

Very-long-period seismicity over the 2008-2018 eruption of Kilauea Volcano

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Abstract

Very-Long-Period (VLP) volcano seismicity often represents subsurface magma resonance, and thus provides insight into magma system geometry and magma properties. We develop a signal processing workflow using wavelet transforms to detect and assess period, decay rate, and ground displacement patterns of a wide variety of VLP signals. We then generate and analyze a catalog of VLP seismicity over the 2008-2018 open vent eruptive episode at Kilauea Volcano, Hawaii USA. This eruption involved a persistent lava-lake, multiple intrusions and rift zone eruptions, and a climactic caldera collapse, with VLP seismicity throughout. We characterize trends in two dominant magma resonances: the fundamental mode of the shallow magma system is a vertical oscillation of the magma column in the conduit/lava-lake, and higher frequency modes largely consist of lateral lava-lake sloshing. VLP seismicity was mainly triggered by lava-lake surface perturbations, and less commonly from depth. Variation in event period and decay rate occurred on timescales from hours-years. On timescales of months or less these changes were often correlated with other datasets, such as ground tilt, SO₂ emissions, and lava-lake elevation. Variation in resonant properties also occurs over days-months preceding and/or following observed intrusions and eruptions. Both gradual and abrupt changes in ground displacement patterns indicate evolution of shallow magma system geometry, which contributes to the variation in resonant modes. Much of the variation on timescales of months or less likely reflects changing magma density and viscosity, and thus could inform a variable shallow magmatic outgassing and convective regime over the ten year eruptive episode.

1 **Very-long-period seismicity over the 2008-2018 eruption of**
2 **Kīlauea Volcano**

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5 **Key Points:**

- 6 • Cataloging very-long-period volcano seismicity with wavelet transforms
7 • 2008-2018 Kīlauea Volcano magma resonance
8 • Comparing Kīlauea Volcano very-long-period seismicity with ground tilt, GPS,
9 lava-lake, and SO₂ data

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Abstract

Very-Long-Period (VLP) volcano seismicity often represents subsurface magma resonance, and thus provides insight into magma system geometry and magma properties. We develop a signal processing workflow using wavelet transforms to detect and assess period, decay rate, and ground displacement patterns of a wide variety of VLP signals. We then generate and analyze a catalog of VLP seismicity over the 2008-2018 open vent eruptive episode at Kīlauea Volcano, Hawaii USA. This eruption involved a persistent lava-lake, multiple intrusions and rift zone eruptions, and a climactic caldera collapse, with VLP seismicity throughout. We characterize trends in two dominant magma resonances: the fundamental mode of the shallow magma system is a vertical oscillation of the magma column in the conduit/lava-lake, and higher frequency modes largely consist of lateral lava-lake sloshing. VLP seismicity was mainly triggered by lava-lake surface perturbations, and less commonly from depth. Variation in event period and decay rate occurred on timescales from hours-years. On timescales of months or less these changes were often correlated with other datasets, such as ground tilt, SO₂ emissions, and lava-lake elevation. Variation in resonant properties also occurs over days-months preceding and/or following observed intrusions and eruptions. Both gradual and abrupt changes in ground displacement patterns indicate evolution of shallow magma system geometry, which contributes to the variation in resonant modes. Much of the variation on timescales of months or less likely reflects changing magma density and viscosity, and thus could inform a variable shallow magmatic outgassing and convective regime over the ten year eruptive episode.

1 Introduction

Volcano seismicity provides vital information for studying processes inside volcanoes and for monitoring changes in volcanic activity that inform hazards [Chouet, 1996; Ripepe *et al.*, 2015]. Amongst the rich variety of seismic signals that are commonly observed at volcanoes, so-called very-long-period (VLP) seismic events are of particular interest for magmatism as they likely represent fluid oscillations in magmatic transport structures [B.Chouet, 2013; McNutt and Roman, 2015]. VLP seismicity is typically defined as having a disproportionate amount of energy at periods greater than ~ 2 s, often focused into one or more discrete spectral peaks. This type of seismicity can provide otherwise unobtainable in situ insight into magma properties and magma plumbing system geometry [Chouet *et al.*, 2011, 2013; Karlstrom *et al.*, 2016; Liang *et al.*, 2019a], and can be sensi-

42 tive to different properties of the system than the longer timescale deformation observed
43 with geodesy. Here we develop a signal processing workflow for cataloging VLP seismic-
44 ity, and then apply this workflow to generate and analyze a catalog of VLP seismicity at
45 Kīlauea Volcano.

46 **1.1 Cataloging VLP seismicity**

47 Several studies have created catalogs of long or very-long period seismicity at vol-
48 canic settings [Battaglia, 2003; Aster *et al.*, 2008; Chouet *et al.*, 2010; Dawson *et al.*, 2014;
49 Knox *et al.*, 2018; Wech *et al.*, 2020], with a variety of approaches demonstrating that de-
50 tecting these signals robustly requires different approaches than detecting standard tec-
51 tonic earthquakes. Time-domain moving short-term-average/long-term-average (STA/LTA)
52 type detectors will miss many signals that do not stand above the background noise level
53 [Schaff, 2008]. Cross-correlation based template matching techniques can be much more
54 sensitive [Schaff, 2008] and have been used to detect some types of long-period seismicity
55 [Aster *et al.*, 2008; Wech *et al.*, 2020]. However, template matching is better suited to de-
56 tecting repeating events than signals that exhibit a continuum of variation (i.e., in resonant
57 periods, decay rates, and trigger mechanisms), and is computationally slow [Yoon *et al.*,
58 2015]. Approaches using feature-extraction to create and cluster waveform ‘fingerprints’
59 are computationally faster, but still best suited to detecting repeating events [Yoon *et al.*,
60 2015].

61 Supervised machine learning approaches can also be effective for detecting earth-
62 quakes [Perol *et al.*, 2018; Jennings *et al.*, 2019; Bergen and Beroza, 2019] and have been
63 used to detect very-long-period seismicity [Chouet *et al.*, 2010]. However, supervised
64 learning methods can require lots of pre-selected training examples, may not detect types
65 of signals they were not trained on robustly, will generally need at least partial re-design
66 and/or re-training to be applied to new networks/volcanoes, and their ‘black box’ nature
67 can make predicting when or why they fail difficult [Bell, 2014; Goodfellow *et al.*, 2016].
68 Unsupervised learning methods have been used to cluster seismic data [Kohler *et al.*, 2010;
69 Mousavi *et al.*, 2019], but have not yet been demonstrated to generate accurate/comprehensive
70 event catalogs. They will also generally require reanalysis/reinterpretation of the output
71 clusters when new data is added [Bell, 2014], and may thus be more promising as a tool
72 to help interpret variability in already cataloged events.

73 Accurately categorizing VLP signals is also important, since the resonant periods,
 74 decay rates (quantified by quality factor Q , a ratio of energy stored to energy lost per cy-
 75 cle), and source motions (from ground displacement patterns) can encode the underlying
 76 resonant mechanism [Liang *et al.*, 2019a,b]. Q is often difficult to calculate robustly, and
 77 several methods have previously been used. The simplest is to calculate the full width at
 78 half the maximum amplitude (FWHM) of peaks in the power spectrum, though this is of-
 79 ten not effective in the presence of noise, complicated signal shapes, or multiple signals
 80 with similar frequency components [Kumazawa *et al.*, 1990; Zadler *et al.*, 2004]. For this
 81 reason autoregressive (AR) methods that fit decaying sinusoids to the coda of signals were
 82 developed [Kumazawa *et al.*, 1990; Nakano *et al.*, 1998; Lesage *et al.*, 2002; Dawson *et al.*,
 83 2014]. When the coda of a signal can be appropriately isolated these methods work well
 84 for classifying dominant resonant modes. However, they often do not accurately detect or
 85 estimate Q of secondary resonant modes or modes with coda interrupted by other signals
 86 (Fig. S.7). Bandpass filtering can help isolate secondary signals, but often a narrow pass-
 87 band would be required which will artificially increase Q [Kumazawa *et al.*, 1990].

88 We use continuous wavelet transforms (CWTs) to detect and classify T , Q , and
 89 ground displacement patterns of VLP seismic signals. CWTs are a method for determin-
 90 ing the frequency content of signals over time [Alsberg *et al.*, 1997; Selesnick *et al.*, 2005]
 91 that have been previously used to analyze volcano seismicity [Lesage, 2009; Lapins *et al.*,
 92 2020]. Our methods are able to robustly determine T and Q in the presence of high noise,
 93 multiple resonant frequencies, and overlapping signals. Although not the focus here, these
 94 methods are readily extendable to characterizing VLP tremor [Chouet, 1996; Dawson
 95 *et al.*, 2014] and gliding-frequency signals. Our approach does not depend upon training
 96 data or templates, and thus can be applied to any seismic network or volcano with mini-
 97 mal configuration.

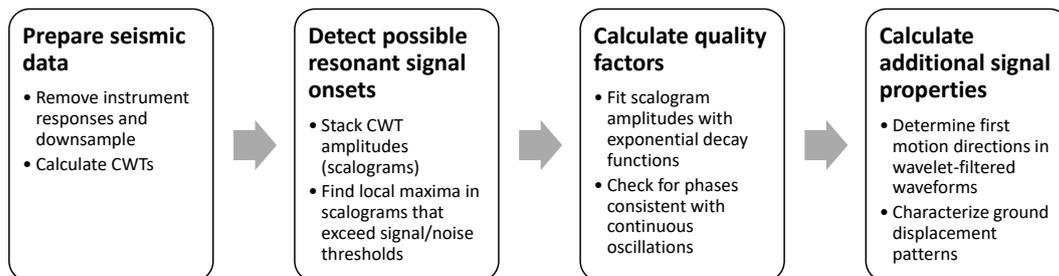
98 **1.2 2008-2018 eruption of Kīlauea Volcano**

99 Kīlauea Volcano is an excellent study location due to the dense broadband seismic
 100 network operated by the Hawaii Volcano Observatory, which has recorded thousands of
 101 VLP events over the past two decades [Dawson *et al.*, 2014; Liang *et al.*, 2019b]. There is
 102 also a wealth of other available data including direct observations of the Halema‘uma‘u
 103 summit lava-lake during these events [Orr *et al.*, 2013; Dawson *et al.*, 2014]. We examine
 104 the 2008-2018 eruptive episode, the most recent period of continuous summit activity fol-

105 lowing decades of quiescence or sporadic events largely focused along the East-Rift-Zone
 106 (ERZ) [Wright and Klein, 2014]. Over this timespan a summit lava-lake persisted at the
 107 surface, then drained as part of a caldera collapse eruption sequence in May-August 2018
 108 [Neal et al., 2019; Patrick et al., 2019a,b]. VLP seismicity at Kilauea has previously been
 109 cataloged up to 2013 using a hidden Markov model to detect events and the Sompi AR
 110 method to determine T and Q of these events [Dawson et al., 2014]; this existing catalog
 111 provides an important benchmark for our methods.

112 We find prevalent VLP seismicity over the whole 2008-2018 timespan, with VLP
 113 T , Q , and ground displacement patterns varying over timescales from hours to years. We
 114 compare our VLP catalog to other datasets such as lava-lake elevation, tilt and GPS (which
 115 measure summit reservoir inflation), SO_2 emissions, and observations of rift zone erup-
 116 tions and inferred intrusions. This yields insights into how known changes in the magma
 117 system are reflected in seismicity, and indicates additional changes on a variety of timescales.

118 2 Methods



119 **Figure 1.** Signal processing workflow for VLP detection and characterization.

2.1 Seismic data

Near-field broadband seismometers are best suited for picking up the often low amplitude long-period signals of interest to VLP studies. We use waveforms from 3-component broadband seismometers in the Hawaii Volcano Observatory (HVO) network [USGS, 1956] that are within ~ 3 km of the vent. We use available data from the following stations: NPB, NPT, SRM, OBL, WRM, SDH, UWE, UWB, SBL, KKO, and RIMD (Fig. 2, 10). Some other stations in the area were not used due to low signal/noise ratios. Data from 2008-2011 was obtained from the USGS, subsequent data is publicly available from IRIS (Incorporated Research Institutions for Seismology). We download and process data in 6 hr time windows. There are gaps in data availability for many of these stations; data gaps of less than 2 s duration are filled by linear interpolation and waveforms with larger gaps are discarded.

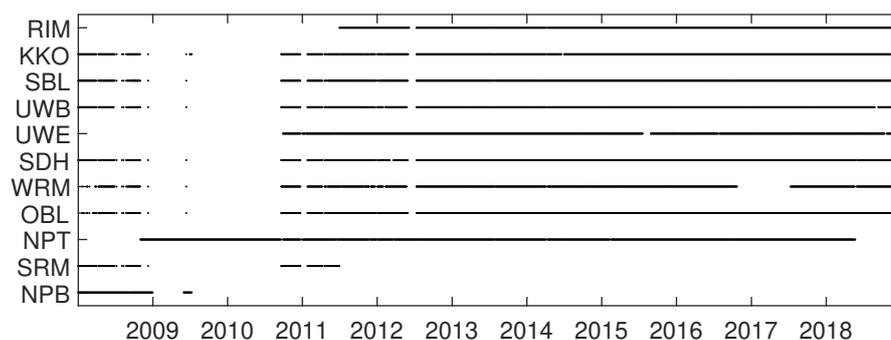


Figure 2. Timeline of data availability at the HVO broadband seismic stations used in this study.

In order to combine data from different instruments, we must deconvolve the instrument responses. A standard 'water level' is first applied to these instrument responses so that the maximum amplification is 10 times the base amplification. This prevents over-magnification of noise at periods longer than the instrument sensitivity ranges. We note that this process is not causal and can introduce artificial tapers around discontinuities (i.e., step functions); an effect included in the synthetic seismograms we use to test our methods (Appendix A:). To facilitate stacking and faster processing, all waveforms are then smoothed with a 'lowess' moving linear regression and resampled at 6 Hz. Lowess smoothing conformed to sharp discontinuities without introducing artificial oscillations better than other smoothing we tested such as FIR and IIR filters or moving quadratic regressions.

2.2 Continuous wavelet transforms

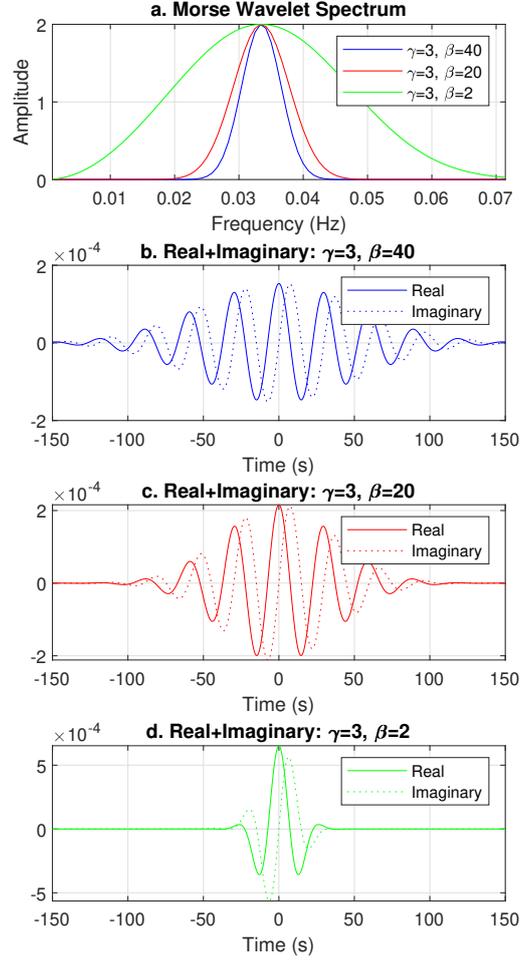
Time-frequency representations of data are well suited to identifying resonant signals [Köcher *et al.*, 2014]. A spectrogram is the simplest such representation, obtained from the amplitudes of a short-time Fourier transform (STFT) which consists of discrete Fourier transforms (DFTs) calculated over sequential time windows. However, there are disadvantages to STFTs that make continuous wavelet transforms (CWTs) better for our purposes. Other methods for time-frequency analysis such as the Wigner-Ville distribution and Hilbert-Huang transform have been applied to seismic data [Lesage, 2009], but we found them less useful than CWTs for our purposes.

CWTs involve specifying a base wavelet that can be stretched or ‘scaled’ to different frequencies and cross-correlated with data to determine frequency content as a function of time [Alsberg *et al.*, 1997; Selesnick *et al.*, 2005]. Plots of CWT amplitudes are termed scalograms. For a given wavelet, CWTs provide increasing temporal resolution with increasing frequency. This is one advantage over STFTs, which for a given window length provide the same temporal resolution for all frequencies, introducing an unnecessary trade-off between temporal resolution of high frequencies and spectral resolution of low frequencies.

Useful wavelets for time-frequency analysis are often sinusoids scaled by some function with symmetric, compact support so as to decay in both directions from a central point (Fig. 3). Wavelets with more gradual decay (i.e., more oscillations) will provide better frequency resolution but worse temporal resolution (Fig. 3), analogous to increasing window length in a STFT. An arbitrary number of ‘stretches’ of a wavelet can be used to sample at any desired frequencies, though there is a limit to the effective frequency resolution possible with a given wavelet width. The gradual onset of wavelets introduce less artificial temporal ‘jaggedness’ than a standard STFT, since a STFT uses sinusoids that terminate abruptly at the edges of each time window. This smoothness allows for more accurate determination of signal decay rates.

The convolution between a wavelet and an impulsive signal (such as a single peak) will have a duration and decay rate similar to the wavelet itself (Fig. S.6), analogous to how STFTs will cause impulsive signals to appear spread in time over the window length used. This means that the wavelet duration and decay rate will determine the minimum

175 signal duration and Q that can be distinguished from an impulsive signal, with narrower
 176 wavelets being able to resolve shorter and lower Q oscillations.



177 **Figure 3.** Morse wavelets used in this study (in this case scaled to a period of 30 s). The $\beta = 40$ (plot b)
 178 and $\beta = 20$ (plot c) wavelets are both used to make combined scalograms from which potential VLP signals
 179 are detected. The $\beta = 20$ wavelet is also used for calculating Q of signals. The $\beta = 2$ (plot d) wavelet is used
 180 for detecting first motions of signals.

181 We use Morse wavelets which are given in the spectral domain (for angular fre-
 182 quency ω) by:

$$\Psi_{\beta,y}(\omega) = U(\omega)a_{\beta,y}\omega^{\beta}e^{-\omega^{\gamma}} \quad (1)$$

183 where $U(w)$ is the Heaviside step function, β is a parameter that governs wavelet duration
 184 (number of oscillations), γ is a parameter that governs wavelet symmetry, and $a_{\beta,\gamma}$ is a
 185 normalizing constant [Lilly and Olhede, 2009]. We set $\gamma = 3$ which yields wavelets that
 186 are symmetric in the frequency domain [Lilly and Olhede, 2009].

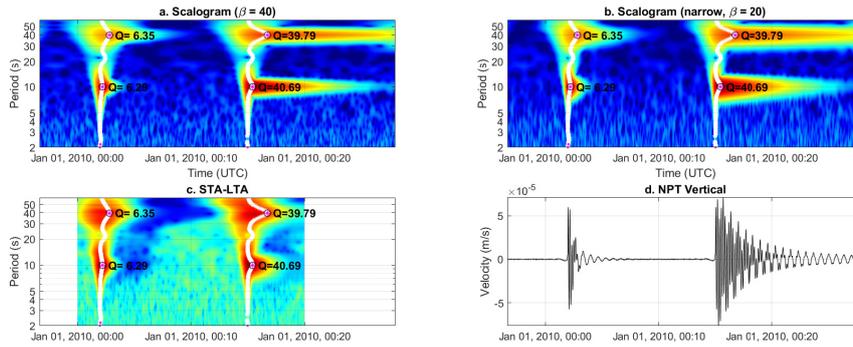
187 **2.3 Detecting potential resonant signal onsets**

188 To mitigate the inherent trade-off between spectral and temporal resolution we make
 189 combined scalograms using wavelets with two different values of β , 40 and 20 (Fig. 3).
 190 The $\beta = 40$ wavelet provides higher frequency resolution which helps more accurately de-
 191 termine resonant signal period. The $\beta = 20$ wavelet provides better time resolution while
 192 still providing enough frequency resolution to isolate typical VLP signals (Fig. S.4). Bet-
 193 ter temporal resolution helps determine onset times, reveal gaps in a signal which could
 194 indicate that it is not continuous resonance (Fig. S.5), and distinguish low Q resonance
 195 from impulsive signals (Fig. S.6).

196 We then stack the scalograms from all available stations to increase the signal/noise
 197 ratio. We exclude periods less than 10 s in this study because of the strong oceanic micro-
 198 seism at these periods [Berger et al., 2004; Dawson et al., 2014]. Given the proximity of
 199 our stations, delays from seismic wave propagation will be minimal relative to the periods
 200 of interest. For reference, at wave-speeds of 1800 m/s (a reasonable estimate for shallow
 201 s-wave speed at Kīlauea [Dawson et al., 1999; Lin et al., 2014]) a wave with a 10 s pe-
 202 riod will have a wavelength of 18 km, roughly four times the distance across our array
 203 (~ 5 km). There is also no concern about destructive interference from stacking scalograms
 204 since they contain no phase information.

205 To detect potential resonant signal onsets in a stacked scalogram, we first calculate
 206 moving long-term averages (LTA) and moving standard deviations of each frequency com-
 207 ponent with 200 s windows (Fig. 4). We then introduce a frequency-dependent delay of
 208 four cycles to the LTA and standard deviation values to account for non-causality intro-
 209 duced by the wavelets. Next we identify all local maxima in the stacked scalogram sep-
 210 arated by at least 200 s in each frequency band (Fig. 4). Finally, we keep only the local
 211 maxima with amplitudes that are above some chosen multiple of the LTA (which we refer
 212 to as the STA/LTA threshold), and that are also more than some threshold number of
 213 standard deviations above the LTA. We select a threshold of 3 for both; chosen to mini-

214 minimize noise (or false positives) while still keeping most desired signals in both synthetic
 215 tests and real data (Fig. S.3, S.10, S.11). Where local maxima occur at adjacent periods
 216 or with periods separated by less than a factor of 1.07 (the minimum separation in periods
 217 that can be robustly resolved with the wavelets we use), we keep the maxima correspond-
 218 ing to the highest energy integrated over the following two cycles, which is more robust
 219 than just keeping the highest maxima (Fig. 4).



220 **Figure 4.** Example scalograms and detected resonant signals from synthetic scalograms (Appendix A:
 221). This synthetic seismogram (plot d) consists of four VLP signals with [start time, T , Q] = [00:05, 40, 6],
 222 [00:05, 10, 6], [00:15, 40, 40], [00:15, 40, 40], plus white noise from a standard normal distribution scaled
 223 by 0.1% of the signal amplitude. We note that the slight precursory oscillations that arise from removing the
 224 instrument response. White dots in scalograms (plots a, b, and c) indicate temporal local maxima that meet
 225 the minimum STA/LTA criteria, and magenta dots indicate points that are spectral local maxima (integrated
 226 over two cycles). Black circles and text indicate the final selected resonant signal onsets and corresponding
 227 calculated Q . Here T and Q of all resonant signals are recovered accurately.

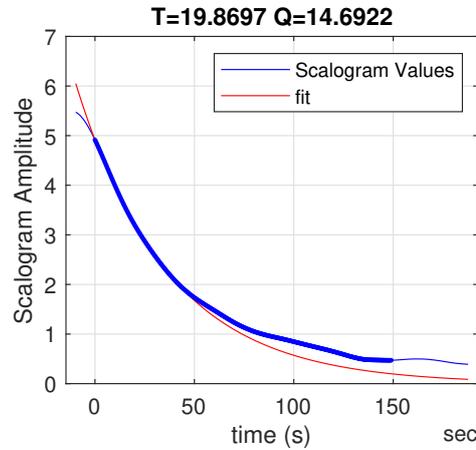
228 2.4 Calculating the quality factor (Q) of resonant signals

229 We calculate Q by fitting decaying exponentials to stacked scalogram amplitudes fol-
 230 lowing each detected potential resonant signal onset (Fig. 5). We use only the narrower
 231 $\beta = 20$ CWTs that have better temporal resolution (Fig. 3); the minimum Q that this
 232 wavelet can robustly resolve is around 6. We extract scalogram amplitudes at the target
 233 frequency over one to eight cycles after the identified signal onset. The one cycle delay
 234 avoids the region near the onset of an impulsively initiated signal where amplitudes will
 235 be inherently underestimated (since part of the wavelet will not be overlapping the signal),
 236 and also helps avoid artifacts that might be present from a resonance trigger mechanism.

237 A standard least-squares exponential regression could underestimate decay rate in the
 238 presence of noise or where another signal starts within the fitting window (Fig. S.8). We
 239 instead solve for the exponential curve with initial amplitude fixed to the initial scalogram
 240 amplitude $A(t_1)$ and with the slowest decay rate g that remains under all of the scalogram
 241 amplitudes in the timespan being fit (t_1 to t_2) (Fig. 5, S.8):

$$g = \min_{t=t_1}^{t_2} \left(\frac{\ln(A(t)) - \ln(A(t_1))}{t - t_1} \right) \quad (2)$$

242 which then yields quality factor: $Q = -\pi/(Tg)$. This fitting method is also less sensitive
 243 to the choice of fitting timespan than a least-squares regression would be. Extending the
 244 timespan will have no effect unless the added amplitudes fall beneath the current fit.



245 **Figure 5.** Example estimation of Q by scalogram exponential fit from a synthetic seismogram. This
 246 seismogram consists of a VLP signal with $[T, Q] = [20 \text{ s}, 15]$, plus white noise from a standard normal distri-
 247 bution scaled by 0.1% of the signal amplitude. The bold part of the blue line shows the part of the scalogram
 248 data that is being fit.

249 Since this method does not account for phase, non-continuous oscillations that are
 250 close in time might not be distinguished from true resonance. To mitigate this we also
 251 extract the phases of the $\beta = 20$ CWTs at each channel and check for consistency over
 252 the timespan being fit. For a continuous oscillation, the phase (θ) of a wavelet stretched to
 253 the oscillation frequency f will increase steadily as it is convolved with the signal (Fig. 6,
 254 S.9):

$$\theta_{\text{expected}}(t) = 2\pi ft + \theta(0) \quad (3)$$

255 A signal that is not a continuous sinusoid can exhibit deviations from this expected phase
 256 (Fig. 6). To quantify how ‘continuous’ a signal is, we calculate the mean deviation from

257 the expected phase over the timespan $(t_0 - t_1)$ being fit and over all N channels:

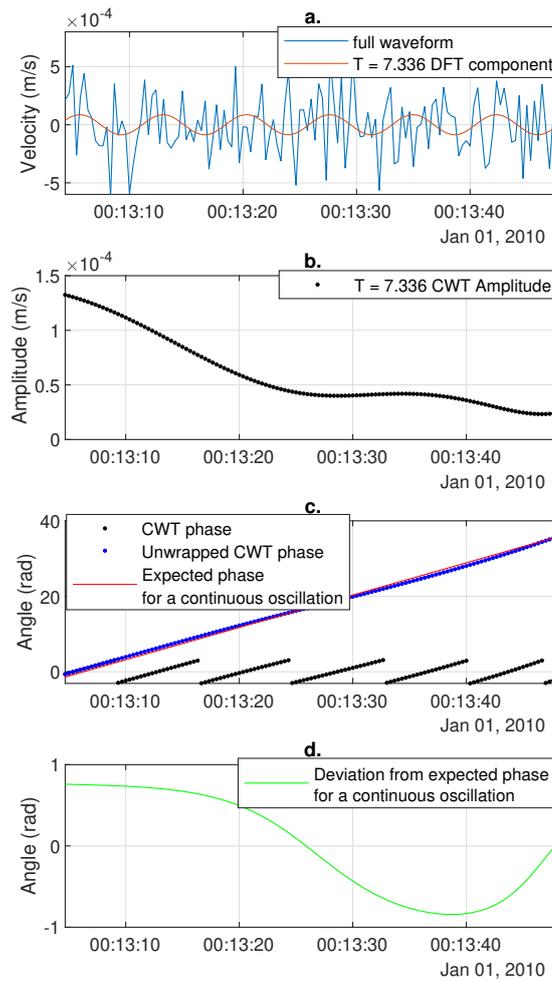
$$\text{mean phase deviation} = \frac{1}{N} \frac{1}{t_1 - t_0} \sum_{n=1}^N \int_{t_0}^{t_1} |2\pi f t + \tilde{\theta}_n - \theta_n(t)| dt \quad (4)$$

258 where $\tilde{\theta}_n$ is the constant phase offset that minimizes phase deviation at channel n . We use
 259 this this phase offset instead of the actual initial phase $\theta_n(t_0)$ in case there are source ef-
 260 fects or strong noise present at the start of the timespan. We then keep only signals with
 261 a mean phase deviation of less than a threshold value of 0.1 radians. This threshold min-
 262 imized noise or other discontinuous signals while still keeping most continuous resonant
 263 signals in tests on both synthetic and real data (Fig. 6, S.9, S.10, S.11).

267 We note that these methods are not designed for detecting or characterizing gliding-
 268 frequency signals. However, the methods introduced here could be readily modified to
 269 characterize gliding-frequency signals, since time-frequency analysis is the most intuitive
 270 way to examine such signals [Köcher *et al.*, 2014]. This would involve first tracing T over
 271 time from scalograms, which while straightforward in concept would need to be imple-
 272 mented in a manner that is robust in the presence of complicated signals and noise. The
 273 exponential fit could then be applied to these traces to calculate decay rates, and the ex-
 274 pected phase at each time could be adjusted according to changing T to check whether the
 275 gliding-frequency signal is likely a continuous oscillation.

276 **2.5 Comparison with previous Kīlauea VLP catalog**

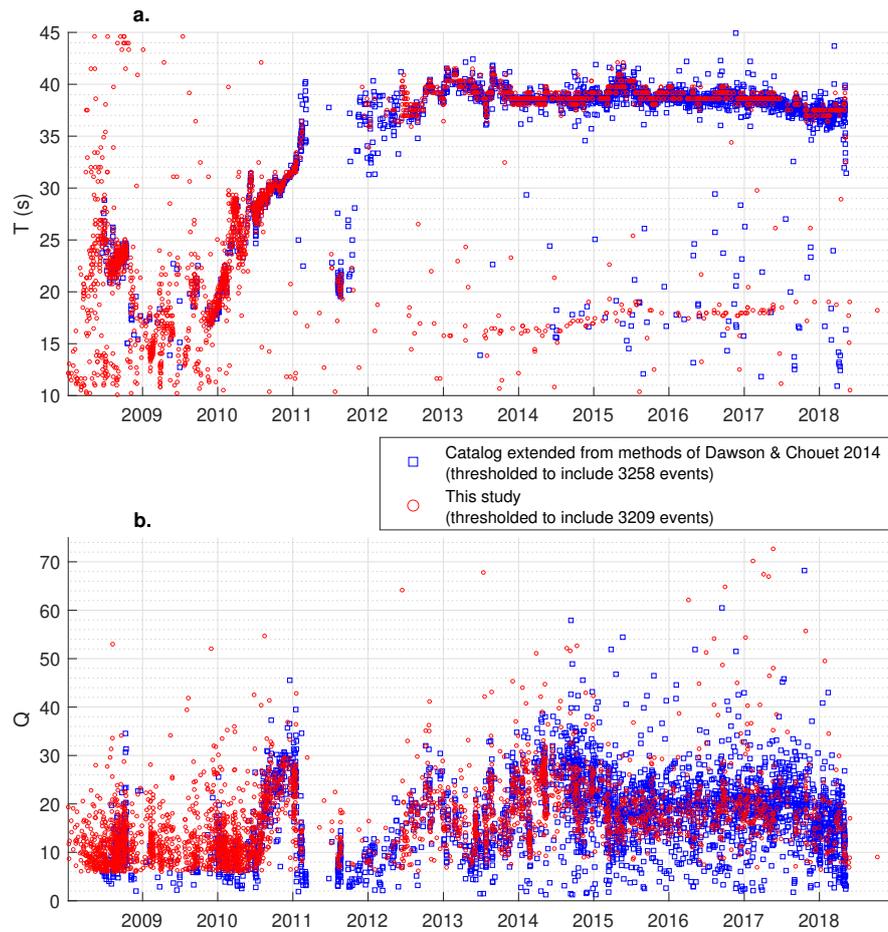
277 We compare our catalog to one produced using the automated detection (via a hid-
 278 den Markov model trained on example events [Dawson *et al.*, 2010]) and classification
 279 (via the Sompi AR model Kumazawa *et al.* [1990]) methods of Dawson *et al.* [2014], ex-
 280 tended through 2018. For both catalogs adjustment of various ‘quality thresholds’ is re-
 281 quired to exclude excessive amounts of likely false picks. In our catalog we use thresh-
 282 olds: STA/LTA > 3 , standard deviations above LTA > 3 , and mean phase deviation < 0.1
 283 radians. In the catalog extended from Dawson *et al.* [2014] the most useful parameters
 284 to threshold are event amplitude at station NPB or NPT and the standard deviation of Q
 285 from the Sompi fits. We set these thresholds to 400 counts and 0.25 so that this catalog
 286 contains a similar number of events to our catalog (~ 3200); but note that stricter thresh-
 287 olds would result in lower apparent scatter. In both catalogs changing these thresholds will
 288 greatly vary the number of events included, and less strict thresholds will include tens of
 289 thousands of additional events (Fig. S.10, S.11). For the thresholds used, the two catalogs



264 **Figure 6.** Example phase continuity from a spectral peak in synthetic random noise. In a scalogram (or
 265 frequency spectrum) this signal appears to contain a potential VLP event, but the high phase deviation (plot d)
 266 correctly indicates that it is not a continuous oscillation.

290 include around 1000 overlapping signals (Fig. 7, 8). Both catalogs also include a similar
 291 number of events that appear likely to be false detections, based on visual inspections of
 292 events in various parts of the parameter space.

297 Both catalogs detect a similar trend of signals with T of ~ 20 s in 2010, increas-
 298 ing to ~ 40 s by 2012 and remaining around 40 s until 2018. Since Dawson *et al.* [2014]
 299 use a Markov model that was trained specifically for events in this trend, it might be ex-
 300 pected to detect some of these events with lower signal/noise ratios than our more general



293 **Figure 7.** Comparison of detected VLP events from this study with a catalog extended from *Dawson et al.*
 294 [2014]. Event detection thresholds were chosen for the catalog extended from *Dawson et al.* [2014] that pro-
 295 duced a similar number of events to our catalog; orders-of-magnitude more or less events would be present in
 296 either catalog depending upon the thresholds chosen (Section 4.1, Fig. S.10, S.11).

301 STA/LTA based approach can detect without also introducing excessive false detections.
 302 There are indeed a number of these events unique to the catalog extended from *Dawson*
 303 *et al.* [2014], but also many of these events unique to our catalog. This may be partly be-
 304 cause our approach leverages data from multiple stations to increase signal/noise ratios,
 305 and partly due to limitations of the Markov model detection approach.

306 Our catalog also includes some additional unique clusters of signals. These include
 307 a clear cluster with $T \sim 15$ s in early 2009, and some other more isolated clusters between

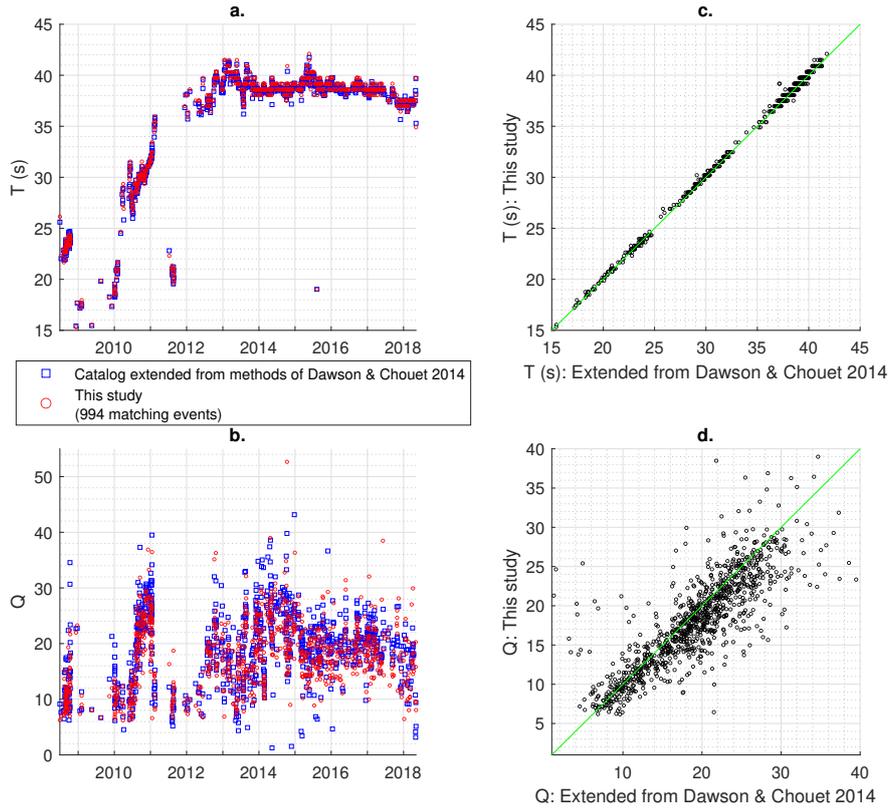
2008 and 2010 (Fig. 7). Most prominently, our catalog also includes a band of signals
 with $T \sim 10\text{-}20$ s between 2010 and 2018 (Fig. 7). Some of these events that coincide
 with a 40 s event are picked up by the Sompi AR method [Dawson *et al.*, 2014], but even
 where they are detected the Sompi AR method often does not produce accurate estimates
 of Q for such secondary signals.

Our catalog appears to exhibit more scatter in T prior to 2010, but many of these
 values do appear to represent real signals. Both catalogs show a number of isolated sig-
 nals after 2011 with T from $\sim 10\text{-}15$ and $\sim 20\text{-}35$ s. Most of these signals in our catalog
 appear to be from gliding-frequency VLP events; some in the catalog extended from *Daw-*
son et al. [2014] also are related to gliding-frequency events whereas some appear to be
 noise.

A final notable difference between the two catalogs is in estimates of Q . As dis-
 cussed in section 2.3, our method cannot robustly detect events with $Q < 6$ given the
 wavelets we are using. However, low Q signals cannot be as accurately characterized any-
 ways, since T cannot be very accurately determined for a small number of oscillations.
 The large scatter in T from late 2011-early 2012 in the catalog extended from *Dawson*
et al. [2014] likely reflects this limitation. Estimates of Q often differ between the two
 methods even for matching events (Fig. 8), though neither method shows a bias for higher
 or lower values than the other. Where the two methods estimate appreciably different val-
 ues of Q we find that there is often some complication (such as overlapping signals or
 strong noise) that causes the Sompi AR method to be inaccurate where our method still
 produces reasonable estimates of Q .

2.6 Determining first motion directions

The first motions of a signal are not well defined for signals without impulsive on-
 sets. Even for impulsive onsets, picking first motions for a particular frequency component
 is difficult to do robustly because band-pass filtering a signal will distort the onset of that
 signal regardless of the filter used (i.e., causal or acausal, FIR or IIR) (Fig. 9). We use a
 ‘wavelet filter’: we compute the CWT of a signal, then reconstruct the signal using an in-
 verse CWT but keeping only the period of interest. This still produces artificial precursory
 oscillations in front of signals with impulsive onsets (Fig. 9), but the size of these oscilla-
 tions are predictable for a given wavelet (Fig. 3), even when the signal onset involves step

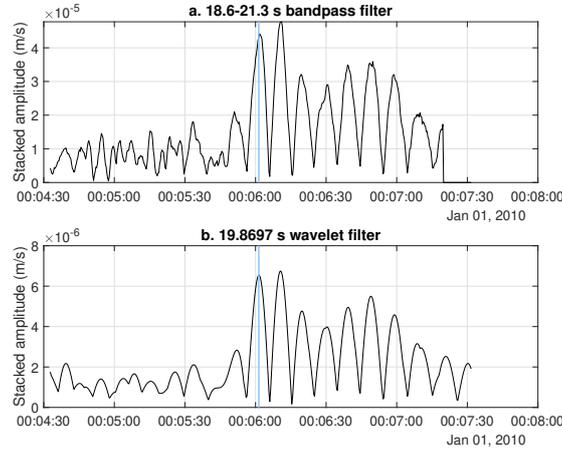


330 **Figure 8.** VLP events from this study that correspond to events in a catalog extended from *Dawson et al.*
 331 [2014]. Corresponding events have start times within 3 minutes of each other, and T ratios within 4/5-5/4 of
 332 each other. Green lines in plots c and d indicate 1:1 values.

342 displacements. We use a very narrow Morse wavelet ($\beta = 2$) in order to minimize pre-
 343 cursory oscillations, though such a narrow wavelet will be more sensitive to surrounding
 344 frequencies (Fig. 3). This method will thus only work well for signals that are the domi-
 345 nant oscillations in their frequency band.

346 We then stack the amplitudes of the wavelet-filtered signals from all channels, and
 347 identify local maxima around the signal onset time that exceed thresholds for both STA/LTA
 348 and number of standard deviations above the LTA (Fig. 9). We discard local maxima that
 349 are less than half of the maximum amplitude, which will exclude precursory oscillations
 350 caused by the wavelet filter for impulsive onset signals. If one or more maxima remain
 351 we select the first of these as the first motion time, and select corresponding first mo-
 352 tion directions at each channel from the wavelet filtered waveforms (Fig. 9). We store the

353 STA/LTA ratio and standard deviations above the LTA for this local maximum as indica-
 354 tors of pick confidence. If no suitable local maxima are found, which occurs if the signal
 355 has a gradual onset or is contaminated by other signals/noise, we label the first motions
 356 undetermined.



357 **Figure 9.** Example first motion pick from a synthetic seismogram for an impulsive onset oscillation with
 358 [start time, T , Q] = [00:06, 20, 20], plus a step displacement (velocity spike) at time 00:06, plus two other
 359 equal-amplitude resonant signals with [start time, T , Q] = [00:05, 80, 20] and [00:05, 5, 20], and plus white
 360 noise from a standard normal distribution scaled by 0.1% of the signal amplitude. Plot a shows stacked am-
 361plitudes from waveforms filtered with an FIR bandpass filter; this is not used for picks and is just shown
 362 for comparison. Plot b shows stacked amplitudes from waveforms filtered with the wavelet filter we use for
 363 picking first motions. The cyan line is the algorithm’s correct first motion pick for the target signal.

364 2.7 Characterizing ground displacement patterns

365 Average phases and amplitudes at each channel are obtained using the Goertzel DFT
 366 algorithm [Proakis and Monolakis, 1990] over a time window between one and five cycles
 367 after each signal onset. Our goal in this study is not to conduct detailed source inversions
 368 for every resonant signal, but rather to quantitatively characterize changes in ground dis-
 369 placement patterns between VLP events. The simplest metric we use is the average verti-
 370 cal/horizontal velocity ratio, defined for a given frequency f as:

$$\text{vertical/horizontal} = \sum_{m=1}^M \frac{|\dot{u}_{Z,m}(f)|}{|\dot{u}_{E,m}(f) + \dot{u}_{N,m}(f)|} \quad (5)$$

371 for vertical (Z), east (E), and north (N) velocities (\dot{u}) at all M stations. This metric is
 372 very simple and requires no assumptions of source location or mechanics, but it is sen-
 373 sitive to tilt which will increase the apparent amplitude of horizontal components at in-
 374 creasing T .

375 We also quantify how radially symmetric horizontal motion vectors are by calcu-
 376 lating the angles from the direction to an inferred source location. We set this location
 377 based on a previous geodetic (InSAR, GPS, and tilt) inversion for the shallow ground de-
 378 flation source in early 2018 [Anderson *et al.*, 2019] (Fig. 10), which is similar to the shal-
 379 low source location inferred by other seismic and geodetic inversions over the past decade
 380 [Chouet *et al.*, 2010, 2011; Anderson *et al.*, 2015; Anderson and Poland, 2016; Liang *et al.*,
 381 2019b]. We then calculate the mean angle between observed \dot{u} and predicted \dot{w} velocity
 382 vectors as:

$$\text{radial misfit} = \frac{1}{M} \sum_{m=1}^M \int_0^{2\pi} \left| \arccos \left(\frac{\dot{u}(t) \cdot \dot{w}(t)}{|\dot{u}(t)| |\dot{w}(t)|} \right) \right| dt \quad (6)$$

383 The final method we use to quantify ground displacement patterns is conducting
 384 source inversions for an inflating/deflating spherical reservoir using a quasi-static ‘Mogi’
 385 model for a point source in an elastic half-space [Mogi, 1958; Segall, 2010]. Multiple
 386 previous seismic and geodetic studies have supported a spherical or ellipsoidal reservoir
 387 geometry [Baker and Amelung, 2012; Anderson *et al.*, 2015; Anderson and Poland, 2016;
 388 Liang *et al.*, 2019b], though some other seismic studies have instead inferred intersect-
 389 ing dikes [Chouet *et al.*, 2011]. Since many studies support a sphere-like reservoir, and
 390 because inversions for these VLP signals with more complex source models such as full
 391 moment tensors or dikes are often not well constrained, we focus only on the spherical
 392 reservoir model. Due to their simplicity, the Mogi source inversions are most useful as a
 393 metric of relative changes in source centroid depth between events rather than as a probe
 394 of detailed reservoir shape. In general changes in inferred Mogi centroid depth could rep-
 395 resent changes in the vertical extents of a spherical/ellipsoidal reservoir, and/or changes in
 396 the geometry or activation of any secondary dike/sill structures that may also be contribut-
 397 ing to the ground displacement patterns. The misfit of predicted and true displacements
 398 from Mogi inversions also provides a second metric for the radial symmetry of ground
 399 displacement patterns.

We include ground tilt (detected as horizontal acceleration by broadband seismometers) in the Green's functions [Maeda *et al.*, 2011] to predict displacements w as:

$$w(f) = \left(\mathbf{G}^{trans} + \mathbf{G}^{tilt} \frac{g}{(i2\pi f)^2} \right) P(f), \quad (7)$$

where \mathbf{G}^{trans} and \mathbf{G}^{tilt} are the tilt and translation Green's function matrices, g is gravitational acceleration, and P is forcing pressure. We can then solve for the P that best fits observed displacements u for a given set of Green's functions using a linear least-squares inversion.

We again fix the east and north source location based on previous geodetic inversions [Anderson *et al.*, 2019] (Fig. 10). We assume a shear modulus of 10 GPa and Poisson's ratio of 0.25. We then conduct a grid search over source depth between 500-2500 m beneath the caldera floor, choosing the depth that minimizes misfit according to:

$$\text{misfit} = \frac{\sum_{n=1}^n |w_n(f) - u_n(f)|}{\sum_{n=1}^n |u_n(f)|} \quad (8)$$

for all N channels.

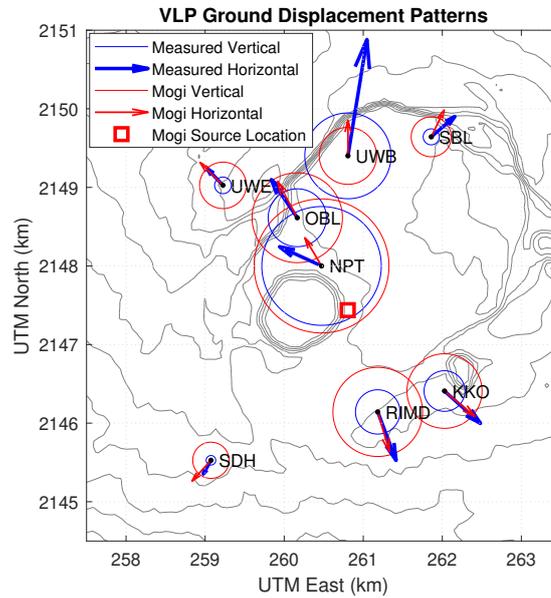


Figure 10. Ground displacements and Mogi inflating spherical reservoir source inversion for an example conduit-reservoir event on 2017-5-21 at the time of peak vertical displacement at station NPT. Displacements are from integrated seismic velocities, so horizontal components in the data and source inversion include both tilt and translation effects.

2.8 Other data

ERZ eruptions prior to 2018 have been compiled in *Patrick et al.* [2019b]: the March 2011 Kamoamoā fissure eruption [*Orr et al.*, 2015], August 2011 Pu‘u ‘Ō‘ō vent opening, September 2011 Pu‘u ‘Ō‘ō vent opening, June 2014 Pu‘u ‘Ō‘ō vent opening [*Poland et al.*, 2016], and May 2016 Episode 61g Pu‘u ‘Ō‘ō vent opening [*Chevrel et al.*, 2018]. Timing of the 2018 eruption is given in *Neal et al.* [2019]. Documented summit intrusions have been compiled in *Patrick et al.* [2019b]: October 2012, May 2014, and May 2015 [*Johanson et al.*, 2016]. Regional slow-slip events (SSEs) have been compiled in *Montgomery-brown et al.* [2015] and *Wang et al.* [2019]: February 2010, May 2012, and October 2015.

To indicate long-term ground deformation we use data from near-field (within ~2 km of the vent) GPS stations (vertical displacements from station HOVL and horizontal line-lengths between stations UWEV and CRIM [*Miklius*, 2008]) and tilt-meters (east and north tilt from station UWE [*Johanson*, 2020]). To infer ground inflation-deflation trends, we combine the GPS and tilt-meter data. We first smooth all four datasets with 30-day moving average filters. We then resample each dataset at 1-day periods and rescale each dataset to have a unit range. Lastly, we flip the sign of UWE east tilt-meter data (since eastward tilt at this station corresponds to ground deflation), and stack the four datasets. We then consider times when the stacked value is positive to represent long-term ground inflation, and negative to represent long-term ground deflation.

Lava-lake elevation data is obtained from webcam images, thermal images, and laser rangefinder data [*Patrick et al.*, 2019b] (data extended through 2018 was obtained from the USGS). We also include estimates of lava-lake surface area from *Patrick et al.* [2019b].

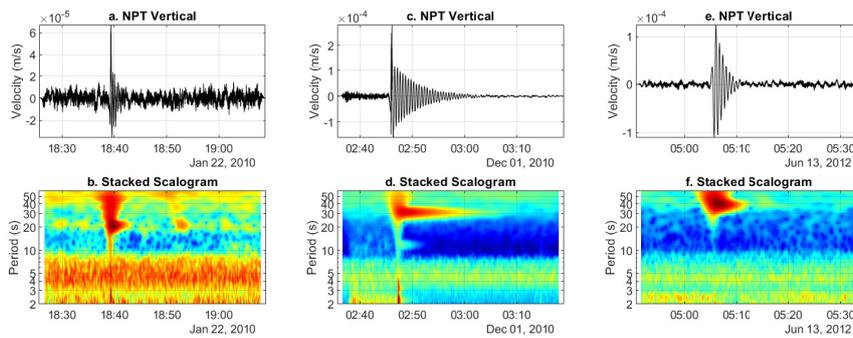
SO₂ is generally the most easily measurable major volcanic volatile species, and is an important indicator of magmatic processes [*Sutton and Elias*, 2014]. SO₂ data from various monitoring stations for the whole timespan does exist *Whitty et al.* [2020], but we only consider data from published studies using direct measurements of the summit plume. We use SO₂ emission data collected by a vehicle-based FLYSPEC UV spectrometer from 2007-2010 [*Elias and Sutton*, 2012]. We also use SO₂ emission data collected by an array of FLYSPEC UV spectrometers from 2014-2017 [*Elias et al.*, 2018]. Both datasets have large uncertainties (Fig. 13, 14) due to spectral fitting limitations and uncertainty in plume speed and location [*Elias and Sutton*, 2012; *Elias et al.*, 2018].

447 We also analyze the time-derivatives of some of these datasets. Comparing time-
 448 derivatives can sometimes better reveal short-term correlations, particularly when gradual
 449 or punctuated changes in the relation between two variables causes the direct correlation
 450 over long timespans to exhibit large scatter. Since derivatives are inherently more sensitive
 451 to high frequency noise, we calculate time-derivatives using FIR differentiator filters with
 452 7-day corner periods.

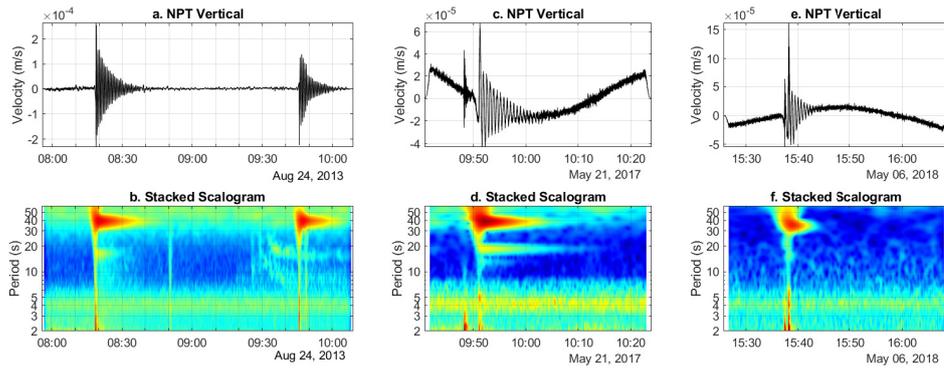
453 3 Results

454 3.1 Types of VLP seismicity at Kīlauea from 2008-2018

455 We will introduce the common types of VLP signals present in the catalog to facili-
 456 tate discussion in the following sections.



457 **Figure 11.** Example VLP signals. (Plots a and b) Normal conduit-reservoir mode event along with back-
 458 ground VLP tremor from January 2010, when the lava-lake became persistent [*Patrick et al., 2019b*]. The
 459 event had an impulsive broadband onset and inflationary first motions, indicative of a rockfall trigger. The
 460 background VLP tremor had the same dominant period as the impulsively triggered VLP event, but often
 461 unclear onsets and no higher frequency triggers. (Plots c and d) Normal conduit-reservoir event with sec-
 462 ondary lava-lake-sloshing mode from December 2010, two months before the March 2011 Kamoamoia fissure
 463 eruption. This event had an impulsive broadband onset and inflationary first motions indicative of a rockfall
 464 trigger. There was also background tremor at periods less than around 3 s that was truncated by this event.
 465 (Plots e and f) Reverse VLP event from June 2012, shortly after the May 2012 SSE. This event had an im-
 466 pulsive onset but no high frequency trigger. There was a small initial inflationary motion but the first large
 467 oscillation was deflationary.



468 **Figure 12.** Example VLP signals. (Plots a and b) Normal and Reverse conduit-reservoir modes and lava-
 469 lake-sloshing mode from August 2013. The Normal conduit-reservoir mode started at around 7:50 with an
 470 impulsive inflationary broadband trigger indicative of rockfall, and with an accompanying lava-lake-sloshing
 471 mode. The Reverse conduit-reservoir event occurred 90 minutes later, with no concurrently triggered lava-
 472 lake-sloshing, and appears to be partially truncated around 5 minutes after it's onset. A gliding-frequency
 473 VLP signal started about 20 minutes before the second event, with no apparent trigger and a final period
 474 similar to the previous lava-lake-sloshing mode. (Plots c and d) Normal conduit-reservoir event with two lava-
 475 lake-sloshing modes from May 2017. A higher frequency impulsive signal occurred about 2 minutes before
 476 these resonant modes that may have been related to their triggering. (Plots e and f) Normal VLP event from
 477 May 2018, 4 days after the lava-lake began draining. This event exhibited a distinctly lower T than preceding
 478 events (35 s as compared to 37-40 s), and is the last event conduit-reservoir event recorded in our catalog.
 479 This event started with an impulsive inflation, though with minimal broadband energy. Another larger broad-
 480 band impulse occurred a minute later that corresponded to increased oscillation amplitude, after which the
 481 oscillation decayed exponentially.

482 3.1.1 Conduit-reservoir resonance

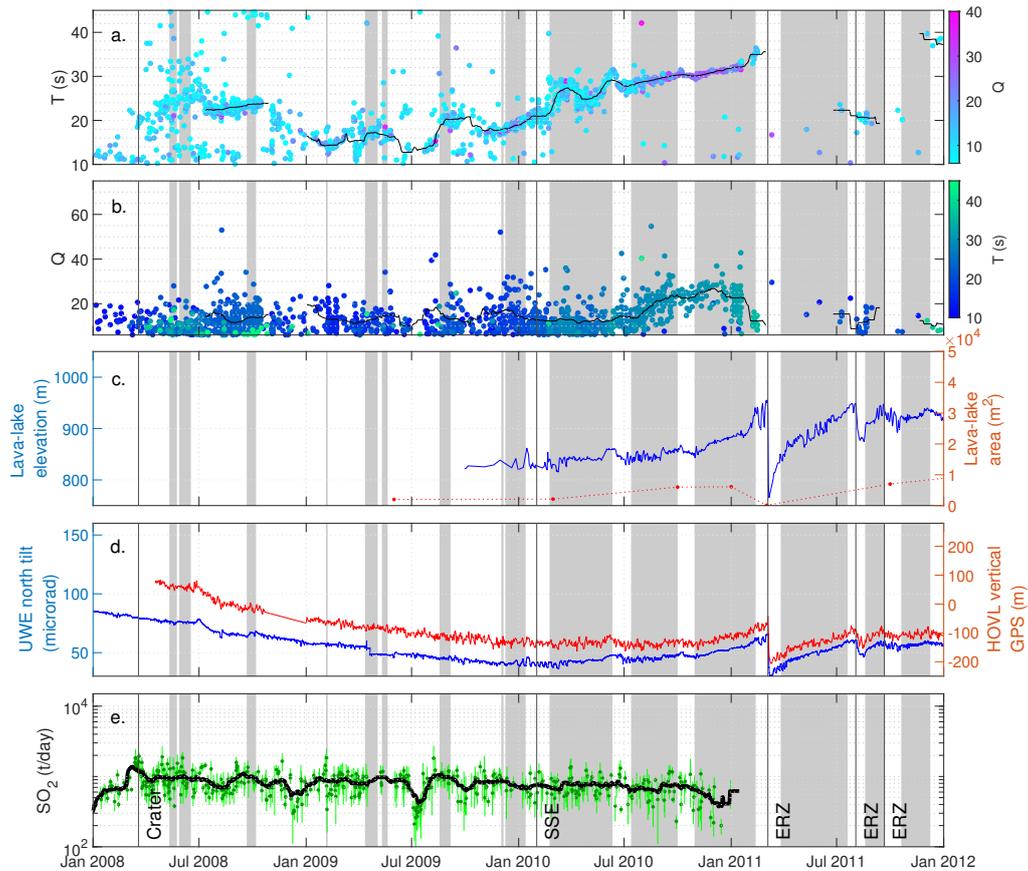
483 The first category of signals we term 'conduit-reservoir modes' *Liang et al.* [2019a].
 484 These modes constitute the main trend of VLPs starting at $T \sim 20$ s in 2010, increasing to
 485 ~ 40 s in early 2011, and fluctuating between 35-43 s from 2012 until the caldera collapse
 486 onset in May 2018 (Fig. 13, 14). Some other signals prior to 2010 and during the series
 487 of lava-lake draining events in 2011 may also fit into this category.

488 The conduit-reservoir oscillation is the fundamental resonant mode of the coupled
 489 conduit and shallow magma reservoir system, in which the magma column in the con-
 490 duct oscillates vertically and pushes magma in and out of the underlying reservoir [*Liang*

491 *et al.*, 2019b]. Other resonant modes such as Krauklis (crack) waves or acoustic resonance
 492 are predicted to generally have higher frequencies and lower amplitudes [*Karlstrom et al.*,
 493 2016; *Liang et al.*, 2019a]. Restoring forces for the conduit-reservoir oscillation come from
 494 magma reservoir compressibility (combined wall rock elasticity and multiphase magma
 495 compressibility) and gravity [*Liang et al.*, 2019a]. Viscous drag along the conduit walls
 496 is probably the primary control of damping for these oscillations, and also impacts res-
 497 onant period. Ground deformation during these events is primarily from uniform infla-
 498 tion/deflation of the magma reservoir; deformation from the conduit is small by compari-
 499 son [*Liang et al.*, 2019b].

500 Conduit-reservoir mode resonance could be triggered/driven by a variety of differ-
 501 ent mechanisms, producing signals with different onset characteristics. We term conduit-
 502 reservoir modes with abrupt onsets and inflationary first motions ‘Normal’ events; this
 503 category includes rockfall or lava-lake surface explosion triggered events and is analo-
 504 gous to ‘type 2’ events in [*Dawson et al.*, 2014]. There is often high-frequency or broad-
 505 band energy present at the onset of Normal events, as well as inflationary steps in tilt
 506 data [*Chouet et al.*, 2013; *Orr et al.*, 2013; *Dawson et al.*, 2014] (Fig. 11, 12, S.23, S.24).
 507 We term conduit reservoir modes with abrupt onsets and deflationary first motions ‘Re-
 508 verse’ modes; analogous to ‘type 3’ events in [*Dawson et al.*, 2014] (Fig. 11). These sig-
 509 nals often do not have obvious high frequency triggers, and some exhibit deflationary tilt
 510 steps [*Dawson et al.*, 2014]. The trigger for Reverse events is not known [*Dawson et al.*,
 511 2014], but could involve impulsive mass injections at depth or bubble rise/collapse. Some
 512 conduit-reservoir events do not fit very clearly into either category, for example those with
 513 gradual onsets or multiple step increases in oscillation amplitude (Fig. 12, S.24).

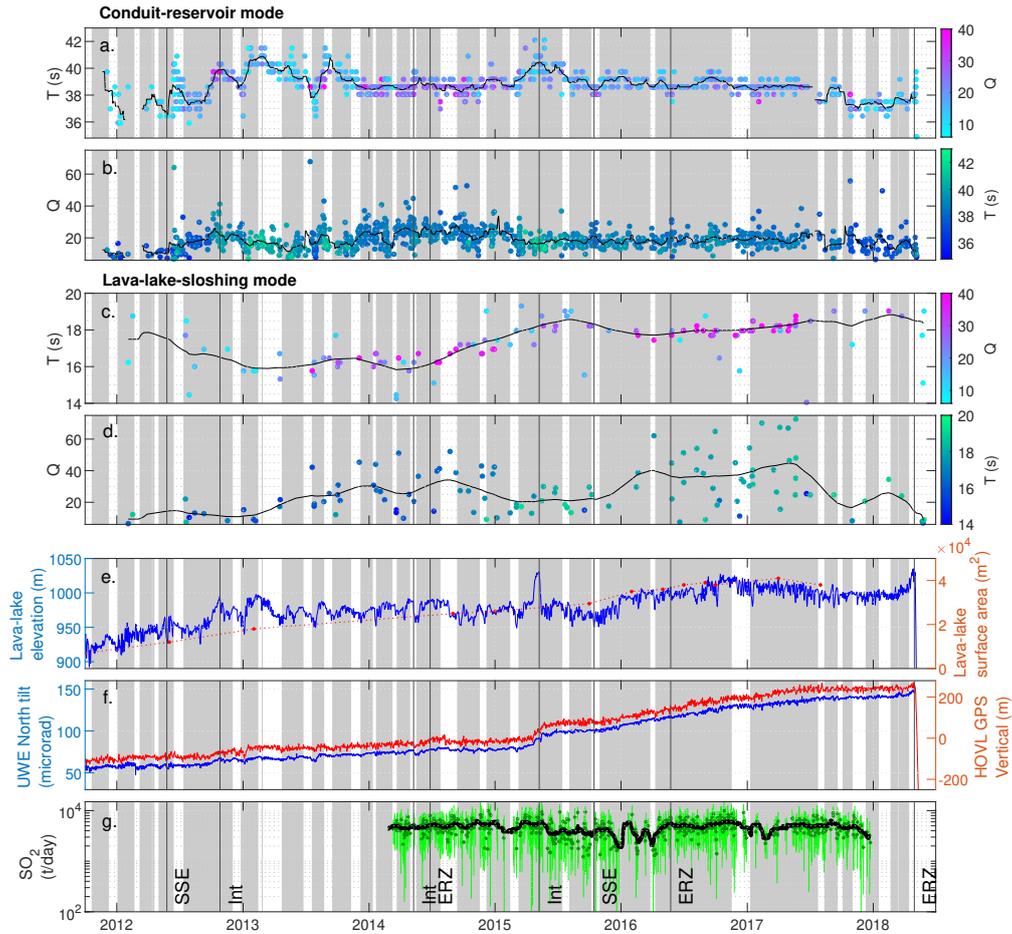
514 Our algorithm classifies ~77% of conduit-reservoir events after 2012 as Normal,
 515 ~17% as Reverse, and the remaining ~6% are undetermined (Fig. 16). Prior to 2012 our
 516 classifications are less reliable due to the prevalence of VLP tremor and shorter resonant
 517 periods (which makes phase offsets between stations less negligible). The mean and me-
 518 dian amplitudes of Normal events are both about twice as large as those of Reverse events,
 519 though both types of events exhibit variation in amplitude over orders of magnitude (Fig.
 520 S.13). We do not find any appreciable differences in distributions of T or Q between Nor-
 521 mal and Reverse events, and also do not find any appreciably different correlations against
 522 other datasets (such as tilt or lava-lake elevation) between the two types of events (Fig.
 523 S.13).



524 **Figure 13.** Section of the VLP catalog from 2008-2011. Black lines in the plots a and b show 30-day
 525 moving averages over the modes we have labeled as potential conduit-reservoir modes, neglecting outliers
 526 or events from times with no consistent dominant period. In plot e dark green dots indicate average daily
 527 SO_2 , light green lines indicates standard deviations, and the black line is a 30-day moving average. ‘Crater’
 528 indicates where the Halema‘uma‘u crater first formed, ‘SSE’ indicates slow slip events, ‘Int’ indicates doc-
 529 umented summit intrusions, and ‘ERZ’ indicates eruptions along the East-Rift-Zone. Grey bars in all plots
 530 indicate times of long-term ground inflation (Section 2.8).

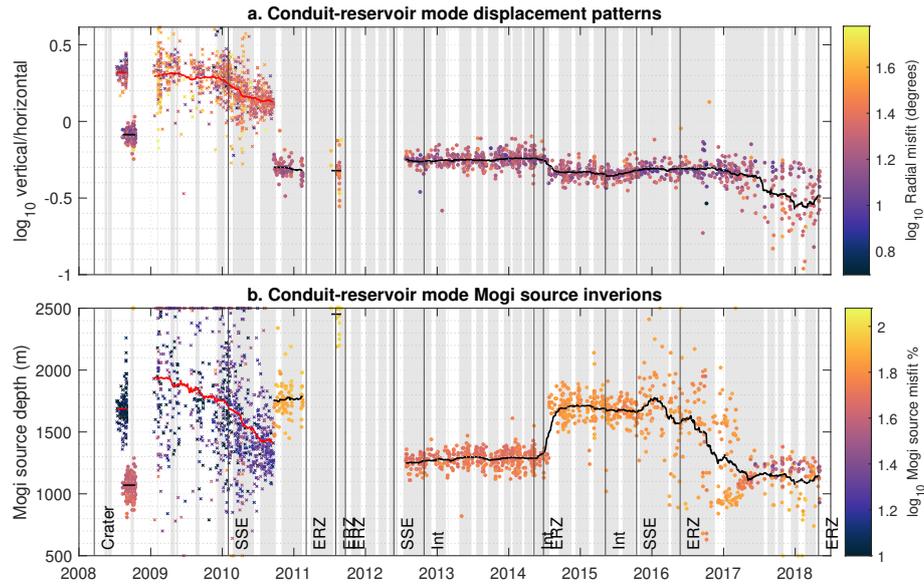
554 3.1.2 Lava-lake sloshing

555 The second category of signals we term ‘lava-lake-sloshing modes’ Dawson *et al.*
 556 [2014]; Liang and Dunham [2020]. These have T of 10-20 s, and are recognizable from
 557 2010-2018 in our catalog (Fig. 13, 14).



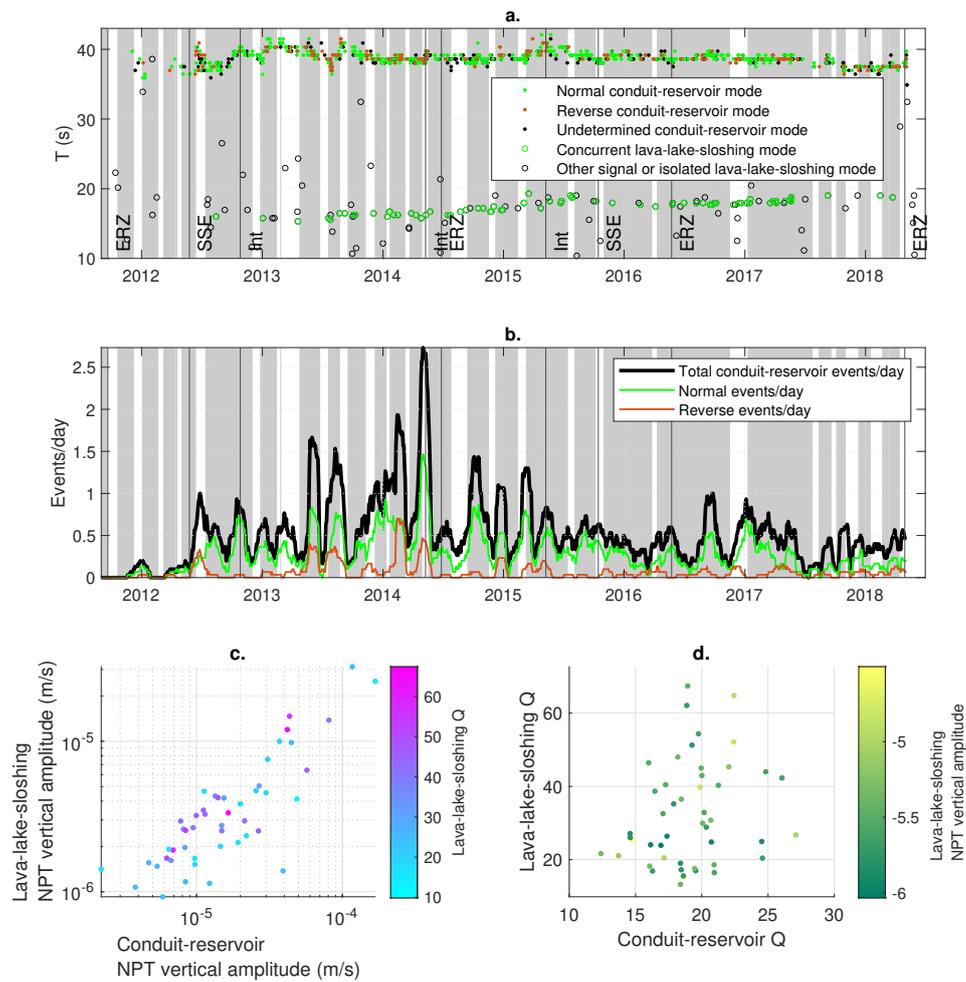
531 **Figure 14.** Section of the VLP catalog highlighting conduit-reservoir and lava-lake-sloshing resonance
 532 from 2012-2018. Black lines in plots a and b show 30-day moving averages, and in plots c and d show 120
 533 day moving averages. ‘SSE’ indicates slow slip events, ‘Int’ indicates documented summit intrusions, and
 534 ‘ERZ’ indicates eruptions along the East-Rift-Zone. Grey bars in all plots indicate times of long-term ground
 535 inflation (Section 2.8).

558 Modeling/inversions for select examples of lava-lake-sloshing events [*Liang and*
 559 *Dunham, 2020*] supports earlier suggestions [*Dawson et al., 2014*] that they are likely
 560 caused by lateral surface gravity wave resonance in the lava-lake (i.e., ‘sloshing’). The
 561 resulting pressure perturbations at the top of the conduit may also force magma flow down
 562 the conduit causing a forced oscillation in the conduit-reservoir system [*Liang and Dun-*
 563 *ham, 2020*].



536 **Figure 15.** Ground displacement patterns and Mogi spherical reservoir source inversions for conduit-
 537 reservoir modes. Dots and black lines indicate events and 120-day moving averages for times with more
 538 than 6 stations available. Crosses and red lines indicate events and 120-day moving averages for times with
 539 only one station available, so ground displacement patterns are poorly constrained and should not be directly
 540 compared to events with more stations. Depths are relative to the caldera floor. ‘Crater’ indicates where
 541 the Halema‘uma‘u crater first formed, ‘SSE’ indicates slow slip events, ‘Int’ indicates documented summit
 542 intrusions, and ‘ERZ’ indicates eruptions along the East-Rift-Zone. Grey bars in all plots indicate times of
 543 long-term ground inflation (Section 2.8).

564 Around 75% of these modes appear alongside Normal conduit-reservoir modes; the
 565 rest appear in isolation (Fig. 11, 12, 16, S.22, S.24). We found no examples occurring
 566 alongside Reverse modes. There are some times where at least two distinct lava-lake-
 567 sloshing modes occur (Fig. 12, S.24); likely representing sloshing in different directions
 568 with an irregular lava-lake geometry [*Liang and Dunham, 2020*]. These do not appear to
 569 be very prevalent in our catalog, though such modes with low signal/noise ratio or very
 570 close period to a larger lava-lake-sloshing mode may have been missed.



544 **Figure 16.** Plot a shows the onset polarity (Normal or Reverse) of conduit-reservoir events, and lava-lake-
 545 sloshing modes that occurred alongside a detected conduit-reservoir event. Plot b shows conduit-reservoir
 546 event density calculated over 30-day windows. We note that event density will vary by orders-of-magnitude
 547 depending upon the event detection thresholds used (Section 2.5), so is most useful for comparing relative
 548 event densities through time. ‘Crater’ indicates where the Halema‘uma‘u crater first formed, ‘SSE’ indi-
 549 cates slow slip events, ‘Int’ indicates documented summit intrusions, and ‘ERZ’ indicates eruptions along
 550 the East-Rift-Zone. Grey bars in plots a and b indicate times of long-term ground inflation (Section 2.8).
 551 Plot c compares amplitudes (from vertical velocity at station NPT) of conduit-reservoir modes with cor-
 552 responding lava-lake-sloshing modes. Plot d compares Q of conduit-reservoir modes with corresponding
 553 lava-lake-sloshing modes.

571 **3.1.3 VLP tremor**

572 We use the term ‘VLP tremor’ to refer to signals with clearly elevated energy in one
 573 or more relatively focused periods, but that are not obviously isolated in time and lack
 574 clear onsets and/or exponential decays. These signals occur throughout the study timespan
 575 (Fig. 11, S.15, S.16, S.17, S.18, S.19, S.21).

576 Many of these signals have the same dominant periods as nearby impulsively-triggered
 577 conduit-reservoir or lava-lake-sloshing modes. We therefore hypothesize that they repre-
 578 sent the same resonant modes with more continuous rather than discrete forcing. Con-
 579 tinuous forcing could occur via superposition of discrete impulses such as rockfalls [*Orr*
 580 *et al.*, 2013], surface explosions/bubble bursts [*Chouet et al.*, 2010; *Richardson and Waite*,
 581 2013], or rock fracture/slip [*Aki et al.*, 1977; *Chouet*, 1996]. Continuous tremor has also
 582 been hypothesized to arise from magma flow through irregular channels [*Julian*, 1994],
 583 bubble-cloud oscillations [*Matoza et al.*, 2010; *Unglert and Jellinek*, 2015], or turbulence
 584 [*Hellweg*, 2000; *Unglert and Jellinek*, 2015].

585 If VLP tremor amplitude is constant our method will not detect it. However, in
 586 this dataset VLP tremor amplitude is almost always variable on timescales ranging from
 587 seconds-minutes, in which case our method detects events corresponding to local maxima.
 588 Q of such signals could be controlled by the forcing time-function rather than damping of
 589 the initial resonance, so may not be sensitive to the same magma system properties as Q
 590 of impulsive-onset decaying resonant signals.

591 **3.1.4 Gliding-frequency VLP signals**

592 We use the term ‘gliding-frequency’ to refer to VLP signals with dominant periods
 593 that change over the duration of a single event (over timescales from seconds to tens of
 594 minutes). These signals are present at various times and with various starting and end-
 595 ing periods throughout the studied timespan (Fig. 12, S.18, S.20). While not designed to
 596 categorize gliding-frequency VLP signals, our method does detect a multitude of them.

597 We are not aware aware of any published analysis of these signals at Kīlauea, though
 598 gliding in higher-frequency tremor has been previously identified [*Unglert and Jellinek*,
 599 2015]. In some cases the gliding-frequency VLP signals appear to start or end at simi-
 600 lar periods to nearby non-gliding conduit-reservoir or lava-lake-sloshing resonances, in-

601 dicating that at least some of the gliding-frequency signals may be related to these other
602 modes.

603 Some gliding-frequency VLP signals may represent rising bubble slugs, which could
604 create a varying oscillation period during ascent and then possibly trigger standard decay-
605 ing conduit-reservoir resonance after bursting at the surface [*James et al.*, 2008; *Chouet*
606 *et al.*, 2010]. Alternately, some gliding-frequency VLP signals may represent examples of
607 either conduit-reservoir or lava-lake-sloshing resonance where magma properties change
608 over the course of the resonance. This could occur if the perturbation that induces reso-
609 nance destabilises some aspect of the shallow magma system, such as by causing collapse
610 of a foam layer in the lava-lake, or by causing release and upward movement of a bubble
611 slug or cloud.

612 **3.2 Timeline of Kīlauea VLP Seismicity**

613 Here we present a brief chronological overview of Kīlauea activity and VLP seis-
614 micity from 2008-2018. We break the timeline into one or two year long time-segments
615 based on where notable changes in VLP seismicity occur.

616 ***3.2.1 January 2008-January 2010: Overlook Crater formation and intermittent*** 617 ***lava-lake***

618 The Overlook Crater first began forming inside the Halema‘uma‘u summit crater in
619 March 2008, following months of elevated SO₂ emissions and seismicity [*Patrick et al.*,
620 2011; *Dawson et al.*, 2014; *Patrick et al.*, 2019b]. Two years of elevated seismicity, long-
621 term ground deflation, and occasional explosive events led to the establishment of a per-
622 sistent lava-lake in early 2010 [*Patrick et al.*, 2011; *Dawson et al.*, 2014; *Patrick et al.*,
623 2019b] (Fig. 13).

624 Our method finds more VLP signals in early 2008 and in 2009 than previous stud-
625 ies [*Dawson et al.*, 2010, 2014], defining a more continuous sequence of VLPs to outline
626 this dynamic early phase of the summit eruption sequence (Fig. 13). Average T increased
627 and decreased significantly multiple times during this interval, from a maximum of around
628 25 s in July 2008 to minima of around 13 s in February and August of 2009. While mea-
629 surements of lava-lake level are limited during this time, the local minima in 2009 corre-

sponds with low reported lava-lake levels and the local maxima around July 2008 corresponds with higher reported lava-lake levels [Patrick *et al.*, 2019b].

Much of the VLP seismicity during this time was tremor (Fig. S.15, S.17), though there were times where discrete events were apparent (Fig. S.14, S.16) [Chouet *et al.*, 2011; Dawson *et al.*, 2014; Liang *et al.*, 2019b]. Q was mostly less than 20.

3.2.2 January 2010-March 2011 Kamoamoia fissure eruption: inflation and lava-lake filling

A more continuous trend of conduit-reservoir events began in November 2009 and continued until the March 2011 Kamoamoia fissure eruption (Fig. 13) [Dawson *et al.*, 2014]. In early 2010 the lava-lake became persistent and filled from an elevation of 820 m to 950 m by early 2011 [Patrick *et al.*, 2019b] (Fig. 13). The previous trend of long-term ground deflation began transitioning to gradual inflation around early 2010, and began inflating more rapidly around November 2010 (Fig. 13).

More distinct VLP events with clear impulsive onsets and decays began occurring during this time, though VLP tremor was also still present (Fig. 11) [Chouet *et al.*, 2011; Dawson *et al.*, 2014]. The discrete events were primarily Normal conduit-reservoir modes. During this time-segment a few likely lava-lake-sloshing modes began to appear alongside some of the Normal conduit-reservoir modes, with T from 10-20 s (Fig. 11, 13).

The general trend of increasing conduit-reservoir T over this time was similar to that previously identified [Dawson *et al.*, 2014], though we find more events in mid-late 2010 that help resolve two pronounced month-long spikes in T ; both are about 2 s above the background trend in T . The onset of the first spike (in March 2010) corresponded to a very subtle shift from ground deflation to inflation, and was followed by a slight rise (by $\sim 20\%$) in average SO_2 emissions. The second spike (in June 2010) corresponded to a pronounced local maxima in ground inflation, local maxima (~ 20 m above the background) in lava-lake elevation, and was followed by a slight decrease (by $\sim 20\%$) in SO_2 emissions. For the remainder of this time-segment, conduit-reservoir mode T was well correlated with both ground inflation and lava-lake elevation. There was a gradual increase in Q starting around August 2010, followed by a rapid drop around February 2011 and leading up to the March 2011 Kamoamoia fissure eruption. This general trend is present in the previous Kīlauea VLP catalog [Dawson *et al.*, 2014], but the lower scatter in Q in our cat-

661 alog reveals that Q was correlated with T , ground inflation, and lava-lake elevation in mid
 662 2010 then becomes anti-correlated with all three datasets by late 2010.

663 Resolution of ground displacement patterns is very limited during this time-segment
 664 due to sparse station coverage. There was a continuous decrease in vertical/horizontal ve-
 665 locity ratios and Mogi source depths from early-mid 2010 (Fig. 15), though the velocity
 666 ratios are likely at least partially influenced by the increasing T , which will cause an ap-
 667 parent increase in horizontal motion due to instrument tilt [Maeda *et al.*, 2011].

668 ***3.2.3 March 2011 Kamoamoā fissure eruption-September 2011 Pu‘u ‘Ō‘ō eruption: 669 multiple East-Rift-Zone eruption and lava-lake draining events***

670 After the March 2011 Kamoamoā fissure eruption, there was a gradual increase in
 671 lava-lake elevation and ground inflation leading up to the August 2011 Pu‘u ‘Ō‘ō erup-
 672 tion, followed by another short stretch of ground inflation and lava-lake refilling before the
 673 September 2011 Pu‘u ‘Ō‘ō eruption (Fig. 13). We do not detect appreciable amounts of
 674 VLP seismicity between the March 2011 Kamoamoā and August 2011 Pu‘u ‘Ō‘ō erup-
 675 tions, despite the lava-lake refilling to pre-eruption levels, though there were a couple of
 676 VLP events that exhibited strong glides in period. There was a cluster of low Q VLP ac-
 677 tivity with T around 20 s between the August and September 2011 Pu‘u ‘Ō‘ō eruptions,
 678 including some events that exhibited strong glides in period (Fig. S.18).

679 ***3.2.4 September 2011 Pu‘u ‘Ō‘ō eruption-October 2012 intrusion: lava-lake filling 680 and reappearance of conduit-reservoir resonance***

681 Until around the time of the May 2012 SSE, conduit reservoir mode resonance had
 682 very low Q , often below our detection threshold (Section 2.3), which contributes to the
 683 apparent sparsity of events (Fig. 14). During this time average lava-lake level increased
 684 from ~930 m to ~960 m, although there was only a very slight net ground inflation. Af-
 685 ter the May 2012 SSE (which also corresponds with a temporary 10-day drop in lava-lake
 686 elevation) average conduit-reservoir mode T , lava-lake elevation, and ground inflation all
 687 decreased until around August, then all continually increased until the October 2012 intru-
 688 sion. Average conduit-reservoir mode Q continually increased following the SSE.

689 Conduit-reservoir seismicity during this time consisted of Normal and Reverse events
 690 (Fig. 11), VLP tremor (Fig. S.19), and gliding-frequency events (Fig. S.20). Analysis of

691 conduit-reservoir mode ground displacement patterns over this time is limited by sparse
 692 station coverage. Lava-lake-sloshing modes were sparse during this time-segment so it is
 693 difficult to determine if any robust trends are present (Fig. 14).

694 ***3.2.5 October 2012 intrusion-June 2014 Pu‘u ‘Ō‘ō eruption: stable lava-lake***

695 Between the October 2012 intrusion and the June 2014 Pu‘u ‘Ō‘ō eruption there
 696 was a long-term ground inflation trend, though average lava-lake level remained fairly con-
 697 stant (Fig. 14). The May 2014 intrusion corresponded to a step ground deflation and drop
 698 in lava-lake elevation. Within this time-segment lava-lake elevation and ground inflation
 699 were generally well correlated (Fig. 18).

700 Conduit-reservoir average event density ranged from 0.2-2.6 events/day during this
 701 time-segment (Fig. 16). Local maxima in event density occurred in May 2013, August
 702 2013, February 2014, and the highest recorded event density in the post-2012 timespan
 703 occurs at the May 2014 intrusion. Conduit-reservoir T was positively correlated with lava-
 704 lake elevation and ground inflation until mid 2013 when the correlation became incon-
 705 sistent, and then negative in the months leading up to the June 2014 eruption (Fig. 18).
 706 Conduit-reservoir Q was positively correlated with T , lava-lake elevation, and ground in-
 707 flation in late 2012, but then was inconsistent for most of the rest of the time-segment and
 708 negatively correlated with T in the months leading up to the June 2014 Pu‘u ‘Ō‘ō erup-
 709 tion. Ground displacement patterns from the conduit-reservoir modes were consistent over
 710 this time-segment (Fig. 15).

711 Lava-lake-sloshing events were sparse until around mid 2013. Average lava-lake-
 712 sloshing T was relatively constant, mostly between 15.5-16.5 s. Q was highly variable
 713 between 6-50, but increased on average over this time-segment (Fig. 14).

714 ***3.2.6 June 2014 Pu‘u ‘Ō‘ō eruption-May 2016 Pu‘u ‘Ō‘ō eruption: changed conduit- 715 reservoir ground displacement patterns***

716 After the June 2014 Pu‘u ‘Ō‘ō eruption there was an abrupt change in conduit-
 717 reservoir mode ground displacement patterns, which then remained stable until around the
 718 October 2015 SSE (Fig. 15). There was fairly steady long-term ground inflation during
 719 this time-segment, with more rapid ground inflation in the months around the May 2015
 720 intrusion, [Patrick *et al.*, 2019b] (Fig. 14). Long-term averaged lava-lake level remained

721 fairly constant through most of this time-segment, with the exception of an overflow in
 722 the month leading up to the May 2015 intrusion, and then a more steady increase between
 723 October-December 2015. The months after the May 2015 intrusion are unique within the
 724 studied timespan for exhibiting a strong anti-correlation between lava-lake elevation and
 725 ground inflation. SO₂ emissions averaged around 5000-6000 t/day from 2014 until the
 726 May 2015 intrusion, then dropped to around 4000 t/day and remained around this level
 727 until increasing in the months leading up to the May 2016 Pu‘u ‘Ō‘ō eruption 14).

728 Conduit-reservoir event density varied from 0.1-1.5 events/day during this time-
 729 segment (Fig. 16). Local maxima in event density occurred during the May 2015 intru-
 730 sion, May 2016 Pu‘u ‘Ō‘ō eruption, and generally near the onset of long-term inflation
 731 periods (for example October 2014, December 2014, and March 2015). Conduit-reservoir
 732 T was remarkably constant around 39 s, until increasing to 41 s in the months leading
 733 up to the May 2015 intrusion, after which it decreased for the remainder of the timespan
 734 (Fig. 14). There was a local minima in T corresponding to the October 2015 SSE. T was
 735 fairly well correlated with lava-lake elevation and ground inflation, except in the months
 736 following the June 2014 eruption (Fig. 18). Conduit-reservoir Q averaged around 25 until
 737 a few months before the May 2015 intrusion, when it dropped to around 18 and remained
 738 stable for the remainder of the time-segment. Q was mostly anti-correlated with T during
 739 this time-segment, and not strongly correlated to lava-lake elevation or ground inflation.

740 Lava-lake-sloshing T increased steadily until a few months after the May 2015 intru-
 741 sion, then decreased until early 2016, then again increased more gradually for the remain-
 742 der of the time-segment (Fig. 14). Lava-lake-sloshing T or Q did not appear to correlate
 743 with any of the other datasets during this time.

744 ***3.2.7 May 2016 Pu‘u ‘Ō‘ō eruption-May 2018 caldera collapse onset: variable*** 745 ***conduit-reservoir ground displacement patterns and climactic eruption pre-*** 746 ***cursors***

747 The months around the May 2016 Pu‘u ‘Ō‘ō eruption heralded a net change in VLP
 748 ground displacement patterns, albeit with significant scatter (Fig. 15). Ground displace-
 749 ment patterns then remained consistent until the May 2018 caldera collapse onset. Long-
 750 term averaged lava-lake elevation increased until late 2016 when small overflows occurred
 751 [Patrick *et al.*, 2019b], then decreased until mid 2017, then remained stable until it be-

752 gan increasing steeply in March 2018 and eventually overflowed on April 26, then began
 753 draining on May 2 [Neal *et al.*, 2019] (Fig. 14). There was consistent long term ground
 754 inflation until mid 2017, then little net inflation or deflation until consistent inflation began
 755 again around March 2018. Lava-lake elevation and ground inflation were mostly correlated
 756 during this time-segment, with the exception of a few months in mid 2017 (Fig. 18). Af-
 757 ter the May 2016 Pu'u 'Ō'ō eruption SO₂ emissions stabilize at around 5000 t/day, and
 758 remained at this level except for drops in early 2017 and late 2017 (when the published
 759 data ends).

760 Conduit-reservoir event density varied from 0-1 events/day during this time-segment
 761 (Fig. 16). decreased in the months following the May 2016 Pu'u 'Ō'ō eruption, exhib-
 762 ited local maxima in September 2016 and January-May 2017, and remained relatively sta-
 763 ble in the year leading up to the 2018 caldera collapse. Conduit-reservoir mode T was
 764 stable around 39 s until October 2017 when it dropped to 37 s; then increased again in
 765 the months leading up to the May 2018 collapse eruptions (Fig. 14). During this time-
 766 segment T was alternately correlated and anti-correlated with lava-lake elevation and ground
 767 inflation (Fig. 18). Conduit-reservoir mode Q remained around 18 until August 2017,
 768 when it became more variable for the remainder of the time-segment. Q was anti-correlated
 769 with T until late 2017, and was alternately correlated and anti-correlated with lava-lake el-
 770 evation and ground inflation.

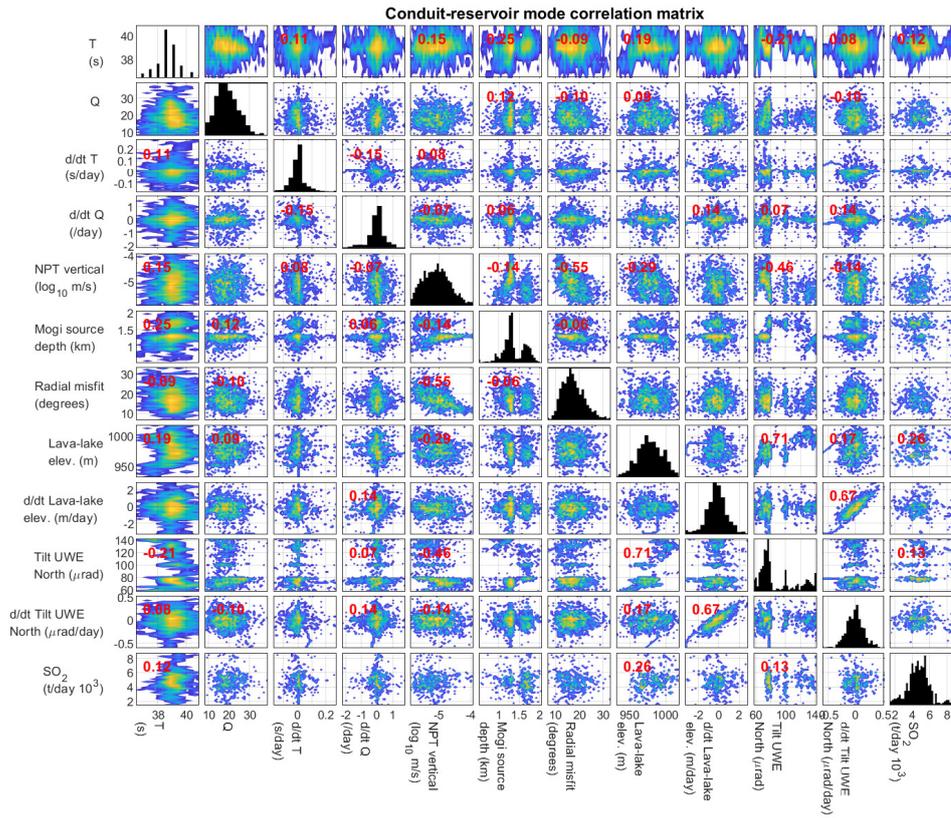
771 Lava-lake-sloshing modes were numerous until around May 2017, then sparse dur-
 772 ing the rest of the time-segment (Fig. 14). Lava-lake-sloshing T increased fairly steadily,
 773 except for a decrease in May 2018. Lava-lake-sloshing Q was highly variable during this
 774 time-segment.

775 3.3 General correlations among datasets

776 Here we analyze correlations between the various geodetic datasets, conduit-reservoir
 777 resonant properties, and lava-lake sloshing properties.

789 3.3.1 Correlations among geodetic datasets

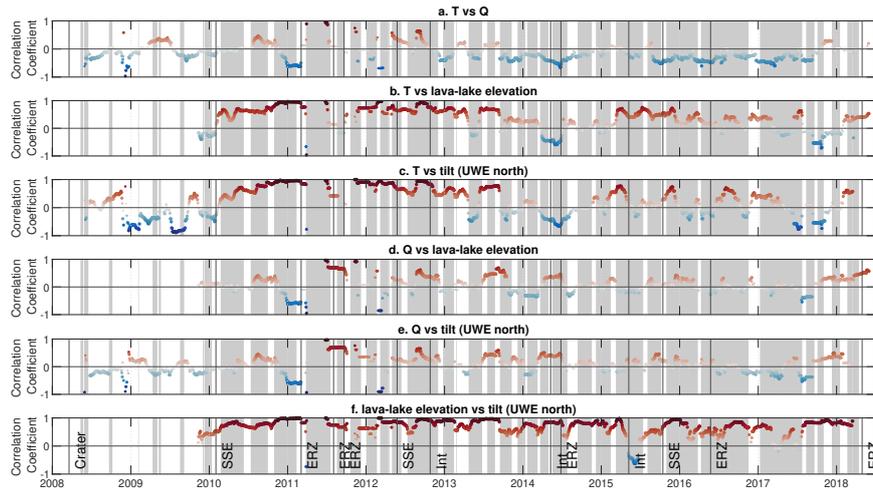
790 Ground surface deformation data from near field tilt-meters and GPS stations indi-
 791 cates the rate of ground inflation/deflation of the Kīlauea summit region. This primarily
 792 reflects pressure in the shallow summit reservoir, but may also be influenced by pressure



778 **Figure 17.** Conduit-reservoir mode correlation matrices from 2012-2018 (see Fig. S.12 for the full 2008-
 779 2018 timespan). Off-diagonal plots are colored by the logarithm of the number of points in a given parameter
 780 bin, and histograms on diagonal plots show the distribution of each parameter. Red numbers are Pearson's
 781 correlation coefficients, only shown for correlations with P-values less than 0.05. 'Lake h' indicates lava-lake
 782 elevation. All time derivatives, notated by 'd/dt', were calculated with a 7-day cutoff-period differentiator
 783 filter (Section 2.8).

793 in the proposed deeper south caldera reservoir [Baker and Amelung, 2012; Anderson et al.,
 794 2015; Anderson and Poland, 2016; Anderson et al., 2019] or along the ERZ [Montagna
 795 and Gonnermann, 2013].

796 Lava-lake elevation is generally positively correlated with ground inflation, particu-
 797 larly on timescales of months or less, as captured by moving correlations (Fig. 18) and
 798 correlations between time derivatives (Fig. 17, S.12). These timescales include the preva-
 799 lent deflation-inflation (DI) events [Patrick et al., 2016a,b; Anderson et al., 2019]. This
 800 correlation implies that lava-lake elevation is analogous to a Pitot tube for the summit
 801 magma reservoir, where the exact relation between lava-lake level and reservoir pres-



784 **Figure 18.** Conduit-reservoir mode Pearson's correlation coefficients calculated over moving 90-day win-
 785 dows. Windows with less than 4 data points were excluded. Larger dots indicate lower p-values; the largest
 786 dot size (encompassing ~30-70% of the values in each plot) corresponds to p-values less than 0.05. 'SSE'
 787 indicates slow slip events, 'Int' indicates documented summit intrusions, and 'ERZ' indicates eruptions along
 788 the East-Rift-Zone. Grey bars in the all plots indicate times of long-term ground inflation (Section 2.8).

802 sure will depend on reservoir stiffness and on the magma density profile [Patrick *et al.*,
 803 2015; Anderson *et al.*, 2015, 2019]. However, there are isolated times where this corre-
 804 lation breaks down (Fig. 14, 18). Additionally, the correlation between ground inflation
 805 and lava-lake elevation over the whole timespan exhibits strong scatter (Fig. 17, S.12), in-
 806 dicated that the relation between ground inflation and lava-lake elevation is not constant
 807 over time. This is partly caused by gradual long-term changes, such as in early 2017 when
 808 ground inflation and lava-lake elevation are positively correlated on day-week timescales
 809 but long-term lava-lake level remains constant despite long-term ground inflation (Fig.
 810 14). There are also abrupt events that change the relation between ground inflation and
 811 lava-lake elevation, such as the May 2015 intrusion (Fig. 14).

812 There was typically an increase in lava-lake elevation and ground inflation over
 813 days-months leading up to ERZ eruptions, followed by an abrupt ground deflation and de-
 814 crease in lava-lake elevation (Fig. 14). The exception was the June 2014 Pu'u 'Ō'ō erup-
 815 tion, around which there were no significant changes in lava-lake elevation or ground in-
 816 flation. Among ERZ eruptions, SO₂ data is only available around the June 2014 and May

2016 Pu‘u ‘Ō‘ō eruptions, but there did appear to be an increase in SO₂ emissions by approximately a factor of 2 over the months leading up to the May 2016 Pu‘u ‘Ō‘ō eruption (Fig. 14).

There was also typically an increase in lava-lake elevation and ground inflation over days-months leading up to intrusions, followed by an abrupt ground deflation and decrease in lava-lake elevation (Fig. 14). However, the ground deflation following intrusions was much less pronounced than the drops in lava-lake elevation. SO₂ data is only available around the May 2014 and May 2015 intrusions, but there did appear to be a decrease in SO₂ emissions by approximately a factor of 1.5 in the months following the May 2015 intrusion (Fig. 14).

Some slow-slip events (SSEs), where aseismic slip on a fault occurs over timescales of hours-days [Schwartz and Rokosky, 2007], have been linked to magmatic activity such as diking events at Kīlauea [Brooks *et al.*, 2008; Montgomery-brown *et al.*, 2015]. The January 2010 and October 2015 SSE do not appear to correspond to changes in lava-lake elevation, ground inflation, or SO₂ emissions (Fig. 13). The May 2012 SSE does correspond to a several day drop in lava-lake elevation and ground inflation (Fig. 14).

3.3.2 Conduit-reservoir resonance correlations

During most of the timespan conduit-reservoir mode T and Q exhibit a weak negative correlation, with an overall Pearson’s correlation coefficient of -0.06 but local correlation coefficients often around -0.7 (Fig. 17, 18, S.12). There are isolated times where T and Q are positively correlated, such as in mid 2010 (correlation coefficient near 1) and mid 2012 (correlation coefficient around 0.7) (Fig. 13, 14, 18).

Conduit-reservoir mode T is positively correlated with lava-lake elevation during most of the timespan, with correlation coefficients mostly between 0.3 and 1 (Fig. 18), and a weak overall correlation coefficient of 0.11 (Fig. 17, S.12). However, there are some times with negative local correlations, such as around the 2014 Pu‘u ‘Ō‘ō eruption (correlation coefficient around -0.6), and in late 2017 (correlation coefficient around -0.7). The correlation between T and ground inflation (i.e., tilt) exhibits a similar trend to the correlation between T and lava-lake elevation after the arrival of a persistent lava-lake in late 2009, and exhibits a variable but mostly negative trend prior to this (Fig. 17, 18, S.12).

847 Conduit-reservoir T is positively correlated with event amplitude, even when considering
 848 only vertical velocity (which should not be sensitive to instrument tilt) (Fig. 17, S.12).

849 We find increases in both conduit-reservoir event density and T around the docu-
 850 mented October 2012 and May 2015 intrusions. There is no obvious change in Q cor-
 851 responding to either intrusion, though the the correlation between T and Q does change
 852 from positive to negative at the October 2012 intrusion (Fig. 7, 18). Neither intrusion ap-
 853 pears to correspond to changes in ground displacement patterns (Fig. 15).

854 ERZ eruptions for which we detect conduit-reservoir modes both before and after
 855 the events (i.e., the June 2014 and May 2016 Pu'u 'Ō'ō eruptions) don't obviously relate
 856 to changes in conduit-reservoir mode T or Q . However, sharp changes in the correlations
 857 between T and Q , T and lava-lake elevation/tilt, and Q and lava-lake elevation/tilt occur
 858 alongside the June 2014 eruption, and more subtle changes in these correlations may also
 859 be present alongside the May 2016 eruption (Fig. 7, 18). There are pronounced changes
 860 in ground displacement patterns following both eruptions that are readily apparent in the
 861 time-series of Mogi source inversions and vertical/horizontal velocity ratios (Fig. 15).
 862 There is an apparent increase in Mogi source centroid depth following June 2014, and
 863 then an apparent decrease following May 2016 (though with more scatter in the inverted
 864 depths around this time).

865 **3.3.3 Lava-lake-sloshing correlations**

866 Due to the smaller amount of lava-lake-sloshing modes present it is more difficult
 867 to determine whether lava-lake-sloshing T or Q are correlated with other datasets. Long-
 868 term average lava-lake-sloshing T increased over most of the timespan, with the exception
 869 of 2012 when lava-lake sloshing events were sparse and exhibited large scatter in T . The
 870 long-term increase in T roughly corresponds to an observed long-term increase in lava-
 871 lake surface area [Patrick *et al.*, 2019b]. There is appreciable scatter (of about 3 s) in T on
 872 timescales of months or less; though much of this appears to be due to a small number of
 873 outlier events. Lava-lake-sloshing Q exhibits large scatter over most of the timespan, with
 874 the exception of 2012 when Q was generally less than 20, and 2015 when Q was generally
 875 between 10 and 30 (corresponding to a local maxima in lava-lake-sloshing T).

876 There is a roughly linear relation between conduit-reservoir mode amplitude and
 877 lava-lake-sloshing mode amplitude, though with an appreciable amount of scatter (Fig.

878 16). Lava-lake-sloshing Q does not appear to be correlated with conduit-reservoir mode Q
 879 (Fig. 16).

880 4 Discussion

881 Very-Long-Period seismic events in our new catalog provide an outstanding tool
 882 both to document the progression of a long-lived (10 year) open vent eruptive episode at
 883 Kīlauea Volcano and probe shallow magma plumbing system geometry and magma prop-
 884 erties through time. In the following discussion we provide a conceptual modeling frame-
 885 work for understanding the physical origin of observations documented in previous sec-
 886 tions, based largely on previously published work. We leave a detailed inversion of these
 887 events and true uncertainty quantification for future studies.

888 4.1 Interpreting changes in conduit-reservoir resonance

889 The conduit-reservoir mode is the most common and also most variable class of
 890 VLP events in our catalog. The reduced conduit-reservoir mode model of *Liang et al.*
 891 [2019a] provides estimates of T and Q assuming a cylindrical conduit and isothermal con-
 892 ditions, neglecting inertia and viscous drag in the overlying lava-lake and compressibility
 893 of magma in the conduit. The inviscid conduit-reservoir resonance period is [*Liang et al.*,
 894 2019a]:

$$894 T_0 = 2\pi \sqrt{\frac{L_c \bar{\rho}_c}{\Delta \rho_c g \sin \alpha + A_c C_t^{-1}}}. \quad (9)$$

895 where L_c is conduit length, $\bar{\rho}_c$ is average magma density in the conduit, $\Delta \rho_c$ is density
 896 difference between the bottom and top of the conduit, α is conduit dip angle, A_c is con-
 897 duit cross-sectional area, and C_t is total reservoir storativity:

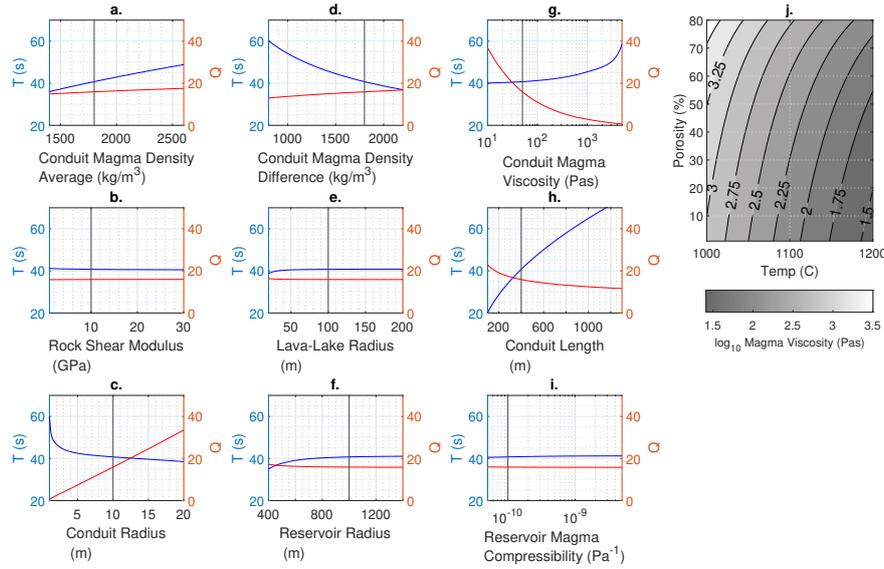
$$897 C_t = (\kappa_m + \kappa_{res})V \quad (10)$$

898 where κ_m and κ_{res} are magma and reservoir compressibility ($\frac{3}{4G}$ for a spherical reservoir
 899 [*McTigue*, 1987]) and V is reservoir volume. With viscous damping included, T and Q
 900 depend upon T_0 as well as a momentum diffusion timescale:

$$900 \tau_{visc} = \frac{R_c^2 \bar{\rho}_c}{\mu_c}, \quad (11)$$

901 where R_c is conduit radius and μ_c is average magma viscosity (Fig. 19).

902 We can use this model to estimate the parameter variation that could cause observed
 903 variation in T and Q (Fig. 19). We focus on short timescales (days-weeks), for which it is



902 **Figure 19.** Plots a-i show predicted variation in T and Q due to varying each model parameter in isolation
 903 in the reduced conduit-reservoir resonance model of *Liang et al.* [2019a] (Eq. 9-11), assuming a spherical
 904 reservoir geometry. Black lines indicate the default value used for each parameter. Plot j shows apparent
 905 magma viscosity as a function of temperature and porosity (Section 4.1).

908 reasonable to assume that the geometry of the system remains constant. Parameters most
 909 likely to cause variation in T and Q on short timescales are properties of the multiphase
 910 magma contained within the conduit-reservoir system: average magma density, density
 911 difference, and apparent viscosity (magma compressibility probably has a comparatively
 912 minimal influence, see Fig. 19). Figure 19 shows that of these magma properties, T is
 913 most sensitive to average magma density and magma density difference. Variation in ei-
 914 ther parameter of up to $\sim 500 \text{ kg/m}^3$ would be required to explain the observed short-term
 915 variability in T of up to $\sim 6 \text{ s}$ (Fig. 14). Q is most sensitive to magma viscosity (Fig. 19).
 916 Variation in magma viscosity of up to an order of magnitude would be required to explain
 917 the observed short-term variability in Q of up to an order of magnitude (Fig. 14, 19). It is
 918 interesting to note that none of the model parameters changing in isolation would produce
 919 a positive correlation between T and Q , as is sometimes observed (Fig. 18).

920 Variation in apparent magma viscosity could be partly due to changing bubble num-
 921 ber and/or size distributions [*Manga and Loewenberg, 2001; Pal, 2003; Llewellyn and*
 922 *Manga, 2005; Huber et al., 2014*] (Fig. 19). Basaltic melt viscosity will also change slightly

923 with dissolved volatile contents, and strongly with temperature [*Giordano et al.*, 2008]
 924 (Fig. 19). We show how magma viscosity might vary in response to temperatures and
 925 porosity in Figure 19 (plot j). Melt viscosity is obtained from the model of *Giordano et al.*
 926 [2008], using the average Kīlauea glass composition from *Edmonds et al.* [2013] and dis-
 927 solved H₂O and CO₂ contents from the solubility model of *Burgisser et al.* [2008], assum-
 928 ing a pressure of 19 MPa. Apparent magma (melt + bubbles) viscosity is then obtained by
 929 using the low capillary-number model from *Llewellyn and Manga* [2005]):

$$\mu = (1 - \phi)^{-1} \mu_l \quad (12)$$

930 where μ_l is melt viscosity and ϕ is porosity.

931 Changes in melt temperature could arise due to due to changes in convective regimes
 932 [*Jones et al.*, 2006; *Witham and Llewellyn*, 2006; *Harris*, 2008] or influx/recharge of deeper
 933 magma. For example, convection extending from the lava-lake surface though the conduit
 934 might result in lower average magma temperatures in the conduit than if there are separate
 935 convective cells in the lava-lake and conduit [*Patrick et al.*, 2016b].

936 While the model of *Liang et al.* [2019a] (outlined in Eq. 9-11) provides an excellent
 937 starting point for interpreting changes in T and Q , it involves a number of simplifications
 938 that would need to be improved to allow for a more detailed analysis of Kīlauea VLP
 939 seismicity. Chiefly, incorporating a background state model for magma density profiles,
 940 whether from a simple magmastic case [*Karlstrom et al.*, 2016] or considering exchange
 941 flow [*Fowler and Robinson*, 2018], would be necessary to assess how average magma den-
 942 sity and density gradient in the conduit vary with lava-lake elevation and/or reservoir pres-
 943 sure. This relation likely plays a role in the observed correlations between T , lava-lake
 944 elevation, and ground inflation. Inertia in the lava-lake and variations in conduit and lake
 945 geometry with depth, which are neglected in the model of *Liang et al.* [2019a] (Eq. 9-
 946 11), may affect both T and Q and also contribute to the observed correlation between T
 947 and lava-lake elevation. Lastly, a more detailed treatment of damping that includes the ef-
 948 fects of bubbles on viscosity [*Manga and Loewenberg*, 2001; *Llewellyn and Manga*, 2005;
 949 *Gonnermann and Manga*, 2013; *Huber et al.*, 2014], bubble growth and resorption [*Karl-
 950 strom et al.*, 2016], and viscous drag in the lava-lake would be necessary to accurately in-
 951 terpret changes in Q .

4.2 Interpreting changes in lava-lake-sloshing

The lava-lake-sloshing mode at Halema'uma'u has been modeled as incompressible surface gravity wave resonance in a cylindrical or wedge-shaped tank [Dawson *et al.*, 2014; Liang and Dunham, 2020]. General studies of incompressible fluid sloshing in various tank geometries indicate that T and Q depend on fluid density and viscosity, tank width, and tank depth (in the case of shallow tanks) [Bauer, 1981; Ibrahim, 2005]. The period for the fundamental sloshing mode of incompressible fluid in a cylindrical tank is given by [Ibrahim, 2005]:

$$T = \frac{2\pi}{\sqrt{1.841 \frac{g}{R_L} \tanh\left(1.841 \frac{h_L}{R_L}\right)}} \quad (13)$$

where R_L is lake radius and h_L is lake depth.

Due to the presence of exsolved volatiles and a solidified surface crust [Karlstrom and Manga, 2006], magma in the Halema'uma'u lava-lake will generally be compressible and stratified [Carbone and Poland, 2012; Carbone *et al.*, 2013; Patrick *et al.*, 2016a; Poland and Carbone, 2016]. Previous inversions [Liang and Dunham, 2020] suggest that the lava-lake sloshing drives magma in and out of the conduit/reservoir, so viscous dissipation from both the conduit and the lava-lake walls needs to be considered. The degree of coupling between lateral fluid motions in the lake and vertical fluid motions in the conduit will depend on the offset of the top of the conduit along the lava-lake sloshing axis, and thus on the direction of lava-lake sloshing [Liang and Dunham, 2020]. The solid crust on the lava-lake surface is likely not static during VLP events, as indicated by videos of rockfall-triggered lava-lake-sloshing events where the crust sometimes disintegrates/overturns following event onsets [Orr *et al.*, 2013; Patrick *et al.*, 2014, 2016a; USGS]. Thus a quantitative interpretation of T and Q for lava-lake-sloshing modes would require modeling that can account for all of these factors, self-consistently coupled to the conduit-reservoir resonator. However, we can still gain some qualitative insights into the lava-lake-sloshing modes with isolated tank models.

If we focus on short timescales (days-months), we can assume that the crater geometry is constant. Lava-lake-sloshing T exhibits variability within ~ 3 s on short timescales (Fig. 14). Part this may be due to sloshing along different axes of the lava-lake with different diameters (Eq. 13). A correlation between lava-lake elevation and T would also be expected if lava-lake walls are inward dipping so that diameter decreases with depth. Such

982 a correlation does not obviously appear in our catalog (Fig. 14), though this may be due
983 to the irregular crater geometry ([Patrick *et al.*, 2019b]).

984 Lava-lake-sloshing Q exhibits order-of magnitude variation on short timescales (days-
985 months) (Fig. 14). We can rule out lava-lake elevation as a sole cause of this variation in
986 Q by noting that many events occurring at similar lava-lake elevations have very different
987 values of Q (Fig. 14). Thus some combination of variation in lava-lake elevation, magma
988 properties, and sloshing direction are likely responsible for observed short-timescale varia-
989 tion in Q .

990 The lack of observed correlation between Q of conduit-reservoir modes and Q of
991 lava-lake-sloshing modes (Fig. 16) suggests that magma properties in the lava-lake and
992 conduit are largely decoupled. Gas volume fraction increases non-linearly as a magma
993 rises [Gonnermann and Manga, 2009] while solubility also decreases [Iacono-Marziano
994 *et al.*, 2012], and the presence of a semi-solid lava-lake crust traps bubbles in the near sur-
995 face. So it is likely that porosity in the lava-lake will be much higher on average than in
996 the conduit, consistent with inferences from gravity [Carbone and Poland, 2012; Carbone
997 *et al.*, 2013; Poland and Carbone, 2016]. If the discrepancy in viscosities between the
998 conduit and lava-lake also requires appreciably different average melt temperatures, this
999 would additionally suggest separate convective cells in the lava-lake and conduit [Patrick
1000 *et al.*, 2016b].

1001 If all lava-lake-sloshing modes had the same forcing (for example rockfall) loca-
1002 tion and satisfied appropriate small-amplitude assumptions, we would expect a linear re-
1003 lationship between lava-lake-sloshing amplitude and conduit-reservoir mode amplitude.
1004 A roughly linear relationship is observed, though with appreciable scatter (Fig. 16). This
1005 scatter could be partly explained by variable forcing location, which might affect the am-
1006 plitude of lava-lake sloshing induced by a given pressure perturbation, the coupling of
1007 lava-lake sloshing to ground displacements [Liang and Dunham, 2020], and/or the cou-
1008 pling of the pressure perturbation at the lava-lake surface to pressure at the top of the con-
1009 duit, which controls conduit-reservoir mode amplitude [Liang *et al.*, 2019a,b].

1010 **5 Conclusions**

1011 We have presented a workflow for using wavelet transforms to both detect and cat-
1012 egorize VLP seismic signals. These methods effectively detect multiple distinct spectral

1013 peaks in impulsive events and provide robust estimates of quality factors. These methods
1014 do not rely upon any training data, are fast to implement, and are readily transferable. The
1015 ability to robustly detect new types of resonant signals, in a fully automated and compu-
1016 tationally efficient manner, makes our method potentially useful for near-real-time volcano
1017 monitoring.

1018 We then used these methods to generate a catalog of ~3000 VLP events that oc-
1019 curred between 2008-2018 during a prolonged open vent eruptive episode at Kīlauea Vol-
1020 cano, Hawaii USA. This catalog expands upon earlier VLP catalogs by characterizing
1021 more types of signals and refined estimates of quality factors, revealing new a rich and
1022 structured timeseries of events that documents changes to the shallow magma plumbing
1023 structures and to multiphase magma properties.

1024 We characterize changes in period, quality factor, and ground displacement pat-
1025 terns over timescales ranging from hours to decades for the ‘conduit-reservoir’ oscillation,
1026 which is prevalent over most of this timespan and represents the fundamental resonant
1027 mode of the shallow magma plumbing system. These likely indicate changes in magma
1028 properties such as density and viscosity in the conduit, and/or changes in magma plumb-
1029 ing system geometry over the course of the eruptive episode. Auxiliary geophysical data
1030 such as tilt, lava lake elevation, and SO₂ emissions corroborate these inferences and help
1031 place the conduit-reservoir resonant mode amongst a rich suite of existing data available
1032 to understand the 2008-2018 eruptive episode.

1033 We also characterize a trend of secondary ‘lava-lake-sloshing’ resonant signals be-
1034 tween 2010 and 2018. These exhibit a relatively consistent increase in period over time,
1035 but wide variability in quality factors. This variability likely indicates changes in lava-lake
1036 geometry, magma density, and magma viscosity. There is no strong correlation between
1037 lava-lake-sloshing and conduit-reservoir mode quality factors, suggesting some decoupling
1038 between magma properties in the conduit and lava-lake. We do not attempt to co-invert
1039 VLP modes with other data, but see this as a rich opportunity for future work on an ex-
1040 ceptionally well documented eruptive episode at Kīlauea volcano.

1041 **A: Synthetic Waveform Tests**

1042 We construct synthetic seismograms to test the resonant signal detection and clas-
1043 sification methods described in the methods section. Displacements are calculated from

1044 an isotropic point source in an elastic half space model [Aki and Richards, 1993], with the
 1045 source located 1 km beneath the Halema‘uma‘u vent. The synthetic source-time functions
 1046 consist of combinations of step displacements and exponentially decaying sinusoids with
 1047 impulsive onsets. We apply a sinusoidal taper to the signal onsets to prevent sharp discon-
 1048 tinuities and create signals with continuous first derivatives (Fig. S.1). The sinusoid used
 1049 as a taper has the same period as the signal, amplitude equal to the initial signal ampli-
 1050 tude divided by $\sqrt{2}$, and is joined at the location where the derivative and position of the
 1051 taper match those of the signal. Where step displacements are also added, we taper the
 1052 step displacement over the same wavelength used to taper oscillation onsets (Fig. S.2). We
 1053 then add white noise from a standard normal distribution, scaled to various fractions of
 1054 the signal amplitude as listed in each test figure. We then calculate displacements and tilts
 1055 at each station location using the point source Green’s functions, and convolve these with
 1056 the instrument responses [Maeda *et al.*, 2011; Liang *et al.*, 2019b].

1057 **Acronyms**

1058 **AR** Auto-Regressive
 1059 **CWT** Continuous Wavelet Transform
 1060 **DFT** Discrete Fourier Transform
 1061 **ERZ** East Rift Zone
 1062 **FIR** Finite Impulse Response
 1063 **FWHM** Full Width at Half Maximum
 1064 **GPS** Global Positioning
 1065 **HVO** Hawaiian Volcano Observatory
 1066 **IIR** Infinite Impulse Response
 1067 **LTA** Long Term Average
 1068 **SSE** Slow Slip Event
 1069 **STA** Short Term Average
 1070 **STFT** Short Time Fourier Transform
 1071 **UV** Ultra-Violet
 1072 **VLP** Very-Long-Period

1073 **Notation**1074 T period1075 Q quality factor1076 f frequency1077 ω angular frequency1078 u measured ground surface displacement1079 w modeled ground surface displacement1080 \dot{u} measured ground surface velocity1081 \dot{w} modeled ground surface velocity1082 **Acknowledgments**

1083 Additional figures S.1-S.24 are included in the supplement. The Kīlauea VLP seismic-
 1084 ity catalog is available at (*included as a spreadsheet with this submission, and will also be*
 1085 *uploaded to a data repository consistent with the Enabling FAIR data Project guidelines*
 1086 *prior to publication*). Codes used to make and analyze the VLP catalog are available at
 1087 `crozierjosh1@bitbucket.org/crozierjosh1/vlp-seismicity-catalog-codes.git`, and the authors
 1088 will provide updated versions and/or assistance upon request.

1089 Seismic data from 2008-2011 was obtained from the USGS, subsequent seismic data
 1090 is publicly available from IRIS. GPS data is publicly available from UNAVCO. Tilt-meter
 1091 data is available at *Johanson [2020] (this citation references a DOI that has been reserved*
 1092 *for a planned data release, and will be updated prior to publication)*. Lava-lake elevation
 1093 data was obtained from the USGS, and is published up to 2018 in *Patrick et al. [2019b]*.
 1094 SO₂ data from 2007-2010 is available at *Elias and Sutton [2012]*. SO₂ emission from
 1095 2014-2017 is available at *Elias et al. [2018]*. The VLP seismicity catalog extended from
 1096 the methods of *Dawson et al. [2014]* was obtained from the USGS.

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Supporting Information for “Very-long-period seismicity over the 2008-2018 eruption of Kīlauea Volcano”

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Contents

Figures S.1 to S.24, arranged in sections 1: Synthetic waveform test figures, 2: Catalog figures, and 3: Example event figures

Additional Supporting Information (Files uploaded separately)

1. *Kilauea_2008-2018_resonant_signal_catalog_presented.csv*

A version of our event catalog thresholded to include 3209 events, as presented in the text. The first row contains descriptions of each variable, and the second row contains the names of each variable.

2. *Kilauea_2008-2018_resonant_signal_catalog_full.csv*

A version of our event catalog thresholded to include 33084 events. The thresholds used in this version are: $STA/LTA > 2$, standard deviations above the LTA > 1 , $Q > 4$, and mean phase deviation < 0.25 radians. The first row contains descriptions of each variable, and the second row contains the names of each variable.

1 Synthetic waveform test figures

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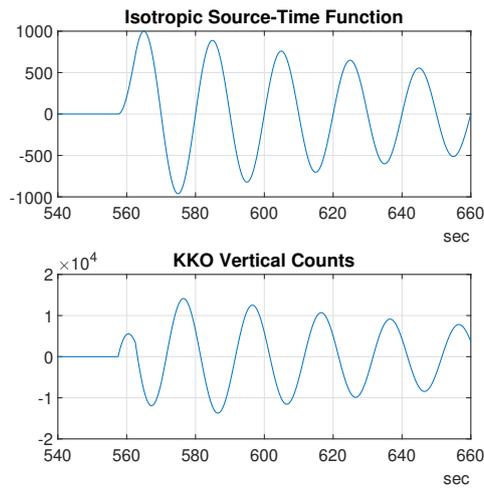


Figure S.1. Example synthetic source-time function and corresponding synthetic seismogram (which has been convolved with the elastic Green’s functions and instrument response), zoomed in around the signal onset to show the tapers used (see appendix). This source-time function is for an impulsive onset oscillation with $T = 20$ s, $Q = 20$, and no added noise.

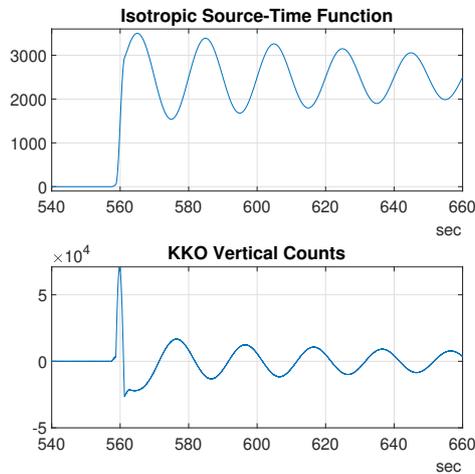


Figure S.2. Example synthetic source-time function and corresponding synthetic seismogram (which has been convolved with the elastic Green’s functions and instrument response), zoomed in around the signal onset to show the tapers used (see appendix). This source-time function is for an impulsive onset oscillation with $T = 20$ s, $Q = 20$, an added step displacement, and no added noise.

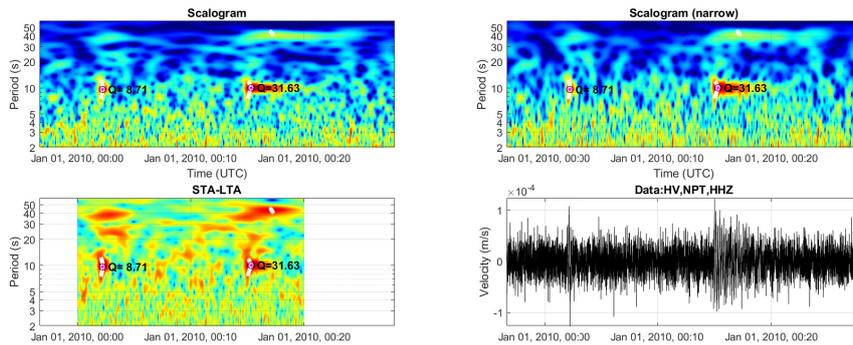


Figure S.3. Example scalograms and detected resonant signals from a synthetic seismogram consisting of four resonant signals with [start time, T , Q] = [00:05, 40, 6], [00:05, 10, 6], [00:15, 40, 40], [00:15, 40, 40], plus white noise from a standard normal distribution scaled by 5.0% of the signal amplitude. At this noise level only two of the signals are found at the detection thresholds used, and the quality factor estimates are less accurate (off by $\sim 25\%$).

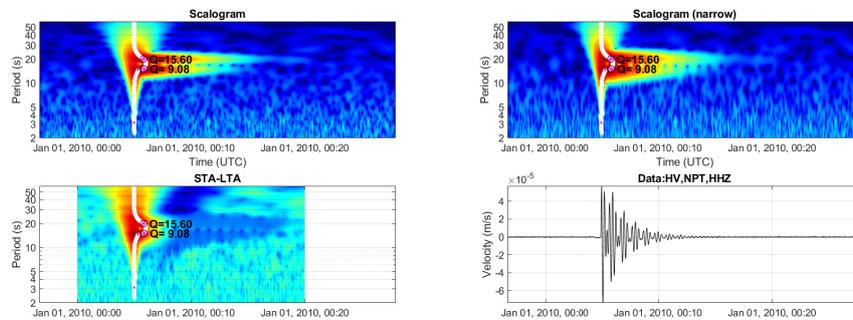


Figure S.4. Example scalograms and detected resonant signals from a synthetic seismogram consisting of two resonant signals with [start time, T , Q] = [00:05, 20, 20], [00:05, 15, 20], plus white noise from a standard normal distribution scaled by 0.1% of the signal amplitude. In this case the spectral proximity of the two signals means that wavelets at the period of one signal are influenced by the other signal, which causes both quality factors to be under-estimated (by 22-54%).

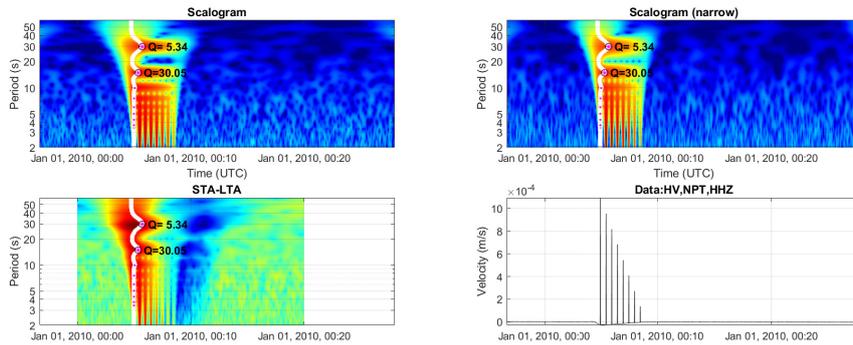


Figure S.5. Example scalograms and detected resonant signals from a synthetic seismogram consisting of eight step displacements (velocity spikes) spaced 30 s apart, plus white noise from a standard normal distribution scaled by 1.0% of the signal amplitude. The closely spaced spikes create a Dirac comb effect, where the spectrum would indicate apparent resonances at 15 s, 7.5 s, 3.25 s, and etc. The temporal resolution of our narrow ($\beta=20$) wavelet, which is used for calculating Q , is high enough that apparent resonances with T less than 15 s are not picked.

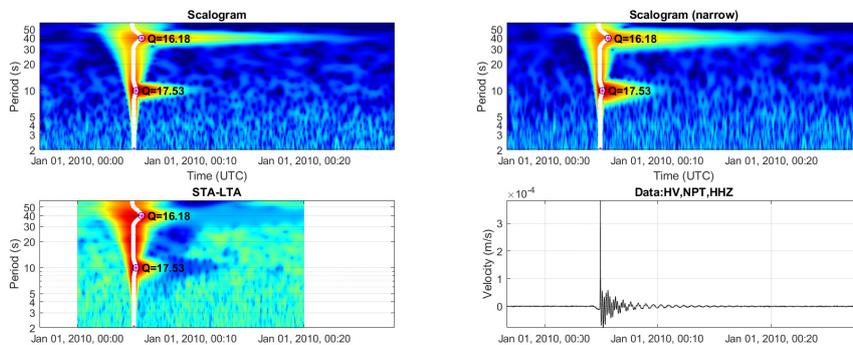


Figure S.6. Example scalograms and detected resonant signals from a synthetic seismogram consisting of a large step displacement (velocity spike) at time 00:05 plus two resonant signals with [start time, T , Q] = [00:05, 40, 20] and [00:05, 10, 20] plus white noise from a standard normal distribution scaled by 0.1% of the signal amplitude. The presence of the step function decreases the estimated quality factors by 12-19% due to the increased energy at the start of the signals, but otherwise does not appreciably impact the results.

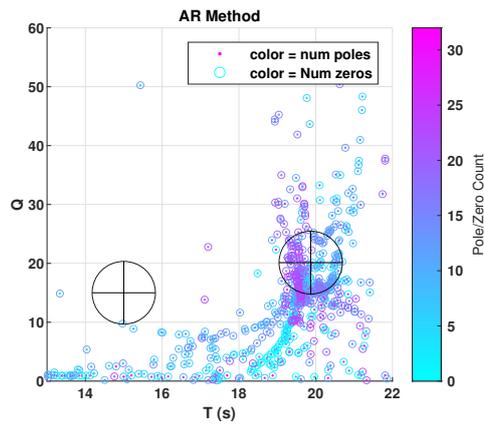


Figure S.7. Example ‘Sompi’ AR method for estimating T and Q applied to a synthetic seismogram. Code used from Lesage 2009. In this case the method was applied to a data window from 10-200 s following the onset of a 20 s oscillation with $Q = 20$ and a smaller (by a factor of 4) 15 s oscillation with $Q = 15$ (indicated by black crosses/circles). Results from filters with 4-32 poles and 0-32 zeros are shown to test a wide parameter space; for practical use narrower ranges would likely be used. A cluster near the actual T and Q of the 20 s oscillation does occur, though mean T and Q values within this cluster are offset from the correct value and exhibit significant scatter. No cluster occurs near the smaller 15 s oscillation, so it would be missed entirely by this AR method.

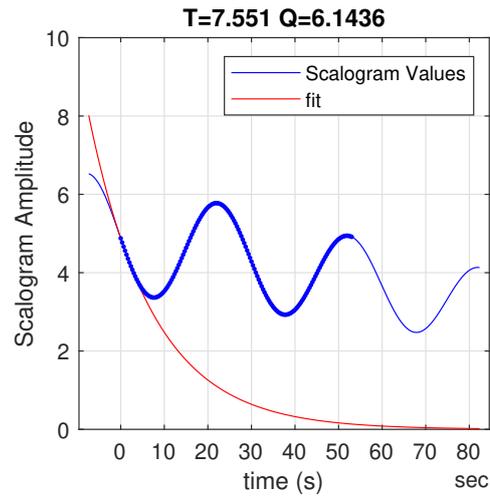


Figure S.8. Example estimation of Q by scalogram exponential fit. Applied to synthetic seismograms consisting of a series of tapered step displacements (velocity spikes) spaced 30 s apart, plus white noise from a standard normal distribution scaled by 0.1% of the signal amplitude. The closely spaced spikes create a Dirac comb effect, where the frequency spectrum would indicate apparent resonances at 15 s, 7.5 s, 3.25 s, and etc. The time resolution of the $\beta=20$ wavelet we use for calculating Q is sufficient to distinguish gaps in this apparent 7.5 s resonance, so our fit avoids overestimating Q as a standard least-squares exponential regression would.

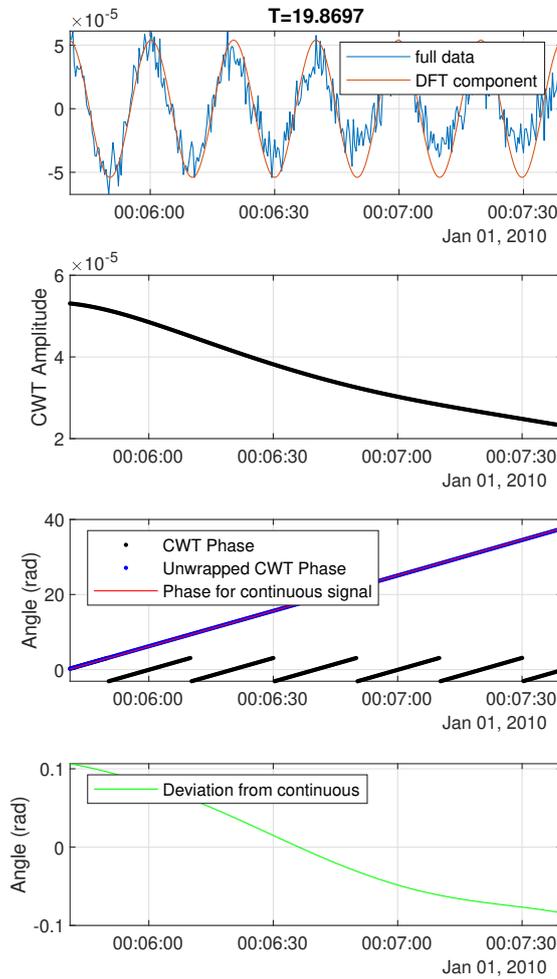


Figure S.9. Example phase continuity from a synthetic seismogram consisting of a resonant signal with $T=20$ s and $Q=20$, plus white noise from a standard normal distribution scaled by 0.1% of the signal amplitude. In this case the phase deviation is small (mean of around 0.05 radians), correctly indicating that this is likely a continuous oscillation.

2 Catalog figures

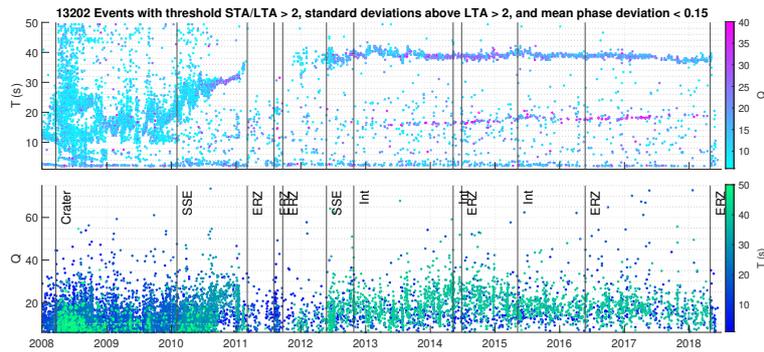


Figure S.10. Resonant signal catalog from 2008-2018 with less strict event detection thresholds than presented in the main text. ‘Crater’ indicates where the Halema‘uma‘u crater first formed, ‘SSE’ indicates slow slip events, ‘Int’ indicates documented summit intrusions, and ‘ERZ’ indicates eruptions along the East-Rift-Zone.

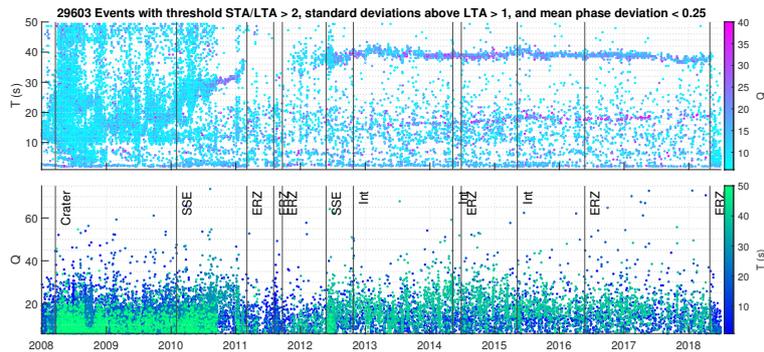


Figure S.11. Resonant signal catalog from 2008-2018 with less strict event detection thresholds than presented in the main text. ‘Crater’ indicates where the Halema‘uma‘u crater first formed, ‘SSE’ indicates slow slip events, ‘Int’ indicates documented summit intrusions, and ‘ERZ’ indicates eruptions along the East-Rift-Zone.

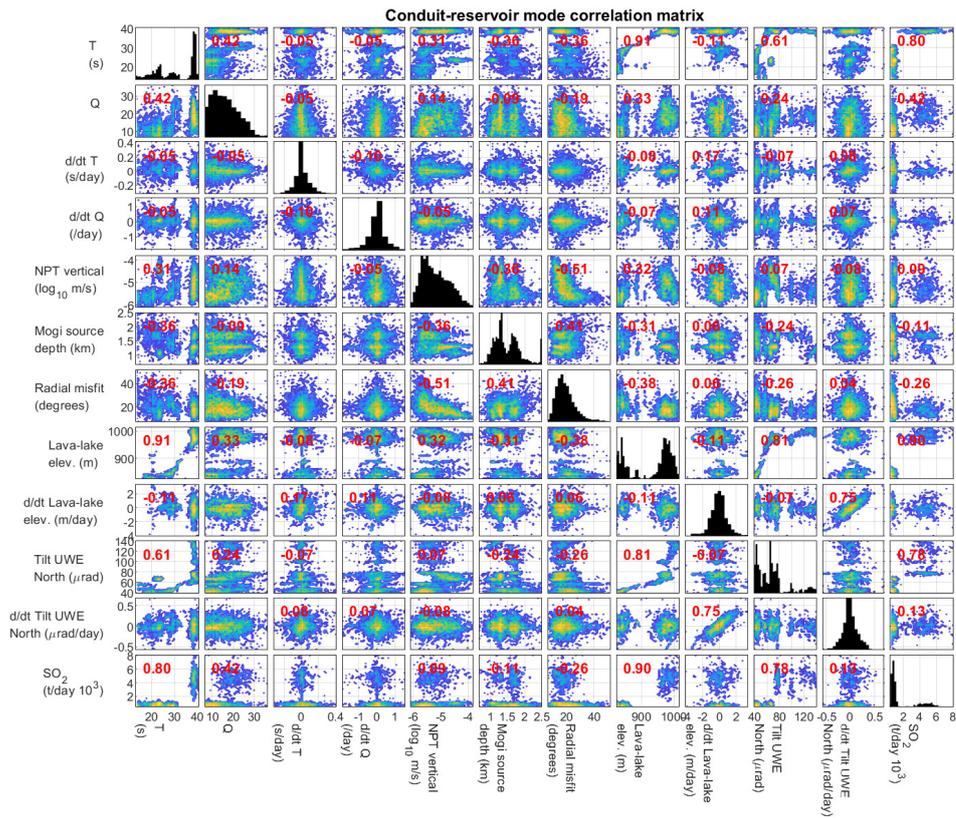


Figure S.12. Conduit-reservoir mode correlation matrices from 2008-2018. Off-diagonal plots are colored by the logarithm of the number of points in a given parameter bin, and histograms on diagonal plots show the distribution of each parameter. Red numbers are Pearson's correlation coefficients, only shown for correlations with P-values less than 0.05. All time derivatives were calculated with a 7-day cutoff-period differentiator filter (see Methods section).

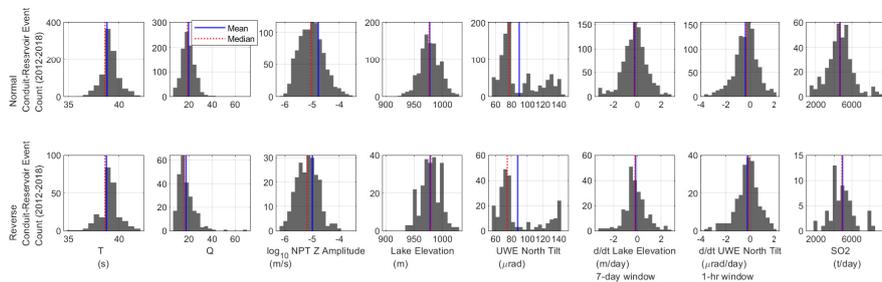


Figure S.13. Histograms of Normal and Reverse conduit-reservoir mode event parameters from 2012-2018.

3 Example event figures

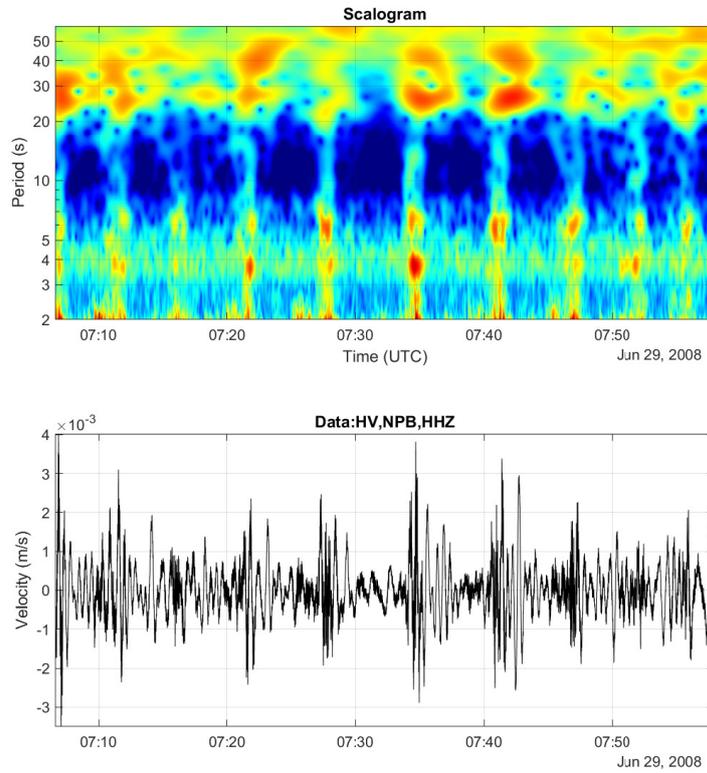


Figure S.14. VLP events with regular recurrence interval from June 2008, a few months after the Overlook Crater began forming. These events occurred roughly every 5 minutes and contained broadband energy with spectral peaks at around 3.5 s, 6 s, 25 s, and possibly 40 s. These events exhibited less clear onsets and exponential decays than typical rockfall-triggered events.

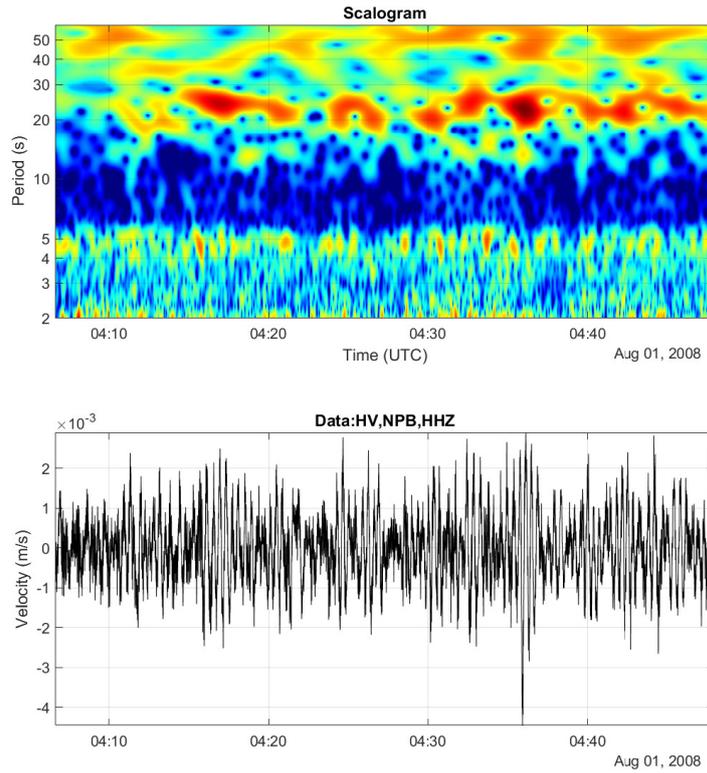


Figure S.15. VLP tremor from August 2008, in the first focused cluster of VLP signals. There was elevated energy at periods from 15-30 s and 4-5 s, though the dominant periods were not clearly focused and were variable over time. The signal cannot readily be separated into distinct events, and exhibited no clear high frequency triggers.

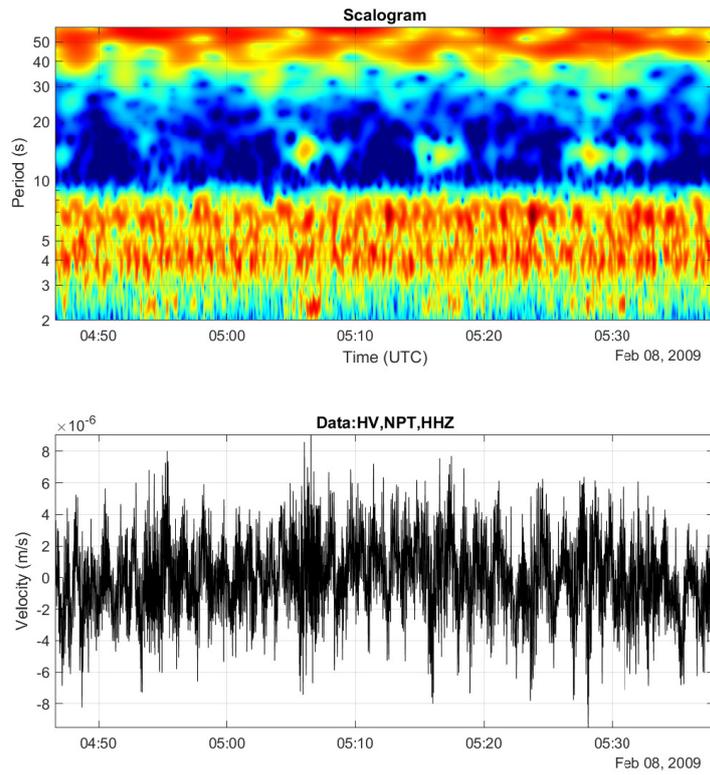


Figure S.16. VLP events from February 2009, around the time where dominant VLP period is at a minimum. These appear to be distinct VLP events, though onsets of some were gradual and first motions were not well defined. Elevated energy at periods < 2 s occurred alongside these signals, but did not appear to represent the more broadband impulsive trigger mechanisms that occur at the onset of typical rockfall events.

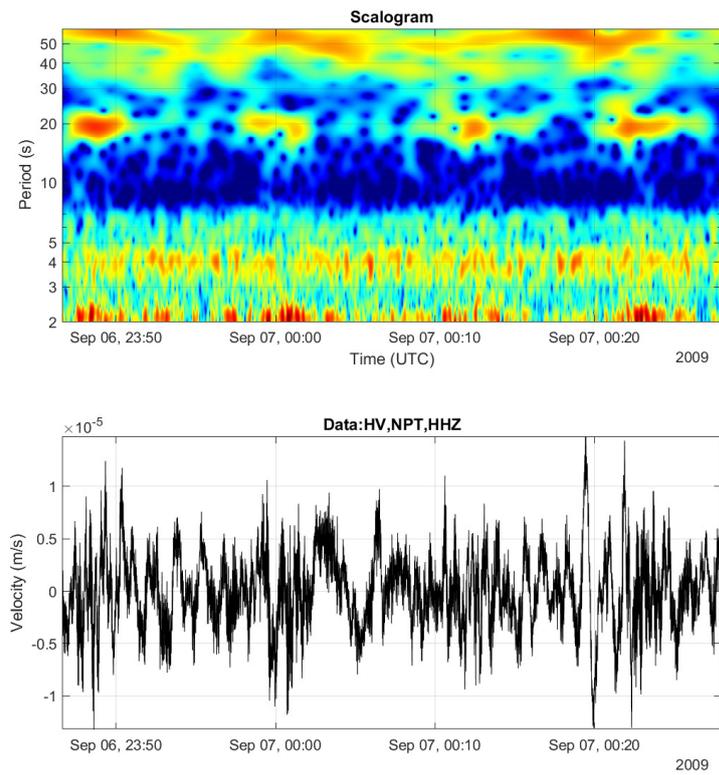


Figure S.17. VLP tremor from September 2009, in a signal cluster that seems to represent a local maxima in VLP period (around 20 s).

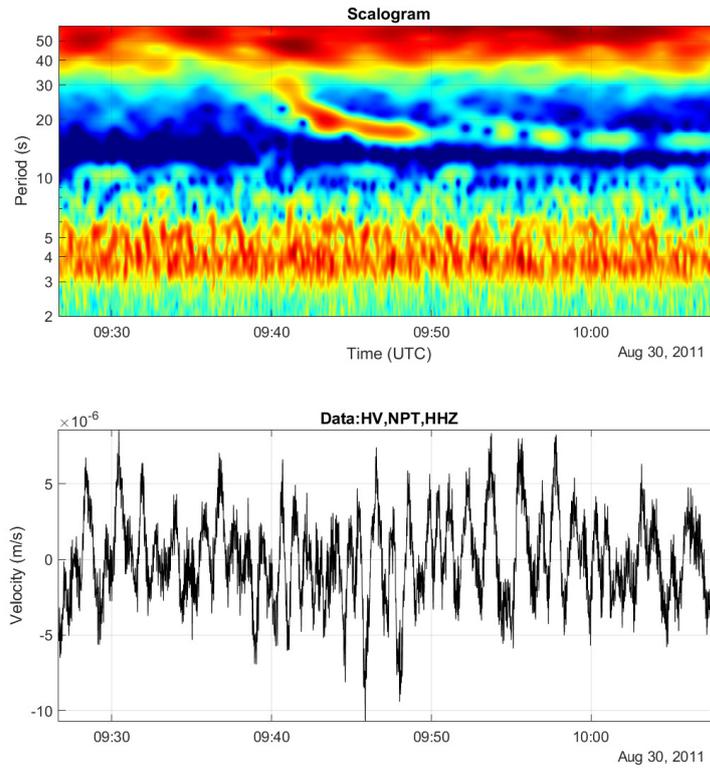


Figure S.18. Gliding-frequency VLP signal from August 2011, part of a small cluster of VLP seismicity following the August 2011 Pu‘u ‘Ō‘ō eruption. This event had no apparent high frequency trigger. VLP energy remained elevated for 10s of minutes after the event, though this energy did not appear to represent continuous decay of the initial resonance but rather continued intermittent forcing, perhaps partly by what may be a second smaller gliding-frequency signal around 10 minutes after the first. There was also background VLP tremor present with a period of around 11 s that does not appear to have been effected by the gliding-frequency event.

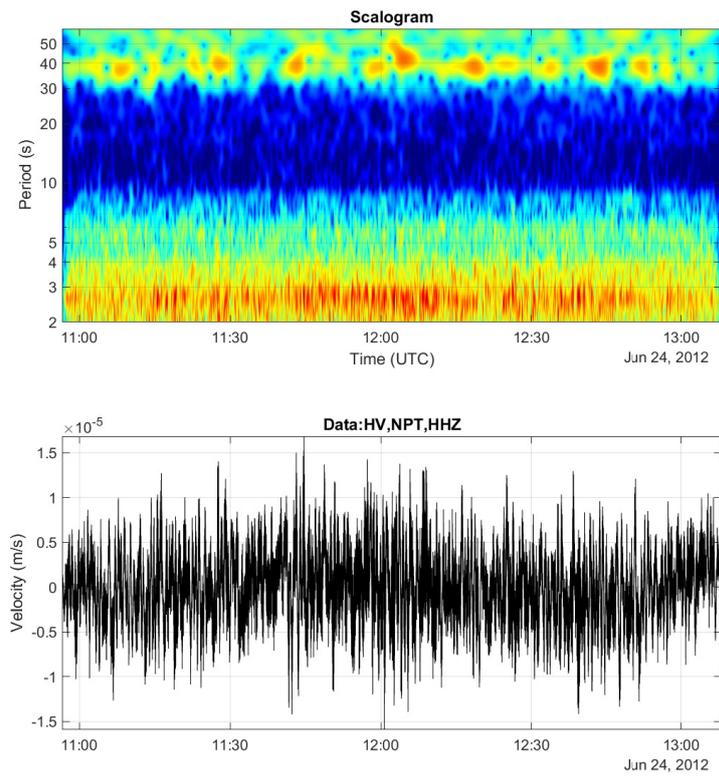


Figure S.19. VLP tremor from June 2012, shortly after the May SSE and around when higher Q VLP events start occurring again after a year with minimal VLP seismicity.

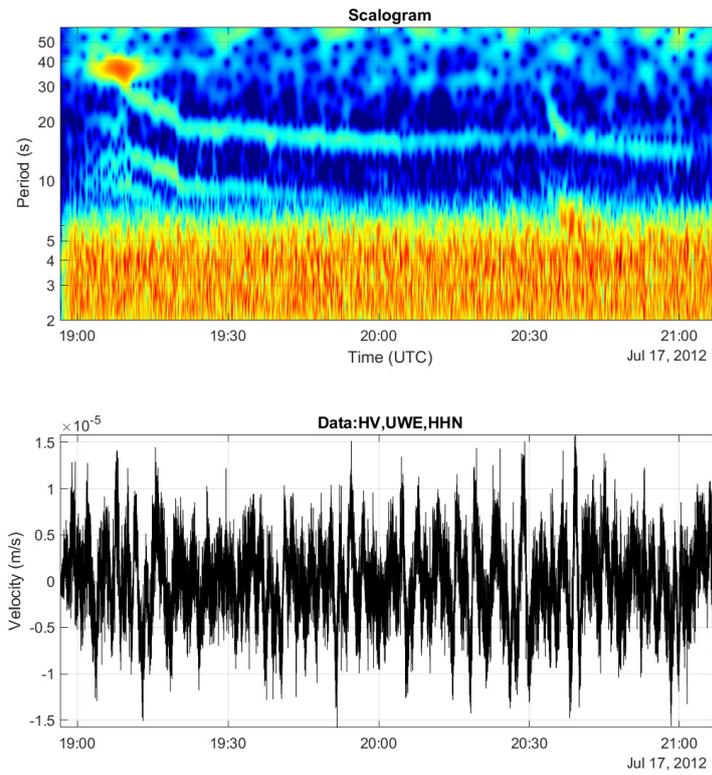


Figure S.20. Gliding-frequency VLP signals from July 2012. There was a set of three resonant modes starting around 19:10, and a single resonant mode that started about 90 minutes later. No high frequency triggers were apparent. The first 3 modes all exhibited a similar glide to lower periods over about 10 minutes, then maintained more stable periods. The later mode had a more rapid initial glide to lower periods (over about 5 minutes) but then continued more slowly gliding for another 20 minutes.

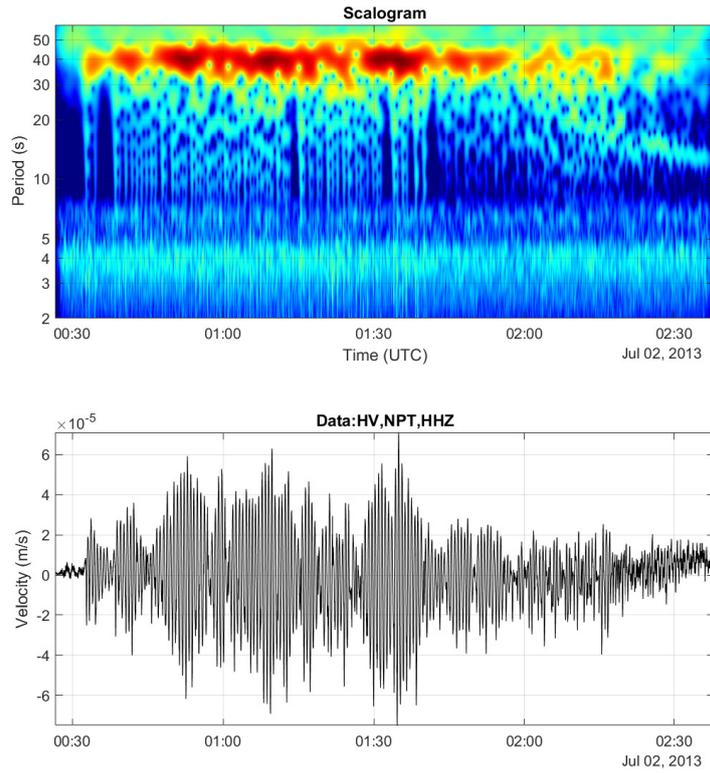


Figure S.21. VLP event/tremor from July 2013. This signal consisted of sustained 40 s oscillations at varying amplitudes and irregular bursts of higher frequency energy. These bursts were much weaker relative to the main VLP oscillation than typical rockfall trigger signals. The main VLP signal had an impulsive onset with deflationary first motions.

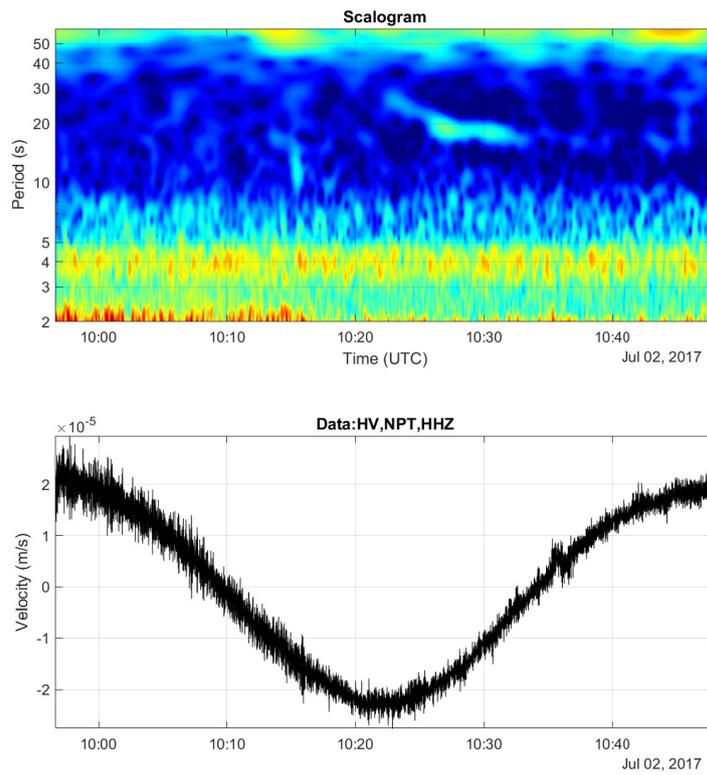


Figure S.22. Isolated lake sloshing mode with possible gliding-frequency onset from July 2017.

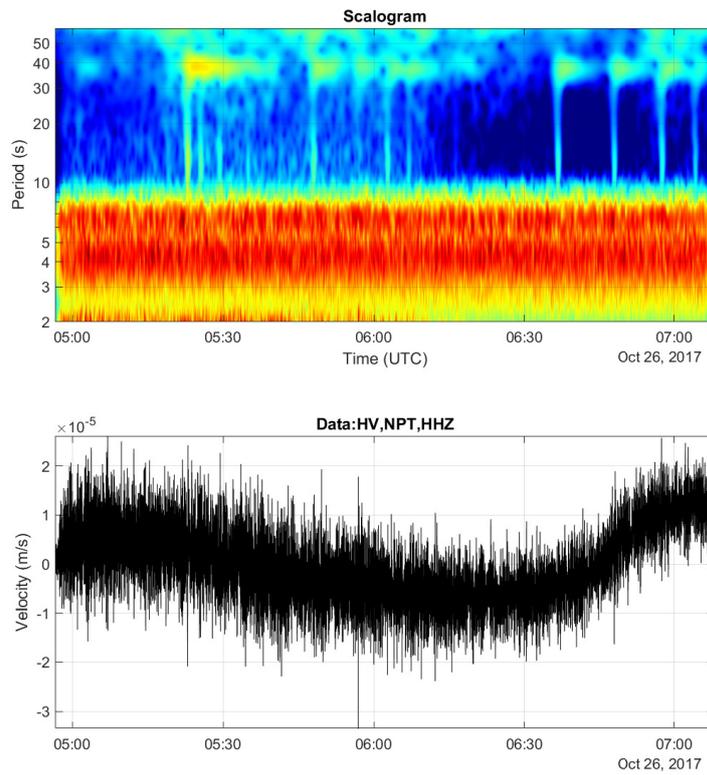


Figure S.23. Closely spaced Normal conduit-reservoir events from October 2017. These may have represented a series of small rockfalls.

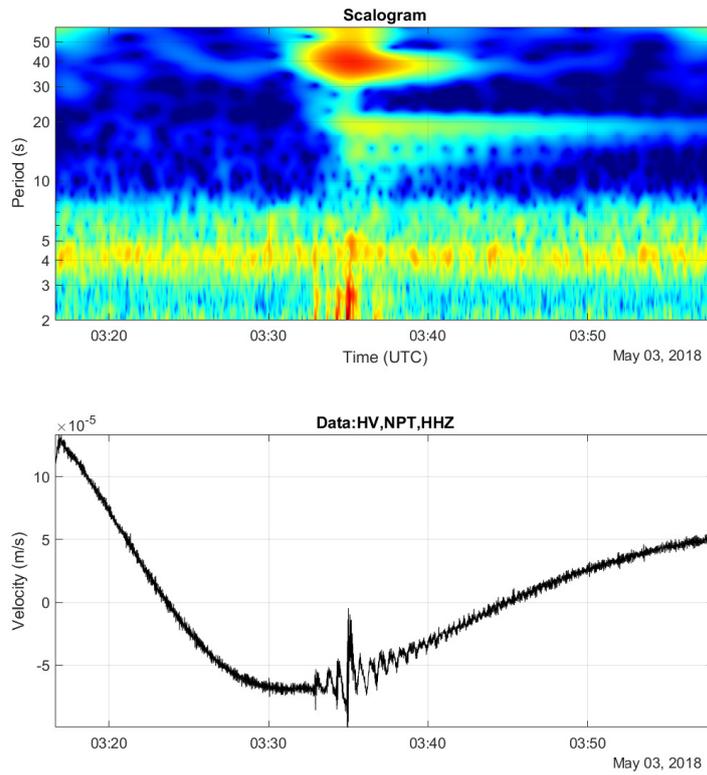


Figure S.24. VLP event with two clear lava-lake-sloshing modes from May 2018, a day after the lava-lake began draining. The dominant 40 s mode for this event started with impulsive inflationary motions, though with only a very faint high frequency trigger, but then grew for several minutes until a second impulse occurred and exponential decay began. The lava-lake-sloshing modes appeared alongside this second impulse.