# Multi-Decadal Measurements of UTLS Gravity Waves Derived from Commercial Flight Data

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#### Abstract

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# Key Points:

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6	•	We exploit in-situ aircraft data from 1994 to date to produce a 26-year UTLS grav-
7		ity wave dataset
8	•	Northern Hemisphere UTLS GWs show a statistically-significant $25\%$ amplitude
9		change over ENSO cycle
10	•	Strong and consistent GW activity seen above and poleward of the upper tropo-
11		spheric jet.

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#### 12 Abstract

Gravity waves (GWs) are key drivers of atmospheric dynamics, with major impacts on 13 climate and weather processes. However, they are challenging to measure in observational 14 data, and as a result no large-area multi-decadal GW time series yet exist. This has pre-15 vented us from quantifying the interactions between GWs and long-timescale climate pro-16 cesses. Here, we exploit temperatures measured by commercial aircraft since 1994 as part 17 of the IAGOS atmospheric chemistry research programme to produce a novel 26-year 18 time series of upper troposphere/lower stratosphere (UTLS) GW measurements across 19 most of the northern hemisphere. We analyse 90342 flight-hours (76.2 million flight-kilometres) 20 of data, typically at a temporal resolution of seconds and with high temperature preci-21 sion. We show that GW activity in the northern-hemisphere UTLS is consistently strongest 22 north of and above the upper tropospheric jet. We also show that GW sources not typ-23 ically observed in stratospheric data but assumed in model schemes, such as the Rocky 24 Mountains, are visible at these altitudes, suggesting that wave momentum from these 25 sources is deposited specifically between  $\sim 200-50$  hPa. Our data shows no significant im-26 pact of the Quasi Biennial Oscillation, the Northern Annular Mode, or climate change. 27 However, we do see strong evidence of links with the El Niño-Southern Oscillation, which 28 modulates the measured GW signal by  $\sim 25\%$ , and weak evidence of links with the 11-29 year solar cycle. These results have important implications for atmospheric process mod-30 elling and for understanding large-scale climate teleconnections. 31

## 32 1 Introduction

In-Service Aircraft for a Global Observing System (IAGOS) is a European research infrastructure programme designed to observe the upper troposphere and lower stratosphere (UTLS) using instruments fitted as an additional payload to commercial aircraft. IAGOS data are well-validated, and have been used in over 400 scientific studies to date. Including flights operated under the earlier MOZAIC and CARIBIC programmes, the dataset contains >60 000 flights, covering the period from 1994 onwards.

While the primary focus of IAGOS is on atmospheric chemistry, the metadata recorded 39 to contextualise these measurements are also of significant scientific use. These meta-40 data provide a long-term high-resolution record of atmospheric temperatures and winds 41 at small spatial scales, with high precision and good quality control. They have previ-42 ously been used in a number of dynamical studies (Lindborg, 1999; Cho & Lindborg, 2001; 43 Scott et al., 2001; Skamarock, 2004; Callies et al., 2014, 2016; Berkes et al., 2017; Li & 44 Lindborg, 2018), which have primarily focused on the bulk statistical nature of the data 45 rather than geographically-resolved features, and in particular on the fine structure of 46 horizontal winds. 47

Here, we exploit these metadata to study geographically-resolved small-scale at-48 mospheric gravity waves (GWs) in temperature data. GWs are near-ubiquitous in the 49 middle and upper atmosphere, and play a vast range of vital dynamical roles, including 50 controlling the stratospheric jets and the large-scale equator-to-pole and pole-to-pole cir-51 culations, driving chemical transport and ozone depletion, and acting to help stratospheric 52 clouds form (Murgatroyd & Singleton, 1961; Fritts & Alexander, 2003; M. J. Alexander 53 et al., 2010; S. P. Alexander et al., 2013; Geller et al., 2013). However, due to their small 54 physical scales (at most tens of kilometres and usually much less in the vertical, and tens 55 to hundreds of kilometres in the horizontal) they are technically challenging to measure, 56 and as a result few long GW time series exist. 57

The challenges of measuring GWs are particularly acute in the UTLS. Ground-based instruments and radiosonde balloons provide limited spatial coverage, particularly over the open ocean, while scientific aircraft and superpressure balloon campaigns are limited in number and coverage and do not take routine measurements. To add to this, the satellite methods that have revolutionised our stratospheric gravity wave knowledge in the last two decades perform extremely poorly in the UTLS, due to their relatively coarse
spatial resolution and the vertical structure of the temperature-tropopause itself, which
heavily complicates use of profile-based detrending methods for extracting GW signatures from background variations. As such, IAGOS data can provide a unique long-term
window on UTLS GWs, particularly at remote locations.

Callies et al. (2016) applied spectral decomposition methods to MOZAIC wind data 68 from 2002–2010, demonstrating that the data contained signatures consistent with in-69 ertia-gravity waves in the lower stratosphere. However, they did not identify such a re-70 71 lationship in the upper troposphere, a conclusion they ascribed to mesoscale variances in measured along-track wind components. Here, we use temperature data from a 1994-72 2019 superset of IAGOS data; these temperatures do not appear to share the along-track 73 variance issues, and contain clear GW signatures amenable to spectral analysis in both 74 the upper troposphere and lower stratosphere. 75

In Section 2 we describe the data used, and in Section 3 the analyses performed. 76 Section 4 then examines seasonal maps of the northern hemisphere, considering both directly-77 measured geophysical variables (i.e. wind and temperature) and our analysis outputs. 78 In Section 5 we study time series for 18 selected regions, which are later tested against 79 multiple climate indices in Section 7, while Section 6 examines the relationship between 80 GWs and the upper-tropospheric jet. Section 8 compares our results to previous UTLS 81 studies of GWs, and Section 9 discusses a key implication of our results. Finally, we dis-82 cuss systematic biases and deficiencies in our dataset in Section 10, before summarising 83 and drawing conclusions in Section 11. 84

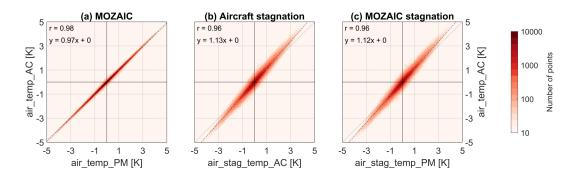
## **2 IAGOS Data**

We use data from the IAGOS Data Portal (Boulanger et al., 2019), incorporating all final-version recorded flights from the 1st of August 1994, the start of the dataset, to the end of 2019. A significant lag often exists between data being recorded and finalised, and accordingly from October 2017 onwards our dataset does not include all IAGOS flights.

We specifically use aircraft-instrument-derived temperature data ('air\_temp\_AC'). These data are recorded by the aircraft's standard instrumentation rather than the IA-GOS package and thus may be less accurate, but are available for every flight. Using these data approximately doubles available data relative to temperatures measured by the instrument packages alone, with the increase largest in later years.

Since we detrend the data to identify GWs, absolute biases and drifts over long pe-95 riods of flight are not important to our analysis, thus potentially justifying use of these 96 data. To test this, Figure 1 compares 'air\_temp\_AC' against three other IAGOS temper-97 ature products, computed over all level-flight measurements for which each variable-pair 98 exists in January 2000, detrended and preprocessed as described below. We see close agree-99 ment, with correlations of >0.95 and linear fits. There is a bias between stagnation-temperature 100 and direct-temperature products; this bias is symmetric about zero perturbation, and 101 thus will lead to a reduction  $\sim 10\%$  in our measured wave amplitudes relative to if stag-102 nation products were used. 103

Other randomly-tested months were analysed in the same way, with equivalent results.



**Figure 1.** Comparison of perturbations in aircraft-measured temperature ('air\_temp\_AC') against perturbations in (a) MOZAIC-measured temperature ('air\_temp\_PM') (b) aircraft-measured stagnation temperature ('air\_stag\_temp\_AC') and (c) MOZAIC-measured stagnation temperature ('air\_stag\_temp\_PM') for corresponding measurements in each dataset for the month of January 2000. Each panel shows a scatter-density plot of (vertical axis) 'air\_temp\_AC' against (horizontal axis) each other product, with a bin size of 0.05 K on each axis. Close agreement is seen, justifying use of the aircraft-measured temperature data for our GW analyses.

## 106 3 Methods

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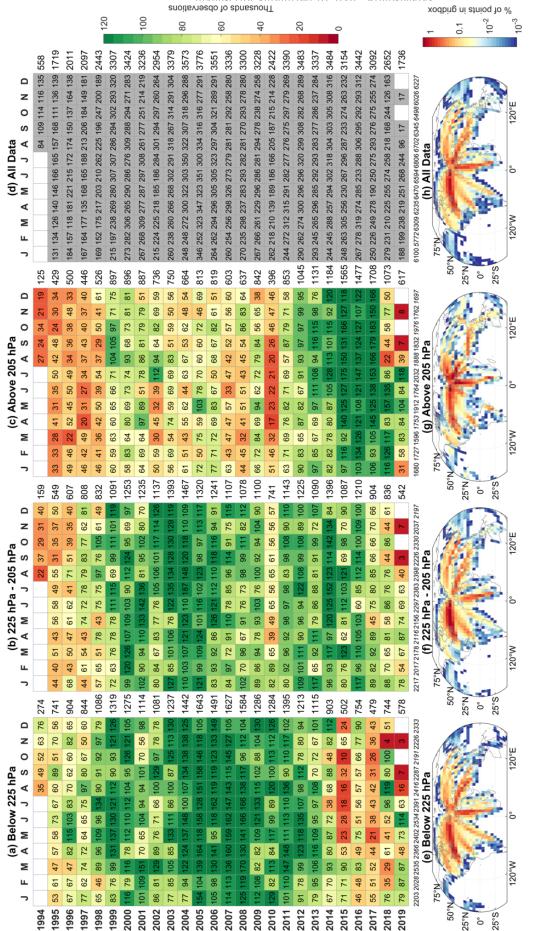
## 3.1 Preprocessing of IAGOS Data

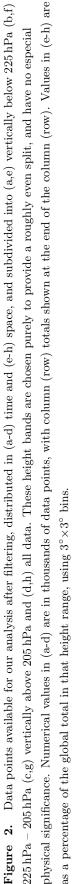
IAGOS data include ascents and descents as well as level cruising. The data also
 vary in resolution and occasionally include minor errors. As such, some preprocessing
 is required to provide consistent data for analysis.

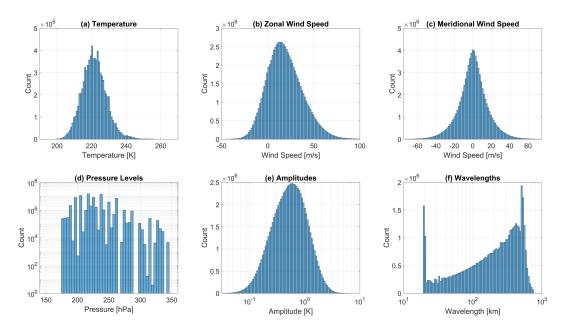
We first apply all error flags, removing data entirely if any variable we subsequently use is flagged as bad. After this, some records contain either negative timesteps or repeated latitude/longitude pairs. For these records, we re-order the data to ascend monotonically in time and replace duplicated pairs via linear interpolation.

We next split the data into sections of straight near-level flight, removing any parts 115 where height or direction changes quickly. These include ascents and descents, which we 116 do not use. We label these straight stable sections as "cruises", which we define as hav-117 ing barometric altitude changes of <100 m and directional changes of  $<45^{\circ}$  within any 118 rolling fifteen-minute time window. The first check ensures we measure a consistent phys-119 ical regime, while the latter check is to avoid potential measuring a laminar feature twice 120 at different angles and misidentifying it as a wave. Each cruise is subsequently analysed 121 independently. 122

We then linearly interpolate the data to a 1 km along-track resolution. This is rep-123 resentative of the original data, which have a mean spacing of 1.09 km (median 1.01 km, 124 standard deviation  $0.50 \,\mathrm{km}$ ). We discard any cruises with discontinuities greater than 125 20 km, linearly interpolating over gaps smaller than this. We then remove cruises which, 126 after these steps, are <1000 km long. After additional removal of out-of-band data (Sec-127 tion 3.2), we are left with 76.2 million data points, distributed as shown in Figure 2 and 128 representing 90342 flight-hours of data. Our data are biased towards the northern hemi-129 sphere, especially the northern Atlantic Ocean, and skew towards higher altitudes in later 130 years. We divide the data into three pressure bands (>225 hPa, 225 -205 hPa and <205 hPa)131 respectively) to contextualise relative coverage as a function of height; these bands are 132 chosen to split the data into three approximately equal subsets, rather than on the ba-133 sis of physical differences, and are not used below. 134







**Figure 3.** Histograms of basic data properties, integrated across the entire dataset. (a-d) show directly measured temperature, zonal wind speed, meridional wind speed and pressure. (e) and (f) show derived wave amplitudes and wavelengths.

Figures 3a-d show the broad characteristics of the resulting dataset. Temperatures and zonal winds show smooth distributions with positive skew, while meridional winds show a near-symmetric distribution focused around zero. Pressures show a spiky distribution (note the logarithmic ordinate), consistent with preferred flight levels.

<sup>139</sup> 3.2 Gravity Wave Analysis

To identify GWs, we use a one-dimensional Stockwell Transform (ST) method. Such methods have been widely used in previous GW research (e.g. M. J. Alexander et al., 2008; Wright et al., 2010; Hindley et al., 2015; Moss et al., 2016), and allow us to localise measured GWs in both location and wavenumber.

We first smooth each cruise with a 900 km second-order Savitzky-Golay filter, and 144 subtract this from the original data to produce perturbation series. Figure 6 of Hindley 145 et al. (2015) shows the transfer function for this filter, which here provides good trans-146 mittance up to  $\sim 500$  km, dropping off sharply above this. Callies et al. (2014) saw a change 147 in horizontal dynamical scales at  $\sim 500$  km, supporting this choice. We then spectrally 148 analyse each cruise, using the ST implementation of Hindley et al. (2019). From this we 149 estimate wave amplitude A and along-flight horizontal wavenumber k for the strongest 150 signal present at each data point (see e.g. M. J. Alexander et al., 2008). We restrict the 151 analysis to wavelengths between 10 km and 600 km. 152

Figures 3e-f show the resulting amplitudes and wavelengths. Both distributions are 153 smooth, with amplitudes centred at 0.7 K and wavelengths peaking at 500 km. A max-154 imum is seen in the lowest two wavelength bins at  $\sim 20$  km, arising from small-scale noise 155 aliased into the analysed range. To prevent this spurious peak contaminating our results, 156 we remove data with wavelengths below 25 km. We also remove wavelengths longer than 157 500 km, as they will be attenuated by the detrending filter and in any case are expected 158 to be dynamically different. Finally, we remove points with measured wavelengths more 159 than twice the distance between the measurement and the start or end of the cruise; this 160

reduces edge-truncation effects in our results, but will introduce a short-wavelength bias
 near the start and end of flights. This effect will be largest over Europe, where most flights
 commence or terminate. The removed data are distributed approximately evenly across
 the other distributions in Figure 3, and removing them does not affect their form.

Sensitivity tests suggest our results are only weakly affected by the choice of a 1 km 165 interpolant or the definition of a cruise. The choice of detrending filter also has only a 166 small effect: boxcar, Lanczos and Gaussian filters were also tested, with very little dif-167 ference for the boxcar or Lanczos and, for the Gaussian, a shift to slightly higher am-168 plitudes and longer wavelengths but with otherwise similar results. However, our results 169 are strongly affected by the 900 km filter length. In particular, tests with significantly 170 longer filters produce large peaks in wavelength over the midpoints of typical long-haul 171 flights, i.e. the mid-Atlantic and central Asia, a bias which is avoided by using our smaller 172 value. 173

## <sup>174</sup> 3.3 Cluster Mapping

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## 3.3.1 Hierarchical Cluster Analysis

Due to the uneven distribution of our measurements, balancing the limited data in many regions with dense coverage in others is non-trivial. Using a high-resolution grid leaves large gaps, while a low-resolution grid underrepresents available resolution, leading to the discard of large areas when we later impose a minimum number of measurements per gridbox.

To balance these competing scales, we use an approach combining hierarchical clustering and statistical bootstrapping. This is based on the work of Wright et al. (2012, 2013), and produces maps with good spatial coverage over most of the northern hemisphere, while still exploiting the full resolution where measurement density is high. We carry out this analysis at the seasonal level, using combined data from all years for boreal winter (DJF), spring (MAM), summer (JJA) and autumn (SON) respectively.

<sup>187</sup> We first produce a list of latitude/longitude pairs. To make calculations compu-<sup>188</sup> tationally tractable, we assume that points within the same  $0.2^{\circ} \times 0.2^{\circ}$  box represent the <sup>189</sup> same location and de-duplicate at this level of precision (this does not affect our results, <sup>190</sup> as the data values are restored once the grid is produced). This reduced dataset still rep-<sup>191</sup> resents the data distribution at scales well above this, and is therefore suitable for clus-<sup>192</sup> ter generation, since we later impose a minimum area of  $0.5^{\circ} \times 0.5^{\circ}$ .

We next apply a k-means algorithm to the geolocation data, dividing the data into a fixed number of spatially-compact clusters. For each season, we do this 10 000 times, selecting the iteration with the lowest total distance between measurements and cluster centres. We use a Euclidian distance metric in latitude/longitude space - this is strictly incorrect for the Earth, but in practice introduces only a small error at the scale of our clusters, and provides significant runtime benefits.

We initially define 4 000 clusters; this is arbitrary, and is chosen as the highest round number to which we can apply our subsequent bootstrapping analysis in reasonable computer time (hours) and memory (tens of gigabytes) on the system used. Systematically varying this number suggests that using 4 000 clusters gives statistically similar results to maps produced using lower values, but providing finer geographic resolution. Tests using regular latitude-longitude grids instead of the cluster analysis also produce similar results, but without geographically-variable resolution.

We then re-assign each measurement we removed in the de-duplication step to the geographically-closest cluster. Measurements >300 km from a cluster centre are discarded. Finally, we require at least 500 measurements within each cluster, discarding those with fewer and reassigning their data to the next-nearest cluster, while retaining the 300 km limit. This reduces the number of clusters used to slightly below 4000 - for example, Figure 4 actually contains 3764 clusters.

#### 212 3.3.2 Bootstrap Statistics

Due to the inhomogenous data distribution, this produces clusters which, while rep-213 resenting both large and small scales well, contain uneven numbers of measurements, with 214 a spread of three orders of magnitude in population (compared to five orders of magni-215 tude using a regular  $3^{\circ}$  grid). While suitable for mapping, this may introduce spurious 216 inter-cluster variations in the statistics we compute, namely the median and Gini's co-217 efficient of concentration G (used previously for GW studies in e.g Plougonven et al., 2012; 218 Wright et al., 2013; Hindley et al., 2019; Kuchar et al., 2020). Means were also calcu-219 lated for all clusters and showed spatially consistent results but with spike-values in some 220 clusters due to outliers. 221

To correct for these possible variations, we bootstrap our statistical calculations 222 within each cluster. We first sample 500 sets of 1000 randomly-chosen measurements 223 from each cluster, allowing replacement of the same value. We then compute our statis-224 tics independently for each set, recording the median value of this 1000-point distribu-225 tion. This median is used as our overall statistical estimate for the cluster properties. 226 This allows us to compute comparable statistics for each cluster which are not massively 227 biased by their different populations. To test sensitivity, these numerical values (i.e. 500 228 and  $1\,000$ ) were varied systematically across an order of magnitude in each direction, with 229 only small changes in the final results. 230

# 3.3.3 Final Mapping

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Finally, we bin the statistics onto a  $0.5^{\circ} \times 0.5^{\circ}$  grid to facilitate mapping. For each gridbox, we use the statistics assigned to the geographically-closest cluster centre. If the nearest cluster centre is >300 km away, then we leave the gridbox blank.

Figure 4a and 4b) show maps of the cluster-areas generated for DJF using data from all heights. We see a dense patchwork over the North Atlantic and Western Europe, with larger clusters in other regions. This is representative of the original data distribution (Figure 2). Figure 4c compares the area assigned to each cluster with the midlatitude area of a regular gridbox (dashed lines); most clusters represent an area equivalent to a 2° degree grid or smaller, with almost all representing an area less than than a 5° degree grid.

#### **3.4 Tropopause Calculation**

For context, we calculate the tropopause pressure level using daily ERA5 reanal-243 ysis output (Hersbach et al., 2020), at full vertical resolution but downsampled to a spa-244 tial resolution of  $1.5^{\circ} \times 1.5^{\circ}$ . For each gridpoint, we find the lowest altitude where (a) the 245 vertical temperature lapse rate is  $< 2 \,\mathrm{Kkm^{-1}}$  and (b) this remains true for at least  $2 \,\mathrm{km}$ 246 above this point. We require the trop pause to lie between 700 hPa - 10 hPa, linearly 247 interpolating to find the transition below  $2 \,\mathrm{Kkm^{-1}}$  as accurately as possible between model 248 levels. Any gridpoints where these criteria are not met are bilinearly interpolated from 249 surrounding gridpoints. 250

## <sup>251</sup> 4 Cluster-Generated Maps

Figure 5 shows seasonal maps of median temperatures, zonal winds, and meridional winds between 200 hPa-250 hPa, and the number of measurements assigned to each cluster. Figure 6 shows differences between these data and the ERA5 reanalysis over the same pressure range and period. At this altitude and over this large a height range and time

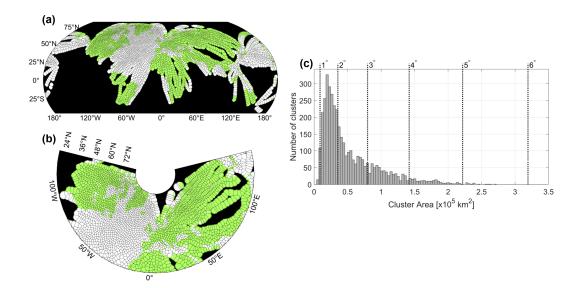


Figure 4. (a,b) Example maps of areas assigned to clusters for (a) all locations with data (b) densely-measured regions of a section of the northern hemisphere. Thin black lines mark perimeters of areas assigned to each cluster; solid black shading indicates no coverage. (c) Histogram showing geographic area assigned to each cluster. Dashed vertical lines indicate the area which would be covered by each cell of a regular latitude-longitude grid (at 45° latitude), indicating that the majority of our clusters represent areas smaller than the average area of a  $2^{\circ} \times 2^{\circ}$  regular grid. These examples are specifically for the cluster maps generated using data from all heights for the DJF composite season, but broadly resemble those generated for other heights and periods.

period, we expect the reanalysis fields to reasonably approximate true values, and therefore allow us to identify systematic biases created by our sampling pattern. Figure 7 shows GW amplitudes (top rows), G (middle rows) and along-track GW wavelengths (bottom rows) over the same pressure range. We focus on the region  $20^{\circ}N-85^{\circ}N$ ,  $130^{\circ}W-140^{\circ}E$ , due to good sampling in this region.

The maps (except Figure 5y-B) have been smoothed using a 2D Gaussian of FWHM 262  $2.5^{\circ} \times 2.5^{\circ}$  to facilitate analysis of large scale features. In the unsmoothed data, not shown, 263 small-area orographic wave sources can be clearly identified, especially in the smaller clus-264 ters over Europe and the North Atlantic, and these will be studied in more detail in fu-265 ture work.

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# 4.1 Temperature, Wind and Measurement Density

4.1.1 Absolute Value Maps

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Figure 5 shows directly-measured variables. These data have been processed identically to the GW data discussed below, to help understand sampling biases.

Temperatures (Figures 5a-d) are low in DJF and high in JJA with a negative meridional gradient from tropics to pole. Standard deviations (Figures 5e-h) are ~4% or less, with higher internal variability in DJF/MAM and lower internal variability in SON/JJA.

Zonal winds, Figures 5i-l, are strongly positive over the western Atlantic all year 273 and over northern Africa/southern Asia for most of the year, with weaker winds in JJA 274 and stronger in DJF. Winds are moderately strong across most of western North Amer-275 ica, especially compared to Eurasia poleward of  $\sim 40^{\circ}$ N. Standard deviations (Figures 276 5m-p) are high over the North Atlantic, consistent with subseasonal variability in jet lo-277 cation. Additional standard deviation peaks are seen east of Asia in MAM, over central 278 Asia in JJA, and over the central USA in DJF and SON; all of these regions also exhibit 279 high absolute values. 280

Meridional winds (Figures 5q-t) are smaller than zonal, consistent with Figure 3c. Except for JJA, they are negative over western North American, Siberia, the edges of Scandinavia and China, and positive over the North Atlantic, the eastern coast of Russia, and to a lesser degree the Middle East and Iran. JJA exhibits a different morphology, with weaker and less spatially consistent winds over the North Atlantic, positive winds over the Eastern Mediterranean, and positive winds over Siberia. Standard deviations (Figures 5u-x) are largest over the North Atlantic and (again, except JJA) North America.

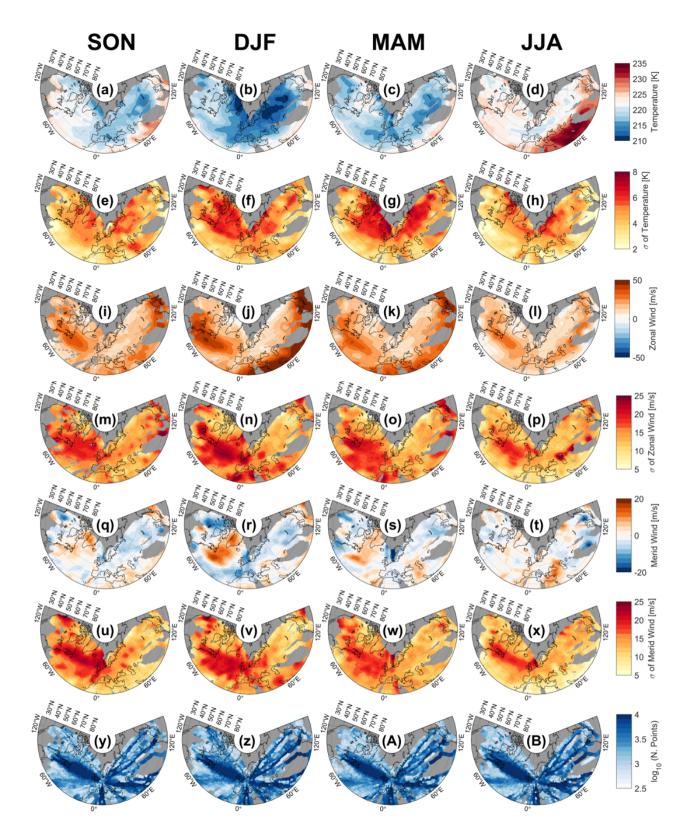
Measurement density (Figures 5y-B) is greatest on direct great-circle flight paths between Europe and major extra-European population centres, particularly in North America. There is a slight reduction over mainland Europe relative to surrounding locations due to our removal of edge-truncated waves in this region.

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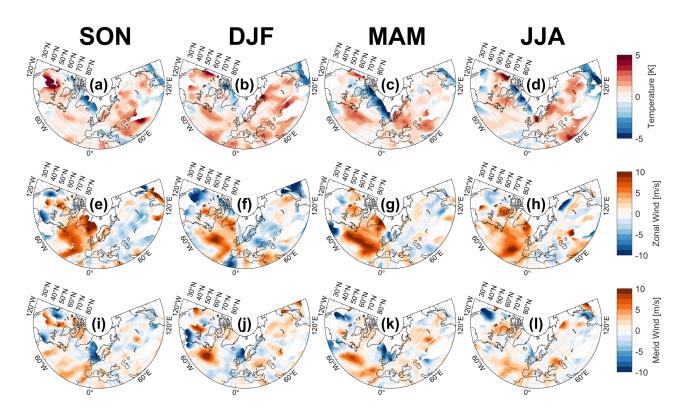
# 4.1.2 Biases Relative to Reanalysis

Figure 6 shows differences from gridded ERA5 output, averaged over the 200–250 hPa range. In general, pointwise differences between IAGOS and reanalysis temperatures and winds should be small (e.g Berkes et al., 2017), and therefore we assume differences represent sampling biases rather than measurement inaccuracies, primarily arising due to aircraft routing optimisations.

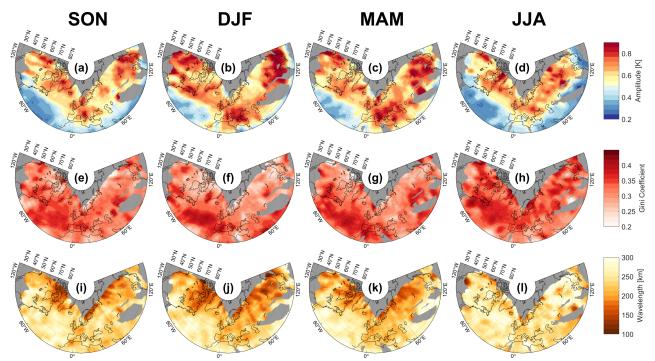
Regional temperature differences (Figure 6a-d) are 5 K or less, except for a lightlysampled region at high latitudes in MAM. Temperatures have a consistent positive bias, except for the far east of Asia and high latitudes near Greenland. Meridional wind differences (Figure 6i-l) are generally patchy, with a positive bias over the North Atlantic



**Figure 5.** Cluster-analysis derived maps of IAGOS-measured (a-d) temperature, (e-h) standard deviation of temperature, (i-l) zonal wind speed, (m-p) standard deviation of zonal wind speed, (q-t) meridional wind speed, (u-x) standard deviation of meridional wind speed and (y-B) number of measurements, computed over the pressure range 150 hPa – 350 hPa. Figure combines data from all years of IAGOS data. Solid grey shading indicates insufficient data density for analysis, according to the criteria outlined in Section 3.3



**Figure 6.** Difference between ERA5 reanalysis and our measured (a-d) temperature, (e-h) zonal wind speed and (i-l) meridional wind speed. Data are computed over the pressure range 150 hPa – 350 hPa for the period August 1994 to December 2019, and are plotted such that a positive value indicates a positive bias in IAGOS relative to ERA5.



**Figure 7.** Cluster-analysis derived maps of IAGOS-measured mean GW (a-d) amplitudes (e-h) *G* and (i-l) along-track wavelengths for (a,e,i) SON (b,f,j) DJF (c,g,k) MAM (d,h,l) JJA, for all data. Figure combines data from all years of IAGOS data.

jet and the central-eastern Atlantic off Africa coast, and a negative bias over Scandinavia and the North Sea. Zonal winds (Figure 6e-h) are the most different field, with patchy biases in most places but with a persistent and large positive bias in the North Atlantic jet stream. This bias is consistent with aircraft flying in the peak of the jet travelling eastwards and avoiding the jet going westwards, leading to a net positive bias when averaged.

#### 4.2 Gravity Waves

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GW parameters are shown in Figure 7. Amplitudes (Figures 7a-d) are generally largest in DJF and MAM, and lower in SON and JJA, but with numerous exceptions. They are consistently low over the mid-Atlantic, and consistently high over the North Atlantic and most land poleward of ~40°N. These regions represent a mix of possible wave sources.

Local maxima over southeastern Greenland, western Mongolia, the Sikhote-Alin mountains of Eastern Russia, and the Rocky Mountains of North America are located close to meridionally-aligned ridges, which would be expected to generate wave activity when exposed to strong near-zonal surface winds (e.g. Bacmeister, 1993). *G* is generally higher than the surrounding area at these locations (except Sikhote-Alin) - this is consistent with orographic wave sources, which tend to be intermittent in nature (Wright et al., 2013).

Many of these orographic signatures are not routinely observed in lower-stratospheric satellite data (e.g Geller et al., 2013). This may suggest that wave momentum generated here is deposited specifically in the lowermost stratosphere. We further discuss the specific case of the Rockies in Section 9.

The strongest clearly non-orographic feature in DJF and MAM is a band of high 326 GW amplitude stretching east from Newfoundland. This band closely mirrors the North 327 Atlantic winter storm track, and is well-resolved due to the high flight density over this 328 region. Measured amplitudes in this band are likely to be skewed slightly high, due to 329 the routing bias discussed above. This makes sense in the context of the dataset: if GWs 330 have the largest amplitudes near the jet, and flights benefit in fuel and travel time from 331 flying nearby on eastward flight legs, then this will skew the measured value anywhere 332 the jet is present during our dataset high relative to other locations. This bias does not 333 imply that there is not a maximum in this band, but merely that we overestimate its rel-334 ative magnitude. As such, this bias likely enhances the apparent importance of this non-335 orographic peak with respect to fixed-location orographic hotspots such as southeast Green-336 land. 337

G (Figures 7e-h) is generally highest around the Atlantic jet, consistent with the varying jet location. There is also higher G over some orographic regions in winter, consistent with orographic source intermittency, and over large fractions of the observed area in general in JJA.

Wavelengths (Figures 7i-l) are relatively short near the pole in all seasons. We believe this to be a bias arising from the sampling direction of the aircraft trajectories. Aside from this, some regional features can perhaps be seen, with short wavelengths over mountain ranges in DJF and MAM, and (to a minor degree) in patches over the Atlantic Ocean corresponding to regions of high G. In general however, the known difficulties of accurately sampling GW wavelengths from randomly-oriented 1D or 2D cuts through the atmosphere (e.g. Ern et al., 2004) make interpretation of these maps challenging.

We see no evidence of the strong subtropical GW maxima over the Carribean, Africa and South-East Asia usually observed in stratospheric GW observations in JJA (e.g Geller et al., 2013). We believe that, rather than these maxima not being present in the real

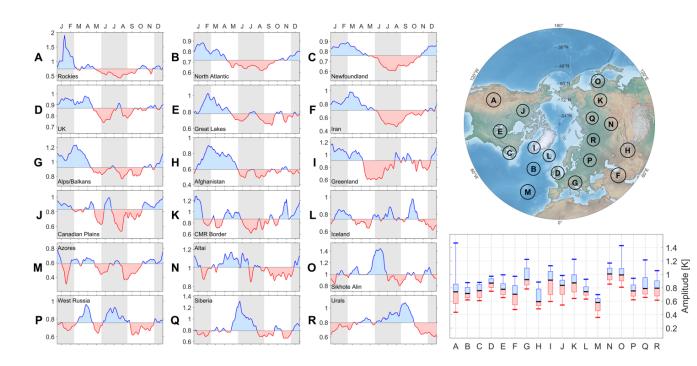


Figure 8. (top right) Map indicating sample regions selected as described in the text. Black circles surrounding each identifying letter show the area averaged. (A-R) Individual composite-year time series at each lettered location. Blue indicates above-median GW amplitudes, red below. Ordinates show wave amplitudes in Kelvin; data have been averaged over a rolling 31-day window. Grey shading indicates the seasons plotted in Figure 7. 'CMR Border' indicates the region at the junction of China, Mongolia and Russia, between the Yablonovy (Russia) and Dà Xīng'ān Lǐng (China) mountain ranges. (bottom right) Box-and-whisker plots showing range of GW amplitudes measured in each region. Upper and lower horizontal lines show the 2.5th and 97.5th percentiles of the data, shaded boxes cover the region between the 18th and 82nd percentiles, and the black solid line indicates the median.

atmosphere, this absence instead is due to aircraft routing biases. Specifically, if these maxima are to a significant extent driven by large-scale storm activity, as seems likely from theoretical, model and satellite evidence, then planes will be routed around rather than through such activity. This will lead to a sharp reduction in relative magnitude in our dataset. However, our available flight metadata do not provide sufficient information to test this hypothesis. The low amplitudes measured over the mid-Atlantic may also be related to such routing choices.

## **5 Regional Time Series**

Figure 8 shows time series of GW amplitude for 18 selected regions, mapped at top right. Most of these regions have been selected visually from Figure 7a-d as local maxima in at least one season, with J, M and P added to fill remaining large spatial gaps in regions of good data coverage.

For each region, median GW amplitude is computed over a rolling 31-calendar-day window, composited over all years and over a geographic circle of radius 500 km indicated by the black circles surrounding the letters. This is done using the original flight-track data, rather than the cluster-analysis output. Within each region, values above the annual median are shown in blue and below in red. The order of display (i.e. A-R) is selected to to emphasise systematic changes in seasonality: specifically, they are ordered by linear correlation with the Rocky Mountains (region A), selected as (a) arbitrarily, it is the westernmost region and (b) subjectively, it has a clear seasonal cycle. Time series most similar to A are assigned earlier letters alphabetically. Sequences of alphabetically-close letters on the map therefore indicate consistent GW-amplitude seasonality at supraregional geographical scales.

At bottom right, the spread of amplitudes in panels A-R is illustrated, allowing comparison between the scaled time series. For each region, a box-and-whisker plot is shown, highlighting the range between the 18th and 82nd percentiles (box, equivalent to the first standard deviation of a normal distribution) and between the 2.5th and 97.5th percentiles (whiskers, equivalent to the the two standard deviation range).

Regions A-H show clear seasonal cycles with maxima in winter/early spring, and minima in summer/early autumn. Regions I-P then show multiple minima and maxima, while regions Q and R are largely in antiphase with A-H and peak in summer.

Spatial groupings with strong similarities in seasonal cycle and median amplitude 383 can be seen in the data. These include a broad curve sweeping from the Rocky Moun-384 tains across the Atlantic to central Europe (A-E,G), a Central Asian grouping (N-Q), 385 and pairs in Greenland/Iceland (I, L) and the Middle East (F,H). These groups are dis-386 tributed in zonal bands, with relatively short meridional distances leading to much larger 387 differences in seasonal cycle. For example, regions I,J,L and M lie spatially close to in-388 dividual members of the A-E, G grouping, but exhibit large differences from this group. 389 These strong meridional gradients are consistent with the seasonal-level view seen in Fig-390 ure 7, but are shown here to be applicable at the subseasonal level. 391

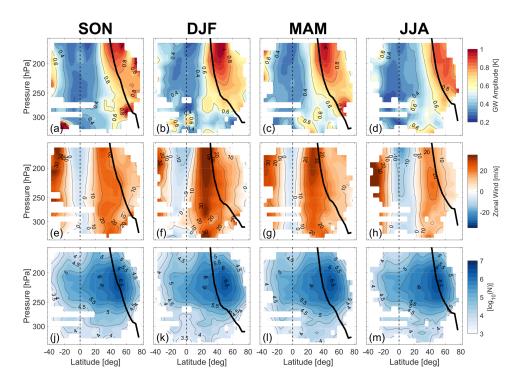
<sup>392</sup> 6 The Effect of the Upper Tropospheric Jet on Observed Gravity Waves

In Section 4 we saw a clear pattern over the North Atlantic of large GW amplitudes to the north of the wind jets. In this Section, we investigate this feature in more depth by examining the height dependence of GW amplitude and wind in the zonal mean.

Figures 9a-d show zonal-mean GW amplitudes, binned onto a grid of height 5 hPa and width 4° latitude. Mean tropopause pressure is indicated for context (thick black lines). This has been weighted spatially and temporally in the same way as the observations, although importantly that there is significant variability in this level. Figures 9e-h show zonal mean zonal winds derived in the same way, while Figures 9j-m show the zonal sum of individual measurements, peaking at around three million (10<sup>6.5</sup>) data points per gridbox.

These figures have been adjusted for the uneven longitude distribution of the data 403 by first gridding onto a  $30^{\circ}$  grid in longitude and then taking the zonal mean (or sum, 404 for Figures 9j-m). However, some bias is unavoidable, and in particular the Pacific Ocean 405 and the southern hemisphere away from routes between Europe and a small number of 406 major airports are significantly underrepresented. Due to highly limited data coverage 407 south of the Equator, we restrict the following discussion to the northern hemisphere; 408 however, we note that based on the limited data available, it appears likely that most 409 of our northern hemisphere conclusions are mirrored at least as far as  $\sim 40^{\circ}$ S. 410

We consider first zonal mean zonal winds, Figures 9e-h. We see a clear jet spanning all measured heights in all four seasons, centred at ~25°N in DJF and ~40°N in JJA. The jet is strongest in DJF and weaker in JJA, consistent with the expected climatological state of the UTLS, and in general increases in speed with height across our data range.



**Figure 9.** Seasonal plots of IAGOS-derived (a-d) zonal mean GW amplitude (e-h) zonal mean zonal wind (j-m) zonal-sum number of measurements for (a,e,j) DJF (b,f,k) MAM (c,g,l) JJA (d,h,m) SON, as a function of (vertical axis) pressure level and (horizontal axis) latitude. Thick black line indicates mean tropospheric pressure level, derived from ERA5 data computed at IAGOS measurements locations.

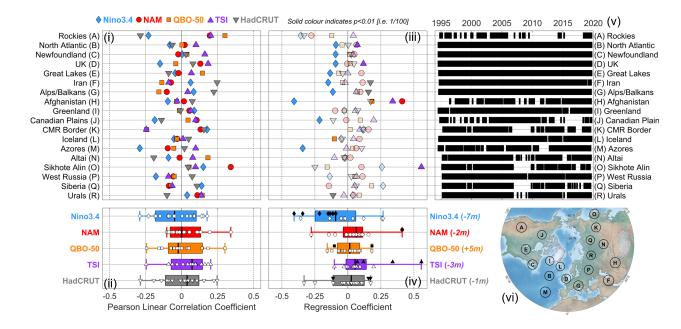


Figure 10. Analyses of possible gravity wave driving mechanisms. (i,ii) Pearson linear correlation coefficient between GW amplitudes for each region defined in Figure 8 and the El Nino 3.4 index (Nino3.4), Northern Annular Mode index (NAM), 50 hPa QBO zonal winds (QBO-50), Total Solar Irradiance (TSI), and the Hadley Central/ Climate Research Unit surface temperature index (HadCRUT). (i) shows each region individually, (ii) shows statistics for all regions combined. Box-and-whisker plots are as described in Figure 8; individual white symbols correspond to the points in panel (i). (iii) Results of a linear regression applied to each region based on the described indices, at a lag described by the y-axis of panel (iv) for each dataset in months. Solid (semi-transparent) symbols show significant (non-significant) results at the p=0.01 level. (iv) As (iii), for all regions combined. White (black) markers indicate (non-) significant results. (v) Temporal coverage for each region, at the monthly scale. Solid black indicates data coverage over the region that month. (vi) Region map, reproduced from Figure 8.

We next examine zonal mean GW amplitudes, Figures 9a-d. Larger GW ampli-416 tudes are clearly visible in the stratosphere, i.e. above the thick black line, and these am-417 plitudes generally increase with height (with minor exceptions). Maximal GW ampli-418 tudes at each latitude are seen slightly poleward of and above the jet in all four seasons. 419 Amplitudes are generally lower in the troposphere and higher in the stratosphere, and 420 outside the tropics they in general grow with altitude, with the exception of some small 421 patches at high latitudes in DJF and a disconnect between a low-altitude near-pole re-422 gion and a higher-altitude midlatitude peak in MAM. There is also evidence of lower-423 altitude maxima at around 40°N and 300 hPa in DJF and MAM, which lie within the 424 jet but are disconnected from the larger regions of high amplitude in the stratosphere. 425

426

# 7 Climate Drivers of Gravity Wave Variability

A unique benefit of these data is the combination of long duration and wide spa-427 tial coverage. Long-term records capable of resolving gravity wave activity over such large spatial regions are rare, and we believe this to be the largest and longest such dataset. 429 As such, our data provide a unique opportunity to assess the role of long-term climate 430 processes in driving or modulating GW activity. 431

In Figure 10 we compare GW amplitude time series from the 18 regions defined 432 in Figure 8 to five climate indices - 'Nino3.4', representing the El-Nino Southern Oscil-433 lation (ENSO), 'NAM', representing the Northern Annular Mode, 'QBO-50', represent-434 ing Quasi Biennial Oscillation-driven equatorial wind speeds at 50 hPa, 'TSI', represent-435 ing the time-varying output of the Sun, and 'HadCRUT', representing long-term changes 436 in surface temperature (Trenberth & Stepaniak, 2001; Ogi et al., 2004; Morice et al., 2012; 437 Coddington et al., 2015). Our HadCRUT4 time series is computed over the Northern 438 Hemisphere only, since all regions lie here and relationships in the southern hemisphere 439 may differ due to the very different relationship between GWs and the southern polar 440 vortex. 441

We first bin our data onto a monthly timescale over the 26 years of data, averaging in all cases over the 500 km radius areas shown in Figure 8 and reproduced in Figure 10vi for context. Figure 10v (top right) shows the temporal coverage available after this binning - most regions are well-covered over the full duration of the dataset, with the possible exception of the Rocky Mountains (region A).

Figures 10i-iv then compare our GW amplitude time series to the five climate indices, using Pearson linear correlation (Figures 10i-ii) and multiple linear regression (Figures 10iii-iv). All time series, both GW and index, have been boxcar-smoothed by three months, except for the rapidly-varying NAM and GW series which we compare to it.

451 7.1 Correlations

We first discuss linear correlations, Figure 10i-ii. Figure 10i shows correlations between each index and GW amplitudes within individual regions. Coloured markers, described by the key at top, show linear correlation between GW amplitudes and each climate index in the region. 10ii compresses these correlations into a single box-and-whisker plot for each index, with individual regions overlaid with markers.

We see extreme scatter, with little consistency between the results and significant 457 inter-region variability for correlations with every metric. Except for the relatively-poorly-458 measured Rocky Mountains (A) and one other point (the NAM and the Sikhote Alin, 459 Region O) we see no correlations exceeding an absolute value of 0.25. There is perhaps 460 a very slight positive skew to the set of correlations between GW amplitudes and TSI, 461 but this is weak and nearly a quarter (5/24) of these correlations are negative. Daily data 462 were also examined, with broadly similar results; these daily data were also lagged over 463 the range -70 to +70 days, but no lag value was found to systematically improve the cor-464 relation with any index over a majority of regions. 465

466

# 7.2 Linear Regression

Figures 10iii-iv show the results of a multiple linear regression applied to the same combinations of GW measurements and indices. As with the correlations, the upper panel shows results for individual regions and the lower panel combined results.

To compute comparable regression coefficients, we normalise each index. For all 470 except HadCRUT, we convert to relative ranges, i.e. a mean of zero and range of one 471 over the period August 1994 to December 2019. For HadCRUT, which is non-cyclical, 472 this conversion is less meaningful, and we keep the original units of Kelvin. This pro-473 duces similar numerical values to the other comparisons since the full range is of order 474 1 K. Thus, for all indices except HadCRUT, a coefficient of +1 suggests that a change 475 476 from the minimum index value over this period to the maximum leads to a 1 K increase in monthly-mean gravity wave amplitude, while for HadCRUT it implies that a 1 K in-477 crease in global-mean surface temperature leads to the same change. 478

We repeat the regression at time lags from -11 months to +11 months, then sum the absolute regression coefficient for each index and maximise to identify the lag which provides the best estimated fit for that index across all regions. This best fit lag is the one displayed in our Figure, and the lag period in months is indicated in the y-axis labels on Figures 10iv, with a positive value indicating that the gravity wave signal leads the index and vice-versa.

We encode the *p*-value of the *t*-statistic for each fit using colour density in the up-485 per panel. Solid-coloured markers are significant at the p < 0.01 level, while semi-transparent 486 markers (the vast majority) are not. In the lower panel this information is encoded with marker colour, where black markers are significant at the 0.01 level. A threshold of 0.01488 is chosen due to the nature of the comparison. A typical threshold of 0.05 would imply 489 that the null hypothesis was false in 1/20 cases just by chance; since we here make 90 490 (5x18) comparisons we would therefore expect 5 of these to be marked as significant re-491 gardless of the actual truth. By using a much lower threshold, we reduce this random 492 effect. This test assumes independence of the measured amplitude series, which is an in-493 herently poor assumption for geophysical data and becomes increasingly less likely to be the case as locations become spatially closer; therefore, these values must be consid-495 ered sceptically. 496

This choice to use a single lag for each index over all regions is made to simplify the presented results, and it may well be the case that some regions have significant and physically meaningful effects at different lags to others; this may therefore be a fruitful pathway for future work.

The most interesting results are for ENSO, shown as blue diamonds. In nine of our 501 eighteen regions we see a statistically-significant link, with an increase in ENSO asso-502 ciated with a drop in GW amplitudes seven months earlier. This reduction is of 0.15-503  $0.4 \,\mathrm{K/cycle}$ , i.e. a maximal El Niño leads to GW amplitudes 25-60% higher than a min-504 imum La Niña, with a seven-month teleconnective lag. All regions which do not exhibit 505 this lag-relationship, excluding the poorly temporally sampled Rockies, are not identi-506 fied as statistically significant and also fall within two distinct geographic groups: Green-507 land/Iceland and Russia (recall from Figure 8 that most Russian regions also have a very 508 different seasonal cycle to the majority of our data). These data therefore represent ev-509 idence of strong teleconnective links between the Pacific Ocean and gravity wave gen-510 eration, propagation and/or filterig in the troposphere and UTLS over North America, 511 Europe, South-Central Asia and the non-Arctic Atlantic Ocean. 512

Other indices produce less clear results, but still present some interesting conclusions, although not as strong as for ENSO. Aside from ENSO, we see ten statisticallysignificant linkages - this is higher than would be expected by chance assuming independence, but since the atmospheric system is inherently highly linked needs to be treated with scepticism.

For the NAM, only one region shows a significant relationship, Afghanistan (H), at a lag of two months. This region is spatially distant from the NAM foci in the North Atlantic, Arctic and Pacific, and a physical mechanism which could link the NAM to this region and no surrounding ones is not immediately obvious. Therefore, we are disinclined to believe this is a meaningful or real teleconnection.

The QBO shows two significant results, but these are in antiphase with each other and spatially disparate. All other results are scattered across regression coefficient space, mostly at relatively small values, and it is thus difficult to conclude that these are truly meaningful signals. Similarly, three regions show a significant relationship with HadCRUT, but again with both positive and negative significant points and with a range of regional values, most clustered near zero.

TSI is slightly more interesting. At a lag of three months, there are four significant 529 points, all positive in sign. 10 of the 14 non-significant points are also positive, as are 530 most correlations presented in Figure 10ii. Finally, a three-month lag is a physically rea-531 sonable time for a teleconnective signal such as this to operate, and since solar input is 532 global a mechanistic pathway to a particular region is not needed, as with the NAM. There-533 fore, although we have insufficient data to draw any firm conclusions, our results may 534 present weak evidence of a positive link between the solar cycle and NH UTLS GWs, 535 with solar maximum corresponding to increased GW amplitudes. 536

#### 537 8 Validation

The height range, spatiotemporal coverage and in-situ nature of our dataset is unusual, and also has several inherent systematic biases. Therefore, it is important to validate our results against other long-term UTLS GW time series.

#### 8.1 United States Radiosondes

Wang and Geller (2003) (WG03) examined GW energy densities derived from high-542 resolution radiosondes over the United States for the period 1998–2001 in two height bands 543 - a tropospheric (2-8.9 km) band below and a stratospheric  $(\sim 18-24.9 \text{ km})$  band above 544 our analysis ( $\sim 10-12 \,\mathrm{km}$  altitude). In their tropospheric band (their Figure 5), WG03 545 saw strong GW activity over the Rockies and western USA throughout the year, over 546 the northwestern US in DJF, and low activity otherwise. In the stratospheric band (their 547 Figure 6), they saw only weak GW activity over the Rockies (except for a slight enhance-548 ment over Wyoming in DJF), but significant activity in the southern and eastern US in 549 DJF, in the south only in MAM, and over Oklahoma in SON. 550

Our maps (Figure 7) share many features with these results. In DJF, we see good 551 agreement, with strong GW activity over the US Rockies and as much as our data cover 552 of the eastern seaboard and southeastern US, consistent with a mixture of the two lay-553 ers of WG03. In MAM, we see GW activity in the western US consistent with their tro-554 pospheric band, but do not see major activity over the southern US (although our data 555 are truncated near this region, and data density is often poor in the areas our data do 556 cover). In JJA/SON our results are less consistent with WG03, with good agreement in 557 most places but lacking their strong signal over the Rockies. 558

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## 8.2 North Atlantic Radiosondes

Plougonven et al. (2003) describe GW measurements from 224 radiosondes launched 560 from ships in the North Atlantic  $(30-70^{\circ}N, 50-0^{\circ}W)$  during the FASTEX campaign in 561 January/February 1997. They focused on short vertical wavelength features  $(<5 \,\mathrm{km})$ , 562 which they studied in the context of the jet up to 20 km. They saw GW activity cen-563 tred on the jet axis in both the stratosphere and troposphere, peaking on the north side 564 in the stratosphere, with a nominal maximum 300 km from jet centre but with extremely 565 large scatter (their Figures 2 and 5). Our peak GW activity is consistently to the north 566 side of the jet but slightly further away, with a maximum between  $5^{\circ}$  and  $10^{\circ}$  north of 567 the jet core. However, given the very high degree of scatter in the data of Plougonven 568 et al. (2003), these conclusions are not inconsistent. 569

570

# 8.3 Greenland Aircraft Turbulence Encounters

Lane et al. (2009) describe statistics of aircraft turbulence encounters over Greenland, derived from aircraft instrument reports. These events primarily arise due to interactions between GWs emitted from Greenland's orography and background directional wind shear. They see a near-annual cycle in these reports, with peaks in November and

January and a trough in May. The winter peaks and May trough are consistent with our 575 data (Figure 8I), but we also see a secondary maximum in August/September which they 576 do not. 577

8.4 Aberystwyth MST Radar 578

Vaughan and Worthington (2007) studied inertia-GWs using eight years of data 579 from the MST Radar in Aberystwyth, Wales (within our region D). They observed a win-580 ter peak in occurrence rate. While these results carry the important caveat that they 581 explicitly aimed to exclude the orographic waves that likely form a significant propor-582 tion of our data in this region, their results were again consistent with ours. 583

#### 8.5 European Satellite Cloud Imagery

Cruette (1976) studied lee waves patterns in satellite cloud imagery from the mid-585 troposphere in western Europe over three years of data. She saw a strong seasonal pat-586 tern over an area including the UK (region D) and western Alps (edge of our region G), 587 with a peak in winter and trough in summer, again consistent with our results. 588

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# 8.6 CHAMP Estimates of Russian Lower-Stratospheric GW Seasonality

Independent estimates of UTLS GW seasonality over Russia and Central Asia (re-591 gions N-R) have proven challenging to locate in the literature. This is a major issue, be-592 cause the observed GW seasonality in our data in these regions is opposite to most other 593 regions and to a more general assumption that GW amplitudes will be on average higher 594 in winter due to the nature of most source mechanisms. 595

Some very limited evidence to support the observed seasonality is seen in Namboothiri 596 et al. (2008), where CHAMP-derived GW variances at 20 km altitude for 2002 exhibit 597 larger values in JJA than DJF by a factor of 4, but from a low base. However, this ev-598 idence is from an experimental instrument type over only a short duration, and longer-599 term measurements at  $30 \,\mathrm{km}$  altitude from the more modern HIRDLS and SABER in-600 struments (Ern et al., 2018) do not show this, peaking instead in January. 601

We note clearly that the observed seasonality could be correct. A plausible mech-602 anism, for example, would be non-orographic generation from summer storms which, deep 603 in the continent, will not be modulated strongly by the North Atlantic jet stream. The 604 annual cycle of rainfall in this region is consistent with this mechanism, and filtering, hor-605 izontal propagation, or observational filter differences could easily explain the lack of ob-606 served signal at 30 km. However, given the lack of validation, it is important to remain 607 sceptical of these specific results until further evidence is available. 608

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# 9 Implications of Strong GW Activity over the Rockies

In our data (e.g. Figure 7), we see large amplitudes over the Rocky Mountains (re-610 gion A). However, similar activity is not usually observed in stratospheric satellite data, 611 including measurements at altitudes as low as  $20 \,\mathrm{km}$  altitude ( $\sim 50 \,\mathrm{hPa}$ ) made using HIRDLS/Aura 612 (e.g. Geller et al., 2013; Wright et al., 2015; Ern et al., 2018). While poorly temporally 613 sampled in the early 2010s (Figure 10), the total number of measurements over this re-614 gion in earlier years is sufficient to properly characterise the seasonal cycle (Figures 5y-615 B), and thus discussion of this mismatch is a meaningful question. 616

The specific comparison to HIRDLS is important. Over the Rockies, IAGOS air-617 craft usually travel near-meridionally, as long-haul flights to Europe dominate the dataset. 618 As such, their along-track vector is very similar to Aura, which has a polar orbit at  $98^{\circ}$ 619

inclination. Alternate HIRDLS profile-pairs are spaced  $\sim 40$  km apart at 20 km altitude (e.g. Figure 4b of Wright et al., 2015), and thus there should be some degree of observational filter overlap between the shortest waves visible to HIRDLS and longer waves seen by IAGOS, which have mean wavelengths in this region of  $\sim 150-250$  km with a standard deviation (not shown) of  $\sim 100$  km.

However, HIRDLS data show neither enhanced GW activity (e.g. Geller et al., 2013; 625 Wright et al., 2015; Ern et al., 2018) over the Rockies, nor the intermittency character-626 istic of orographic wave sources in the statosphere (Wright et al., 2013), even when fil-627 tered to identify only the shortest accessible lengthscales (Figure 7 of Wright et al., 2015). 628 Assuming the waves seen here are orographic and therefore propagate mostly vertically, 629 which seems likely given their geographic location and relatively large amplitudes, this 630 suggests three possibilities: (a) that the waves break between  $\sim 200 \text{ hPa} - 50 \text{ hPa}$  ( $\sim 10 -$ 631 20 km altitude), (b) that the waves propagate into the stratosphere, but only have ex-632 tremely short horizontal wavelengths and are thus invisible to HIRDLS, or (c) that the 633 waves are invisible to HIRDLS due either to their phase fronts being aligned along the 634 instrument boresight (e.g. Figure 1 of Wright & Hindley, 2018) or due to having verti-635 cal wavelengths  $< \sim 2-3$  km. 636

Option (b) is unlikely as observations with AIRS/Aqua, which measures short-horizontal long-vertical wavelength waves, also show no significant stratospheric wave activity over the Rockies (Hoffmann et al., 2013). This restricts the possible wavelengths of any waves propagating into the stratosphere to small scales in both the vertical and horizontal. Also, our median horizontal wavelength is in the HIRDLS observational filter, so at least some signal would be expected to be visible unless there was a dramatic shift in horizontal wavelength between the two heights.

<sup>644</sup> Option (c) is also unlikely as the HIRDLS boresight, which points at 47° off the <sup>645</sup> orbital track, will be aligned at a high angle to the zonally-aligned waves the relative ori-<sup>646</sup> entations of the winds and mountain ridge in this region would imply (Bacmeister, 1993), <sup>647</sup> while short vertical wavelengths would preclude sharp vertical ascent.

It is therefore likely that the waves we observe here break in the UTLS, driving the winds specifically in this geographic and height region. This contrasts with many models (Geller et al., 2013), which propagate significant quantities of wave momentum flux into the statosphere over the Rockies.

## 652 10 Discussion

The data used have significant, but known, biases and deficiencies. The biggest such 653 deficiency is that the data are one-dimensional, lacking vertical and cross-track spatial 654 information. This prevents us from estimating momentum fluxes and other important 655 GW properties, and complicates interpretation of our results. While we do not do so here, 656 the combination of horizontal wavelength and wind speed can in principle be used to es-657 timate vertical wavelength provided very significant assumptions are made (e.g. M. J. Alexan-658 der & Grimsdell, 2013), and this may be useful for future work. The lack of cross-track 659 information is a problem common to stratospheric limb sounding studies of GWs which 660 is harder to adjust for, and therefore we have in general avoided discussion of horizon-661 tal wavelengths, except in Section 9 where the relative orientations of the measurements 662 platforms being compared are known. 663

Biases include flight routing around storms, which may explain the lack of subtropical amplitude maxima in our summer data, a bias towards shorter wavelengths at the start of our routes due to spectral edge truncation, limited spatiotemporal coverage outside major flight paths, and a preference for flights to fly in the jet eastwards and away from the jet westwards leading to a likely positive skew over the North Atlantic in measured wave amplitudes. These biases are discussed in the text where appropriate. With the exception of the storm routing bias and to a lesser degree the jet-flight bias, these are likely to be systematic across the dataset rather than seasonally-varying, and thus should not significantly affect our time-series analyses for most regions.

<sup>673</sup> 11 Summary and Conclusions

We have analysed 26 years of UTLS GW measurements derived from commercial flight data, primarily over the northern hemisphere. This dataset is uniquely long for a large area dataset suitable for GW research, allowing us to study important effects inaccessible with other tools.

We show regionally-varying seasonal cycles, with regional similarities primarily in the zonal direction. Our results include strong GW signals associated with orographic sources which are not seen in stratospheric satellite data. The strongest and most consistent GW amplitudes are seen above and to the north of the upper tropospheric jet, and over orographic hotspots such as south-east Greenland, East Asia and the Rocky Mountains.

Analysis of the relationship between our long-timeseries data and multiple climatesystem indices show no significant relationship with the QBO, climate change or the NAM. There may be weak evidence of a link with solar output 3 months earlier, although this is not statistically significant in most regions.

We do however see statistically-significant evidence in all studied regions other than 688 Russia and the Arctic of a seven-month-lagged link with El Niño, with positive-phase 689 ENSO leading to an increase in GW amplitudes in these regions of order 25%. A qual-690 itative relationship between ENSO and UTLS GWs was previously suggested by Wang 691 and Geller (2003) on the basis on four years of United States radiosonde data as a pos-692 sible explanation for differences between their annual distributions, but here we advance 693 significantly upon this earlier work by (a) demonstrating the link over a much larger area, 694 (b) showing that it is statistically significant, and (c) and quantifying the change it in-695 duces in the wave field. 696

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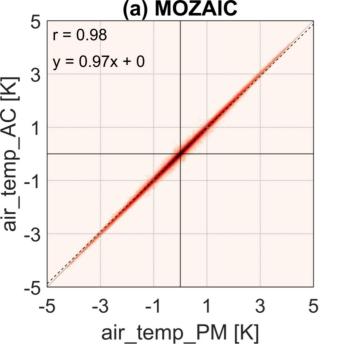
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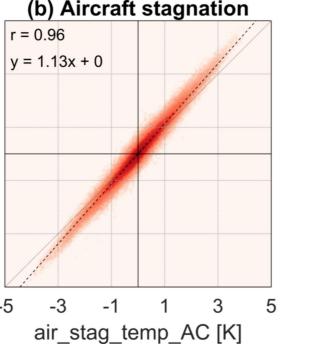
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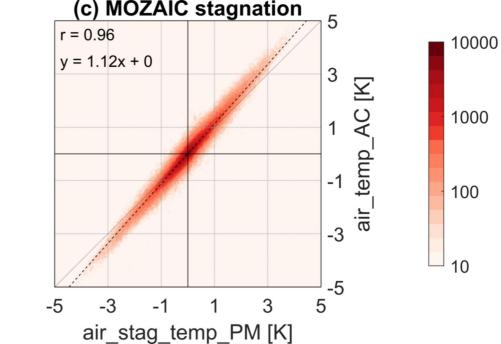
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Figure 1.







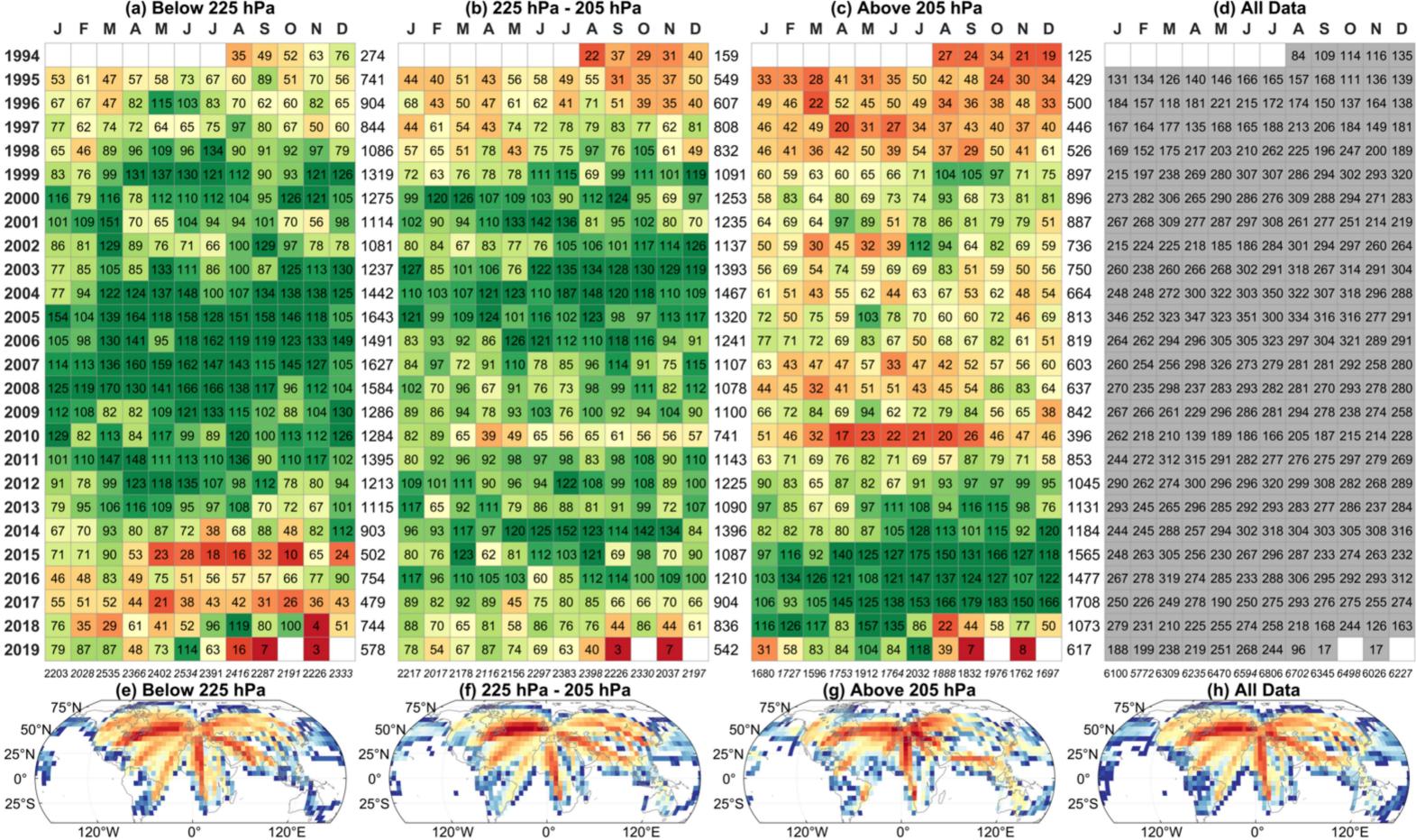
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Figure 2.



(d) All Data																
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97	215	197	238	269	280	307	307	286	294	302	293	320	3307		120	
96	273	282	306	265	290	286	276	309	288	294	271	283	3424			
87	267	268	309	277	287	297	308	261	277	251	214	219	3236		100	
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03	260	254	256	298	326	273	279	281	281	292	258	280	3336		60	s of
37	270	235	298	237	283	293	282	281	270	293	278	280	3300			and
42	267	266	261	229	296	286	281	294	278	238	274	258	3228	_	40	Thousands
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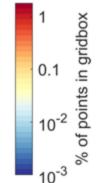
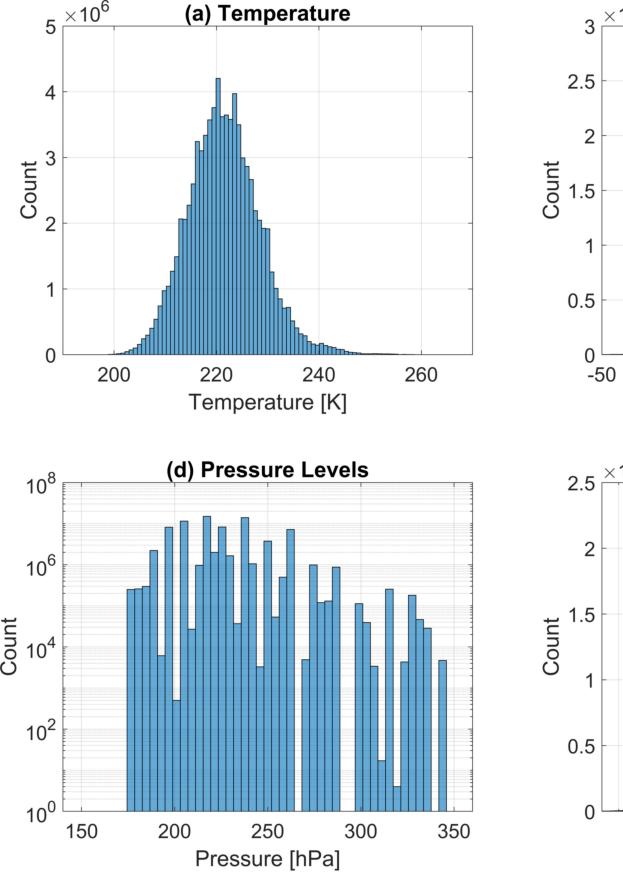
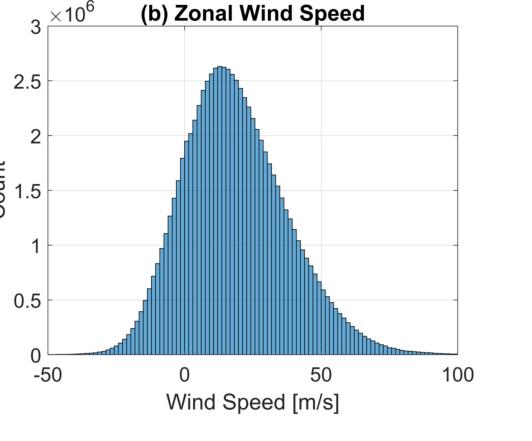
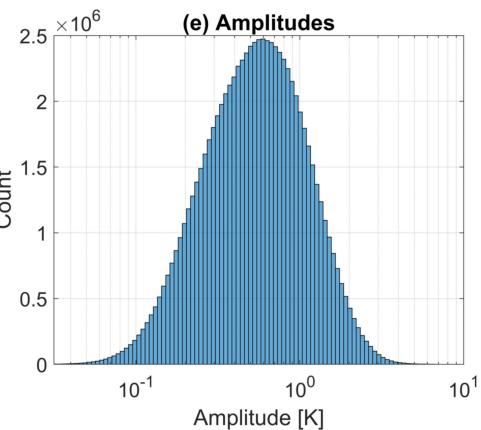
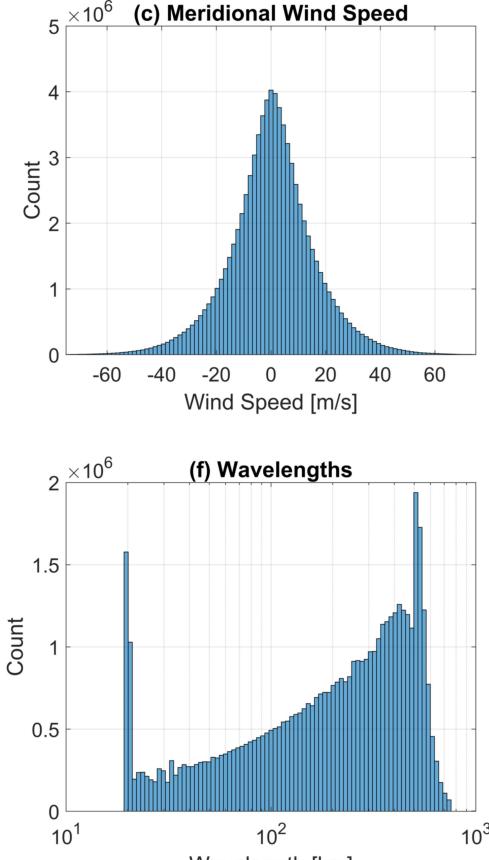


Figure 3.



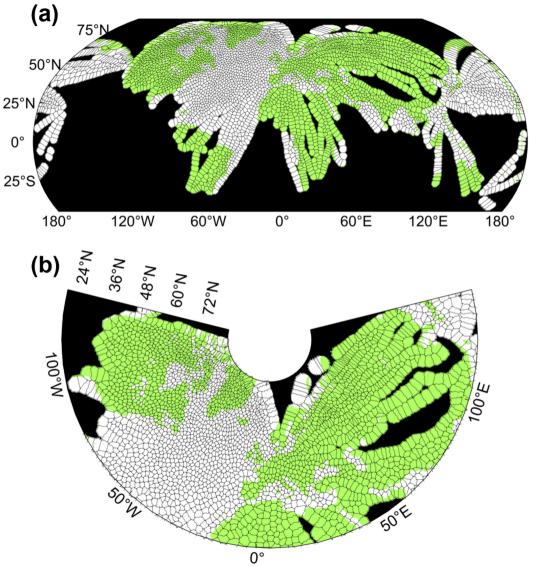






Wavelength [km]

Figure 4.



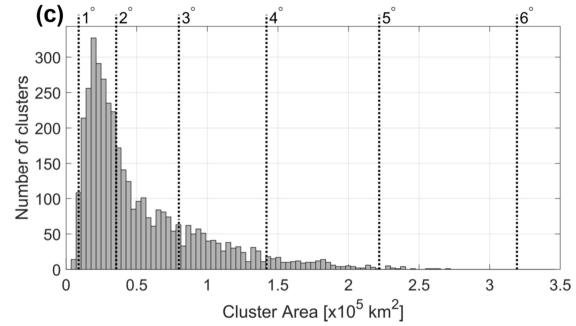


Figure 5.

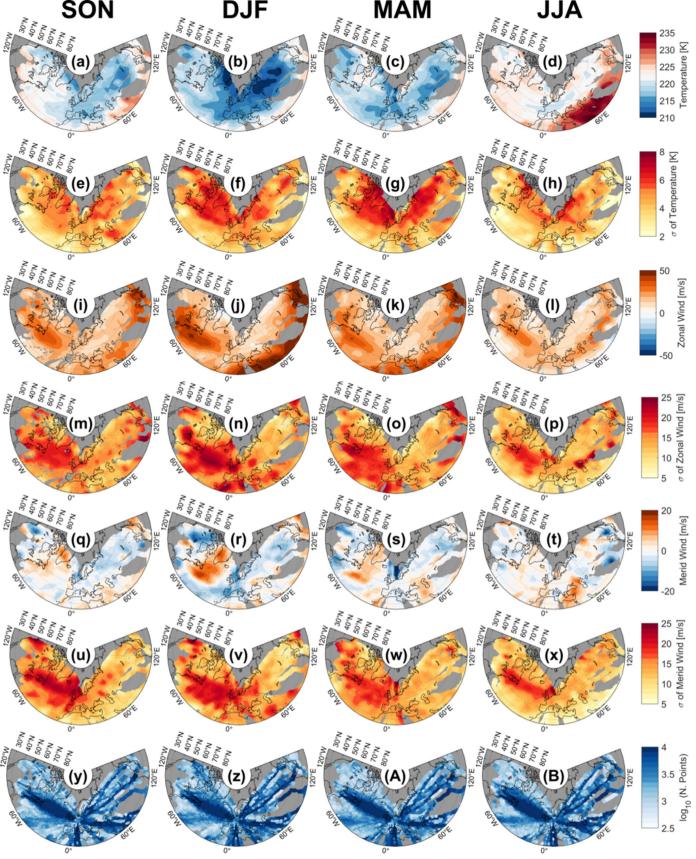


Figure 6.

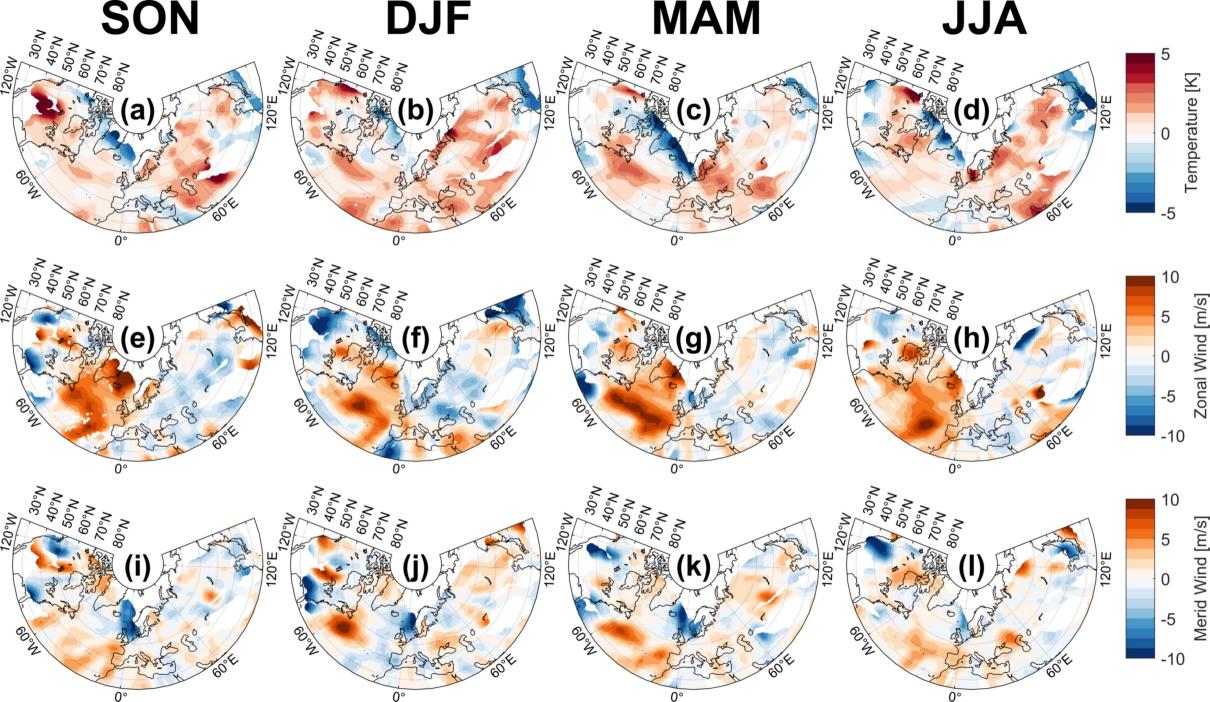


Figure 7.

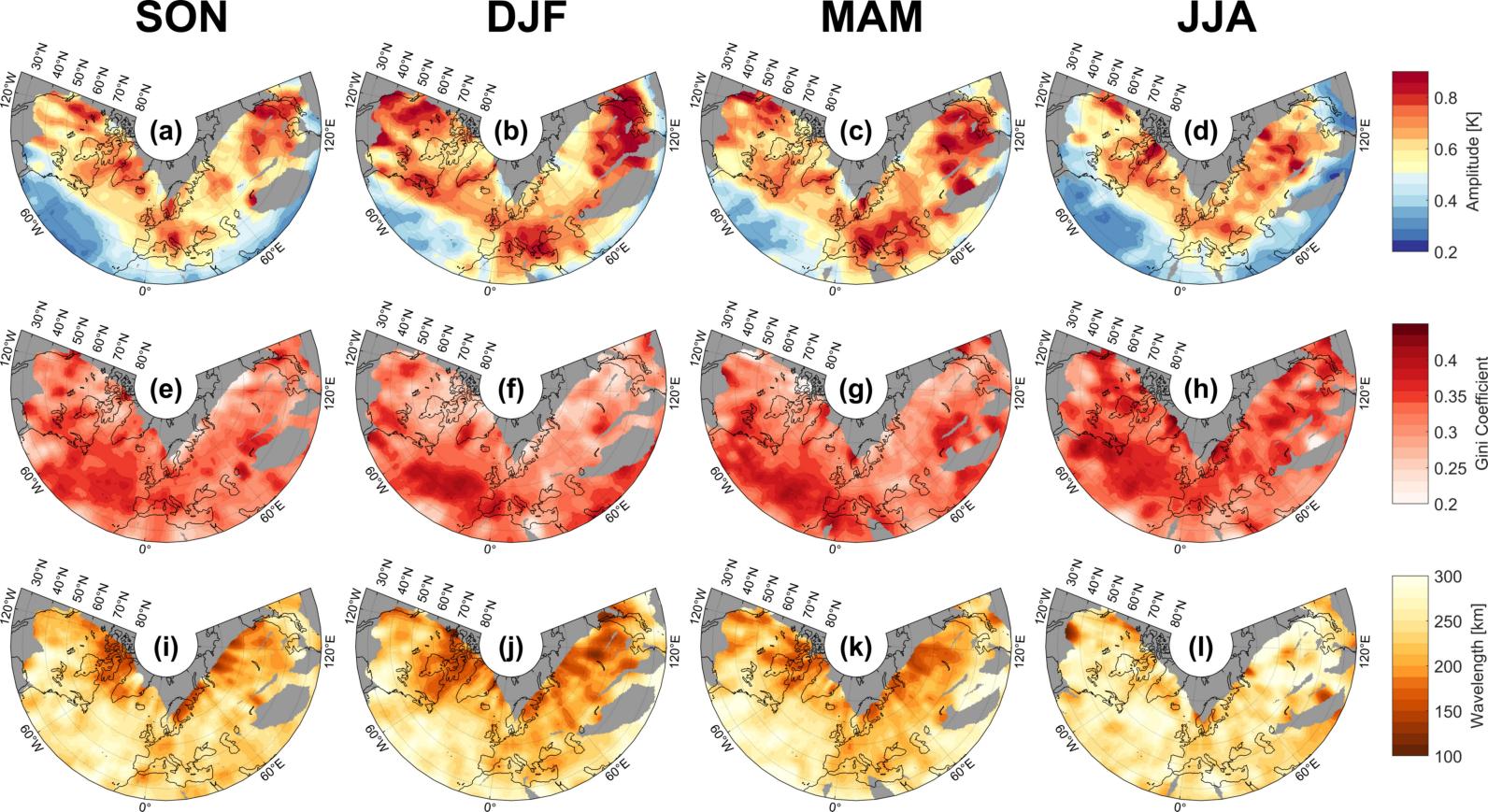
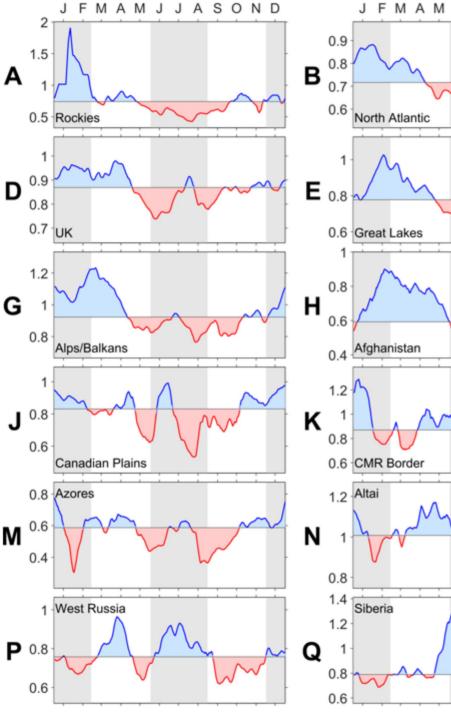
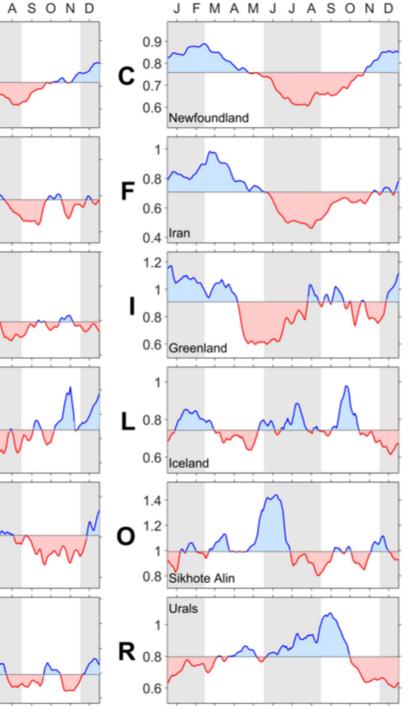




Figure 8.





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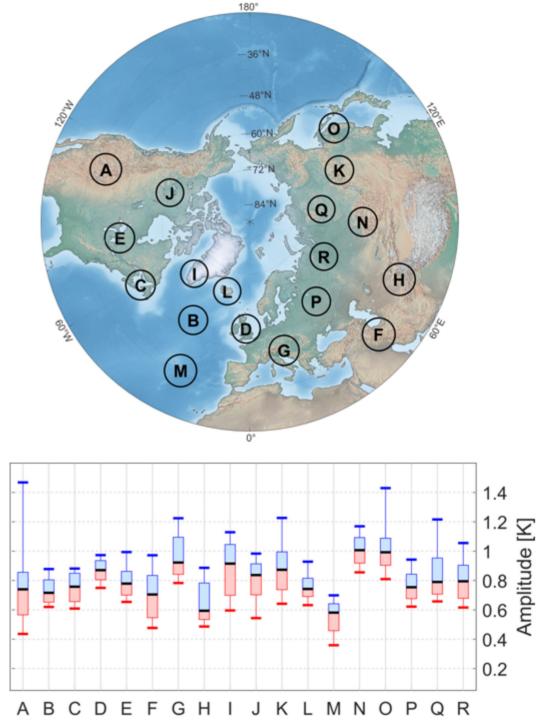


Figure 9.

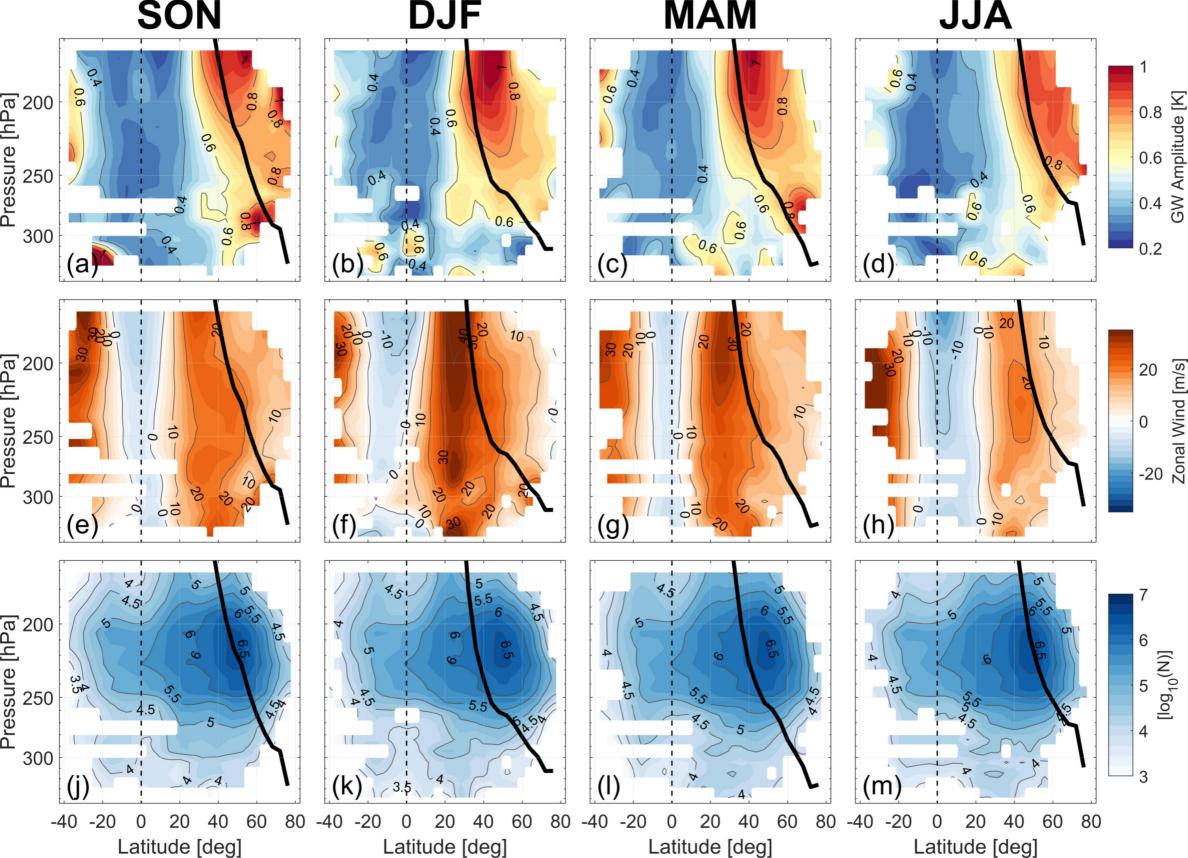


Figure 10.

