

Can magnetic fabrics distinguish between aseismic and seismic slip along faults?

Ran Issachar¹, Tsafrir Levi¹, and Ram Weinberger¹

¹Geological Survey of Israel

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Abstract

A long-standing problem in paleoseismic studies is to distinguish between aseismic and seismic slips along faults. We address this problem by characterizing the magnetic fabrics of rock samples around mesoscale normal faults of aseismic origin, and comparing them to those around co-seismic normal faults. Detailed sampling profiles along traverses, <0.2 m up to ~20 m from both sides of the faults, indicate symmetric and unchangeable magnetic fabrics of deformation origin that are coaxial with the regional strain field. These results are essentially different from magnetic fabrics detected around co-seismic normal faults, which show fault-related fabrics with asymmetric and changeable orientations. The analysis demonstrates two end-member cases of magnetic fabrics in association with aseismic and seismic slips along faults. We suggest that magnetic fabric analysis provides a powerful and efficient tool to characterize inelastic deformation around faults, enabling to distinguish between seismic and aseismic slip histories.



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Supporting Information for

Can magnetic fabrics distinguish between aseismic and seismic slip along faults?

R. Issachar¹, T. Levi¹ and R. Weinberger^{1,2}

¹*Geological Survey of Israel, Jerusalem, Israel*

²*Department of Geological and Environmental Sciences, Ben Gurion University of the Negev, Beer Sheva, Israel*

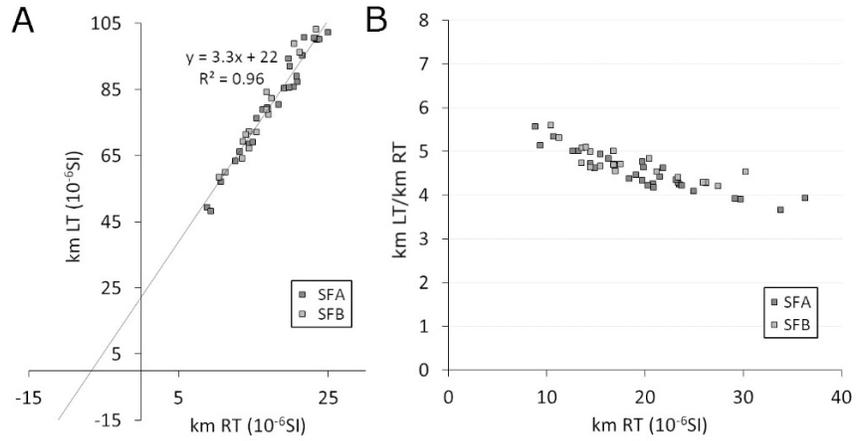


Figure S1. Room-temperature mean susceptibility (k_m^{RT}) and low-temperature mean susceptibility (k_m^{LT}) of chalk samples, indicating that paramagnetic and diamagnetic minerals are the main contributors to the rock mean susceptibility (see Issachar et al., 2018 for details). (A) k_m^{LT} versus k_m^{RT} plot showing a linear correlation with a slope of 3.3. (B) k_m^{LT}/k_m^{RT} versus k_m^{RT} plot showing an average amplification factor of 4.6 ± 0.4 .

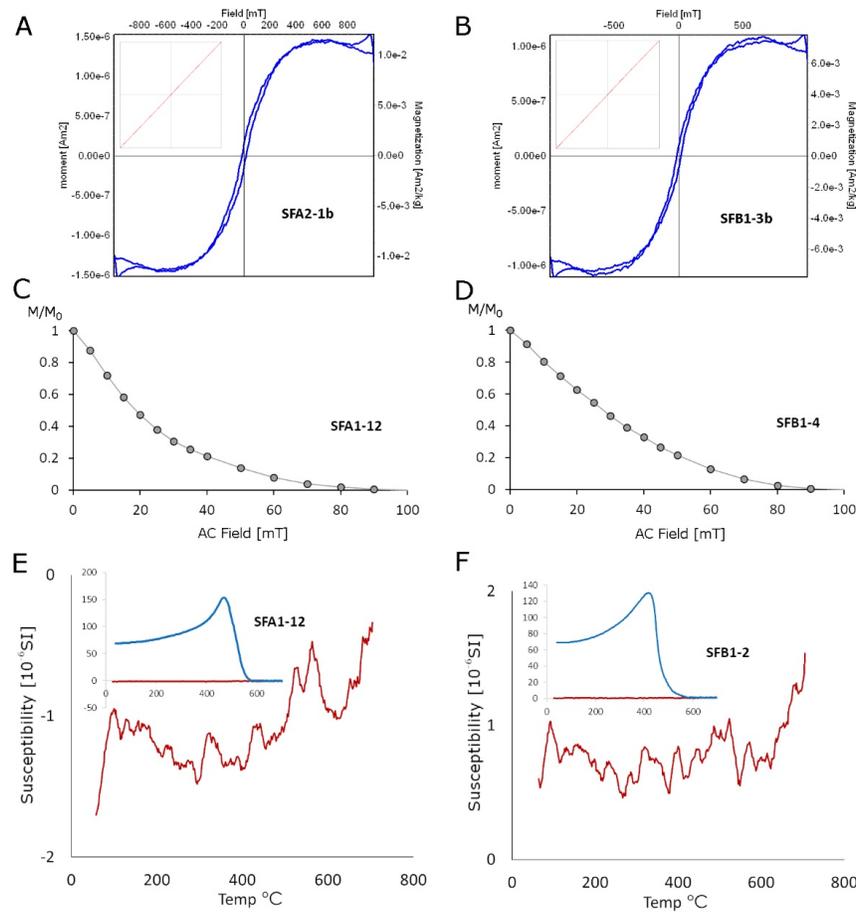


Figure S2. Magnetic mineralogy tests for representing samples. (A) and (B) Vibrating Sample Magnetometer (VSM) hysteresis loops. The raw data (red curves in the small boxes) indicates the dominance of paramagnetic minerals (positive slope) and negligible contribution of ferromagnetic minerals (linear slop). The slope-corrected data (blue curves) indicates the presence of low-coercivity ferromagnetic minerals. (C) and (D) Alternating Field demagnetization curves (AF curves). The samples were first magnetized by AC field of 100 mT and DC field of 500 ?T and then demagnetized in 5 mT steps. The curves indicate that most of the magnetic remanence (up to 90%) is lost by a demagnetization field of 60 mT, suggesting that mostly low-coercivity minerals carry the remanent magnetization. (E) and (F) Temperature-dependent susceptibility curves ($k(T)$). The susceptibility changes during heating (red curves) indicate slight variations, suggesting negligible contribution of ferromagnetic minerals. The cooling curves (blue) indicate substantial susceptibility increase at ~ 580 °C, suggesting the formation of magnetite.

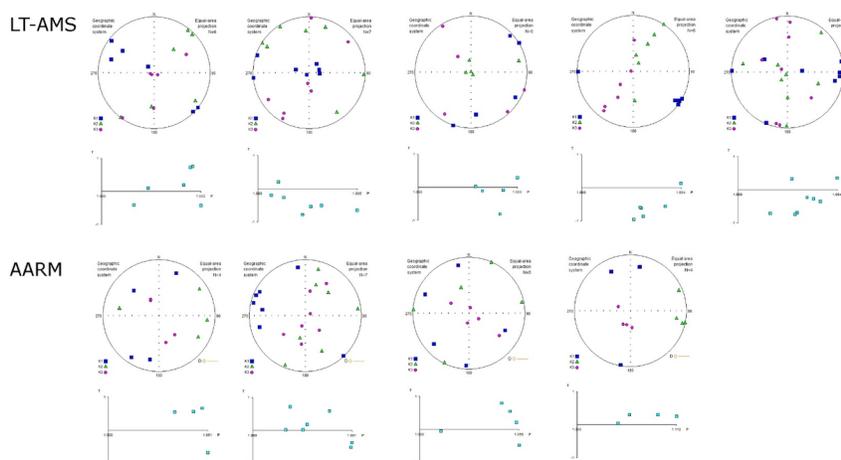


Figure S3. Fault A - LT-AMS and AARM magnetic fabrics. Stereoplots are lower-hemisphere, equal-area projection of principal axes.

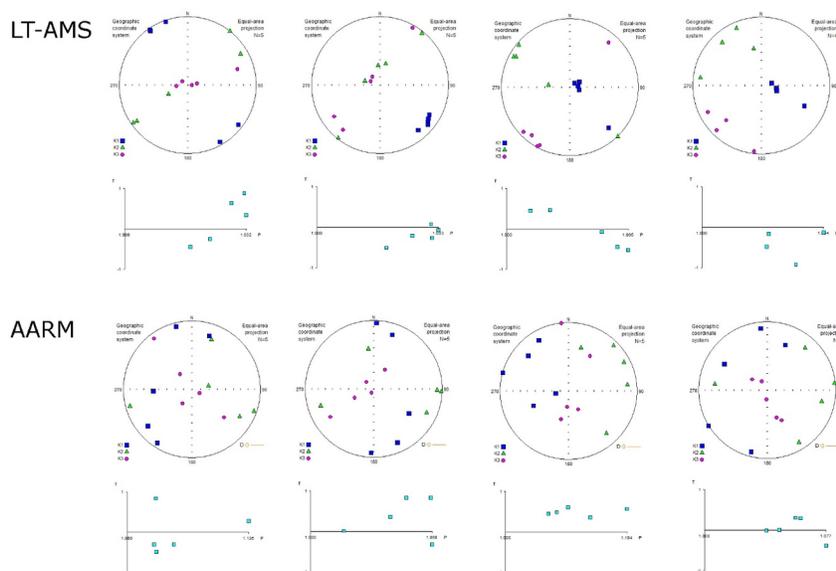


Figure S4. Fault B - LT-AMS and AARM magnetic fabrics. Stereoplots are lower-hemisphere, equal-area projection of principal axes.

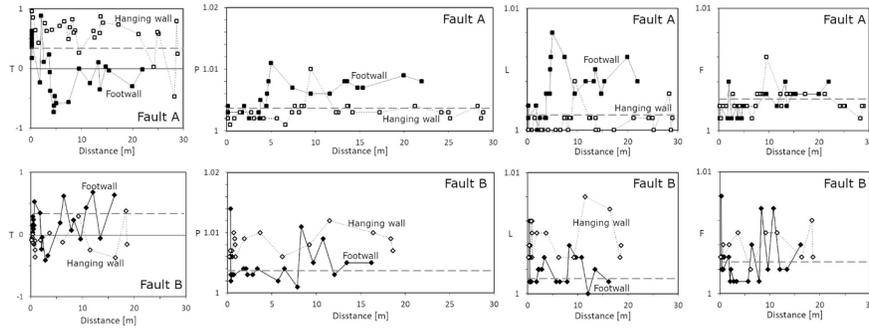


Figure S5. RT-AMS parameters shape of anisotropy (T), degree of anisotropy (P) magnetic lineation (L) and magnetic foliation (F) versus distance from footwall samples (black symbols) and hanging wall samples (empty symbols). Dashed lines represent the average values of the reference samples.

Table S1. Sampling and RT-AMS data including the distance of the sample from fault plane, mean susceptibility (k), shape of anisotropy (T), degree of anisotropy (P) magnetic lineation (L) and magnetic foliation (F). The coordinates of the faults are: Fault A (31.128362, 34.764077); Fault B (31.127435, 34.763659).

Fault	Block	Distance from fault plane (m)	k ($\times 10^{-6}$ SI)	P	T	L	F
Fault A	Hanging wall	28.9	18.1	1.003	0.243	1.008	1.005
		28.7	27.5	1.002	0.792	1.003	1.005
		28.3	16.4	1.004	-0.469	1.003	1.002
		25.2	20.2	1.002	0.581	1.005	1.005
		25.0	17.1	1.003	0.610	1.005	1.004
		24.2	13.9	1.003	0.027	1.003	1.003
		21.2	16.2	1.004	0.577	1.005	1.003
		17.3	17.7	1.003	0.734	1.006	1.003
		14.2	12.0	1.003	0.762	1.006	1.004
		13.7	9.9	1.004	0.869	1.003	1.003
		13.6	9.0	1.004	0.592	1.004	1.003
		12.6	12.6	1.003	0.515	1.003	1.003
		12.4	13.4	1.003	0.625	1.003	1.003
		9.5	6.0	1.01	0.261	1.004	1.003
		8.7	11.7	1.004	0.632	1.003	1.003
		8.5	10.8	1.004	0.599	1.003	1.003
		8.2	11.3	1.003	0.820	1.006	1.008
		7.7	12.6	1.004	0.688	1.001	1.002
		7.4	10.2	1.003	0.494	1.001	1.002
		6.6	24.4	1.001	0.786	1.001	1.002
5.7	16.0	1.002	0.711	1.001	1.002		
4.2	16.8	1.002	0.643	1.001	1.002		
3.1	15.5	1.002	0.630	1.001	1.002		

Fault	Block	Distance from fault plane (m)	km ($\times 10^{-6}$ SI)	P	T	L	F
		2.8	14.5	1.003	0.759	1.001	1.002
		1.5	19.1	1.003	0.424	1.001	1.002
		0.9	23.5	1.002	0.641	1.001	1.003
		0.4	29.8	1.001	0.852	1.002	1.001
		0.2	36.3	1.002	0.956	1.002	1.002
	Footwall	<0.2	14.6	1.003	0.504	1.002	1.001
		<0.2	33.9	1.002	0.418	1.003	1.001
		<0.2	29.2	1.002	0.627	1.001	1.001
		<0.2	20.8	1.002	0.359	1.001	1.004
		<0.2	14.9	1.004	0.171	1.001	1.001
		<0.2	15.8	1.002	0.437	1.004	1.007
		<0.2	21.5	1.002	0.355	1.003	1.002
		<0.2	18.4	1.003	0.566	1.003	1.007
		<0.2	12.6	1.003	0.415	1.000	1.002
		<0.2	8.9	1.003	0.420	1.002	1.002
		1.8	10.7	1.003	-0.233	1.001	1.004
		2.0	16.9	1.004	0.882	1.003	1.002
		2.6	26.2	1.002	0.105	1.002	1.001
		3.6	21.8	1.003	0.230	1.002	1.003
		3.7	19.9	1.002	-0.062	1.001	1.003
		3.8	20.9	1.005	-0.376	1.001	1.001
		4.5	24.9	1.004	-0.729	1.002	1.001
		4.6	23.4	1.006	-0.626	1.001	1.004
		4.7	13.2	1.008	-0.463	1.001	1.003
		5.0	9.4	1.011	-0.582	1.001	1.003
		7.4	16.9	1.007	-0.565	1.000	1.003
		9.5	23.8	1.006	-0.004	1.001	1.004
		11.7	20.4	1.006	-0.246	1.000	1.003
		13.3	21.5	1.008	0.100	1.008	1.005
		13.6	19.8	1.008	-0.351	1.003	1.005
		14.8	23.2	1.007	0.032	1.003	1.002
		15.4	21.2	1.007	-0.033	1.005	1.005
		20.0	16.3	1.009	-0.302	1.005	1.004
		22.0	19.8	1.008	-0.015	1.003	1.003
Fault B	Hanging wall	18.6	26.3	1.007	-0.157	1.004	1.003
		18.4	20.4	1.009	0.383	1.003	1.006
		16.4	27.5	1.010	-0.370	1.007	1.003
		11.5	16.8	1.012	-0.247	1.008	1.005
		9.2	16.6	1.008	0.298	1.003	1.005
		6.2	30.3	1.006	-0.124	1.003	1.002
		3.7	18.3	1.010	0.026	1.005	1.005
		1.9	17.4	1.009	-0.087	1.005	1.004
		0.9	21.3	1.006	-0.098	1.003	1.003
		0.9	18.9	1.008	-0.358	1.005	1.003
		0.8	17.5	1.009	-0.246	1.006	1.003
		0.7	14.5	1.010	-0.157	1.006	1.004
		0.5	24.6	1.006	-0.145	1.003	1.003

Fault	Block	Distance from fault plane (m)	km ($\times 10^{-6}$ SI)	P	T	L	F
		0.5	21.2	1.007	-0.123	1.004	1.003
		0.4	26.0	1.006	0.001	1.003	1.003
		0.3	23.3	1.006	-0.042	1.003	1.003
		0.3	23.4	1.007	-0.123	1.004	1.003
		0.2	25.9	1.006	-0.074	1.003	1.003
		0.2	25.7	1.006	-0.090	1.003	1.003
	Footwall	0.3	3.8	1.014	0.150	1.006	1.008
		0.4	14.5	1.003	0.034	1.001	1.002
		0.4	15.5	1.002	0.290	1.001	1.002
		0.5	13.6	1.003	0.097	1.001	1.002
		0.5	13.7	1.003	0.239	1.001	1.002
		0.5	14.5	1.003	0.084	1.001	1.002
		0.5	16.3	1.003	0.241	1.001	1.002
		0.7	17.0	1.003	0.148	1.001	1.002
		0.7	16.8	1.003	0.529	1.001	1.002
		1.8	13.6	1.004	0.349	1.001	1.003
		2.1	10.4	1.004	-0.234	1.002	1.001
		2.2	11.3	1.003	-0.043	1.002	1.002
		2.8	14.0	1.003	-0.409	1.002	1.001
		3.3	12.2	1.004	-0.338	1.003	1.001
		5.7	14.8	1.002	0.190	1.001	1.001
		6.4	4.2	1.004	0.621	1.001	1.004
		7.9	26.7	1.001	0.069	1.001	1.001
		8.3	6.0	1.011	0.234	1.004	1.007
		9.7	17.7	1.005	-0.074	1.003	1.002
		10.8	8.8	1.009	0.434	1.003	1.007
		12.0	28.6	1.003	0.683	1.000	1.002
		13.4	32.3	1.005	-0.058	1.002	1.002
		16.2	10.0	1.005	-0.157	1.001	1.004
Reference samples	Reference samples	>50	13.3	1.005	0.551	1.001	1.004
		>50	15.3	1.004	0.488	1.001	1.003
		>50	25.7	1.004	0.657	1.001	1.003
		>50	35.1	1.003	0.766	1.000	1.003
		>50	18.0	1.005	0.529	1.001	1.004

1 Can magnetic fabrics distinguish between 2 aseismic and seismic slip along faults?

3 R. Issachar¹, T. Levi¹ and R. Weinberger^{1,2}

4 ¹Geological Survey of Israel, Jerusalem, Israel

5 ²Department of Geological and Environmental Sciences, Ben Gurion University of the Negev,
6 Beer Sheva, Israel

7 **Key Points**

8 (1) Magnetic fabrics are analyzed near faults in order to distinguish between aseismic and seismic
9 slips.

10 (2) Analysis demonstrates two end-member types of magnetic fabrics - symmetric and
11 unchangeable vs. asymmetric and changeable.

12 (3) Magnetic fabrics near aseismic and seismic faults are different and enable to distinguish
13 between slip histories.

14 **Abstract**

15 A long-standing problem in paleoseismic studies is to distinguish between aseismic and seismic
16 slips along faults. We address this problem by characterizing the magnetic fabrics of rock
17 samples around mesoscale normal faults of aseismic origin, and comparing them to those around
18 co-seismic normal faults. Detailed sampling profiles along traverses, <0.2 m up to ~20 m from
19 both sides of the faults, indicate symmetric and unchangeable magnetic fabrics of deformation
20 origin that are coaxial with the regional strain field. These results are essentially different from
21 magnetic fabrics detected around co-seismic normal faults, which show fault-related fabrics with
22 asymmetric and changeable orientations. The analysis demonstrates two end-member cases of

23 magnetic fabrics in association with aseismic and seismic slips along faults. We suggest that
24 magnetic fabric analysis provides a powerful and efficient tool to characterize inelastic
25 deformation around faults, enabling to distinguish between seismic and aseismic slip histories.

26 **Plain Language Summary**

27 A long-standing problem in the study of Earth's deformation is to differentiate between faults
28 that originate during earthquake events (seismic faults) and those that are creeping due to long-
29 lasting, regional stresses (aseismic faults). We address this problem by characterizing the
30 magnetic fabrics of rock samples around faults of aseismic origin and compare them with
31 magnetic fabrics around faults of seismic origin. The magnetic fabrics mimics the distribution of
32 minerals and grains in the rocks, and, hence, are commonly served as a tool to detect "order" and
33 "disorder" in rocks. The results indicate that around the aseismic faults the magnetic fabrics are
34 uniform and reflect the large-scale regional deformation, whereas around the seismic faults the
35 magnetic fabrics are changeable and reflect the local deformation along the faults. Since all the
36 faults were formed under similar sedimentary conditions, we demonstrate two end-member cases
37 of magnetic fabrics in association with aseismic and seismic slips along faults. We suggest that
38 magnetic fabric analysis provides a powerful and efficient tool to characterize inelastic
39 deformation around faults, enabling us to distinguish between seismic and aseismic slip histories.

40 **1. Introduction**

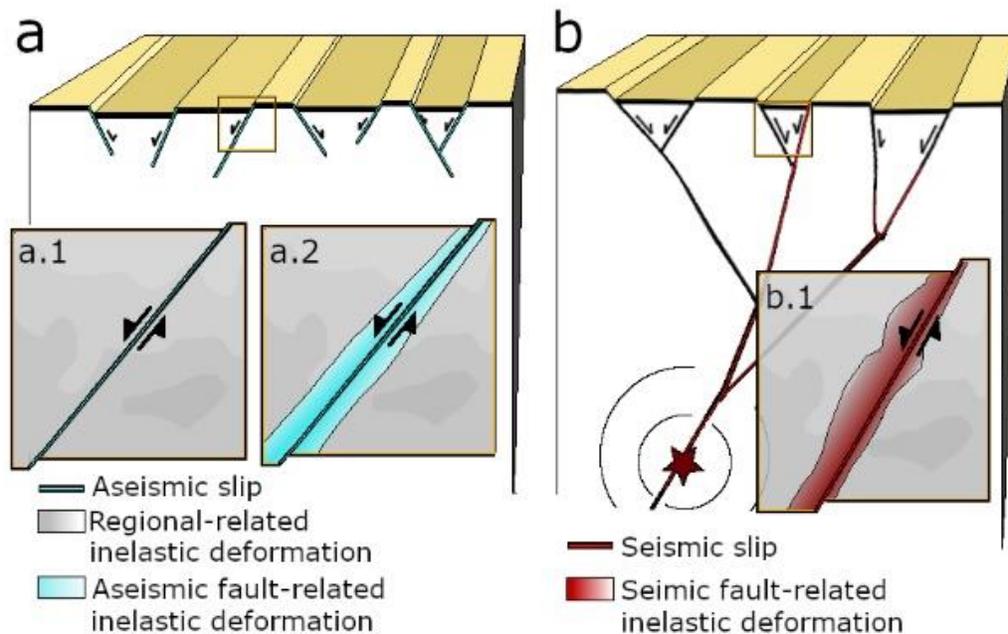
41 Faults are common geological features reflecting deep and shallow crustal processes and, as
42 such, are used to decipher different geological environments. Faults accumulate displacement by
43 slow aseismic slip (creep) at plate tectonic rates (10^{-12} to 10^{-10} m/s) or by fast seismic slip of
44 earthquake rates (10^{-4} to 10 m/s). There are several suggested criteria to recognize seismic slip
45 along faults, including pseudotachylytes that form during frictional heating, pulverized rocks that
46 form by transient stresses at the rupture tip (Rowe and Griffith, 2015), and fault-related texture
47 of hematite (Ault, 2020). Nevertheless, seismic slip does not necessarily leave distinct
48 petrological traces in the host rock, especially in cases where the rupture reaches close to the
49 surface. The characteristics of host-rock deformation around faults have the potential to
50 categorize their associated slip histories (Faulkner et al., 2010). Inelastic deformation within the
51 host rock is expected to form already before and during the early stages of fault formation and
52 growth (Anderson, 1951; Scholz, 2002). Field observations (e.g., Kim et al., 2004; Crider and
53 Peacock, 2004; Di Toro et al., 2005; Faulkner et al., 2010), models of quasi-static (e.g., Pollard
54 and Segall, 1987) and dynamic rupturing (Ben-Zion and Shi, 2005; Johri et al., 2014) show that
55 zones of inelastic deformation develop around propagating faults for both aseismic and seismic
56 slips, but in different ways (e.g., Faulkner et al., 2010). In the aseismic case, inelastic
57 deformation would develop on both sides of the fault (i.e., symmetrical deformation), and might
58 be related to the regional strain field (Figure 1a.1; e.g. Peng and Johnson, 1972). Alternatively,
59 the aseismic slip would lack a prominent fault-related deformation zone (Figure 1a.2). In the
60 seismic case, inelastic deformation would form mainly on one side of the fault (i.e.,
61 asymmetrical deformation), with relation to changeable stress field due to fault directivity and
62 strain perturbations (Figure 1b; Ben-Zion and Shi, 2005; Ma and Andrews, 2010; Johri et al.,

63 2014). The potential of distinguishing between seismic and aseismic slips by characterizing the
64 geometry and distribution of the deformation around faults was previously explored mainly by
65 the distribution of brittle deformation markers such as veins, joint sets or microcracks (Faulkner
66 et al., 2010), yet it is not always clear which marker best portrays the zone of inelastic
67 deformation.

68 In this study, we characterize the zone of inelastic deformation around faults using anisotropy of
69 magnetic susceptibility (AMS)-based magnetic fabric analysis, aiming to distinguish between
70 aseismic and seismic origins of the faults. The AMS analysis has commonly been used for
71 quantifying inelastic deformation in sedimentary rocks (Parés, 2015). The magnetic fabrics
72 reflect the preferred alignment of crystals and grain shapes within a rock sample and, as such,
73 provide a powerful tool to characterize intrinsic deformation on a grain scale (Borradaile and
74 Jackson, 2010; Weinberger et al., 2017). Various intra- and inter-crystalline deformation
75 mechanisms contribute to the AMS, such as grain boundary sliding, dislocation glide, twinning,
76 and kinking (Evans et al., 2003). These mechanisms produce irreversible deformation without
77 destruction of the lattice integrity, and thus, magnetic fabrics provide a sensitive tool to assess
78 inelastic deformation even prior to the development of macroscopic brittle deformation (e.g.,
79 Larrasoña et al., 2011). Another chief advantage of the AMS analysis for petrofabric
80 characterization is the averaging out of large numbers of grains over the whole volume of the
81 sample ($\sim 10 \text{ cm}^3$).

82 Previous works showed that magnetic fabrics in fault zones may developed during slow
83 geological processes as well as short and fast events such as earthquakes (Levi et al., 2006, 2014;
84 Casas-Sainz et al., 2018; Marcén et al., 2019; Elhanati et al., 2020). To test the applicability of
85 AMS to diagnose aseismic and seismic slips, we compare the magnetic fabrics around faults of

86 two end-member origins: (1) faults of known aseismic slip that represent slow and shallow
 87 crustal deformation (Figure 1a), and, (2) faults of proven seismic slip origin that represent
 88 earthquake-driven fast crustal deformation (Figure 1b). Levi et al. (2014) and Elhanati et al.
 89 (2020) show that magnetic fabrics near co-seismic normal faults within the seismically-active
 90 Dead Sea Fault (DSF) system indicate fault-related magnetic fabrics, explained in terms of
 91 dynamic rupturing. In this study, we characterize the magnetic fabrics near normal faults that
 92 represent shallow crustal deformation of aseismic origin. The comparison of the magnetic fabrics
 93 around these two end-member types of faults allows us to test the hypothesis that the zone of
 94 inelastic deformation around aseismic and seismic faults is essentially different and can be
 95 distinguished by AMS analysis.



96

97 **Figure 1.** Structural setting of two end member types of faults and their associated inelastic
 98 deformation: (a) Faults of aseismic origin represent shallow crustal deformation with regional strain-
 99 related inelastic deformation (a.1), or fault-related inelastic deformation distributed symmetrically on
 100 both sides of the fault (a.2). (b) Faults of seismic origin connected to a deep-seated hypocenter. The red

101 lines represent rupture propagation during a single event. Fault-related deformation is distributed
102 asymmetrically along one side of the fault.

103 **2. Magnetic Fabrics**

104 Magnetic fabrics describe the rock magnetic susceptibility tensor (\mathbf{k}) with maximum,
105 intermediate, and minimum principal susceptibility axes - \mathbf{K}_1 , \mathbf{K}_2 , and \mathbf{K}_3 , respectively, which
106 correspond to $k_1 \geq k_2 \geq k_3$ eigenvalues (Rochette et al., 1992; and reference therein). Studies of the
107 magnetic fabrics of weakly deformed sedimentary rocks indicate that the orientation of the
108 principal susceptibility axes (AMS axes) are related and coaxial to the principal strain (or stress)
109 axes. Under sedimentary and compaction processes, \mathbf{K}_3 axes align perpendicular to bedding and
110 \mathbf{K}_2 and \mathbf{K}_1 axes are scattered on a plane parallel to bedding (Type I, deposition and compaction;
111 Parés, 2015). In weakly tectonically deformed rocks, \mathbf{K}_1 axes are often oriented along the
112 intersection of bedding and the tectonic flattening plane, whereas \mathbf{K}_3 axes align perpendicular to
113 bedding (Type II, deformation). During slow progressive deformation, the magnetic fabrics often
114 show clusters of all three principal axes, and the \mathbf{K}_1 , \mathbf{K}_2 , and \mathbf{K}_3 axes are coaxial with the
115 minimum (ϵ_3), intermediate (ϵ_2) and maximum (ϵ_1) shortening axes, respectively (Type III,
116 deformation; Borradaile and Jackson, 2010).

117 We use AMS data measured at room temperature (RT-AMS) as the main method for
118 characterizing the magnetic fabrics of the chalk sample. RT-AMS was measured at a low
119 magnetic field of 450 A/m and a frequency of 976 Hz with a KLY-4S Kappabridge (AGICO
120 Inc.) at the Geological Survey of Israel. To correlate the RT-AMS with specific minerals, we use
121 several experimental techniques, including measurements of AMS at low temperatures of ~ 77 K
122 (LT-AMS), anisotropy of anhysteretic remanence magnetization (AARM), vibrating sample
123 magnetometer (VSM) hysteresis loops, alternating field demagnetization curves (AF curves) and

124 temperature-dependent ($k_{(T)}$) susceptibility measurements. Experiments were performed at the
125 Geological Survey of Israel and at the Institute for Rock Magnetism, University of Minnesota.
126 The parameters used to characterize the magnetic fabrics, including the mean susceptibility
127 ($k_m=[k_1+k_2+k_3]/3$), degree of anisotropy ($P=k_1/k_3$), shape of anisotropy ($T=\ln(F)-$
128 $\ln(L)/\ln(F)+\ln(L)$), magnetic lineation ($L=k_1/k_2$) and magnetic foliation ($F=k_2/k_3$) were
129 calculated according to Jelinek (1981) using the Anisoft 4.2 software package (www.agico.com).

130 **3. Geologic Setting and Sampling**

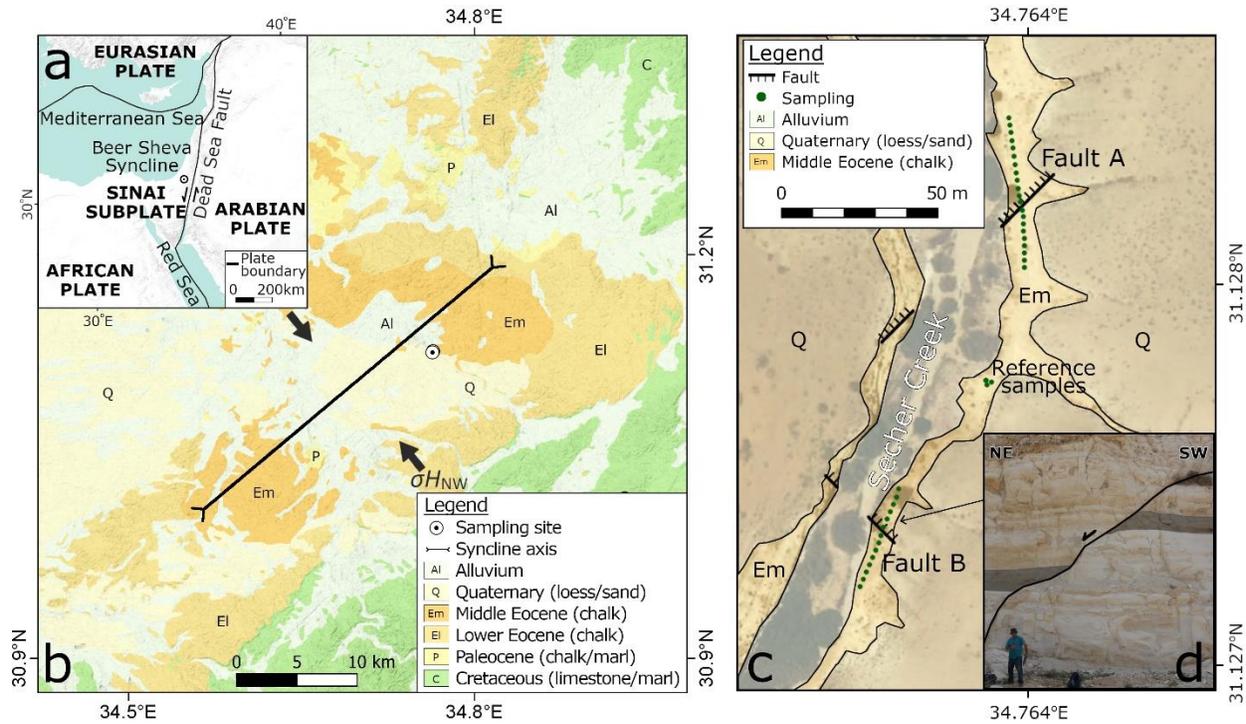
131 We study two normal faults of aseismic origin within the Beer Sheva syncline. The syncline is an
132 approximately 100 km-long, ~35 km-wide that was formed during the Eocene under long-lasting
133 compressional regional stress field of NW maximum horizontal compression, denoted $\sigma_{H_{NW}}$
134 (Figure 2b) (Bahat, 1988; Levi et al., 2019). The Beer Sheva syncline is located within the Sinai
135 subplate and is spatially remote (60 km) from the active DSF system and any other known active
136 faults (Figure 2a) (Kurzon and Wetzler, 2019). Faults and joint sets were formed during burial in
137 the poorly consolidated pelagic chalk sediments at shallow depths (Buchbinder et al., 1988) and
138 represent paleostress regimes that are associated with the formation and uplift episodes of the
139 syncline (Bahat et al., 2005).

140 The studied faults (hereafter Fault A and Fault B) are exposed along the banks of Secher Creek
141 at the Beer Sheva syncline and are located ~120 m apart (Figure 2c). The faults cut chinks of the
142 middle Eocene Maresha Formation (Figure 2d) (Buchbinder et al., 1988). The ~10 m exposed
143 surfaces of the faults are accompanied by a few mm of fine-grained gouge layers. The faults
144 trend differently with an angle of 90° between their strikes: Fault A dips NW ($52^\circ/310^\circ$; dip/dip
145 direction) and is striking subparallel to the syncline axis; Fault B dips NE ($41^\circ/040^\circ$) and is
146 striking perpendicular to the syncline axis (Figure 2c and figure 3). The throw of both faults is

147 approximately 3.5 m. The faults are considered to form under wet conditions in the early burial
148 stages of the chalks during the formation of the Beer Sheva syncline (Buchbinder et al., 1988).

149 The studied chalks of the Maresha Formation have high porosity (>20%) (Palchik & Hatzor,
150 2002) and may contain up to 25% clay fraction (Nathan and Flexer, 1977). A previous AMS and
151 rock magnetic study of this formation demonstrates that the rocks contain pure diamagnetic
152 carbonate coccolith as the rocks matrix, paramagnetic clays (palygorskite and smectite), and
153 minor quantities of low-coercivity ferromagnetic Fe oxides (Issachar et al., 2018).

154 To construct the magnetic fabric profiles, we collected standard (25x25 mm) cylindrical core
155 samples along horizontal cross-sections, <0.2 m up to ~20 m from both sides of the tested faults
156 (Figure 2c). In addition, we sampled a reference outcrop, located ~50 m away from the faults
157 (Figure 2c). By this strategy, we provide a high-resolution “continuous” view of the inelastic
158 strain field around the faults. The sampling strategy allows us to diagnose three different
159 scenarios of magnetic fabrics: (1) regional-related fabrics, that are uniformly distributed on both
160 sides of the tested faults; (2) fault-related fabrics that are symmetrically distributed on both sides
161 of the fault with similar orientations of AMS axes within each block (i.e., unchangeable
162 distribution); and (3) fault-related fabrics that are developed asymmetrically mainly on one side
163 of the fault with varying orientations of AMS axes (i.e., changeable distribution).



164

165 **Figure 2.** (a) General tectonic map of the study area. The circle marks the location of Beer Sheva
 166 syncline. (b) Geological map of the Beer Sheva syncline, showing Cretaceous to Quaternary strata (after
 167 Sneh et al., 1998), the syncline axis (after Bahat et al., 2005) and the direction of the regional stress field
 168 of $\sigma_{H_{NW}}$. (c) Detailed geological map of the study area, showing the fault traces cutting the middle
 169 Eocene strata and the distribution of the samples next to the faults (for more details see supporting
 170 information Table S1). (d) Photograph of Fault B showing an apparent fault dip and throw.

171 **4. Results and Discussion**

172 RT-AMS measurements of 111 chalk samples near the two tested faults indicate a narrow range
 173 of susceptibility values between 4 and 36 ($\times 10^{-6}$ SI) with a mean of 17 ± 7 ($\times 10^{-6}$ SI). The
 174 narrow range of the susceptibility values indicates that the samples has uniform mineralogy. The
 175 LT-AMS mean susceptibility value is 88 ± 21 ($\times 10^{-6}$ SI), indicating amplification factors of $4.6 \pm$
 176 0.4 (supporting information Figure S1), and suggests that both paramagnetic and diamagnetic
 177 minerals contribute to the rocks susceptibility (Issachar et al., 2018). VSM hysteresis loops
 178 indicate positive linear response between -1 and 1 T. Yet, slope-corrected curves suggest minor

179 presence of low-coercivity ferromagnetic minerals. AF demagnetization curves indicate that
180 most of the remanent magnetization (up to 90%) is lost by a demagnetization field up to 60 mT,
181 suggesting that mostly low-coercivity minerals carry the remanent magnetization. Temperature-
182 dependent susceptibility curves indicate minor susceptibility changes during heating, suggesting
183 minor contribution of ferromagnetic minerals to the rocks susceptibility. Hence, we conclude that
184 the rocks predominantly contain diamagnetic carbonate minerals and paramagnetic clay minerals
185 with negligible presence of ferromagnetic minerals (supporting information Figure S2), in
186 accordance with previous magnetic fabrics of chalks from the area (Issachar et al., 2018).

187 LT-AMS and AARM fabrics show insignificant anisotropy and inconsistent orientations of
188 susceptibility axes, suggesting that neither paramagnetic nor ferromagnetic minerals are the
189 carriers of the RT-AMS (supporting information Figure S3 and Figure S4). The insignificant
190 anisotropies of the LT-AMS and AARM, the negligible ferromagnetic contribution and the
191 dominance of diamagnetic minerals (i.e., calcite) suggest that the RT-AMS is controlled by the
192 orientation distribution of the diamagnetic calcite minerals. Previous studies have shown that the
193 orientation of the AMS axes of calcite are sensitive strain indicators with \mathbf{K}_3 axes parallel to the
194 compaction/shortening direction (e.g., Owens and Rutter, 1978; de Wall et al., 2000; Almqvist et
195 al., 2010; Issachar et al., 2018).

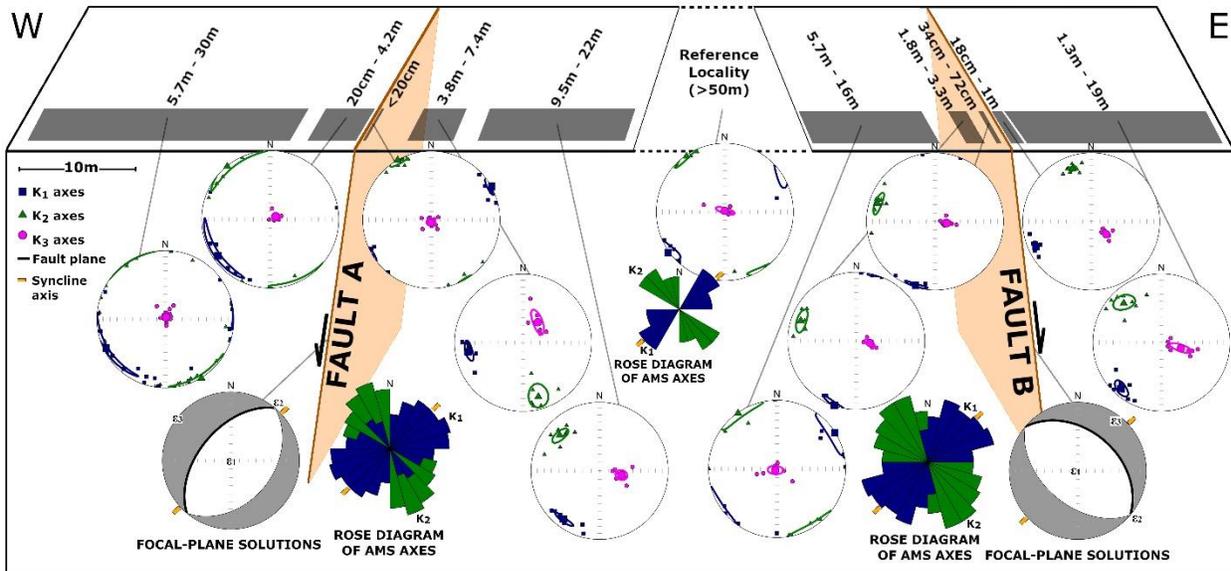
196 The orientation of the AMS axes of the studied chalk samples indicate deformation fabrics of
197 Type II, which is characterized by tight clusters of \mathbf{K}_3 axes oriented sub-perpendicular to
198 bedding, and moderate to tight clusters of \mathbf{K}_1 and \mathbf{K}_2 axes within the bedding plane (Figure 3).
199 The AMS axes of samples collected from distance of <0.2 m up to 20 m from both sides of the
200 tested faults, as well as from the reference outcrop have similar and consistent orientations with
201 NE-SW trending \mathbf{K}_1 axes and NW-SE trending \mathbf{K}_2 axes (Figure 3). The \mathbf{K}_1 and \mathbf{K}_2 axes are

202 syncline parallel and syncline perpendicular, respectively, and are compatible with the syncline-
203 driven σH_{NW} stress field. Notably, irrespective to the attitude of the faults, the AMS axes around
204 them are similar and parallel to the syncline axes. In that sense, the AMS axes indicate uniform,
205 unchangeable distribution of inelastic deformation that is associated with the regional stress
206 field. The anisotropy P and shape T have values of 1.005 ± 0.003 and of 0.2 ± 0.4 , respectively
207 (supporting information Figure S5 and Table S1). The AMS parameters show perceptible
208 differences between the blocks of the faults (supporting information Figure S5). The parameters
209 P and L have higher values in the footwall of Fault A and in the hanging wall of Fault B. The
210 shape of anisotropy is oblate ($T \sim 1$) in the hanging wall of Fault A and is neutral ($T \sim 0$) in its
211 footwall and in both blocks of Fault B. The F parameter shows no significant differences
212 between the blocks. The variations in the AMS parameters may imply that strain magnitudes
213 locally develop differently between the faulted blocks.

214 Fault observations and experiments show that inelastic deformation develops in association with
215 the regional stress field prior to host-rock faulting (e.g., Scholz, 1968; Mollema and Antonellini,
216 1999; Wilson et al., 2003; Crider and Peacock, 2004; Paterson and Wong, 2005; Blenkinsop,
217 2008). The faults of the current study are considered to form under wet conditions in the early
218 burial stages of the chalks (Buchbinder et al., 1988). The low strength of the chalks at the time of
219 the syncline formation could explain the absence of prominent fault-related deformation during
220 aseismic (i.e., creeping) slips on the faults. Hence, we suggest that the AMS axes represent the
221 regional stress (strain) field during the formation of the syncline and were already acquired prior
222 to the formation of the faults.

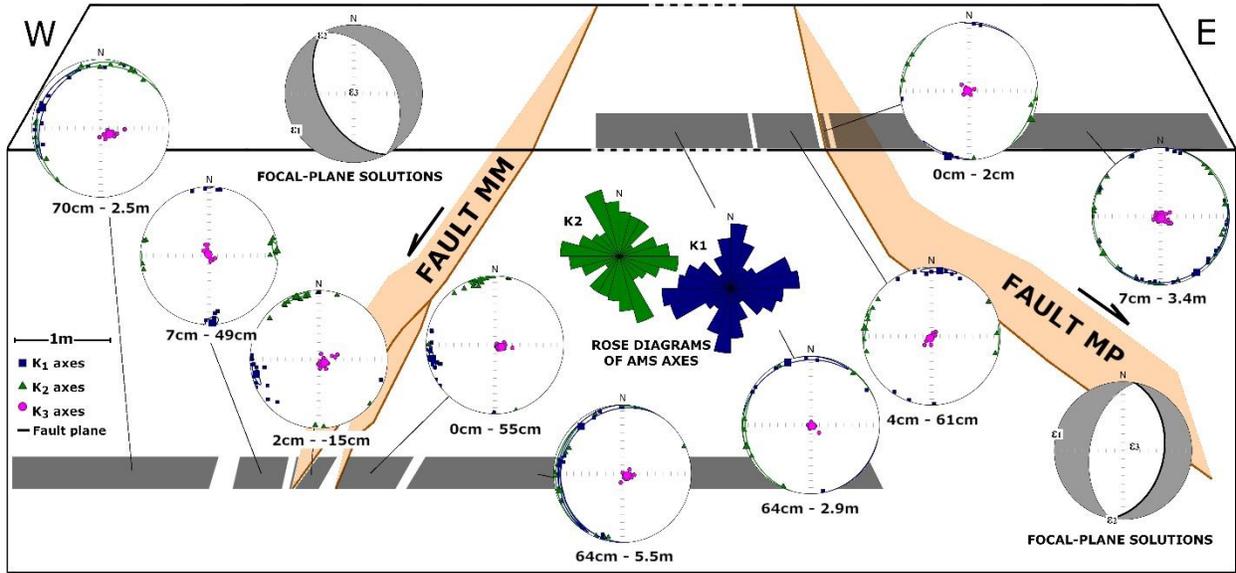
223 Contrary to the present findings, AMS studies of co-seismic faults within the active DSF system
224 infer prominent fault-related magnetic fabrics (Levi et al., 2014; Elhanati, 2019). Figure 4

225 presents the AMS results near two tested normal faults of a proven seismic origin (Marco et al.,
226 1996). The faults cut late Pleistocene soft sediments in the Masada Plain, Israel and are
227 associated with throw of ~2 m that occurred during single earthquake events (Marco and Agnon,
228 2005). The two tested faults were propagated throughout low-strength sediments, close to the
229 surface and under wet conditions (Marco et al., 1996). Several AMS characteristics are well
230 explained in terms of dynamic rupturing: (1) fault-related fabrics are asymmetrically distributed
231 between the fault blocks; (2) the orientation of AMS axes is changeable and occasionally
232 compatible with the principal strain directions defined by the fault focal-plane solutions; and (3)
233 the width of the deformation AMS fabric zone is roughly similar to the amount of displacement
234 during a single slip event. The faults in the Beer Sheva syncline and in the DSF system are
235 comparable as they have similar throw and were developed in low strength carbonate sediments
236 at shallow depth under wet conditions. The comparison of the magnetic fabrics of these two end-
237 member type of faults shows clearly that only the seismic faults form a prominent fault-related
238 magnetic fabrics (Figure 5a and Figure 5c). Nevertheless, we do not dispute the possible
239 formation of fault-related magnetic fabrics around aseismic faults that would differ from those
240 around seismic faults (Figure 5b).



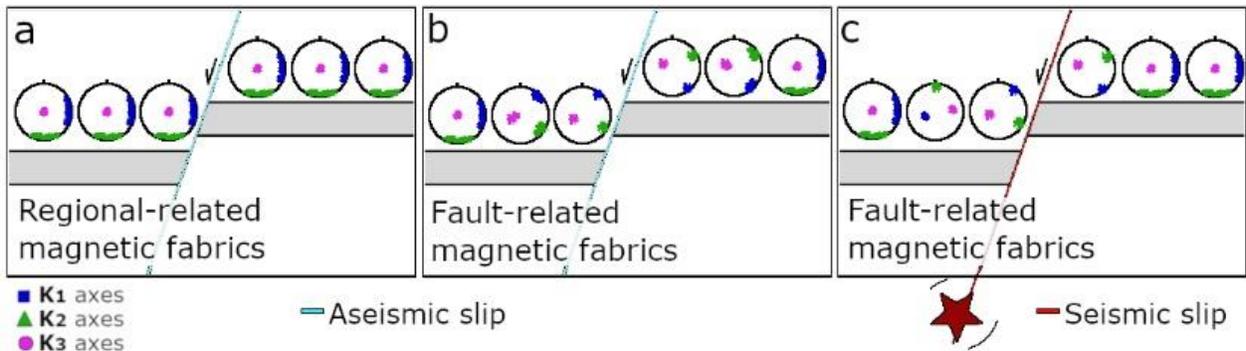
241

242 **Figure 3.** RT-AMS magnetic fabrics from sites at varied distances from aseismic faults (Fault A and Fault
 243 B) in the Beer Sheva syncline. Stereoplots are lower-hemisphere, equal-area projection of AMS principal
 244 axes, and the 95% confidence ellipses. The AMS axes show clustering of three principal axes indicating
 245 deformation origin. Rose diagrams show that the K_1 axes trend NE-SW parallel to the syncline axis
 246 (yellow line), and the K_2 axes trend NW-SE perpendicular to the syncline axis. Focal-plane solutions for
 247 the studied normal faults are indicated with ϵ_1 , ϵ_2 , and ϵ_3 are the infinitesimal maximum, intermediate
 248 and minimum principal strain axes, respectively.



249

250 **Figure 4.** RT-AMS magnetic fabrics from sites at varied distances from co-seismic faults (Fault MM and
 251 Fault MP) in the Dead Sea Fault system, after Levi et al. (2014). The AMS axes show changeable
 252 orientations and asymmetrically distribution between the fault blocks. See Figure 3 for definition of
 253 stereoplots and focal-plane solutions.



254

255 **Figure 5.** Conceptual model for AMS axes around faults. (a) Fault of aseismic origin showing AMS axes
 256 that are related only to the regional strain field. (b) Fault of aseismic origin with fault-related AMS axes
 257 on both sides of the fault, indicating an unchangeable strain field in each block. (c) Faults of seismic
 258 origin showing fault-related AMS axes mainly on one side of the fault, which indicate a changeable
 259 strain field. Note that away (>50 m) from the faults in (b) and (c), the magnetic fabrics could be of type I
 260 (deposition, compaction).

261 **5. Conclusions**

262 We measured the magnetic fabrics around two mesoscale faults of aseismic origin. The results
263 show robust and consistent magnetic fabrics of deformation origin, implying that AMS provides
264 a sensitive deformation marker to study inelastic deformation in the host rock. The orientation of
265 the AMS axes indicate symmetric and unchangeable magnetic fabrics of deformation origin on
266 both sides of the studied faults, showing a strong association with the syncline axis and the
267 regional strain field. The symmetrical and unchangeable regional-related magnetic fabrics are
268 compatible with observations and models of strain field around aseismic fault formation. The
269 present results were compared with previous magnetic fabric around co-seismic faults that also
270 formed in low strength sediments at shallow depths under wet conditions. Next to these fault, the
271 magnetic fabrics are changeable and asymmetrically distributed. In light of the current results
272 and previous aforementioned magnetic fabric studies, we suggest that magnetic fabric analysis
273 provides a powerful and sensitive tool to characterize inelastic deformation around faults. We
274 highlight the new possibility of using magnetic fabric analysis to distinguish between aseismic
275 and seismic slip histories.

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279 The data used in this paper are available for download at:

280 <https://figshare.com/s/c6331844fa5cdf0580c5>.

281

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