolution of 128 the vertical component from the radial component, thus only includes P-to-S phases, and 129 the PmP phase has been removed by deconvolution (Figure 1). As seen from Figure 1, 130 the autocorrelations of both the radial and vertical components not only have the P-to-131 S phases that exist in the RF but also contain the PmP phase, and are able to carry ad-132 ditional information from the subsurface. Therefore, the inverse modeling of autocor-133 relation waveforms provides a means of subsurface characterization. 134



Figure 1. Left: schematic representation of a teleseismic plane wave impinging on a layer. Thickness, V_p , V_s , V_p/V_s of the layer are 35 km, 6.65 km/s, 3.69 km/s and 1.8, respectively. Note that the diagram is horizontally exaggerated to highlight the seismic phases. Right: shows the delay times of the main phases and the free-surface related multiples associated with the receiver function (top), the autocorrelation of the radial (middle) and vertical (bottom) components which were obtained using the model shown on the left. Capital P denotes upcoming P-wave in the half-space/mantle. All the lower case letters denote p or s waves in the crust (upgoing and downgoing waves).

135 **2.2 Bayesian Inversion Framework**

In the Bayesian inverse problem the general solution in the model space **m** is a task of combining pieces of information (expressed in the forms of probability density functions-PDF) provided by i) some measurements (observations/data space **d**), ii) a model or theory that can predict observations [forward problem $g(\mathbf{m})$], and iii) some a priori knowledge about the model $\rho(\mathbf{m})$ into the *a posteriori* knowledge, the posterior PDF $\sigma_M(\mathbf{m})$ (Tarantola, 2005; Debski, 2010). The data space is described as the space of all conceivable observations (e.g., instrumental responses) and is given by a probability density function (*PDF*) with the mean value at the center of this distribution and a width (dispersion, standard deviation) that quantifies the uncertainty of the observed data and their correlations (Tarantola, 2005; Mosegaard & Hansen, 2016). It has been shown that this posterior PDF under as-

sumption that the data space is a linear space can be proportional to (Tarantola, 2005)):

$$\sigma_M(\mathbf{m}) \propto \rho(\mathbf{m}) L(\mathbf{m}),$$
 (1)

¹⁴⁷ where $L(\mathbf{m})$ represents the likelihood function and gives a measure of how well a ¹⁴⁸ model \mathbf{m} explains or fits the data (**d**). If all uncertainties of the problem are approxi-¹⁴⁹ mated by the Gaussian model, $\sigma(\mathbf{m})$ can be given by (Tarantola, 2005; Gouveia & Scales, ¹⁵⁰ 1998):

$$\sigma_M(\mathbf{m}) \propto \exp(-S(\mathbf{m})),$$
 (2)

where $S(\mathbf{m})$ is sum of squares:

$$S(\mathbf{m}) = \frac{1}{2} \left[\left(g(\mathbf{m}) - d_{obs} \right)^t C_D^{-1} \left(g(\mathbf{m}) - d_{obs} \right) \right]$$
(3)

 C_D in equation (3) represents covariance matrix which is sum of the modelization and observational uncertainties (Gouveia & Scales, 1998; Tarantola, 2005). The modelization uncertainties are often neglected because they are difficult to quantify. In the case of independent uncertainties (the data covariance matrix is diagonal), equation (3) can be generalized to the following form (Tarantola, 2005) for joint inversion of autocorrelations of the vertical and radial components:

$$\sigma_M(\mathbf{m}) \propto \exp\left(-\frac{1}{2}\left[w_z \sum_{i=1}^N \left(\frac{d_i - g_i(\mathbf{m})}{\sigma_d^i}\right)^2 + w_r \sum_{j=1}^N \left(\frac{d_j - g_j(\mathbf{m})}{\sigma_d^j}\right)^2\right]\right)$$
(4)

The term inside the exponential is the misfit function, the total number of data points is N, i and j refer to data points in the autocorrelations of the vertical and radial components, respectively, g(m) denotes calculated or predicted data at each point i and j (forward problem) and σ_d^i and σ_d^j are 1σ uncertainties for each data point of the autocorrelation waveforms along the time axis, w_z and w_r are weighting values that can be used for the optimization of the inversion. In this study, w_z and w_r are set to one.

In practice, the Markov Chain Monte Carlo (MCMC) techniques are usually em-164 ployed to estimate the general solution of the inversion. MCMC methods are a class of 165 techniques used to sample a model space and can estimate a range of models that fit the 166 data well in the form of the posterior distribution. We adopted the Delayed Rejection 167 Adaptive Metropolis algorithm (DRAM) (Haario et al., 2006). DRAM is a Markov Chain 168 Monte Carlo (MCMC) approach and has been widely used in many hydrological and geo-169 physical model parameter estimation problems (Smith & Marshall, 2008; Ball et al., 2014; 170 Afonso et al., 2016; Tork Qashqai et al., 2018). It combines the Adaptive Metropolis (AM) 171 (Haario et al., 2001) with Delayed Rejection (DR) (Mira, 2001) algorithms. DR provides 172 a mechanism for drawing alternative samples when the current sample is rejected in the 173 standard Metropolis-Hastings algorithm (Tierney, 1994), thus improving the efficiency 174 of the sampling by reducing the number of rejections. In the AM, to reduce the sampling 175

¹⁷⁶ of low probability areas in the model space, the algorithm updates the covariance ma-¹⁷⁷ trix of the chain (after a non-adaption period) using all the previously accepted samples.

This process can be iterated at regular intervals.

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In this study, the total number of simulations per station is 400,000. The non-adaptation time includes 150,000 iterations (or samples), and the proposal distribution of the model is updated after the non-adaptation time every 3,000 samples.

2.3 Parameterization and Forward Problem

The 1-D crust below each seismic station is parameterized with a stack of ten plane-183 parallel homogeneous crustal layers bounded at the base of the crust by a homogeneous 184 upper mantle as a half-space. We use a step function parameterization, including ten crustal 185 layers for joint inversion of the field data. This allows the inversion to capture the po-186 tential gradational transition across crustal discontinuities, including the crust-mantle 187 boundary (Moho) that exist in central Australia (Kennett. et al., 2011). Each crustal 188 layer is described by three parameters: density (ρ), thickness variation (ΔH) and V_p/V_s . 189 Upper mantle parameters $(V_p/V_s$ and density). As we applied a move out correction to 190 the observed autocorrelations assuming a reference slowness of 0.065 s/km (see section 191 4.2), the slowness parameter is allowed to vary only between 0.064 and 0.066 km/s (fixed 192 around the reference slowness). 193

The first step in the MCMC inversion is to randomly draw values for the main un-194 known parameters from their prior distributions (Table 1). Note that for the perturba-195 tion of crustal thicknesses in our inversion framework, the MCMC algorithm needs an 196 initial crustal thickness model to perturb. Here, we assume that the initial crustal thick-197 ness below each seismic station is 40 km and includes ten layers, each with an equal thick-198 ness of 4 km. To perturb the initial crustal thickness model, the initial thickness of each 199 layer is randomly perturbed (by the MCMC algorithm) using the random value drawn 200 in the earlier step from the range of ΔH given in Table 1. In the next step, P-wave ve-201 locities for each layer (including the half-space) are estimated from the sampled random 202 density values using the empirical Nafe-Drake curve (Ludwig et al., 1970; Brocher, 2005). 203 The V_s value for each layer is then calculated from the computed V_p and the V_p/V_s val-204 ues drawn by the MCMC algorithm. 205

The synthetic seismograms are computed using the reflectivity method of Kennett (1983) as implemented by Randall (1989). Autocorrelations of the vertical and radial components are calculated to estimate the synthetic reflection responses of the 1-D Earth beneath each station. A first-order Butterworth band-pass filter in the frequency range of 1 and 2 Hz is applied before and after the autocorrelation and amplitudes are normalized to unity.

Main parameters	Minimum	Maximum
Density in layer one (g/cm^3)	2.4	2.7
Density in layer two to layer five (g/cm^3)	2.65	2.85
Density in layer six to layer ten (g/cm^3)	2.7	3.0
V_p/V_s (in all crustal layers)	1.65	1.95
Δh in layer one to layer nine (km)	-4.0	4.0
Δh in layer ten (km)	-5.0	15.0
Density in the upper mantle/half-space (g/cm^3)	3.25	3.5
Slowness (s/km)	0.064	0.066

Table 1. The list of all unknown parameters and priors.

3 Synthetic Inversion Tests

To demonstrate the feasibility of the joint inversion framework, we run a series of 213 inversion tests on synthetic autocorrelations and RFs generated using a synthetic 1-D 214 Earth model consists of four crustal layers and a half-space. The aim is to invert these 215 data using a ten-layer crust to recover the synthetic model (Figure 2 and Figure 3). We 216 generate synthetic seismograms using the model and fifty random events with horizon-217 tal slowness values varying between 0.04 and 0.08 s/km. A small amount of Gaussian 218 noise is added (with mean zero and standard deviation of 0.01) to the synthetic seismo-219 220 grams. Then, RFs, radial and vertical autocorrelations are generated from synthetic seismogram, the move-out correction is applied and all waveforms (for each data type) are 221 stacked to obtain the stack of the data and their uncertainties. We performed the fol-222 lowing three different inversion tests: 1) inversion of only RF, 2) inversion of only au-223 to correlation of the vertical component, and 3) the joint inversion of the vertical and ra-224 dial component autocorrelations. Results are displayed in Figure 2 and Figure 3. 225

A ten-layered parameterization is chosen because we invert the field data using a 226 ten-layered crust (next section), which improves data fitting. The mean values of the pos-227 terior distributions of V_p , V_s and V_p/V_s , as well as the true models, are given in Figure 228 2. Fits to the data are displayed in Figure 3, where the fit to synthetic data is reason-229 ably well. Results suggest that the inversions of RFs or autocorrelation of the vertical 230 component alone can provide reasonable estimates of the true V_s or V_p structures in a 231 probabilistic sense, but they cannot fully recover the true V_p/V_s ratio. This to a large 232 extent is because they are not sufficiently sensitive to either V_p or V_s and consequently 233 a small deviation from the true values can lead to significant changes in the V_p/V_s ra-234 tio. The joint inversion of the radial and vertical components' autocorrelations, on the 235 other hand, resolve a much better estimate of crustal properties, especially the V_p/V_s 236 ratio, and reduces the non-uniqueness of the inversion solution. This is an important as-237 pect of this joint inversion approach that we would like to emphasize here. 238



a) Inversion of the receiver function









Figure 2. Synthetic inversions of a four-layered model with a ten-layered parameterization. Density plots of the posterior distribution of the V_p/V_s , V_p and V_s structures for a) inversion of the RF, b) inversion of the autocorrelation of the teleseismic P-wave coda recorded on the vertical component, c) joint inversion of autocorrelations of the teleseismic P-wave coda recorded on the vertical and radial components. High probability areas are shown with hot colors. Magenta profiles are true (synthetic) velocity, and V_p/V_s structures and the dark blue profiles indicate the mean of the posterior distribution.



c) Joint Inversion of the vertical and radial component autocorrelations



Figure 3. Synthetic inversions of a four-layered model with a ten-layered parameterization. Fits to data for a) inversion of the RF, b) inversion of the autocorrelation of the teleseismic P-wave coda recorded on the vertical component, c) joint inversion of autocorrelations of the teleseismic P-wave coda recorded on the vertical and radial components. Magenta waveforms are synthetic data and the blue waveform shows the mean of the best 2000 accepted predicted data (gray waveforms). The black dashed lines are the three standard deviation bounds (3σ) for the data. The Pms, Pppmp and Ppsms phases associated with the Moho are also marked on the top of waveforms.

4 Joint Inversion of Field Data (central Australia)

Following the synthetic tests, we applied the new joint inversion method on field 240 data recorded on a series of broadband seismic sensors deployed along a north-south pro-241 file (BILBY) in central Australia, comprised of 25 broadband seismic sensors operated 242 between August 2008 and February 2011. This profile spans over 1000 km from the south 243 of Australia to the north of the Australian continent, cutting several east-west trending 244 geological domains of central Australia including the Gawler Craton, Eromanga and Of-245 ficer Basins (in the Nawa Domain of the Gawler Craton), Musgrave Province, Amadeus 246 Basin, Arunta Block and Georgina Basin (Figure 4). In the following sections, we first 247 summarize the data processing and then discuss the results. 248



Figure 4. a) The BILBY experiment seismic stations are marked on the topographic map of Australia. The blue box shows the area shown in the right sub-figure (b) where the locations of the seismic stations and the GOMA and 09GA active seismic reflection lines are superimposed on the main geological units of central Australia. The crustal properties (shown in Figure 8) are extracted along the red dashed line.

4.1 Data Processing and Selection Criteria

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In this study, we follow the approach of Tork Qashqai et al. (2019) for data selec-250 tion and processing. We use teleseismic P-wave coda associated with events with mag-251 nitude greater than 5.5 and epicentral distances between 30° and 90° recorded by the 252 vertical and radial components of broadband seismic stations along the BILBY profile. 253 The chosen time window for P-wave coda is 10-70 s (tapering is applied) after the on-254 set of the theoretical P-wave arrivals predicted by the ak135 model (Kennett et al., 1995). 255 Only the first 30 seconds of this time window is used in the optimization of the joint in-256 version (when comparing the predicted and observed autocorrelations). Previous stud-257 ies showed that there is coherent, prominent and stable scattered energy with compa-258 rable phase velocity to the direct P or PP phases in teleseismic P-wave coda (Aki, 1982; 259 Dainty, 1990; Sens-Schönfelder et al., 2015; Kennett & Fichtner, 2020 [and see text S3 260 and Figures S8-S57 in the supporting information]. We remove the mean and trend of 261 each record, and reject bad quality records (extremely noisy records, those with data gaps 262 and short recording length), and then resample the data to 10 Hz. 263

To retrieve the equivalent local reflection responses below each seismic station, we 264 apply a first-order Butterworth band-pass filter in the frequency range of 1 and 2 Hz be-265 fore and after the autocorrelation, and then normalize amplitudes to unity. For both the 266 radial and vertical components, we stack the autocorrelation waveforms from multiple 267 events (after the move-out correction; see section 4.2) to increase the signal-to-noise ra-268 tio and suppress the source effects on reflection seismograms (Claerbout, 1968). We note 269 here that one may need at least 30 events (for a simple structure) to 150 events (for a 270 complex structure) to estimate a consistent velocity structure to that estimated from an 271 inversion whose input (stacked autocorrelation) is calculated from all available events 272 (Tork Qashqai et al., 2019, see supporting information). Tork Qashqai et al. (2019, see 273 supporting information) showed that phases such as reflections from the core-mantle bound-274 ary (PcP) can contaminate our selected time window. They showed via synthetic cal-275 culations that the PcP reflections arrive in our selected time window of the P-wave coda 276 at epicentral distances between 63° and 77° (slowness of 0.059-0.05 s/km). These core-277 mantle reflections can be considered as crustal reflections by the inversion. However, in 278 this study, the average slowness for each seismic station is greater than 0.06 s/km. Fur-279

thermore, PcP reflections are suppressed by the move-out correction and stacking processes. Therefore, we believe their effect in our inverse modeling is negligible.

We note that for inversion to be robust, one needs to be cautious in describing the 282 data uncertainties in the probabilistic inversion framework because the solution of the 283 inverse problem is not unique. In the Bayesian joint inversion, for each station, we used 284 (1σ) bounds of the variability (observational uncertainties/standard deviations) of the 285 raw (observed) autocorrelations as 1σ values (σ_d^i and σ_d^j in equation (4)). One may ar-286 gue that the observed data are the stacked autocorrelations and the standard deviations 287 of them (can be estimated using the bootstrap resampling (Efron & Tibshirani, 1991)) 288 should be used instead of the standard deviation of the observed autocorrelations. In the 289 supporting information (text S1 and text S2, Figures S2-S7), we demonstrate with syn-290 thetic inversion tests that it will be near impossible to recover the true or synthetic model 291 if the standard deviation of the stacked (mean) autocorrelation is used as data uncer-292 tainty in the probabilistic inversion framework. 293

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4.2 Pre-stack move-out correction of the autocorrelation waveforms

Traveltimes of converted and reverberation phases (e.g., direct Ps or Pms, Pppmp, 295 Pppms and Ppsms) in the autocorrelations of teleseismic P-wave coda are a function of 296 the incident angle (slowness) of the impinging seismic wave. Tauzin et al. (2019) showed 297 that for slowness values between 0.04 and 0.08 s/km, and for a simple model consists of 298 a horizontal one-layer crust over a half-space mantle, the downgoing P and S-waves in 299 the crust impinging on the Moho with an incident angle vary between 15 and 35° for P-300 waves, and between 10 and \sim 18 ° for S-waves. These incident angles over the range of 301 slownesses (e.g., 0.04-0.08 s/km) introduce a move-out (or shift) in the traveltime of the 302 converted and reverberation phases on the radial and vertical autocorrelations. There-303 fore, the converted and reverberated phases in the autocorrelations of the teleseismic P-304 wave coda need to be corrected for their move-out before the stacking. This is an im-305 portant step because if no move-out correction is applied to the autocorrelations prior 306 to the stacking, the misalignment of the phases can potentially weaken and broaden the 307 amplitudes of the stacked waveforms. Thus, the signal-to-noise ratio will not be improved 308 and the stacking becomes destructive which is particularly more severe when working 309 with high-frequency signals (Wang et al., 2020). In the presence of a thick crust, and in 310 the case of having a broad range of epicentral distances, this move-out is larger. This 311 is shown in Figures 5a and 5c with synthetic autocorrelations with slowness values be-312 tween 0.04 and 0.08 s/km. The autocorrelations are calculated using a simple synthetic 313 (horizontal) one-layer crust (with thickness of 50 km) over an upper mantle half-space. 314

Since there might be more than one layer in the crust, we use the AuSREM crustal model (Salmon et al., 2012) and discretize it in fine layers below each seismic station. Then for each autocorrelation trace, the relative traveltimes of the direct P-to-S conversion (Pms), and the three crustal reverberations phases (Pppmp, Pppms and Ppsms) with respect to the primary P-wave are approximated by the following equations (Chen & Niu, 2013; Sun & Kennett, 2016; Zhu & Kanamori, 2000):

$$t_{0p1s} = \sum_{i=0}^{N} \left(\sqrt{\frac{1}{\left(V_s\right)_i^2} - p^2} - \sqrt{\frac{1}{\left(V_p\right)_i^2} - p^2} \right) dz_i \tag{5}$$

$$t_{3p0s} = \sum_{i=0}^{N} 2\left(\sqrt{\frac{1}{(V_p)_i^2} - p^2}\right) dz_i \tag{6}$$

$$t_{2p1s} = \sum_{i=0}^{N} \left(\sqrt{\frac{1}{\left(V_s\right)_i^2} - p^2} + \sqrt{\frac{1}{\left(V_p\right)_i^2} - p^2} \right) dz_i \tag{7}$$

$$t_{1p2s} = \sum_{i=0}^{N} 2\left(\sqrt{\frac{1}{(V_s)_i^2} - p^2}\right) dz_i \tag{8}$$

Naming of the arrival times of the above phases, t_{mpns} , follows the convention used by Chen and Niu (2013) and Niu and James (2002), where **m** and **n** denote numbers of P and S-wave legs in the crust, *i* denotes layer's number, N is the total number of layers in the crust, p is the apparent horizontal slowness of the primary incident plane Pwave, and dz_i , $(V_p)_i$ and $(V_s)_i$ represent the thickness, P-wave and S-wave velocities of each fine layer in the AuSREM model, respectively.

For the move-out correction of Pms, Pppmp, Pppms and Ppsms phases, a refer-327 ence slowness needs to be chosen, at which the time scale of the radial and vertical au-328 to correlations remains unchanged. Following equations (2) to (5) and using the AuSREM 329 crustal model, the time scales of the radial and vertical autocorrelations are calculated 330 for all slowness values including the reference slowness. The calculated time scales of au-331 to correlations are either stretched and compressed with respect to the time scale of the 332 reference slowness. In this study, the mean values of all slowness values for all of the seis-333 mic stations vary between 0.064 and 0.066 km/s. Therefore, a fixed slowness value of 0.065334 s/km is chosen as a reference slowness for the move-out corrections. We adopt and mod-335 ify the "four-pin" approach developed by Chen and Niu (2013) to perform the move-out 336 corrections. Our move-out correction differs from the "four-pin" approach in that we cut 337 the autocorrelation traces into the following five segments (a "five-pin" approach): 1) 338 $0 < t \le t_{0p1s}, 2$ $t_{0p1s} < t \le t_{3p0s}, 3$ $t_{3p0s} < t \le t_{2p1s}, 4$ $t_{2p1s} < t \le t_{1p2s}, and 5$ $t > t > t_{1p2s}$ 339 t_{1p2s} . For each segment, we stretch/contract the time scale of each autocorrelation seg-340 ment with respect to the corresponding time segment in the time scale of the reference 341 slowness. 342

The stacked radial and vertical autocorrelations for all stations with and without applying the move-out correction are given the supporting information (Figures S106-S155).



Figure 5. Synthetic vertical autocorrelations a) before and b) after the move-out correction. Synthetic radial autocorrelations c) before and d) after the move-out correction. The approximate locations of the expected seismic phases are also marked in a) and c). Autocorrelations are in the range of slowness between 4.4-8.8 s/deg (0.04 and 0.08 s/km) but the Y-axis is re-scaled for better visualization of the move-out of the converted and reverberated phases.

4.3 1-D Examples

346

This section represents two example results obtained from the joint autocorrelation inversions for stations BL22 (Figure 6) and BL05 (Figure 7). Moho depths and their σ uncertainties inferred from our approach for all the stations are given in Table S1 in the supporting information. Locations of all the stations, including BL22 and BL05 are given in Figures 4,6, and Figure 7. Station BL22 is located in the younger Officer Basin. Figure 6 shows the posterior distributions of the V_p/V_s ratio, V_p and V_s in the forms of density plots, as well as the fits to the observed data (Figures 6a and 6b) for this station. Figure 6a shows that both the observed vertical and radial component autocorrelations are well fitted by the joint inversion. In density plots, the high probability areas are displayed with hot colors, and show five to six major velocity jumps in the crust at depths of approximately 4 km, 10 km, 15 km (only for V_p), 20-27 km, and 35 km (only for V_s) and 38 km where the latter is associated with the Moho depth. The V_p/V_s ratio plot shows values < 1.7 for the first and third crustal layers and values between 1.75 and 1.95 for the V_p/V_s in the second crustal layer and the layers below 10 km.

Results of the inversion for station BL05 are displayed in Figure 7. The thinnest 361 crust (37 km) in this study is found beneath this station in the Georgina Basin-Davenport 362 Range region (Table S1). As can be seen, the crustal models contained in the posterior 363 distribution explain well the autocorrelations of the radial and vertical components. Over-364 all, the probability density plots of the V_p and V_s structures show that the crust can be 365 characterized with four major layers, where the velocity jumps at the top and the lower 366 crust are relatively sharp. The shear-wave structures (V_s) for both stations BL05 and 367 BL22 show a gradational transition across the Moho depth. 368

The models contained in our posterior distributions of the crustal properties generally explain well most of the observed data along the BILBY transect. However, the observed data for stations BL23 and BL24 are not well fitted by the joint inversion. We provide the results of the joint inversion including fit to the data for the other stations in the supporting information (Figures S58-S80). The evolution of the misfit throughout the MCMC joint inversion for each of the stations is also given in the supporting information (Figures S81-S105).

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Figure 6. Results of the joint inversion of field measurements for station BL22. a) fit to the autocorrelation of the vertical and radial components, b) Density plot of V_p , V_s and the V_p/V_s ratio, c) The location of station BL22 is highlighted with a red box on the map of the study area. In the density plots, the mean values of the best 2000 accepted models in the posterior distributions are displayed with blue colors. In plots showing the fits to autocorrelations of the radial and vertical components, magenta represents the observed data and the mean of the best 2000 accepted predicted data (grays) is given by blue. The black dashed lines in a) are the one standard deviation bounds obtained during the stacking of the data. The Pms, Pppmp and Ppsms phases associated with the Moho are calculated using the velocity models in b) and marked on the radial and vertical autocorrelation waveforms.



Figure 7. Results of the joint inversion of field measurements for station BL05. a) fit to the autocorrelation of the vertical and radial components, b) Density plot of V_p , V_s and the V_p/V_s ratio, c) The location of station BL05 is highlighted with a red box on the map of the study area. In the density plots, the mean values of the best 2000 accepted models in the posterior distributions are displayed with blue colors. In plots showing the fits to autocorrelations of the radial and vertical components, magenta represents the observed data and the mean of the best 2000 accepted predicted data (grays) is given by blue. The black dashed lines in a) are the one standard deviation bounds obtained during the stacking of the data. The Pms, Pppmp and Ppsms phases associated with the Moho are calculated using the velocity models in b) and marked on the radial and vertical autocorrelation waveforms.

³⁷⁶ 5 2-D Results and Discussion

377

5.1 Sections of Crustal Properties

The mean values of the posterior distributions of the V_p , V_s and V_p/V_s associated 378 with individual stations are interpolated to create sections of the crustal properties along 379 the BILBY transect (Figure 8). The transect which is used to extract these sections of 380 the crustal properties is shown by a red dashed line in Figure 4. The Moho depth esti-381 mated directly from the joint inversion is also marked with a continuous black line in Fig-382 ure 8 and compared with the previous studies (section 5.2). Our modeling results include 383 many detailed features but only the large-scale and most robust features are discussed 384 here. Before we point out some systematic variations of the crustal properties, it is use-385 ful to briefly describe the importance of having knowledge of both V_p and V_s structures. 386

Understanding of both crustal V_p and V_s (e.g., V_p/V_s ratio) is of great importance 387 to discriminate between different crustal materials. The V_p/V_s ratio in the crust usually 388 varies from ~ 1.50 for felsic rocks (e.g., quartzite) to ~ 2.12 for unconsolidated sediments 389 and mafic or ultramafic materials (e.g., serpentinite) (Christensen, 1996). Any major in-390 crease or decrease in abundance of either end-member has a dominant effect on the V_p/V_s 391 ratio of the crustal layers. However, we note that several mafic and ultramafic litholo-392 gies have relatively low V_p/V_s ratios (Christensen, 1996). In addition to the above caveats, 393 cracks with small volumes of aqueous fluids and the presence of gas/air bubbles in melt-394 bearing materials can also reduce the V_p more efficiently than V_s , thus decreasing the 395 V_p/V_s ratio (Takei, 2002; Caricchi et al., 2008; Nakajima et al., 2001; Tripoli et al., 2016) 396 The existence of fluid and/or melt, however, increase the V_p/V_s ratio (Takei, 2002; Naka-397 jima et al., 2001). 398

Several important features can be inferred from the sections of the Vp, Vs and the 399 V_p/V_s ratio, which we believe they are likely to be robust. The crustal V_p and V_s sec-400 tions show dominant low velocities at depths less than 30 km along the line whereas depths 401 greater than 30 km are mainly characterized with higher velocities. The V_p/V_s section 402 shows a very complex pattern of high and low values, where the high V_p/V_s values at depths 403 greater than 10 km are mostly localized beneath locations where the crustal blocks in-404 teract (e.g., across the transition from one domain to another). Another observation is 405 that the dominant crustal V_p/V_s ratio variation is consistent with geological units. For 406 example, the shallow crust below the Georgina and Amadeus Basins is dominated by large 407 V_p/V_s ratios. This correlation is also observed in the average V_p/V_s ratios estimated us-408 ing the inversion of receiver functions and the H-K approach (Sippl, 2016). In the mid-409 dle crust of the Nawa Domain (Officer Basin) and the Gawler Craton, a high-velocity 410 layer (which has a low V_p/V_s ratio) lies between the low-velocity layers (featuring high 411 V_p/V_s ratios). 412

5.2 Moho structure

413

Recently, several studies have discussed the crustal structure of central Australia 414 using either passive seismic data recorded by the stations along the BILBY experiment 415 (Sippl, 2016; Thompson et al., 2019), or deep seismic reflection data. Geoscience Aus-416 tralia conducted several deep seismic sounding profiles with 20 s recording time across 417 Australia (Kennett & Saygin, 2015; Kennett. et al., 2016) to better map the crustal struc-418 ture and geodynamic evolution of the major geological provinces of Australia. Korsch 419 and Doublier (2016) interpreted the major crustal boundaries and Moho across several 420 regions in Australia, including along the GOMA and 09GA-GA seismic lines in central 421 Australia (see Figure 4 for their locations). Among them, the GOMA deep seismic lines 422 approximately overlap with parts of the BILBY transect (between $31^{\circ}S$ and $25^{\circ}S$). 423

In Figure 8, we compare our Moho estimates (continues black line) with the Moho depths estimated from RFs (Sippl, 2016), the AusMoho model (Kennett. et al., 2011),

and the inversion of the autocorrelation of the vertical component (Tork Qashqai et al.,
2019). In Figure 9, we also compared our Moho estimates (dashed black lines) with those
interpreted from the deep seismic reflection (dark green line) method (Korsch & Doublier, 2016, GOMA line). The Moho values obtained from this study are converted to
their corresponding two-way travel time using our inverted P-wave velocity model (Figure 8a) to be comparable with the Moho depth interpreted from the migrated seismic
section obtained from the deep seismic reflection method.

Moho varies between 37 km and 53 km along the BILBY transect. The crust be-433 neath the Musgrave Province, Officer Basin and Arunta Block is characterized with a Moho uplift, which might be due to the presence of crustal-scale faults and shear zones 435 [e.g., Redbank shear zone in Arunta block (Sippl, 2016; Goleby et al., 1989; Lambeck & 436 Penney, 1984; Lambeck et al., 1988)]. A distinct and systematic pattern of the Moho vari-437 ation is recognizable along the transect across the crustal domains. For example, the Moho 438 is deep in the Georgina Basin and gets shallower beneath the Arunta block. It then gets 439 deeper towards the Amadeus Basin, uplifted in the Musgrave province and the Officer 440 Basin, reaches 41 km and 38 km beneath stations BL19 and BL22, respectively. The Moho 441 gets deeper below stations BL23 (53 km), BL24 (51 km) and BL25 (46 km) in the Gawler 442 Craton at the southern part of the BILBY passive seismic array. 443

Sippl (2016) reported a new Moho offset between stations BL09/BL08 and BL07, 444 and imaged a crustal thickening of about 10 km from south to north (red horizontal bars 445 in Figure 8). Sippl argued that this offset can be caused by a large-scale crustal fault that 446 has not been imaged previously. We also image such a Moho offset from station BL09 447 to stations BL08 and BL07. This is also consistent with results obtained with the vir-448 tual deep seismic sounding method (Thompson et al., 2019). Since our approach relies 449 on the 1-D Earth assumption, any existing crustal-scale fault cannot be clearly imaged 450 when creating a pseudo 2-D profile by interpolation of the 1-D structures. However, the 451 velocity models exhibit some large-scale planar structures dipping southward or north-452 ward which might be related to some deep large-scale penetrating faults/structures im-453 aged with the active seismic method (Korsch & Doublier, 2016). 454

From Figure 9, the estimated traveltimes of the seismic waves reflected back from 455 the Moho in our study are mostly less than those interpreted from the active seismic re-456 flection method which can be due to a different velocity model being used to produce 457 the migrated section. A large station interval in the passive seismic recording of the BILBY 458 profile and the higher spatial resolution of the deep seismic reflection line can be other 459 reasons for this difference. In this study, the Moho values between stations are estimated 460 by interpolation of the 1-D estimates. However, the general trend of our Moho model 461 in the time domain in the south of the Amadeus Basin follows the long-wavelength pat-462 tern of Moho interpreted from the deep seismic reflection sounding along the GOMA seis-463 mic line (Figure 9). The trend of the Moho structure from the active seismic, AusMoho 464 model, the inversion of the autocorrelation of the vertical component and from our ap-465 proach beneath the Officer Basin shows southward thinning of the crust (from stations 466 BL20 to station BL22 in Figures 8-9). In contrast, the crust thickens toward the Gawler 467 Craton in the south of station BL22 and reaches ~ 50 km below station BL23, where 468 a crustal-scale fault has been interpreted to cut and displace the Moho (Korsch & Dou-469 blier, 2016). The traveltimes of the seismic waves reflected from Moho, estimated from 470 our velocity model, are closer to the change of the reflectivity at the base of the crust 471 below most of the stations (e.g., station BL17, BL18, BL19, station BL23, and between 472 stations BL21 and BL22). 473



Figure 8. Sections of crustal properties inferred from combining the 1-D joint inversion results along the red transect shown in Figure 4. a) V_p , b) V_s , and c) V_p/V_s . The locations of seismic stations along the transect (red dashed line in Figure 4) are marked by blue triangles. The main geological units and crustal blocks are also indicated on top of the station codes. The Moho structure obtained in this study is shown by a black line. The Moho depth values estimated by the H- κ stacking method (Sippl, 2016, red thick horizontal bars) and the inversion of the vertical component autocorrelation (Tork Qashqai et al., 2019, light brown thick horizontal bars) are also plotted on the V_p and V_s sections for comparison. Moho estimates from the AusMoho model (Kennett. et al., 2011) are represented by green thick horizontal bars.



Figure 9. The Moho structure imaged by our joint inversion approach (black horizontal bars) is superimposed on the Gawler-Officer-Musgrave-Amadeus (GOMA) migrated seismic reflection section (Maher, 2010), and compared with the Moho structure interpreted from the deep seismic reflection method (Korsch & Doublier, 2016, shown by the green color). The major crustal-scale faults penetrating to the base of the crust are indicated by the brown color (Korsch & Doublier, 2016).

474 6 Conclusion

Unlike the P-to-S RFs, where main P-waves are attenuated by the deconvolving 475 vertical component (P) seismogram from the radial (S_v) component, autocorrelations 476 of the radial and vertical components retain P-waves as well as all other phases such as 477 P-to-S phases (as in RFs). Therefore, the joint inversion of the radial component auto-478 correlation with autocorrelation of the vertical component can account for the variabil-479 ity of both V_p and V_s structures. We introduce a new joint inversion approach to simul-480 taneously estimate V_p , V_s , and V_p/V_s ratio structures below a seismic station utilizing 481 both the vertical and radial component autocorrelations of the teleseismic P-wave coda. 482 We design and run a series of synthetic inversion tests to examine the feasibility of ap-483 plying this approach. Results of the tests suggest that this approach can provide a sig-484 nificant improvement in the estimation of the crustal properties compared to the inver-485 sion of either the teleseismic radial RFs or the autocorrelation of the vertical component. 486

We successfully applied this technique on passive seismic data recorded along a north-487 south transect in central Australia to image and characterize the crustal blocks and their 488 properties $(V_p, V_s, Moho and V_p/V_s ratio)$ simultaneously. This study provides the first 489 comprehensive joint estimates of all crustal properties for the BILBY seismic transect. 490 The joint inversion of the radial and vertical autocorrelations of the teleseismic P-wave 491 coda has significant implications for characterizing the V_p/V_s ratio, which is a good in-492 dicator of crustal composition. Velocity models show large-scale structures dipping north-493 ward and southward at some locations that might have caused the Moho offsets reported 494 along the line. The overall trend of the Moho structure along the profile is quite com-495 patible with the Moho depth estimated from the previous studies, including those inter-496 preted from the migrated seismic reflection section (GOMA), with the exception in the 497 north of Amadeus Basin, where inferred Moho depths shows some level of variability. The 498 trend of the traveltimes of the reflection phases associated with the Moho, estimated from 499 our velocity model, to a large extent, follows the general reflectivity character of the Moho 500 reflections at the base of the crust. This demonstrates the feasibility of our joint inver-501 sion method to provide complementary information on the crustal structure. The new 502 joint inversion approach is more cost-effective than the deep reflection profiling method 503 and can be used to obtain additional information about the deep crust, especially at depths 504 where the deep seismic reflection method has penetration problems. 505

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