

# 3-D Synthetic Modeling of Anisotropy Effects on SS Precursors: Implications for Mantle Flow in the Transition Zone

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## Abstract

The Earth's mantle transition zone (MTZ) plays a key role in the thermal and compositional interactions between the upper and lower mantle. Seismic anisotropy provides useful information about mantle deformation and dynamics across the MTZ. However, seismic anisotropy in the MTZ is difficult to obtain from surface wave or shear wave splitting measurements. Here, we investigate the sensitivity to anisotropy of a body wave method, SS precursors, through 3-D synthetic modeling. Our study shows that the SS precursors can distinguish the anisotropy originating from three depths: shallow upper mantle (80-220 km), deep upper mantle above 410-km, and MTZ (410-660 km). Synthetic resolution tests indicate that SS precursors can resolve 3% azimuthal anisotropy where data have an average signal to noise ratio (SNR=7) when azimuthal coverage is sufficient. To investigate regional sensitivity, we apply the stacking and inversion methods to two densely sampled areas: Japan subduction zone and a central Pacific region around the Hawaiian hotspot. We find evidence for a trench-perpendicular fast direction ( $\Theta=87^\circ$ ) of MTZ anisotropy in Japan, but the strength of anisotropy is poorly constrained due to limited azimuthal coverage. We attribute the azimuthal anisotropy to lattice-preferred orientation of wadsleyite induced by trench-parallel mantle flow near the stagnant slab. In the central Pacific study region, there is a non-detection of MTZ anisotropy, although modeling suggests the data coverage should allow us to resolve up to 3% anisotropy. Therefore, the Hawaiian mantle plume does not produce detectable azimuthal anisotropy in the MTZ.

1           **3-D Synthetic Modeling of Anisotropy Effects on SS Precursors:**  
2           **Implications for Mantle Flow in the Transition Zone**

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19          **Key Points:**

- 20  
21           ● We investigated the sensitivity of SS precursors to azimuthal anisotropy using  
22           3-D synthetics computed from SEPCFEM3D\_GLOBE;  
23           ● The SS precursors can resolve  $\geq 3\%$  azimuthal anisotropy with average level  
24           of noise (signal-to-noise ratio = 7);  
25           ● Japan subduction zone is dominated by trench-perpendicular fast direction in  
26           the MTZ, whereas the central Pacific anisotropy is undetectable.

27 **Abstract:**

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29 The Earth's mantle transition zone (MTZ) plays a key role in the thermal and  
30 compositional interactions between the upper and lower mantle. Seismic anisotropy  
31 provides useful information about mantle deformation and dynamics across the MTZ.  
32 However, seismic anisotropy in the MTZ is difficult to obtain from surface wave or shear  
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35 SS precursors can distinguish the anisotropy originating from three depths: shallow upper  
36 mantle (80-220 km), deep upper mantle above 410-km, and MTZ (410-660 km).  
37 Synthetic resolution tests indicate that SS precursors can resolve  $\geq 3\%$  azimuthal  
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39 coverage is sufficient. To investigate regional sensitivity, we apply the stacking and  
40 inversion methods to two densely sampled areas: Japan subduction zone and a central  
41 Pacific region around the Hawaiian hotspot. We find evidence for a trench-perpendicular  
42 fast direction ( $\Theta=87^\circ$ ) of MTZ anisotropy in Japan, but the strength of anisotropy is  
43 poorly constrained due to limited azimuthal coverage. We attribute the azimuthal  
44 anisotropy to lattice-preferred orientation of wadsleyite induced by trench-parallel mantle  
45 flow near the stagnant slab. In the central Pacific study region, there is a non-detection of  
46 MTZ anisotropy, although modeling suggests the data coverage should allow us to  
47 resolve up to 3% anisotropy. Therefore, the Hawaiian mantle plume does not produce  
48 detectable azimuthal anisotropy in the MTZ.

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50 **Key words:** Mantle Transition Zone; Seismic Anisotropy; Subduction Zone; Mantle  
51 Flow; Hotspots.

## 52 **1. Introduction**

53

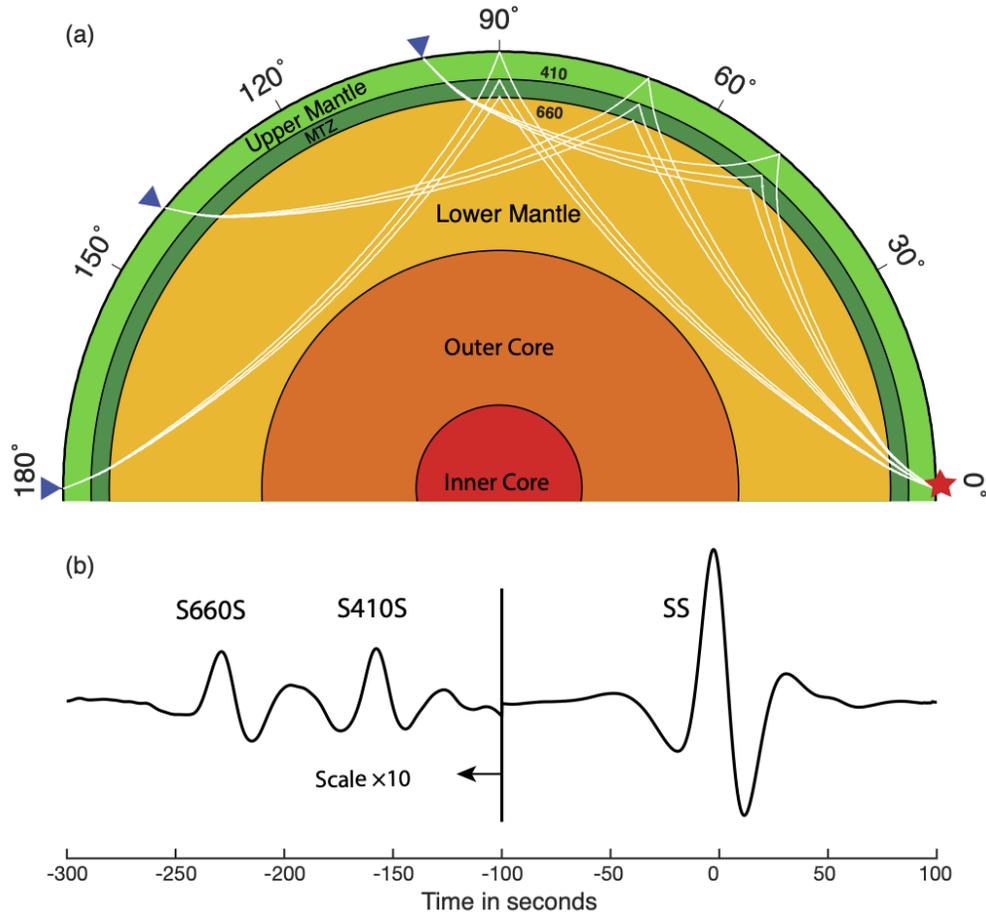
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55 The Earth's mantle convection is strongly influenced by the properties of the mantle  
56 transition zone (MTZ), a distinct layer that controls the thermal and compositional  
57 exchange between the upper and lower mantle (e.g., Bercovici & Karato, 2003; Morgan  
58 & Shearer, 1993). The boundary of the MTZ is defined by two sharp seismic  
59 discontinuities at 410-km and 660-km depths. The formation of these discontinuities is a  
60 consequence of the pressure-induced phase changes of upper mantle mineral olivine, and  
61 to a lesser extent the garnet phase transitions. Mineral physics experiments show that the  
62 phase change of olivine to wadsleyite occurs at 410 km depth, and the dissociation of  
63 ringwoodite to bridgmanite + ferropericlase occurs at 660 km depth (Ita & Stixrude,  
64 1992; Ringwood, 1975). The opposite Clapeyron slopes of the olivine phase changes (Ito  
65 & Takahashi, 1989; Katsura & Ito, 1989) make them useful for studying the mantle  
66 thermal and compositional heterogeneities (Bina & Helffrich, 1994; Helffrich, 2000;  
67 Stixrude, 1997) via mapping of MTZ topography (e.g., Flanagan & Shearer 1998).

68

69 The SS precursors are seismic body waves that manifest as shear wave reflections  
70 occurring at the underside of the 410-km and 660-km discontinuities (Figure 1). The SS  
71 precursors have served as a primary tool to investigate topography on these  
72 discontinuities at both regional (e.g., Schmerr et al., 2010; Thomas & Billen 2009; Yu et  
73 al., 2017) and global scales (e.g., Deuss & Woodhouse, 2002; Flanagan & Shearer, 1998;  
74 Gu & Dziewonski, 2002; Houser et al, 2008; Huang et al., 2019; Lawrence & Shearer,  
75 2008). The SS precursors that reflect from the 410-km and 660-km discontinuities are  
76 named as S410S and S660S respectively, or generally referred as SdS where d is the  
77 depth of discontinuity within the Earth. Several studies of the SS precursors have  
78 detected seismic anisotropy in the upper mantle and MTZ (Huang et al., 2019; Rychert et  
79 al., 2012, 2014). Seismic anisotropy, the dependence of seismic velocity on direction and  
80 polarization, is a useful tool to constrain mantle deformation and dynamics, and it is  
81 primarily produced by two key mechanisms: the lattice-preferred orientation (LPO) of  
82 intrinsically anisotropic minerals under a dislocation creep regime, or the shape-preferred  
83 orientation (SPO) of isotropic materials with distinct elastic properties (e.g., due to

84 compositional layering or lenses of melt). Here we further evaluate the sensitivity of the  
 85 SS precursors to mantle anisotropy, to demonstrate how these seismic phases can provide  
 86 insights into mantle deformation and dynamics in the MTZ.  
 87



88

**Figure 1.** (a) The ray paths of SS phase and SS precursors at the epicentral distances of 100, 140 and 180 degrees. The red star and blue triangles represent the source and receivers respectively. (b) An example of stacked waveform of SS phase and SS precursors. The amplitudes of S410S and S660S are amplified by 10 times to facilitate comparisons with the SS phase.

89

90 Observations of upper mantle anisotropy are traditionally obtained from shear wave  
 91 splitting (e.g., Long & van der Hilst, 2005; Marone & Romanowicz, 2007; Silver &  
 92 Chan, 1988), surface wave dispersion (e.g., Anderson, 1962; Montagner & Nataf, 1986;  
 93 Nettles & Dziewonski, 2008) and global tomography models (e.g., Chang et al., 2015;  
 94 French & Romanowicz, 2014; Moulik & Ekström, 2014). Upper mantle anisotropy is  
 95 typically interpreted as the LPO of olivine (e.g., Karato et al., 2008) caused by the current

96 pattern of mantle flow in the asthenosphere or the preservation of paleo-flow directions in  
97 the lithosphere (i.e., “fossil anisotropy”). At MTZ depths, evidence for seismic anisotropy  
98 is more limited, but consistently reported from multiple methods: shear wave splitting  
99 (Chen & Brudzinski, 2003; Foley & Long, 2011; Fouch & Fischer, 1996; Tong et al.,  
100 1994), surface wave measurements (Debayle et al., 2016; Trampert & van Heijst, 2002;  
101 Yuan and Beghein 2013, 2014, 2018), coupling of normal models (Beghein et al., 2008),  
102 and inversion of deep earthquake focal mechanisms (Li et al., 2018). The surface wave  
103 models that incorporate higher mode surface waves (Debayle et al., 2016; Schaeffer et  
104 al., 2016; Yuan and Beghein, 2013, 2014) suggest that ~1% azimuthal anisotropy exists  
105 in the MTZ globally, despite regional discrepancies amongst these models. Recently,  
106 Ferreira et al. (2019) discovered ubiquitous radial anisotropy in the MTZ and uppermost  
107 lower mantle in the vicinity of western Pacific subduction zones. Our previous study  
108 using SS precursors (Huang et al., 2019) also found regional evidence for 3% azimuthal  
109 anisotropy in the MTZ beneath subduction zones but detected negligible anisotropy (<  
110 1%) at a global scale.

111

112 Unlike the upper mantle, where deformation is expressed in the LPO of the mineral  
113 olivine, the MTZ may have several mechanisms for accommodating seismic anisotropy.  
114 For example, the MTZ anisotropy in subduction zones has primarily been attributed to  
115 the LPO of wadsleyite (Kawazoe et al., 2013), although the SPO of subducting slabs has  
116 also been proposed (Faccenda et al., 2019). In the upper transition zone (410-520 km),  
117 wadsleyite has up to ~14% single-crystal Vs anisotropy (Sawamoto et al., 1984;  
118 Sinogeikin et al., 1998; Zha et al., 1997), making wadsleyite the main candidate mineral  
119 for accommodating anisotropy at these depths. Below 520 km, ringwoodite is nearly  
120 isotropic with a cubic structure (Kiefer et al., 1997; Li et al., 2006; Sinogeikin et al.,  
121 2003; Weidner et al., 1984). Other minerals such as majorite garnet and clinopyroxene  
122 have either weak single-crystal anisotropy, or not enough mineral fraction abundance to  
123 accommodate the seismic observations of MTZ anisotropy (Bass and Kanzaki, 1990;  
124 Pamato et al., 2016; Sang and Bass, 2014). Slab mineralogy and layering may provide an  
125 alternative mechanism for accommodating anisotropy in the deep transition zone  
126 (Faccenda et al., 2019). Although wadsleyite can accommodate up to 14% anisotropy, it

127 must be aligned by mantle dynamics into a fabric detectable by seismic waves. Numerical  
128 simulations of strain-induced fabric of mantle mineral aggregates are therefore key to  
129 understanding the relationship between mantle flow direction (or strength) and fast  
130 direction (or strength) of seismic anisotropy in the MTZ. Previous modeling has  
131 primarily focused on the upper mantle anisotropy (e.g., Becker et al., 2006), whereas few  
132 studies explore deeper anisotropy in the MTZ and uppermost lower mantle (Faccenda,  
133 2014; Sturgeon et al., 2019). For example, Sturgeon et al. (2019) predicts that up to ~2%  
134  $V_s$  radial anisotropy may form in the MTZ beneath subduction zones. Mineral physics  
135 modeling by Tommasi et al. (2004) predicts that ~1%  $V_s$  azimuthal anisotropy can exist  
136 within the MTZ.

137

138 Although the sensitivities of SS precursors to the topography of 410-km and 660-km  
139 discontinuities have been investigated (Bai et al., 2012; Koroni & Trampert, 2016; Zhao  
140 & Chevrot, 2003), their sensitivity to azimuthal anisotropy at MTZ depths remains  
141 unexplored. Motivated by both geodynamic and mineral physics predictions, we use SS  
142 precursors to better constrain anisotropy at MTZ depths, thereby illuminating the  
143 dynamics of the upper mantle. In this study, we construct 3-D models of anisotropy and  
144 propagate synthetic seismic waves through the model to test the sensitivity of SS  
145 precursors to azimuthal anisotropy. We next compare the results of our modeling to  
146 observations in the central Pacific region and Japan subduction zone to determine the  
147 detectability and sensitivity of the SS precursory phases to MTZ anisotropy. Finally, we  
148 interpret the mantle flow pattern in the MTZ in the context of our observations.

149

## 150 **2. Methods**

151

### 152 2.1 SS Dataset

153

154 We expanded a global hand-picked SS dataset described in Huang et al. (2019) and  
155 Waszek et al. (2018) to include earthquakes in the depth range 0-75 km (previously only  
156 0-30 km) and broadband stations from 1988 to 2017. Any records with signal-to-noise  
157 ratio (SNR), which was computed from the amplitude of SS phase over the maximum

158 amplitude in a noise window (65 to 275 seconds before SS phase), lower than 2.5 were  
159 removed from the dataset. The final SS dataset consists of 58,566 seismograms, which is  
160 ~10,000 more than the previous dataset. We used the transverse component of the data to  
161 study the azimuthal anisotropy. To remove seismic noise, we filtered the data between 15  
162 and 50 s using a Butterworth band-pass filter and aligned the waveform at the peak  
163 amplitude of SS phase. Each SS seismogram was normalized to unity to equalize the SS  
164 arrivals across events and stations.

165

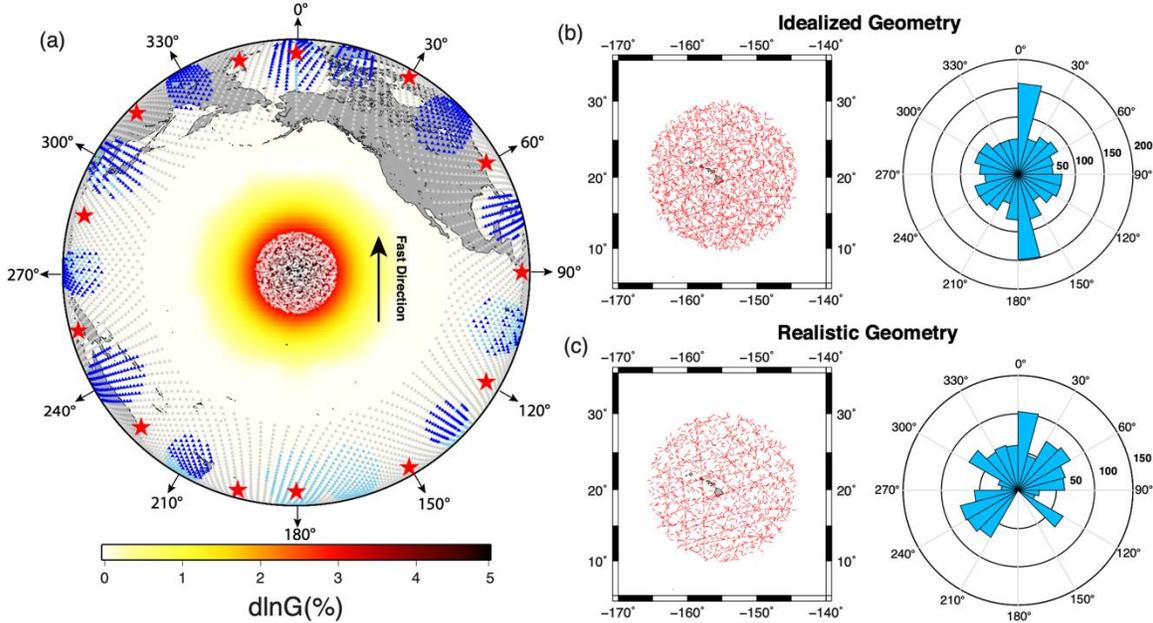
## 166 2.2 3-D Synthetics

167

168 We used the spectral element code SPECSEM3D\_GLOBE (Komatitsch & Tromp, 2002  
169 a, b) to compute 3-D synthetic SS precursor waveforms. The mesh consisted of 6  
170 domains with 320 spectral elements on each side. Therefore, the minimum period of the  
171 synthetics was 13.6 seconds. We created 13 earthquakes around the target region to  
172 provide ideal azimuthal coverage (Figure 2 & 3). The focal mechanism of each  
173 earthquake was set to maximize the SH energy in the receiver direction. We chose a  
174 dense array to guarantee ideal azimuthal coverage as well. After computing the  
175 synthetics, we generated random noise based on the realistic power spectrum of Earth's  
176 noise (Peterson 1993) and added it to the synthetics. The synthetics were then processed  
177 in the same way as the data in section 2.1.

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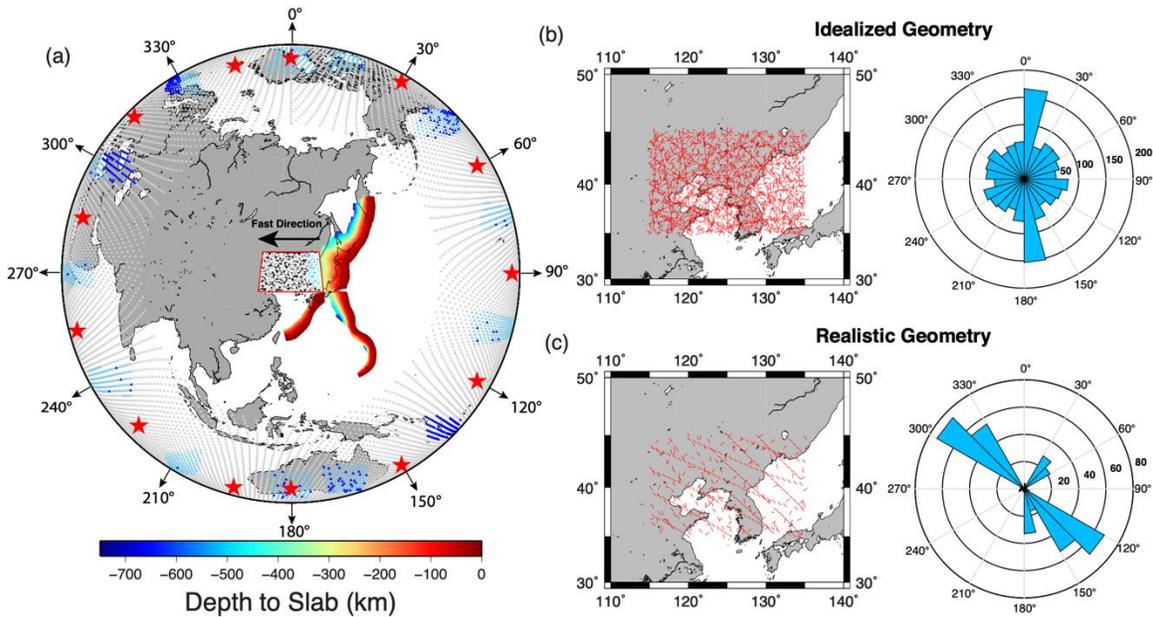
180

**Figure 2.** (a) The source-receiver geometry for the model in the central Pacific region. The red stars denote the earthquake sources. The input fast direction is due north ( $0^\circ$ ). The triangles represent all the stations: the light blue + dark blue stations are the ones used for the idealized geometry; the dark blue stations only are the ones used for the realistic geometry; the grey stations are the unused ones. The white points in the center are the SS bounce points with a bin radius of 10 degrees ( $\sim 1100$  km). (b) The SS bounce points in the idealized geometry (left) and their azimuthal distribution (right). (c) The SS bounce points in the realistic geometry (left) and the azimuthal distribution of the bounce points (right) which mimics the azimuthal coverage of data in this region.

181

182 We chose two study regions that are densely sampled by the SS phase: (1) the central  
 183 Pacific region near Hawaii (Figure 2) and (2) Japan subduction zone (Figure 3). The goal  
 184 was to simulate the azimuthal anisotropy generated by a mantle plume versus a  
 185 subducting slab. In each region, we created two types of source-receiver geometries:  
 186 idealized, and realistic geometry. The idealized geometry provided at least 100 records in  
 187 each  $15^\circ$  azimuthal bin to ensure enough data for the stacking of SS precursors. The  
 188 realistic geometry was a subset of the idealized geometry to mimic the actual azimuthal  
 189 coverage of the data that sample each region. This anisotropic structure was overprinted  
 190 on the Preliminary Reference Earth Model (PREM, Dziewonski & Anderson, 1981) at  
 191 three depth ranges: (1) the shallow upper mantle (80 – 220 km); (2) the deep upper  
 192 mantle (250 – 400 km); (3) the MTZ (400 – 670 km). The boundaries of these layers  
 193 coincide with the discontinuities in PREM such as 220-, 400- and 670-km discontinuities.  
 194 The models at each depth included three strengths of anisotropy: 1%, 3% and 5%. The

195 input fast directions were due north ( $0^\circ$ ) in the central Pacific bin and trench  
 196 perpendicular ( $270^\circ$ ) in the Japan bin. We set the radius of the central Pacific bin to 10  
 197 degrees ( $\sim 1100$  km) and the size of Japan bin was  $1500 \times 1000$  km. The choice in size of  
 198 the central Pacific structure was controlled by the standard deviation of normal  
 199 distribution, and we explored the effects of lateral size of anisotropic structures on  
 200 resolution using 5- and 2.5-degrees radius.  
 201



202

**Figure 3.** (a) The source-receiver geometry for the model in Japan subduction zone. The source, receiver and bounce point legends are the same as Figure 2. The study region is highlighted by the red box and the size is  $1500 \text{ km} \times 1000 \text{ km}$ . The strength of anisotropy is a uniform value thus represented in black color and the fast direction is trench-perpendicular ( $270^\circ$ ). The values of slab depths are from Slab 1.0 model (Hayes et al., 2012). (b) The SS bounce points in the idealized geometry (left) and their azimuthal distribution (right). (c) The SS bounce points in the realistic geometry (left) and their azimuthal distribution (right).

203

204

### 205 2.3 Stacking and Corrections

206

207 SS precursors are typically similar in amplitude to background noise or less, thus their  
 208 retrieval requires stacking when noise is present. Here, we followed the stacking methods  
 209 of Schmerr and Garnero (2006) to stack the precursors along the predicted travel-time

210 curves from PREM. We chose 125 degrees as our reference distance and applied distance  
211 exclusion windows ( $0^{\circ}$ - $100^{\circ}$  and  $135^{\circ}$ - $145^{\circ}$  for S410S;  $0^{\circ}$ - $115^{\circ}$  and  $165^{\circ}$ - $180^{\circ}$  for  
212 S660S) to avoid interferences of other seismic phases. We stacked the data by azimuth of  
213 the ray-path at the central bounce-point of SS to study the azimuthal variations of SS  
214 precursor travel-times and amplitudes. The  $2\sigma$  uncertainties of travel-time and amplitude  
215 measurements were estimated from a bootstrapping technique that implemented 300  
216 resamples, allowing replacements within each bin (Efron & Tibshirani, 1986).

217

218 We applied a series of travel-time and amplitude corrections to ensure that precursor  
219 travel-times and amplitudes were not contaminated by factors other than anisotropy, and  
220 then inverted for azimuthal anisotropy. Full details regarding travel-time and amplitude  
221 correction methods are contained in Huang et al. (2019). (1) The travel-times of SS  
222 precursors are affected by the lateral heterogeneities of crustal and upper mantle  
223 structures. We used CRUST 2.0 model (Bassin et al., 2000) for the crustal corrections  
224 and S40RTS model (Ritsema et al., 2011) for the tomography corrections. We computed  
225 travel-time residuals with respect to PREM based on 1-D ray tracing as pre-stacking  
226 travel-time corrections for each individual record. (2) We also corrected for the travel-  
227 time perturbations caused by topography of 410-km and 660-km discontinuities using the  
228 MTZ topography measurements by Huang et al. (2019). We computed the topography  
229 corrections from the differences between the local 410-km, 660-km depths and their  
230 global mean depths. (3) The amplitudes of SS precursors were corrected for attenuation,  
231 geometrical spreading, and focusing and defocusing effect using the 1-D synthetics  
232 generated by GEMINI code (Friederich and Dalkolmo, 1995). We calculated the  
233 amplitude ratios between the stacking results of data and corresponding 1-D synthetics  
234 and multiplied by the amplitudes at reference distance ( $125^{\circ}$ ) to remove these effects on  
235 amplitudes. The observed data were corrected for both travel-time and amplitude, but the  
236 3-D SPECFEM synthetics were only corrected for amplitude since no 3-D velocity  
237 structures existed in PREM.

238

## 239 2.4 Inversion for Azimuthal Anisotropy

240

241 After applying the amplitude and travel time corrections, we inverted for the strength and  
 242 fast direction of azimuthal anisotropy from SdS travel-times and amplitudes. In a  
 243 transversely isotropic medium with a horizontal symmetry axis, the velocity of vertically  
 244 propagating SH wave is expressed as the following equations (Crampin, 1984;  
 245 Montagner and Nataf, 1986):

246

$$\rho V_{qSH}^2 = L - G_c \cos 2\psi - G_s \sin 2\psi \quad (1)$$

247

$$L = \rho V_{SV}^2 = \frac{1}{2}(C_{44} + C_{55}) \quad (2)$$

248

249

250 where  $\rho$  is density,  $V_{qSH}$  is the velocity of quasi-SH wave,  $\psi$  is the azimuth of wave  
 251 propagation direction,  $L$  is a function of the isotropic SV wave velocity and can be  
 252 expressed as elastic parameters  $C_{ij}$ ,  $G_c = \frac{1}{2}(C_{55} - C_{44})$  and  $G_s = C_{54}$  are the  $2\psi$   
 253 azimuthal terms of  $L$  (Montagner et al., 2000). The strength of anisotropy ( $G$ ) and fast  
 254 direction ( $\theta$ ) are derived from the  $G_c$  and  $G_s$  parameters:

$$G = \sqrt{G_s^2 + G_c^2} \quad (3)$$

$$\theta = \frac{1}{2} \arctan \left( \frac{G_s}{G_c} \right) \quad (4)$$

255 First, we built simple 1-D anisotropy models with constant  $G_c$  and  $G_s$  values at the three  
 256 depth ranges mentioned in section 2.2. Then, we computed the predicted travel-times and  
 257 amplitudes based on equation (1) and (2). Finally, we used a grid-search method to find  
 258 the best-fitting  $G_c$  and  $G_s$  values and used equation (3) and (4) to compute the best-fitting  
 259 fast direction and strength of anisotropy. The uncertainties of the anisotropy fit are  
 260 estimated from the chi-squared statistics using p-values for 2 standard deviations.

261

### 262 **3. Results**

263 We first investigated the effect of depth, strength, and size of anisotropy on measurement  
 264 resolution using clean synthetics in the central Pacific region. Next, we examined the data

265 in the central Pacific region and Japan subduction zone from which we inverted  
266 azimuthal anisotropy and quantified the uncertainties. Finally, we added realistic noise to  
267 the 3-D synthetics for direct comparison to the data, and we also explored the effect of  
268 source-receiver geometry on the resolution.

269

### 270 3.1 The Effects of Depth, Strength and Size of Anisotropy

271

272 Unlike shear wave splitting measurements, SS precursors can distinguish the depths of  
273 anisotropy structures based on the combinations of five differential travel-time and  
274 amplitude ratio measurements: S410S-SS time, S660S-SS time, S660S-S410S time,  
275 S410S/SS and S660S/SS amplitudes. In order to understand the effect of depth, we fixed  
276 the size of anisotropy to be 10-degrees in radius and varied the depths of anisotropy in the  
277 model and performed the following synthetic tests.

278

279 Test (1): Shallow Upper Mantle, Fixed Size. The first experiment simulates anisotropy in  
280 the asthenosphere (80-220 km), which is often attributed to the LPO of olivine (Figure  
281 4a). The SS travel-times are sensitive to the asthenosphere anisotropy and their variations  
282 are mapped to S410S-SS (Figure 4b) and S660S-SS travel-times (Figure 4c) since SS is  
283 our reference phase. The S660S-S410S time (Figure 4d) and amplitudes (Figure 4e)  
284 remain constant because their ray paths do not encounter the anisotropic layer.

285

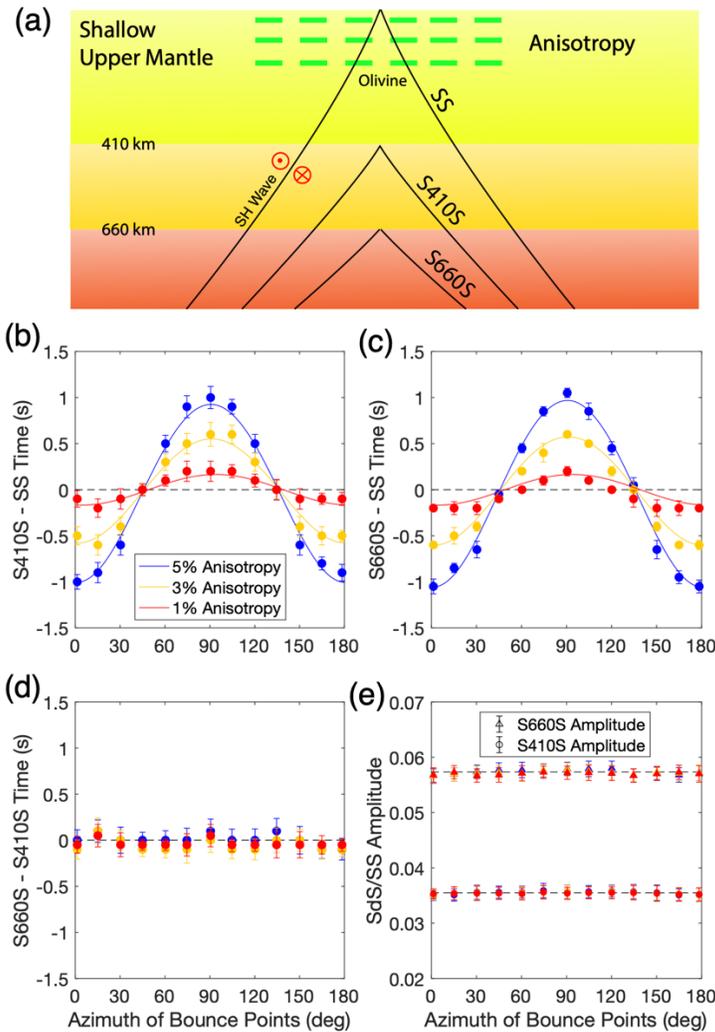
286 Test (2): Deep Upper Mantle, Fixed Size. The second case creates the scenario where  
287 anisotropy is present in the deep upper mantle (250-400 km, Figure 5a), which can still  
288 be caused by the fabric of olivine (e.g., Mondal and Long, 2020). In this model, the  
289 S410S/SS amplitude starts to vary with azimuth due to the change of reflection  
290 coefficients at 410-km (Figure 5e), whereas the S660S-S410S time (Figure 5d) and  
291 S660S/SS amplitudes (Figure 5e) remain constant.

292

293 Test (3): MTZ anisotropy, Fixed Size. The final scenario models an anisotropic layer in  
294 the MTZ where the LPO of wadsleyite and ringwoodite are formed (Figure 6a). The

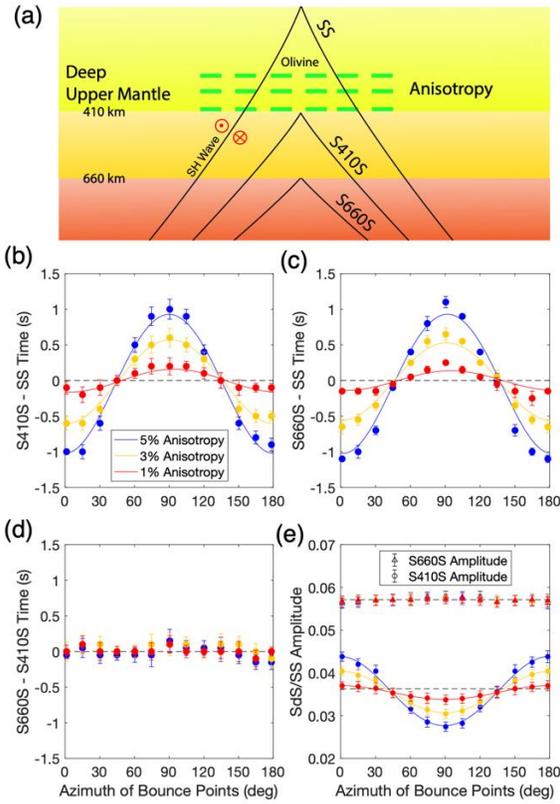
295 S660S-S410S time becomes an independent measurement for MTZ anisotropy (Figure  
 296 6d) which is not affected by the upper mantle anisotropy. This model includes anisotropy  
 297 throughout the whole MTZ so both S410S/SS and S660S/SS amplitudes (Figure 6e)  
 298 display variations with azimuth. These two amplitudes have opposite trends because the  
 299 MTZ anisotropy is below the 410-km but above the 660-km discontinuity, therefore  
 300 changing the signs of reflection coefficients.

301  
 302



**Figure 4.** (a) The anisotropy model in the shallow upper mantle (80-220 km). The green bars represent the LPO of olivine. The black curves are the ray paths of SS precursors beneath the bounce point region. The measurements of (b) S410S – SS time, (c) S660S – SS time, (d) S660S – S410S time, (e) S410S/SS and S660S/SS amplitudes from the azimuthal stacking of SPECFEM3D synthetics as a function of bounce point azimuths. The solid lines are the best-fitting models for 1%, 3% and 5% input anisotropy. The dashed lines denote the mean values of each measurement.

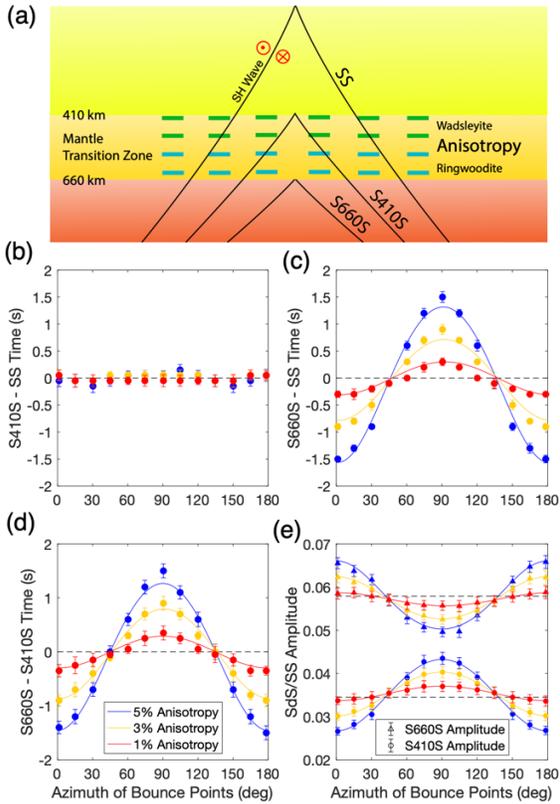
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**Figure 5.** (a) The same as Figure 4 but for anisotropy model in the deep upper mantle (250 – 400 km). The measurements of (b) S410S – SS time, (c) S660S – SS time, (d) S660S – S410S time, (e) S410S/SS and S660S/SS amplitudes from the azimuthal stacking of SPECFEM3D synthetics as a function of bounce point azimuths.

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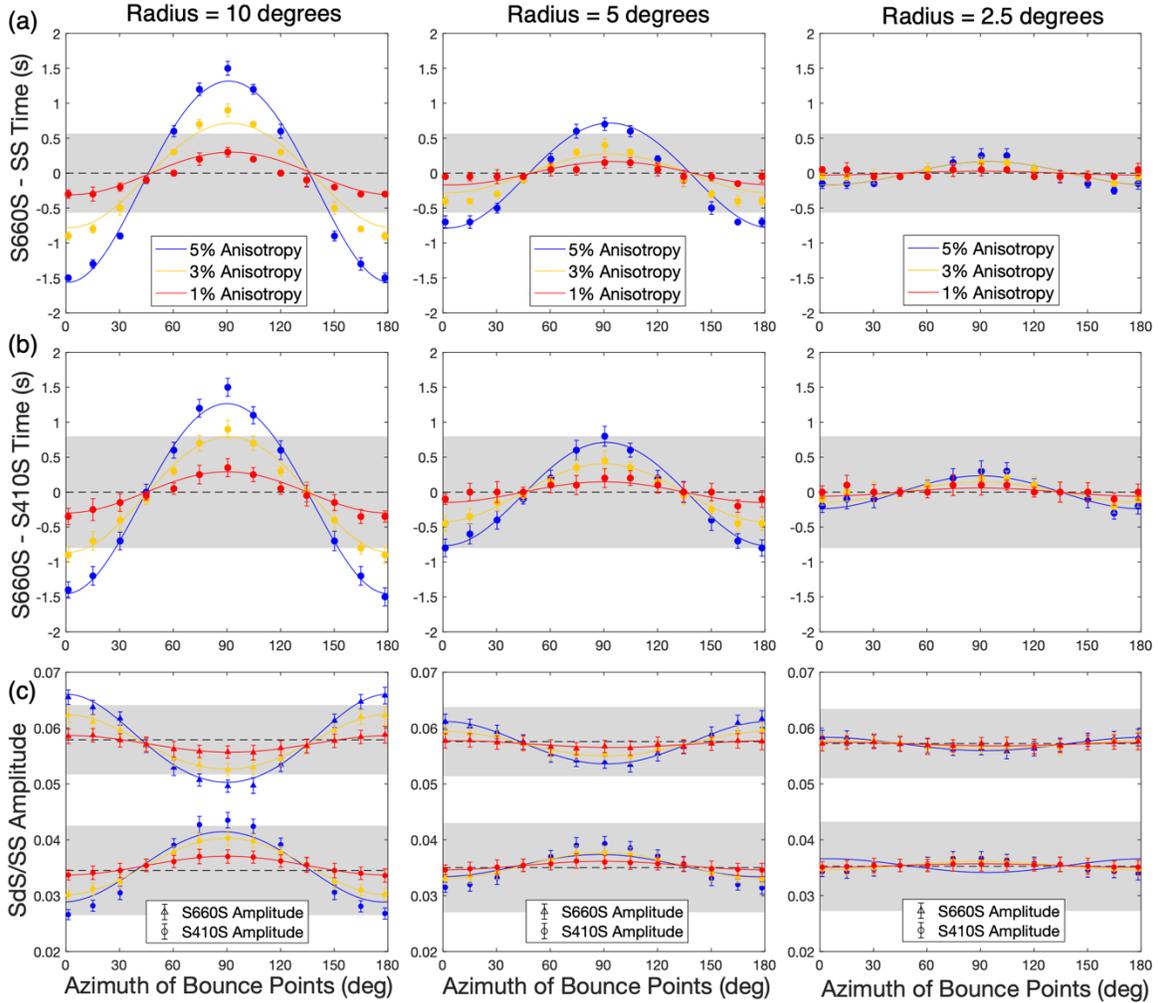
**Figure 6.** (a) The same as Figure 4 but for anisotropy model in the MTZ (400 – 670 km). The green and blue bars represent the LPO of wadsleyite and ringwoodite respectively. The measurements of (b) S410S – SS time, (c) S660S – SS time, (d) S660S – S410S time, (e) S410S/SS and S660S/SS amplitudes from the azimuthal stacking of SPECFEM3D synthetics as a function of bounce point azimuths.

307

308 Test (4): Fixed Depth in the MTZ, Varied Size: The strength and size of anisotropy  
309 control the peak-to-peak amplitudes of SdS travel-time and amplitude variations. In this  
310 test, the bin radius was fixed at 10 degrees and the anisotropy layer is in the MTZ. Then,  
311 we varied the strength and size of anisotropy and identified four measurements that were  
312 sensitive to MTZ anisotropy: S660S-SS time (Figure 7a), S660S-S410S time (Figure 7b),  
313 S410S/SS and S660S/SS amplitude (Figure 7c).

314

315 The first column of Figure 7 shows that 3-5% MTZ anisotropy with 10-degree radius can  
316 translate into 0.9-1.5 seconds travel-time variations respectively. Generally, the variations  
317 caused by anisotropy need to be greater than the corresponding uncertainties of SdS  
318 travel-times or amplitudes to become detectable. The amplitudes of uncertainties are  
319 directly related to the noise level in the data or synthetics. We estimated the average  $2\sigma$   
320 uncertainties by adding random noise (SNR=7, average noise level of our SS dataset) to  
321 synthetics before stacking (see section 2.2). The average uncertainties, which are shown  
322 as gray shaded regions in Figure 7, can be used as detection thresholds for SdS travel-  
323 times and amplitudes. When the radius is 10 degrees, Figure 7 illustrates that the S660S-  
324 SS and S660S-S410S times can both detect  $\geq 3\%$  anisotropy. The uncertainties of SdS  
325 amplitudes are generally larger in terms of percentage so anisotropy is more difficult to  
326 detect, requiring over 5% anisotropy to be detectable. Moving from left to right in Figure  
327 7, the peak-to-peak amplitudes of SdS travel-times and amplitudes both decrease as the  
328 size of the structure is reduced. When the radius is decreased to 5 degrees, the S660S-SS  
329 and S660S-S410S times can only detect  $\geq 5\%$  anisotropy, whereas the variations of SdS  
330 amplitudes are below the detection thresholds. The 2.5-degrees radius structures are too  
331 small to be detectable because all the variations become much lower than the detection  
332 thresholds.



333

**Figure 7.** The effect of anisotropy size and strength on SS precursors. The measurements of (a) S660S – SS time, (b) S660S – S410S time, (c) S410S/SS and S660S/SS amplitudes as a function of bounce point azimuths. The radius of anisotropy decreases from 10 degrees to 5 degrees and 2.5 degrees from left to right, whereas the bin radius remains 10 degrees. The depth of anisotropy is in the MTZ (400 – 670 km). The solid lines represent the best-fitting models for 1%, 3% and 5% input anisotropy. The dashed lines denote the mean values of each measurement. The gray shaded regions represent the detection thresholds for SdS travel-times and amplitudes estimated from the stacking of synthetics with average noise level (SNR=7).

334

335 **3.2 Central Pacific Data and Resolution Test**

336

337 Following the depth, strength, and size 3-D synthetic tests, we used our modeling to  
 338 study the detectability and resolution of anisotropy for an SS precursor dataset sampling  
 339 the central Pacific region. In Figure 8a, the MTZ thickness beneath the central Pacific bin  
 340 is thinner than average predominantly due to the hot thermal anomalies caused by the

341 Hawaiian hot spot (e.g., Schmerr et al., 2010). This bin has sufficient azimuthal coverage  
 342 and number of records ( $NR > 100$ ) in five azimuthal bins (Figure 8c), therefore both  
 343 S410S and S660S were observed in data and synthetics from azimuthal stacking (Figure  
 344 8b). We measured the S660S-S410S times from the azimuthally stacked data results and  
 345 inverted for azimuthal anisotropy after removing topography variations. As shown in  
 346 Figure 8d, the recovered strength of anisotropy ( $dlnG=2.9\pm2.8\%$ ) is not significantly  
 347 above zero based on the chi-squared statistics, which suggests that the central Pacific  
 348 region has very weak MTZ anisotropy.

349

350 To further test the weak anisotropy hypothesis, we added noise to the 3-D synthetics (see  
 351 Section 2.2) with different SNR values, using both idealized and realistic geometries to  
 352 explore the effect of data coverage in the central Pacific region (Figure 2). The goal was  
 353 to test the resolution of the SS precursors in the central Pacific region and determine the  
 354 minimum strength of anisotropy that would provide a detectable signal in the data. To  
 355 quantify detectability of anisotropy, we define a parameter  $\varepsilon$  as the total misfit of the  
 356 best-fitting model compared to the input anisotropy model:

357

$$\varepsilon_G = \sqrt{\left(\frac{dlnG_{out} - dlnG_{in}}{dlnG_{in}}\right)^2 + \left(\frac{2\sigma_G}{dlnG_{out}}\right)^2} \quad (5)$$

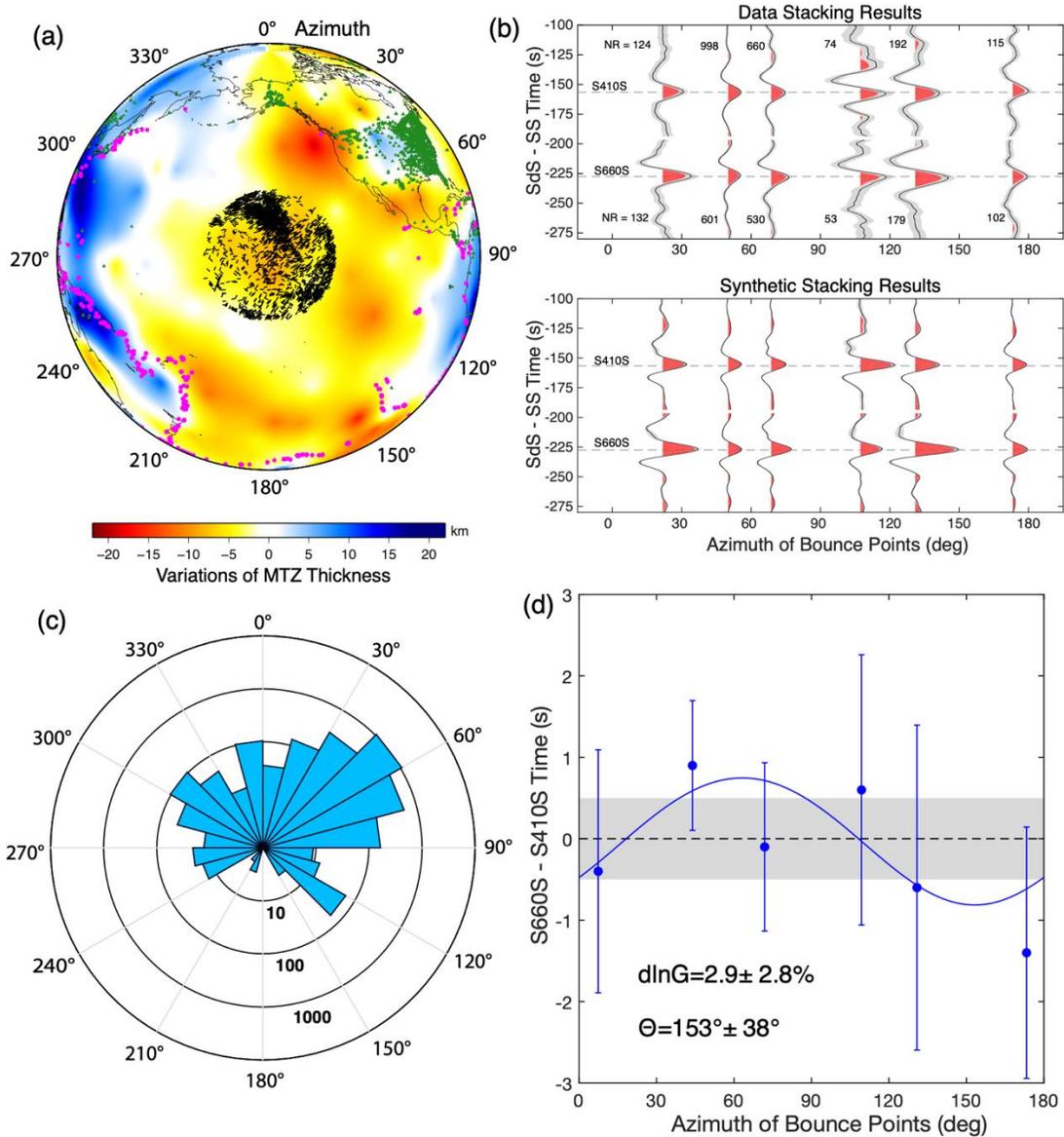
358

$$\varepsilon_\theta = \sqrt{\left(\frac{\theta_{out} - \theta_{in}}{\pi/2}\right)^2 + \left(\frac{2\sigma_\theta}{\pi/2}\right)^2} \quad (6)$$

359

360 where  $\varepsilon_G$  and  $\varepsilon_\theta$  are the total misfit for strength of anisotropy and fast direction  
 361 respectively,  $dlnG_{in}$  and  $\theta_{in}$  are the input strength of anisotropy and input fast direction  
 362 respectively,  $dlnG_{out}$  and  $\sigma_G$  are the best-fitting strength of anisotropy and  $1\sigma$  error from  
 363 inversions respectively,  $\theta_{out}$  and  $\sigma_\theta$  are the best-fitting fast direction and  $1\sigma$  error  
 364 respectively.  $\varepsilon$  can quantify the resolution as it takes into account the misfit between the  
 365 input and best-fitting anisotropy parameters, and the uncertainties of the best-fitting  
 366 model as well.  $\varepsilon$  is a positive value, and if  $\varepsilon < 1$ , we define this scenario as a detectable  
 367 case. Conversely, if  $\varepsilon \geq 1$ , we define this scenario as a non-detectable case. Since  $\varepsilon$

368 represents the misfit of the best-fitting model, the larger this value is, the lower the  
 369 resolution.

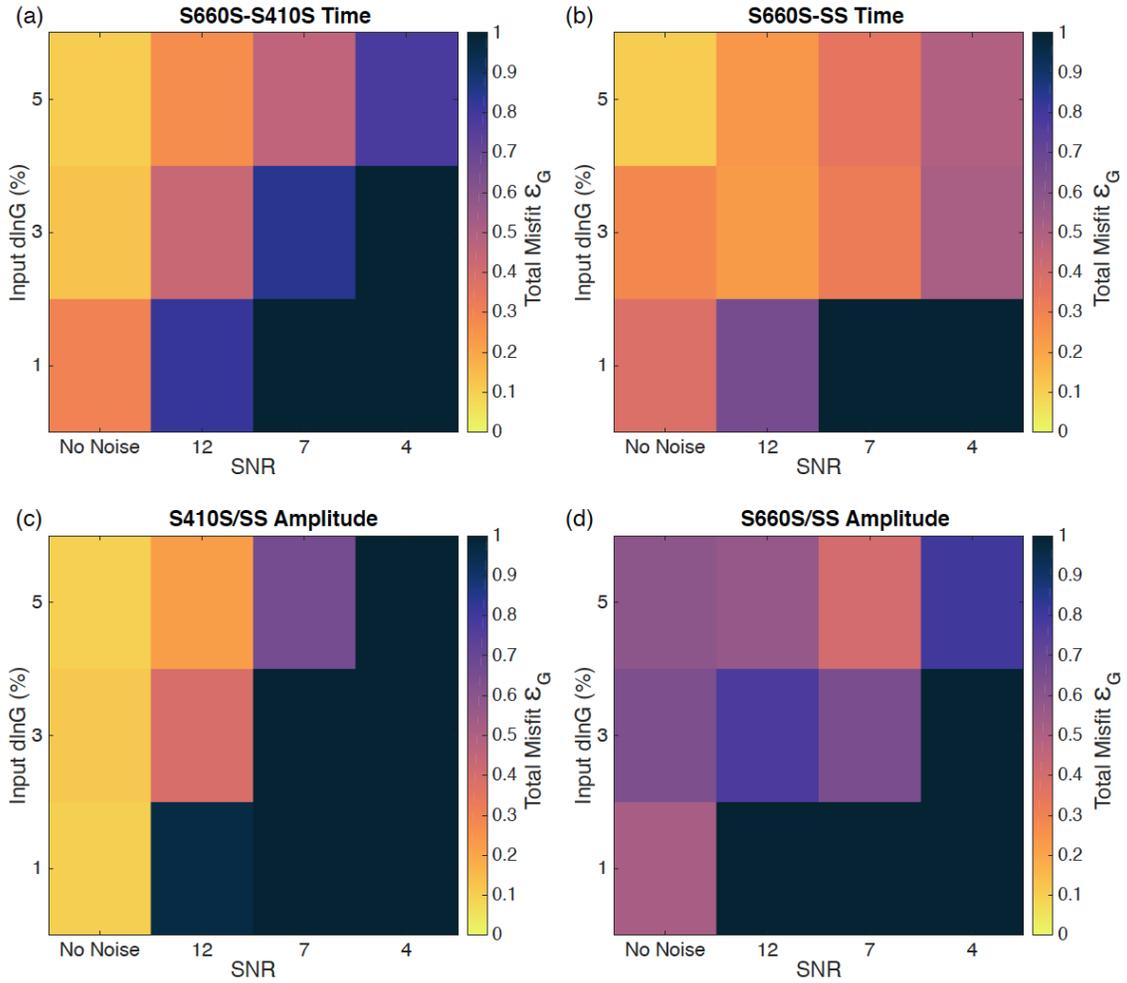


370

**Figure 8.** (a) The central Pacific bin of SS precursor data superimposed on the MTZ topography map. The MTZ thickness from Huang et al. (2019) is expressed as the variations with respect to the mean value 244.4 km. The black bars represent the SS bounce points and the azimuths of SS ray paths. The radius of the bin is 2000 km. The pink circles denote the earthquakes and the green triangles represent the stations. (b) The azimuthal stacking results of the data (top) and synthetics (bottom) in the central Pacific bin. The number of records (NR) of each azimuthal bin is labelled beside the waveform. The dashed lines highlight the average SdS travel times. (c) Rose diagram showing the azimuthal coverage of SS bounce points in log scale. (d) The S660S – S410S times and  $2\sigma$  errors shown as a function of azimuth. The solid line is the best-fitting model. The  $d\ln G$  and  $\Theta$  values are the best-fitting strength of anisotropy and fast direction respectively. The dashed line is the average S660S-S410S time from the stacking of all azimuthal bins, and gray shaded box is the corresponding  $2\sigma$  errors.

371 For the idealized geometry models, the resolutions of SdS travel times and amplitudes are  
372 shown in Figure 9. Generally, the detectability of anisotropy increases as the input  
373 anisotropy increases or the noise level decreases. Figure 9a illustrates that the S660S-  
374 S410S time can resolve 3% anisotropy with intermediate level of noise (SNR=7). Figure  
375 9b suggests that the S660S-SS time has better resolution and can resolve 3% anisotropy  
376 even with higher levels of noise (SNR=4). However, our tests with shallow anisotropy  
377 demonstrates that S660S-SS time is also potentially affected by the upper mantle  
378 structure (Figure 4c), so it is not a unique indicator for MTZ anisotropy. The test also  
379 indicates that S410S and S660S amplitudes have lowered resolutions compared to the  
380 travel-time metrics. The S410S amplitude can only resolve 5% anisotropy with  
381 intermediate level of noise (Figure 9c), and the detectability of anisotropy with the S660S  
382 amplitude is always low even when noise is absent (Figure 9d). When using a more  
383 realistic geometry model, detectability is further degraded due to the lack of stations in  
384 the southern Pacific (Figure 10). However, despite the incomplete azimuthal coverage,  
385 our tests prove that the S660S-S410S time should present a detectable travel-time  
386 anomaly where there is 3% anisotropy in the central Pacific region (Figure 10a). The  
387 conclusion is that the central Pacific data have the potential to resolve 3% or greater  
388 anisotropy but the data failed to detect an anomaly of this magnitude, and as a result, the  
389 azimuthal anisotropy in this region is likely to be smaller than 3%.

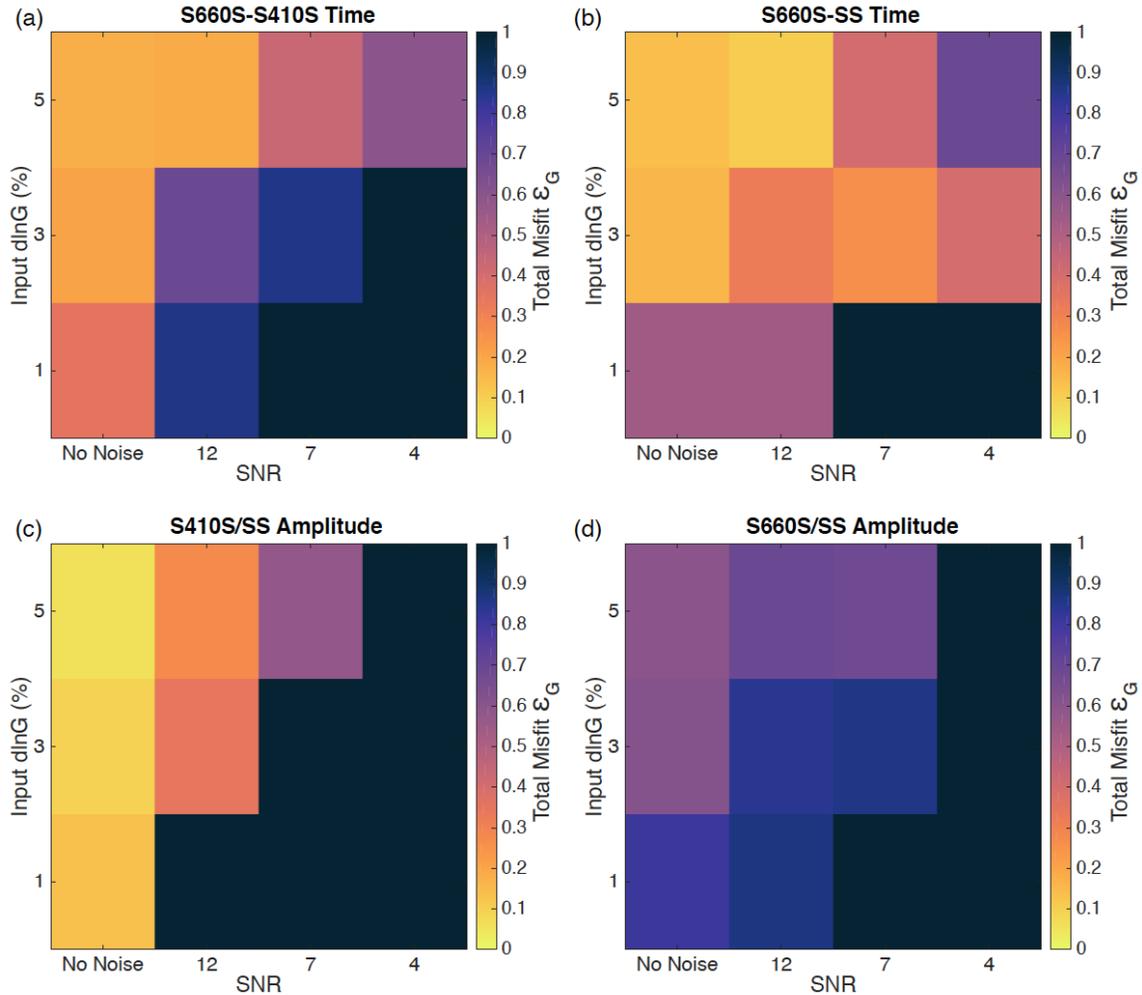
390



391

**Figure 9.** The resolution matrix for the central Pacific bin using an idealized source-receiver geometry. The dlnG resolution matrix for (a) S660S-S410S time, (b) S660S-SS time, (c) S410S/SS amplitude, and (d) S660S/SS amplitude. SNR=7 is the average noise level of our SS dataset. The total misfit  $\epsilon_G$  of dlnG is inversely correlated with the resolution. When  $\epsilon_G$  is greater than 1 (saturated in the plot), the model is considered as non-detectable.

392



393

**Figure 10.** The same as Figure 9 but using a realistic source-receiver geometry. The dlnG resolution matrix for (a) S660S-S410S time, (b) S660S-SS time, (c) S410S/SS amplitude, and (d) S660S/SS amplitude.

394

### 395 3.3 Japan Subduction Zone

396

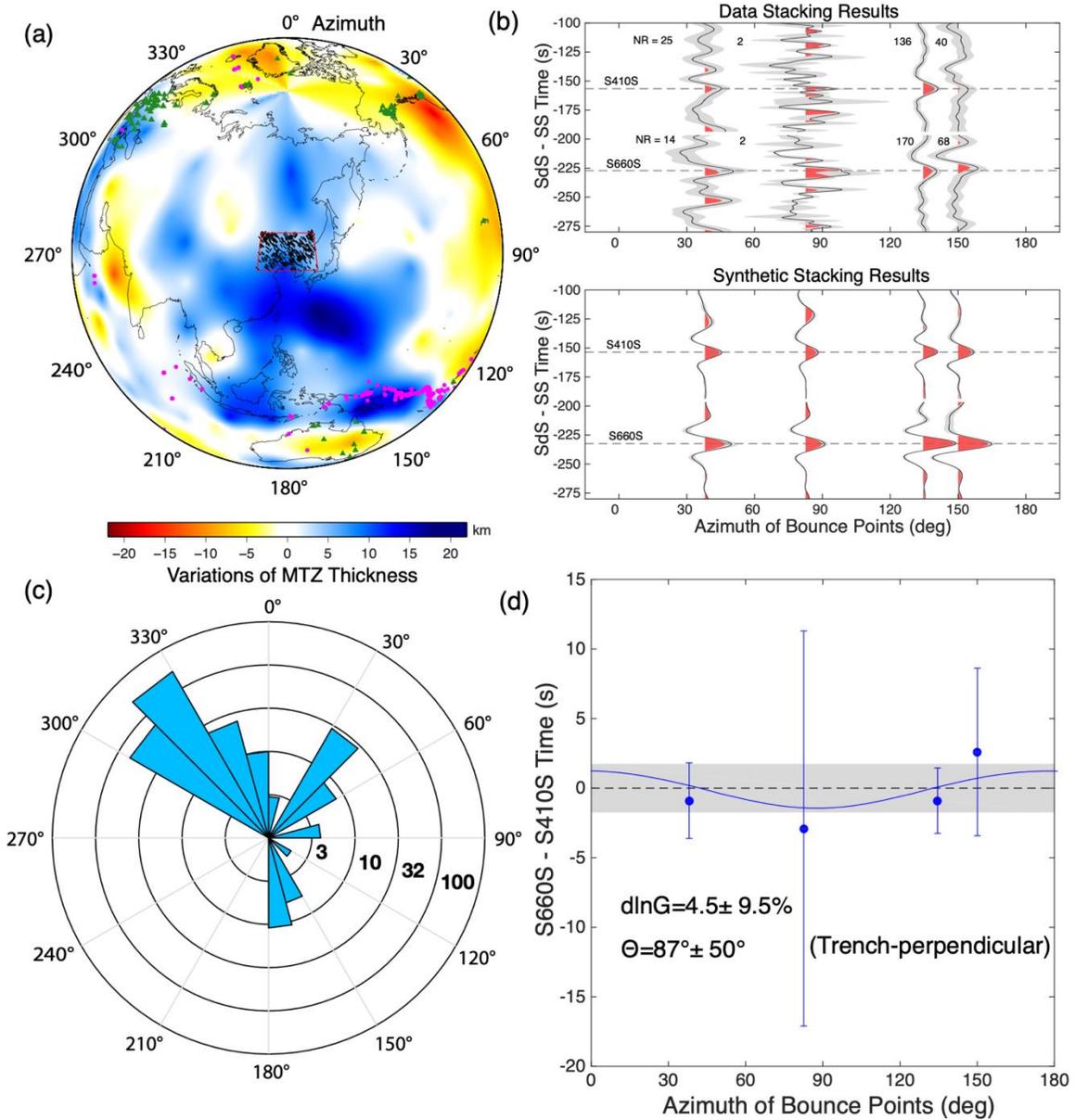
397 In Huang et al. (2019), we found evidence for ~3% MTZ anisotropy beneath the circum-  
 398 Pacific subduction zones. However, the results were based on the stacking of multiple  
 399 subduction zones, so there was ambiguity from which subduction zone originated the  
 400 signal, or if it was ubiquitous amongst all subduction zones. Due to the relatively high  
 401 density of data, we identified the Honshu subduction zone near Japan as a study region to  
 402 uniquely determine the character of MTZ anisotropy. Honshu subduction zone is located  
 403 in a region where the MTZ is thickened primarily due to the cold slab (e.g., Helffrich,

404 2000; Figure 11a). The tomography models also suggest that the Japan slab is stagnant  
405 above the 660-km discontinuity beneath eastern China and Korea (e.g., Fukao and  
406 Obayashi, 2013). We chose this region to study the mantle flow associated with the  
407 stagnant slab, and it is the best-sampled subduction zone in our SS dataset. However, the  
408 azimuthal coverage is relatively poor, with only four usable azimuthal bins for stacking  
409 and especially lacking the data with east-west orientations (Figure 11c). The data and  
410 synthetic stacking results are shown in Figure 11b. Despite limited data coverage, the  
411 S410S and S660S are recovered from stacking in all these four bins, noting that the  
412 S410S in 90 degrees bin (NR=2) is very noisy. We inverted the strength of anisotropy  
413 and fast direction from the S660S-S410S time (Figure 11d). The strength of anisotropy  
414 ( $d\ln G=4.5\pm 9.5\%$ ) is not significantly above zero due to the large uncertainties of travel-  
415 time measurement in 90° bin. The fast direction ( $\Theta=87^\circ\pm 50^\circ$ ) is trench-perpendicular and  
416 shows relatively smaller uncertainties compared to  $d\ln G$ . The fast direction is consistent  
417 with our observations in Huang et al. (2019), which suggests that the structure beneath  
418 Japan is representative of the structure found in the circum-Pacific subduction zones.

419

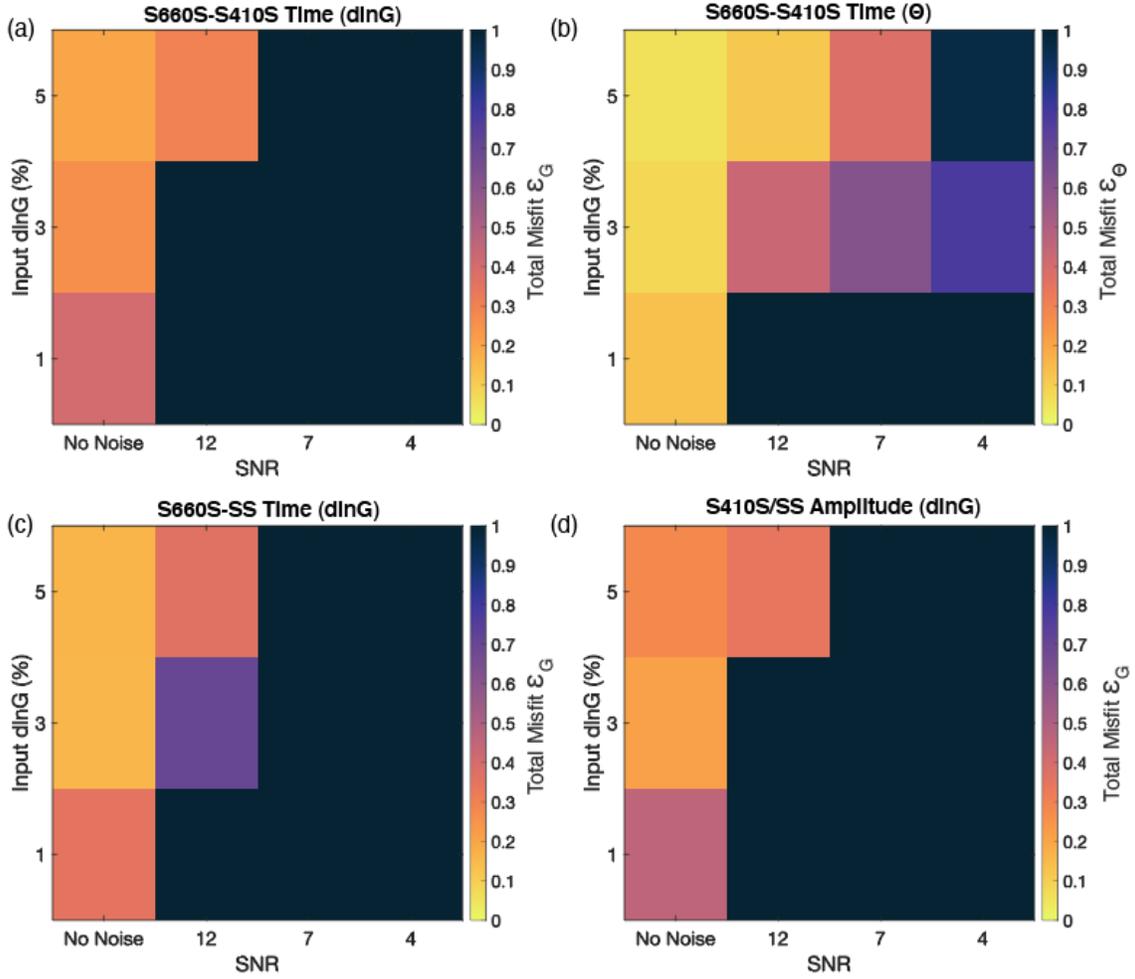
420 Following the Central Pacific study region methodology, we ran a similar resolution test  
421 for Japan subduction zone using both idealized and realistic source and receiver  
422 geometries (Figure 3). The resolution of the idealized geometry model is very similar to  
423 that of the central Pacific region although the bin size is smaller. However, the realistic  
424 geometry model using the actual azimuthal coverage of Japan displays poorer  
425 detectability in Figure 12. The SS precursors generally require tens of data points to  
426 obtain a stable stacking which is not the case here. Figure 12a illustrates that the S660S-  
427 S410S time can only resolve 5% anisotropy when noise level is very low (SNR=12)  
428 which is hardly observed in real data. Surprisingly, Figure 12b reveals that the resolution  
429 of fast directions is relatively higher, and S660S-S410S time can reconstruct the input  
430 fast direction of 3% anisotropy even when noise level is high (SNR=4). This suggests  
431 that SS precursors can still recover the input fast directions even if they cannot resolve  
432 the strength of anisotropy. This also indicates that the trench-perpendicular fast direction  
433 in Japan subduction zone is robust, thus can be used to infer mantle flow directions in the

434 MTZ. The S660S-SS time has slightly better resolution (Figure 12c), whereas the S410S  
 435 amplitude has similar resolution to S660S-S410S time (Figure 12d).  
 436



437

**Figure 11.** (a) The SS precursor data in Japan subduction zone superimposed on the MTZ topography map. The legends are the same as Figure 8. The study region is highlighted by the red box, and the size is 1500 km × 1000 km. (b) The azimuthal stacking results of the data (top) and synthetics (bottom) in Japan subduction zone. The number of records (NR) of each azimuthal bin is labelled beside the waveforms. The dashed lines highlight the average SdS travel times. (c) The rose diagram showing the azimuthal coverage of SS bounce points in log scale. (d) The S660S – S410S times and  $2\sigma$  errors shown as a function of azimuths. The best-fitting fast direction is trench-perpendicular. The dashed line is the average S660S-S410S time from the stacking of all azimuthal bins, and gray shaded box is the corresponding  $2\sigma$



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**Figure 12.** The resolution matrix for Japan subduction zone using a realistic source-receiver geometry. The (a) dlnG and (b)  $\Theta$  resolution matrix for S660S-S410S time. The dlnG resolution matrix for (c) S660S-SS time and (d) S410S/SS amplitude.

## 451 **4. Discussion**

452

### 453 4.1 Proof of concept: the resolution and limitation of SS precursors

454

455 Currently, the two major tools to constrain deep mantle anisotropy are shear wave  
456 splitting and higher mode surface waves. However, the shear wave splitting method has  
457 limited vertical resolution to distinguish the origin depths of the anisotropy structures,  
458 and surface waves have low horizontal resolution (~6500 km in the MTZ from Visser et  
459 al, 2008, Yuan and Beghein, 2018) and cannot detect small-scale anisotropy such as the  
460 structures near subduction zones. We have demonstrated that stacks of the SS precursors  
461 have sensitivity to azimuthal anisotropy in the upper mantle and MTZ through 3-D  
462 synthetic modeling. The travel-times of SS precursors can resolve  $\geq 3\%$  azimuthal  
463 anisotropy in the MTZ with intermediate level of noise (SNR=7). The amplitudes of SS  
464 precursors shed light on the anisotropy change across a seismic discontinuity such as  
465 410-km discontinuity (Saki et al., 2018). Due to the effect of stacking, the uncertainties of  
466 SdS amplitudes are often larger than travel-time measurements so they can only detect  
467  $\geq 5\%$  azimuthal anisotropy in the MTZ. However, we can apply this method to a  
468 shallower upper mantle discontinuity with stronger anisotropy such as the lithosphere-  
469 asthenosphere boundary (LAB) or mid-lithosphere discontinuity (MLD) where the  
470 polarity change of the amplitudes can take place (e.g., Rychert et al., 2014; Wirth and  
471 Long, 2014).

472

473 Sufficient azimuthal coverage is key to successfully applying SS precursors to anisotropy  
474 studies. This method requires at least 4 to 5 different azimuths with NR>100 in each  
475 azimuthal bin to obtain a robust estimate of strength and fast direction of anisotropy.  
476 However, the lack of azimuthal coverage is common in our SS dataset, and we only  
477 identify four suitable candidate locations: (1) the northwestern Pacific, (2) the central  
478 Pacific, (3) the central Atlantic, and (4) Greenland (Huang et al., 2019). This is primarily  
479 due to the uneven distributions of large earthquakes concentrated near plate boundaries  
480 and dense stations mostly in North America (e.g., USArray). This means that although

481 the western Pacific subduction zones have large numbers of records sampling the region,  
482 azimuthal coverage is actually quite limited. The data could be augmented by future  
483 ocean bottom seismometers (OBS) deployed across the Pacific Ocean (e.g., Kawakatsu et  
484 al., 2009).

485

486 A second challenge presented by this approach is the determination of the depth and  
487 thickness of the anisotropic layer. Our tests show that the SS precursors cannot resolve  
488 multiple sub-layers of anisotropy structures in the target depth range. For example, the SS  
489 precursors cannot distinguish whether MTZ anisotropy is located in the upper or lower  
490 MTZ, or the whole MTZ. Thus, in our modeling and data analysis, we only assume  
491 uniform anisotropy across the whole MTZ, which may underestimate the strength of  
492 anisotropy if it is only localized in a sub-layer. The final challenge is that we focus only  
493 on the SH waves to constrain azimuthal anisotropy, but SV waves can also provide useful  
494 information about anisotropy, via the splitting of the SS phase and its precursors (e.g.,  
495 Wolfe & Silver, 1998). Despite these limitations, SS precursors can serve as a new  
496 method to constrain seismic anisotropy in the upper and mid-mantle, especially beneath  
497 oceanic regions where seismic stations are underpopulated for shear wave splitting  
498 measurements.

499

## 500 4.2 Interpretation of mantle flow in the transition zone

501

502 The central Pacific region reveals a non-detection of MTZ anisotropy near the Hawaiian  
503 hot spot. Few shear wave splitting studies have reported evidence for MTZ anisotropy in  
504 this region, either due to the interference of strong lithosphere and asthenosphere  
505 anisotropy (e.g. Collins et al., 2012) or simply lack of data. Therefore, we compare our  
506 results to three higher mode surface wave models at 500-km depth: YB13SVani (Yuan  
507 and Beghein, 2013), SL2016SvA (Schaeffer et al., 2016), and 3D2017\_09Sv (Debayle et  
508 al., 2016). These three models only show 0.5–1.0% azimuthal anisotropy near Hawaiian  
509 hot spot. Ideally, the SS precursors can only detect 1% azimuthal anisotropy with very  
510 clean data (SNR > 12, Figure 9). In this case, the strength of MTZ anisotropy beneath  
511 Hawaiian hot spot is likely below our resolution. This suggests the mantle flow

512 associated with the Hawaiian plume does not produce significant MTZ azimuthal  
513 anisotropy (Figure 13).

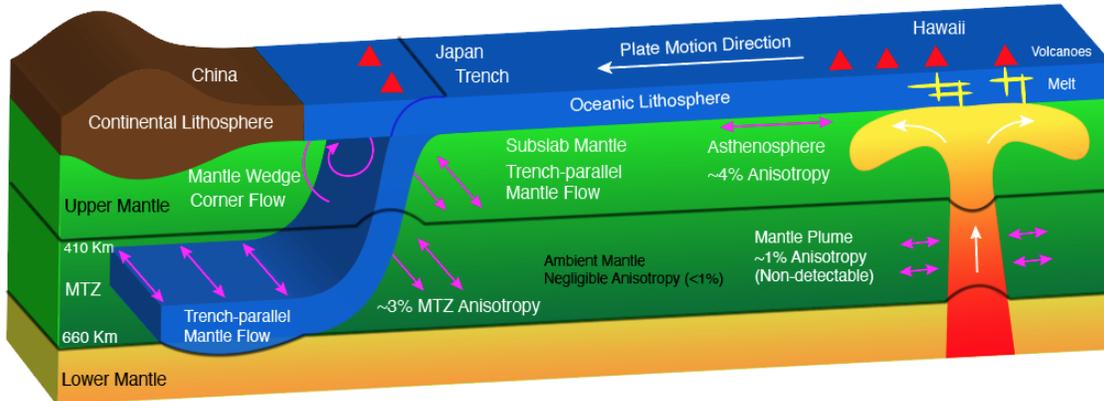
514

515 We find that the fast direction of MTZ anisotropy is trench-perpendicular (parallel to the  
516 Pacific plate motion direction) beneath the Japan subduction zone. The strength of  
517 anisotropy is less constrained compared to the fast direction because inversion of fast  
518 direction only requires azimuthal coverage in the fast and slow directions. The mantle  
519 flow direction can be inferred from the fast direction of azimuthal anisotropy. Our results  
520 are consistent with previous observations using source-side shear wave splitting  
521 measurements of deep earthquakes (e.g., Lynner and Long, 2015; Nowacki et al., 2015).  
522 Lynner and Long (2015) found evidence for azimuthal anisotropy originating from the  
523 MTZ and uppermost lower mantle beneath Japan, and the fast splitting direction is also  
524 trench-perpendicular. The higher mode surface wave model (Yuan and Beghein, 2018)  
525 shows trench-perpendicular fast direction as well at 450 km depth beneath Japan. The  
526 consistency among these three different methods combined with our 3-D synthetic  
527 modeling suggest that trench-perpendicular is a definitive feature of the MTZ beneath the  
528 Japan subduction zone. However, the fast splitting directions in the MTZ display a  
529 heterogeneous pattern across different subduction zones around the Pacific Ocean.  
530 Beneath Tonga subduction zone, the dominant fast direction is trench-parallel (Foley and  
531 Long, 2011; Mohiuddin et al., 2015); whereas the Sumatra and South America  
532 subduction zones show both trench-parallel and trench-perpendicular fast directions  
533 (Nowacki et al, 2015; Di Leo et al., 2012; Lynner and Long, 2015). Our previous study  
534 (Huang et al., 2019) combined all the data in the circum-Pacific subduction zones for  
535 stacking and concluded that trench-perpendicular fast direction is the most coherent  
536 pattern, however our data could be biased by Japan and South America where the data  
537 coverage is densest. The variability of fast directions also suggests heterogeneous flow  
538 patterns associated with different subducting slabs (penetrating or stagnant) in the mid-  
539 mantle.

540

541 The observed trench-perpendicular fast direction of MTZ anisotropy could be attributed  
542 either to the SPO of subducting slab, or the LPO of wadsleyite. Geodynamic modeling

543 suggests that the strain induced LPO under a dislocation creep regime can better fit the  
544 global tomography model than SPO in the mid-mantle (Ferreira et al., 2019; Sturgeon et  
545 al., 2019). Therefore, we assume that LPO of wadsleyite is the primarily contributor to  
546 the MTZ anisotropy beneath the Japan subduction zone. This interpretation of mantle  
547 flow direction from the fast direction of mantle anisotropy requires the experimental data  
548 of wadsleyite's dominant slip systems (Sharp et al., 1994; Thurel et al., 2003; Demouchy  
549 et al, 2011; Kawazoe et al, 2013; Ohuchi et al., 2014). Kawazoe et al. (2013) reported  
550 that the dominant slip system of wadsleyite is [001] (010), therefore the mantle flow  
551 direction is perpendicular to the fast polarization direction of shear wave, which is the  
552 same as B-type olivine (e.g., Jung and Karato, 2001). Under the assumption of a [001]  
553 (010) slip system for wadsleyite, the seismic trench-perpendicular fast direction then  
554 infers trench-parallel flow near the stagnant Japan slab in the MTZ (Figure 13). The  
555 inferred flow also agrees with the trench-parallel flow directions in the sub-slab MTZ  
556 predicted by 3-D geodynamic modeling (Faccenda, 2014). Our previous study (Huang et  
557 al., 2019) reported that trench-parallel fast direction is dominant in the upper mantle  
558 beneath the slab, which is also consistent with trench-parallel flow assuming A-type  
559 olivine fabric (e.g., Karato et al., 2008). The trench-parallel flow around slabs is  
560 consistent with the toroidal flow induced by trench migration (Faccenda and Capitanio,  
561 2012). Consequently, our results suggest the mantle flow direction remains consistently  
562 trench-parallel throughout the entire sub-slab upper mantle and transition zone in Japan  
563 (Figure 13). Alternatively, the flow direction in the transition zone can rotate by 90  
564 degrees and become trench-perpendicular if we assume the dominant slip system of  
565 wadsleyite is [100] (0kl) (Demouchy et al., 2011). Other complexities such as the water  
566 and iron content in the MTZ (e.g. Zhang et al., 2018), mineral recrystallization through  
567 phase changes (e.g. Karato, 1988), and SPO of slabs would alter our interpretations.



568

**Figure 13.** A schematic diagram depicting the mantle flow patterns in the upper mantle and transition zone observed from SS precursors. The pink arrows denote the inferred mantle flow direction from azimuthal anisotropy assuming A-type olivine in the upper mantle and [001] (010) slip system of wadsleyite in the MTZ (Kawazoe et al., 2013). On the left, the trench-parallel flow consistently exists in the sub-slab upper mantle and MTZ beneath Japan subduction zone (Huang et al., 2019). The trench-perpendicular fast direction near the stagnant slab is interpreted as trench-parallel flow in the MTZ as well. The mantle wedge is likely dominated by the trench-perpendicular corner flow which is not observed in our study. On the right, the Hawaiian mantle plume can only produce very weak azimuthal anisotropy (~1%) which is not detectable from SS precursors. The asthenosphere beneath Hawaii show strong azimuthal anisotropy (~4%) with a fast direction parallel to the plate motion direction (Huang et al., 2019). The azimuthal anisotropy is negligible (<1%) in the ambient mantle.

569

## 570 5. Conclusions

571

572 We have investigated the sensitivity of SS precursors to azimuthal anisotropy using 3-D  
 573 synthetics computed from SPECSEM3D\_GLOBE. We tested the following factors that  
 574 affect the sensitivity of SS precursors: the depth of anisotropy, size and strength of  
 575 anisotropy, source and receiver geometry, and noise level. We demonstrate that the SS  
 576 precursors can distinguish between anisotropy in the upper mantle versus the MTZ via  
 577 the azimuthal variations of SdS amplitudes and travel-times. The source and receiver  
 578 geometry and azimuthal coverage both play key roles in constraining the strength of  
 579 anisotropy and fast direction. When data coverage is sufficient (i.e.,  $NR > 100$  in each  
 580 azimuthal bin), the SS precursors can resolve  $\geq 3\%$  azimuthal anisotropy in the MTZ with  
 581 average level of noise (SNR=7). In case of biased or limited azimuthal coverage (e.g.,  
 582 Japan), the fast direction can still be inverted from S660S-S410S time even if the strength  
 583 of anisotropy is poorly constrained.

584 We searched for evidence for mantle flow associated with mantle plume or subducting  
585 slab in the MTZ beneath the central Pacific region and Japan subduction zone. In Japan  
586 subduction zone, we find evidence for trench-perpendicular fast direction ( $\Theta=87^\circ\pm 50^\circ$ ) in  
587 the MTZ, which is consistent with previous shear-wave splitting and surface wave  
588 measurements. We attribute this MTZ anisotropy to the LPO of wadsleyite such that the  
589 trench-perpendicular fast direction in Japan is interpreted as trench-parallel mantle flow  
590 based on the dominant slip system of wadsleyite. We infer that the mantle flow direction  
591 beneath Japan is likely to remain consistently trench-parallel throughout the entire sub-  
592 slab upper mantle and transition zone. In the central Pacific region, the resolution test  
593 suggests that our data can resolve 3% anisotropy, but no anisotropy is detected. The MTZ  
594 anisotropy is probably  $\ll$  3% beneath central Pacific, which is consistent with  $\sim$ 1%  
595 anisotropy from the surface wave models, and therefore below the resolution of SS  
596 precursors. This suggests that the vertical flow caused by the mantle plume beneath  
597 Hawaiian hotspot cannot produce strong azimuthal anisotropy in the MTZ.

598

599

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601

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