Global models from sparse data: A robust estimate of Earth's residual topography spectrum

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Abstract

A significant component of Earth's surface topography is maintained by stresses induced by underlying mantle flow. This 'dynamic' topography cannot be directly observed, but it can be approximated — particularly at longer wavelengths — from measurements of residual topography, which are obtained by removing isostatic effects from the observed topography. However, as these measurements are made at discrete, unevenly-distributed locations on Earth's surface, inferences about global properties can be challenging. In this paper, we present and apply a new approach to transforming point-wise measurements into a continuous global representation. The approach, based upon the statistical theory of Gaussian Processes, is markedly more stable than existing approaches — especially for small datasets. We are therefore able to infer the spatial pattern, wavelength and amplitude of residual topography using only the highest-quality oceanic spot measurements within the database of Hoggard et al. (2017). Our results indicate that the associated spherical harmonic power spectrum peaks at l=2, with power likely in the range 0.46-0.76 km². This decreases by over an order of magnitude to around 0.02 km²2 at l=30. Around 85% of the total power is concentrated in degrees 1-3. Our results therefore confirm previous findings: Earth's residual topography expression is principally driven by deep mantle flow, but shallow processes are also crucial in explaining the general form of the power spectrum. Finally, our approach allows us to determine the locations where collection of new data would most impact our knowledge of the spectrum.

Global Models from Sparse Data: A Robust Estimate of Earth's Residual Topography Spectrum

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Key	Points:
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7	•	We present a new method for spectral analysis of sparse point data and apply this to
8		measurements of residual topography within the oceans;
9	•	Using a smaller, more reliable dataset than earlier studies, we confirm a power spec-
10		trum that is dominated by long-wavelength components;
11	•	We also confirm that the spectrum contains significant power at shorter wavelengths,

• We also confirm that the spectrum contains significant power at shorter wavelengths, reflecting the multiscale nature of global mantle flow.

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13 Abstract

A significant component of Earth's surface topography is maintained by stresses induced 14 by underlying mantle flow. This 'dynamic' topography cannot be directly observed, but 15 it can be approximated — particularly at longer wavelengths — from measurements of 16 residual topography, which are obtained by removing isostatic effects from the observed 17 topography. However, as these measurements are made at discrete, unevenly-distributed 18 locations on Earth's surface, inferences about global properties can be challenging. In this 19 paper, we present and apply a new approach to transforming point-wise measurements 20 into a continuous global representation. The approach, based upon the statistical theory 21 of Gaussian Processes, is markedly more stable than existing approaches — especially for 22 small datasets. We are therefore able to infer the spatial pattern, wavelength and amplitude 23 of residual topography using only the highest-quality oceanic spot measurements within 24 the database of Hoggard et al. (2017). Our results indicate that the associated spherical 25 harmonic power spectrum peaks at l = 2, with power likely in the range 0.46–0.76 km². 26 This decreases by over an order of magnitude to around $0.02 \,\mathrm{km^2}$ at l = 30. Around 85% 27 of the total power is concentrated in degrees 1–3. Our results therefore confirm previous 28 findings: Earth's residual topography expression is principally driven by deep mantle flow, 29 but shallow processes are also crucial in explaining the general form of the power spectrum. 30 Finally, our approach allows us to determine the locations where collection of new data 31 would most impact our knowledge of the spectrum. 32

³³ Plain Language Summary

As the mantle flows, it induces deformation at Earth's surface, pushing it up at some 34 locations and pulling it down elsewhere. This deformation can be quantified by measuring 35 so-called 'residual topography', at specific locations. However, we only have a small number 36 of such measurements, and they are clustered in particular areas on Earth's surface: inferring 37 a global representation of residual topography is therefore a challenge. In this paper, we 38 develop a new method for deriving robust global maps from such data, and apply this to a 39 set of measurements of residual topography. An advantage of our method is that it remains 40 effective for small datasets, enabling a more conservative approach to data selection. Our 41 results complement recent studies on residual topography, highlighting how mantle dynamics 42 shapes the surface of our planet across a range of scales. 43

44 **1** Introduction

Earth's surface topography arises as a balance between processes that create elevation, 45 such as tectonic convergence, and those that destroy it, including erosion, sediment transport 46 and deposition. Most topography is *isostatic*, being maintained by thickness and density 47 variations within the crust and lithospheric mantle. However, there is growing consensus 48 that a substantial proportion is controlled by mantle convection: as the mantle flows, it 49 transmits normal stresses to the lithosphere, and these are balanced by gravitational stresses 50 arising through topographic deflections of Earth's surface (e.g. Pekeris, 1935; Parsons & 51 Daly, 1983; Hager et al., 1985; Hager & Richards, 1989; Mitrovica et al., 1989; Gurnis, 1993; 52 Lithgow-Bertelloni & Silver, 1998; Gurnis et al., 2000; Conrad & Husson, 2009; Braun, 2010; 53 Shephard et al., 2010; Flament et al., 2013; Hoggard et al., 2016; Yang & Gurnis, 2016; Rubey et al., 2017; Eakin & Lithgow-Bertelloni, 2018). This so-called *dynamic* topography 55 is transient, varying both spatially and temporally in response to underlying mantle flow. 56 As a result, it is challenging to isolate — but doing so can provide important constraints 57 upon the flow regime, planform and intensity of mantle convection. 58

⁵⁹ Observational constraints on dynamic topography are indirect: estimates of its spatial ⁶⁰ pattern, wavelength and amplitude are generally equated to so-called *residual* topography, ⁶¹ which is calculated by removing the isostatic contribution of sediments, ice, crust and litho-⁶² sphere from the observed topography (e.g. Panasyuk & Hager, 2000; Kaban et al., 2003;

Flament et al., 2013; Guerri et al., 2016; Hoggard et al., 2016, 2017). In practice, residual 63 topography provides an upper bound on flow-related dynamic topography (e.g. Hoggard 64 et al., 2017). Indeed, whilst the long-wavelength components of residual topography are 65 principally dynamic in origin, a significant proportion of shorter-wavelength oceanic 'resid-66 ual' topography arises from isostatic effects associated with local variations in lithospheric 67 thickness and density, that cannot be accounted for through a simple plate cooling model 68 (Davies et al., 2019). Despite this, measurements of residual topography are currently our 69 best source of information on dynamic topography. 70

71 Accurate measurements of residual topography depend upon a careful synthesis of information from a variety of sources — including models for surface elevation, sediment 72 cover, crustal properties, and thermal effects. Inevitably, the availability and quality of such 73 information is not uniform across Earth's surface. A recent database of residual topography 74 within the world's oceans, compiled by Hoggard et al. (2016, 2017), provides one of the 75 most comprehensive datasets currently available. This has underpinned a series of studies 76 into the spectral character of residual topography and its relationship to underlying mantle 77 dynamics (e.g. Hoggard et al., 2016, 2017; Steinberger, 2016; Yang & Gurnis, 2016; Yang 78 et al., 2017; Watkins & Conrad, 2018; Steinberger et al., 2019). However, the conclusions 79 from these studies have often appeared contradictory. At least in part, this has been due 80 to authors making different subjective choices within their analyses, with the debate being 81 further complicated by a lack of consistency around the spherical harmonic normalisation 82 conventions employed by different groups. 83

In an effort to reduce subjectivity in the analysis, a recent study by Davies et al. 84 (2019) considered the database of Hoggard et al. (2017) using a novel hierarchical Bayesian 85 approach (Valentine & Sambridge, 2018). In particular, a strategy of 'Automatic Relevance 86 Determination' (ARD) was employed to avoid the need to impose any pre-determined notion 87 of 'smoothness' upon the residual topography. Results indicate that the spectrum was 88 dominated by long-wavelength ($\sim 10^4$ km) features, with power dropping by about an order 89 of magnitude at shorter wavelengths ($\sim 10^3$ km). Based upon predictions from instantaneous 90 models of global mantle dynamics, Davies et al. (2019) demonstrate that both deep mantle 91 flow and shallow, lithosphere-controlled processes are important in generating this surface 92 response. 93

However, one drawback in the ARD approach employed by Davies et al. (2019) is that 94 it becomes unstable as the number of data points is reduced. Within the database compiled 95 by Hoggard et al. (2016, 2017), a subset of around 5% of the measurements are considered 96 markedly more robust than the rest. Ideally, it would be possible to analyse those points in 97 isolation — but this is not possible using the ARD strategy. The present paper, therefore, 98 develops an alternative approach to analysis, building on the statistical theory of Gaussian 99 Processes (e.g. Rasmussen & Williams, 2006) and the work of Valentine and Sambridge 100 (2020a, 2020b). This method remains stable for even the smallest datasets, and allows us to 101 convincingly demonstrate that the conclusions of Davies et al. (2019) remain valid when only 102 the highest-quality data points are used. We are also able to obtain maps showing where 103 new measurements would be most valuable in constraining the residual topography profile. 104 While the present paper focusses only on residual topography, our underlying method has 105 much wider application, and may be useful for any attempt to characterise continuous Earth 106 properties from discrete samples. 107

We begin by providing a brief summary of the observational datasets used in this study, which are described more fully elsewhere. We then summarise how Gaussian processes can be used to represent functions on Earth's surface, building on the presentation in Valentine and Sambridge (2020a, 2020b), and develop the mathematical results necessary for converting the recovered model into a spherical harmonic representation. Finally, we show that application of this theory to the observational data yields results that are self-consistent across different data subsets, and in general agreement with earlier studies.

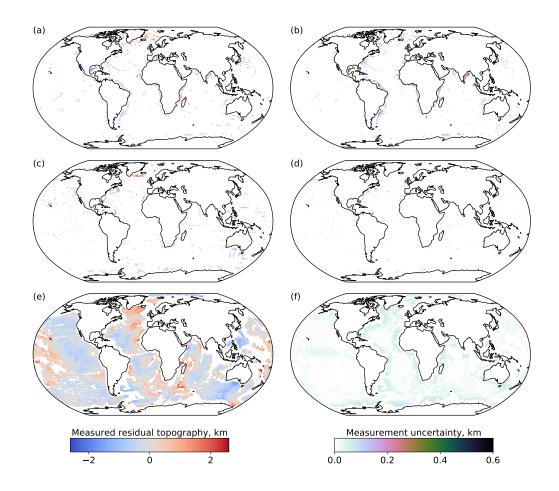


Figure 1. Measurements of residual topography. 1160 high-accuracy spot measurements are shown in (a), with associated uncertainties in (b). A further 870 spot measurements that lack crustal corrections are shown in (c), with uncertainties in (d): these are shown without the additional 200 m uncertainty assessed by Hoggard et al. (2017) to reflect the absence of crustal information. Finally, (e) shows 20767 gridded measurements derived from shiptrack bathymetry; again, the uncertainties depicted in (f) do not include any additional uncertainty to reflect the lack of crustal corrections.

115 2 Datasets

The database employed for this study is described in Davies et al. (2019), and is a 116 compilation of residual topography measurements within the world's oceans. It represents 117 an incremental update to the database of Hoggard et al. (2017), which itself builds on several 118 previous studies (e.g. Winterbourne et al., 2009; Czarnota et al., 2013, 2014; Winterbourne 119 et al., 2014; Hoggard et al., 2016). In comparison to Hoggard et al. (2017), the version 120 accompanying Davies et al. (2019) incorporates information from additional marine seismic 121 surveys, and has undergone further quality control. For the purposes of this paper, it is 122 appropriate to divide the dataset into three classes, and we discuss each in turn. 123

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2.1 High-accuracy spot measurements

These measurements are derived from marine seismic surveys, allowing direct observation of the topography associated with the oceanic crust. Careful processing allows isostatic effects associated with sedimentary loading and crustal thickness to be removed, and agedepth cooling is accounted for using a simple analytical plate model (Hoggard et al., 2017).
A total of 1160 residual topography observations are obtained (see Fig. 1a), with associated uncertainty (Fig. 1b). We consider these points to be the most robust within the database, and the primary goal of this paper is to ascertain the extent to which the power spectrum of residual topography can be constrained using only these measurements.

2.2 Lower-accuracy spot measurements

A further 870 points, shown in Fig. 1(c), were obtained in a similar manner — but a 134 lack of information prevented isostatic correction for variations in crustal thickness. These 135 observations must therefore be assumed to be less accurate. To reflect this, Hoggard et 136 al. (2017) increases the uncertainties reported for these data points by $0.2 \,\mathrm{km}$, a somewhat 137 arbitrary figure chosen based on the median crustal correction associated with the high-138 accuracy data points. In the present paper, we replace this additional 0.2 km component by 139 an unknown correction, Δ . We can then determine the value Δ should take to ensure sta-140 tistical consistency with information from the high-accuracy data points. The un-corrected 141 uncertainties (i.e. without the additional $0.2 \,\mathrm{km}$ component) are shown in Fig. 1(d). 142

2.3 Shiptrack-derived measurements

Finally, the database contains 20767 measurements derived from a global shiptrack 144 bathymetry grid (Smith & Sandwell, 1997), with sedimentary corrections based on the 145 model of Laske and Masters (1997). Again, Hoggard et al. (2017) increases the uncertain-146 ties associated with these measurements by $0.2 \,\mathrm{km}$, to reflect the lack of detailed crustal 147 information; again, we replace this with the unknown correction Δ . These data points are 148 illustrated in Fig. 1(e), with uncorrected uncertainties in Fig. 1(f). Clearly, these points 149 provide excellent spatial coverage within the oceans, but the use of global datasets, which 150 are themselves derived from a variety of sources, raises the possibility of significant un-151 quantified systematic biases within this portion of the dataset. We therefore regard the 152 shiptrack-derived measurements as being the least robust within the database. 153

154 **3** Methodology

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The fundamental challenge in this work — and in many other geoscience questions 155 is to infer a continuous spatial function (i.e. the residual topography at any position on 156 Earth's surface) from a finite set of observations made at discrete locations. A variety of 157 approaches are possible, but one common strategy is to assume that the spatial function 158 can be expanded in terms of a finite set of basis functions. This transforms the inference 159 task into one of determining the expansion coefficients relative to this basis, typically as a 160 least-squares inversion problem. This is the approach adopted by Hoggard et al. (2016) and 161 Davies et al. (2019), using a spherical harmonic basis. However, two key difficulties emerge. 162 First, it is likely that the true spatial function contains features that cannot be represented 163 using the finite basis. This can lead to an aliasing-like effect known as 'spectral leakage' 164 (Trampert & Snieder, 1996), biasing the recovered expansion coefficients away from their 165 correct values. While it is possible to correct for this, it is computationally expensive to 166 do so. Instead, Davies et al. (2019) adopted a pragmatic approach: since spectral leakage 167 mainly affects the shortest-wavelength components within the expansion, these terms were 168 omitted from further analysis. 169

The second difficulty arises because with sparse, unevenly-distributed data, not all components of the basis function expansion are equally-well constrained. Typically, some expansion coefficients can be varied significantly without affecting the fit to data, which can cause the interpolated function to behave unrealistically in regions of low data coverage. To address this, one typically introduces 'regularisation': additional constraints on the basis function expansion, designed to ensure that a well-behaved solution can be found. In

Hoggard et al. (2016), the chosen constraints amounted to a requirement that the residual 176 topography field be 'small' and 'smooth'. In general, these are reasonable requirements-177 but they carry implications for the spectral characteristics of the recovered field. Thus, one 178 could question whether the power spectrum from Hoggard et al. (2016) was truly mandated 179 by the data, or if it had arisen as a consequence of the assumptions inherent to their analy-180 ses. As noted above, to overcome this potential shortcoming, Davies et al. (2019) employed 181 a novel approach to regularisation, developed in Valentine and Sambridge (2018), based on 182 the idea of 'Automatic Relevance Determination' (ARD) (Mackay, 1992). In the present 183 context, this enables appropriate 'smoothness' characteristics to be determined from the 184 data, rather than being imposed from the outset. However, application of this approach 185 requires a relatively large dataset and, as a result, the ARD analysis can only be applied to 186 the full residual topography dataset (i.e. that incorporating all spot- and shiptrack-derived 187 measurements); attempts to apply it to only the high-accuracy spot measurements prove 188 unstable. 189

Subsequently, Valentine and Sambridge (2020a, 2020b) set out an alternative framework for inferring continuous functions from point data, based on the statistical theory of Gaussian Processes. This avoids a number of the difficulties associated with the basisfunction–expansion approach, and remains stable for small datasets. Thus, unlike ARD, it can be used to analyse the high-accuracy spot data from Hoggard et al. (2017) in isolation, as well as their dataset incorporating ship-track derived constraints.

3.1 A Gaussian Process model for residual topography

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Gaussian Processes are a class of statistical model that have become an increasingly popular tool for machine learning. They are discussed at length in, for example, Rasmussen and Williams (2006) or Murphy (2012), and we do not attempt to provide a comprehensive introduction here. Instead, we outline the key features, building on the development and notation adopted in Section 2 of Valentine and Sambridge (2020a). Readers are encouraged to consult that paper if further details are required.

We wish to develop a model for the residual topography field, which we denote $h(\theta, \varphi)$, 203 as a function of latitude (θ) and longitude (φ). To simplify notation, we introduce the posi-204 tion vector $\mathbf{x} = (\theta, \varphi)$, and will use $h(\mathbf{x})$ interchangeably with $h(\theta, \varphi)$. Whereas expressing 205 $h(\mathbf{x})$ using a basis function expansion requires us to choose the set of basis functions and 206 an approach to regularisation, expressing it as a Gaussian Process requires us to specify a 207 'mean function' and a 'covariance function'. The mean function, $\mu(\mathbf{x})$, describes our *a priori* 208 statistical assumptions about the expected value of $h(\mathbf{x})$; for present purposes, we assume 209 that this has a constant value everywhere, 210

$$\mu(\mathbf{x}) = \mu_0 \,. \tag{1}$$

The covariance function, $k(\mathbf{x}, \mathbf{x}')$ describes our *a priori* assumptions about the covariance 211 between $h(\mathbf{x})$ and $h(\mathbf{x}')$. Put in more intuitive terms: $k(\mathbf{x}, \mathbf{x}')$ describes how learning the 212 value of h at some point, x, would influence our beliefs about the value of h at some other 213 point, \mathbf{x}' . In the present case, it is reasonable to assume that the degree of influence should 214 diminish as the distance between \mathbf{x} and \mathbf{x}' grows: we might expect residual topography to 215 be similar for two localities 1 km apart, but knowledge of residual topography in the North 216 Atlantic would not help us constrain residual topography in Australia. In this paper, we 217 adopt the Matérn family of covariance functions, for which 218

$$k(\mathbf{x}, \mathbf{x}') = \sigma_1^2 \frac{2^{1-\nu}}{\Gamma(\nu)} \left(\frac{\sqrt{2\nu} \, d(\mathbf{x}, \mathbf{x}')}{\sigma_2} \right)^{\nu} K_{\nu} \left(\frac{\sqrt{2\nu} \, d(\mathbf{x}, \mathbf{x}')}{\sigma_2} \right) \,, \tag{2}$$

where Γ denotes the Gamma function, and K_{ν} is a modified Bessel function of the second kind. Here, $d(\mathbf{x}, \mathbf{x}')$ is some measure of the distance between the points \mathbf{x} and \mathbf{x}' ; for reasons which will later become apparent, we employ the epicentral angle, Θ , defined such that

$$d(\theta,\varphi;\theta',\varphi') \equiv \Theta(\theta,\varphi;\theta',\varphi') = \arccos\left(\sin\theta\sin\theta' + \cos\theta\cos\theta'\cos(\varphi-\varphi')\right).$$
(3)

Note that this can be converted into the great-circle distance on the surface of the (assumed 222 spherical) Earth by multiplying by Earth's radius. The quantities σ_1 , σ_2 and ν in eq. (2) 223 are hyperparameters controlling the detailed shape of the covariance function. Specifically, 224 σ_1 governs its amplitude, σ_2 represents the spatial length-scale over which we expect an 225 observation to be informative, and ν is an order parameter that, essentially, determines 226 the smoothness of the interpolation. This Matérn family arises as a generalisation of the 227 Gaussian (or squared-exponential) covariance function, which corresponds to the case $\nu =$ 228 ∞ , and is widely-used in machine learning and spatial statistics in cases where a general, 229 localised covariance function is required. Readers seeking greater intuition on the role played 230 by the covariance function may wish to consult Figs. 1 & 2 of Valentine and Sambridge 231 (2020a) and the associated discussion. 232

When we represent residual topography using a Gaussian Process, which we denote by writing $h(\mathbf{x}) \sim \mathcal{GP}(\mu(\mathbf{x}), k(\mathbf{x}, \mathbf{x}'))$, we effectively assert that whenever we consider the value of h at any N distinct locations, $\mathbf{x}_{1...N}$, our overall state of knowledge should be described by an N-dimensional Gaussian distribution,

$$\begin{pmatrix} h(\mathbf{x}_1) \\ h(\mathbf{x}_2) \\ \vdots \\ h(\mathbf{x}_N) \end{pmatrix} \sim \mathcal{N}(\boldsymbol{\mu}, \mathbf{K}) .$$
(4)

The elements of $\boldsymbol{\mu}$ are obtained by evaluating the mean function, $\mu_i = \mu(\mathbf{x}_i)$, and the elements of **K** are derived from the covariance function, $K_{ij} = k(\mathbf{x}_i, \mathbf{x}_j)$. In particular, if we consider a single point \mathbf{x} , we find that

$$h(\mathbf{x}) \sim \mathcal{N}(\mu_0, \sigma_1^2) \,. \tag{5}$$

This represents our prior probability distribution for residual topography at a point. Thus, before observing any data, our presumptions about residual topography, irrespective of location, can be characterised as $h = \mu_0 \pm \sigma_1$.

Now, suppose we are able to measure residual topography at a discrete set of points, 243 $\hat{\mathbf{x}}_{1...N}$ (as in Valentine and Sambridge (2020a), a 'hat' is used to distinguish quantities as-244 sociated with the observational dataset). Our set of N measurements can be represented as 245 the N-dimensional vector $\hat{\mathbf{d}}$, such that \hat{d}_i represents our estimate of $h(\hat{\mathbf{x}}_i)$, and we assume 246 that all measurement errors are Gaussian, described overall by the $N \times N$ covariance matrix, 247 $\mathbf{C}_{\hat{\mathbf{a}}}$. In the present case, this has entries on the leading diagonal only (so that all measure-248 ments are assumed independent), and is treated as a function of the unknown contribution 249 to uncertainty arising from a lack of crustal corrections, Δ . According to the assumptions 250 we have made in setting up our Gaussian Process model, this specific set of observations 251 have likelihood 252

$$\mathbb{P}(\hat{\mathbf{d}} \mid \boldsymbol{\sigma}) = \frac{1}{(2\pi)^{N/2} |\hat{\mathbf{K}}_{\boldsymbol{\sigma}} + \mathbf{C}_{\hat{\mathbf{d}}}(\Delta)|^{1/2}} \exp\left\{-\frac{1}{2} \left(\hat{\mathbf{d}} - \hat{\boldsymbol{\mu}}_{\boldsymbol{\sigma}}\right)^{\mathbf{T}} \left(\hat{\mathbf{K}}_{\boldsymbol{\sigma}} + \mathbf{C}_{\hat{\mathbf{d}}}(\Delta)\right)^{-1} \left(\hat{\mathbf{d}} - \hat{\boldsymbol{\mu}}_{\boldsymbol{\sigma}}\right)\right\}$$
(6)

where $\hat{\mu}_{\sigma}$ and \hat{K}_{σ} are obtained by evaluating the mean and covariance functions at the 253 sample points $\hat{\mathbf{x}}$, and where the subscript $\boldsymbol{\sigma}$ is used to emphasise that these quantities them-254 selves depend on specific choices for the hyperparameters, $\boldsymbol{\sigma} = (\mu_0, \sigma_1, \sigma_2, \nu, \Delta)$. Plainly, 255 we should not adopt assumptions that make our observations inherently 'unlikely', and so 256 it is sensible to select hyperparameter values that maximise eq. (6). This is straightforward 257 to implement using standard computational tools for optimisation, such as the L-BFGS-B 258 algorithm (Byrd et al., 1995), available through Scipy. We remark that this concept can be 259 framed more formally as an approximation to a hierarchical Bayesian estimation of σ ; see 260 Valentine and Sambridge (2018). 261

Having selected optimal values for σ , we are now in a position to predict residual topography at unseen locations. To do this, we note that eq. (4) holds for *any* set of locations: in particular, it must hold if we consider the N observed locations and one additional point, \mathbf{x} . We can then condition this distribution on the values we actually observed, exploiting the correlations between locations to refine our knowledge about $h(\mathbf{x})$. As explained in Valentine and Sambridge (2020a), application of a standard result for the conditioning of Gaussian distributions results in a posterior estimate of residual topography given by

$$\tilde{h}(\mathbf{x}) \sim \mathcal{GP}\left(\tilde{\mu}(\mathbf{x}), \tilde{k}(\mathbf{x}, \mathbf{x}')\right),$$
(7a)

270 where

$$\tilde{\mu}(\mathbf{x}) = \mu(\mathbf{x}) + \hat{\mathbf{k}}(\mathbf{x})^{\mathbf{T}} \left(\hat{\mathbf{K}} + \mathbf{C}_{\hat{\mathbf{d}}} \right)^{-1} \left(\hat{\mathbf{d}} - \hat{\boldsymbol{\mu}} \right)$$
(7b)

$$\tilde{k}(\mathbf{x}, \mathbf{x}') = k(\mathbf{x}, \mathbf{x}') - \hat{\mathbf{k}}(\mathbf{x})^{\mathbf{T}} \left(\hat{\mathbf{K}} + \mathbf{C}_{\hat{\mathbf{d}}} \right)^{-1} \hat{\mathbf{k}}(\mathbf{x}')$$
(7c)

and where $\mathbf{k}(\mathbf{x})$ is the vector-valued function that evaluates the covariance function between point \mathbf{x} and each of the observation locations, such that

$$\left[\hat{\mathbf{k}}(\mathbf{x})\right]_{i} = k(\mathbf{x}, \hat{\mathbf{x}}_{i}).$$
(7d)

²⁷³ Thus, within the Gaussian Process framework, the data allow us to infer that

$$\tilde{h}(\mathbf{x}) = \tilde{\mu}(\mathbf{x}) \pm \sqrt{\tilde{k}(\mathbf{x}, \mathbf{x})} \,. \tag{8}$$

This allows us to map the spatial pattern and amplitude of residual topography, which we return to in Section 4.

3.2 The power spectrum of residual topography

To understand the power spectrum of the recovered residual topography field, and to compare this with previously-published results, we must express $h(\mathbf{x})$ in terms of spherical harmonics. As in Hoggard et al. (2016) and Davies et al. (2019), we employ the real surface spherical harmonics, $\mathcal{Y}_{lm}(\theta, \varphi)$, defined as in Section B6 of Dahlen and Tromp (1998). The infinite set of spherical harmonics is complete and, thus, we can express our Gaussian Process model exactly as an infinite sum

$$\tilde{h}(\theta,\varphi) = \sum_{l=0}^{\infty} \sum_{m=-l}^{l} c_{lm} \mathcal{Y}_{lm}(\theta,\varphi) \,.$$
(9)

Moreover, the spherical harmonics are normalised, and orthogonal, in the sense that

$$\int_{\Omega} \mathcal{Y}_{lm}(\theta,\varphi) \mathcal{Y}_{l'm'}(\theta,\varphi) \,\mathrm{d}\Omega = \delta_{ll'} \delta_{mm'} \,, \tag{10}$$

where δ_{ij} is a Kronecker delta and integration is over a spherical surface. Hence, any coefficient c_{lm} can be found by evaluating

$$c_{lm} = \int_{\Omega} \tilde{h}(\theta, \varphi) \mathcal{Y}_{lm}(\theta, \varphi) \,\mathrm{d}\Omega \,. \tag{11}$$

It is important to recognise that, whereas Hoggard et al. (2016) and Davies et al. (2019) sought to directly fit a spherical harmonic expansion to point data, our approach is different: we have already obtained a global representation of $\tilde{h}(\theta, \varphi)$, and we simply need to express this within the spherical harmonic basis. This distinction underpins the stability of our approach.

Because $h(\theta, \varphi)$ is a Gaussian Process, the integration in eq. (11) results in a probability distribution for c_{lm} . As discussed in Valentine and Sambridge (2020a), c_{lm} is normallydistributed, and the coefficients associated with different values of l and m are jointly normal. If we use **c** to denote some set of these coefficients, then

$$\mathbf{c} \sim \mathcal{N}(\mathbf{y}, \boldsymbol{\Sigma})$$
 (12a)

295 where

$$y_i = \int_{\Omega} \tilde{\mu}(\theta, \varphi) \mathcal{Y}_{l_i m_i}(\theta, \varphi) \,\mathrm{d}\Omega \tag{12b}$$

$$\Sigma_{ij} = \iint_{\Omega^2} \mathcal{Y}_{l_i m_i}(\theta, \varphi) \tilde{k}(\theta, \varphi; \theta', \varphi') \mathcal{Y}_{l_j m_j}(\theta', \varphi') \,\mathrm{d}\Omega \,\mathrm{d}\Omega' \tag{12c}$$

and where we have used (l_i, m_i) to indicate the spherical harmonic degree and order appropriate to the *i*th element of **c**.

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Substituting the expressions for $\tilde{\mu}$ and \tilde{k} from eq. (7) into these integrals, we obtain

$$y_{i} = \int_{\Omega} \mu(\mathbf{x}) \mathcal{Y}_{l_{i}m_{i}}(\mathbf{x}) \,\mathrm{d}\Omega + \sum_{pq} f_{l_{i}m_{i}}(\hat{\mathbf{x}}_{p}) \left[\left(\hat{\mathbf{K}} + \mathbf{C}_{\hat{\mathbf{d}}} \right)^{-1} \right]_{pq} \left(\hat{d}_{q} - \mu(\hat{\mathbf{x}}_{q}) \right)$$
(13a)

$$\Sigma_{ij} = g_{l_i m_i l_j m_j} - \sum_{pq} f_{l_i m_i}(\hat{\mathbf{x}}_p) \left[\left(\hat{\mathbf{K}} + \mathbf{C}_{\hat{\mathbf{d}}} \right)^{-1} \right]_{pq} f_{l_j m_j}(\hat{\mathbf{x}}_q)$$
(13b)

where we have introduced f and g to represent certain integrals of the covariance function,

$$f_{lm}(\theta,\varphi) = \int_{\Omega} k(\theta,\varphi;\theta',\varphi') \mathcal{Y}_{lm}(\theta',\varphi') \,\mathrm{d}\Omega'$$
(14a)

$$g_{lml'm'} = \int_{\Omega} f_{lm}(\theta, \varphi) \mathcal{Y}_{l'm'}(\theta, \varphi) \,\mathrm{d}\Omega \tag{14b}$$

Naïvely, one might anticipate evaluating these numerically. However, this is impractical, since integration over the surface of the sphere requires a double integral, and thus g represents a quadruple integral. For our approach to be tractable, it is essential that we find analytic reductions of these integrals.

By design, we adopted epicentral angle as the distance measure within our covariance function. As a result, the function k can be regarded as a function of epicentral angle, $k(\theta, \varphi; \theta', \varphi') \rightarrow k(\cos(\Theta))$. Recognising this allows us to write it as an expansion in terms of Legendre polynomials,

$$k(\cos(\Theta)) = \sum_{l=0}^{\infty} a_l P_l(\cos(\Theta))$$
(15)

where, again, no approximation is involved provided the sum is infinite. Any coefficient in this expansion can be computed,

$$a_l = \frac{2l+1}{2} \int_0^\pi k(\cos\Theta) P_l(\cos\Theta) \sin\Theta \,\mathrm{d}\Theta \tag{16}$$

and this integral can be evaluated numerically for any l.

However, by the spherical harmonic addition theorem, we can also express the Legendre polynomials as a sum of spherical harmonics,

$$P_l(\cos(\Theta)) = \frac{4\pi}{2l+1} \sum_{m=-l}^{l} \mathcal{Y}_{lm}(\theta,\varphi) \mathcal{Y}_{lm}(\theta',\varphi') \,. \tag{17}$$

Combining eqs. (14–17) and exploiting the orthogonality properties given in eq. (10), we therefore have

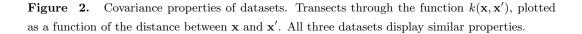
$$f_{lm}(\theta,\varphi) = 2\pi \mathcal{Y}_{lm}(\theta,\varphi) \int_{0}^{\pi} k(\cos\Theta) P_{l}(\cos\Theta) \sin\Theta \,\mathrm{d}\Theta$$
(18a)

$$g_{lml'm'} = 2\pi \delta_{ll'} \delta_{mm'} \int_0^\pi k(\cos\Theta) P_l(\cos\Theta) \sin\Theta \,\mathrm{d}\Theta \tag{18b}$$

³¹⁵ We are now in a position to evaluate \mathbf{y} and $\mathbf{\Sigma}$, and, thus, characterise the distribution ³¹⁶ describing \mathbf{c} . We highlight that no approximations have been made in determining these

	$\mu_0 \ (\mathrm{km})$	Δ (km)	$\sigma_1 \ ({\rm km})$	$\sigma_2 \ (rad)$	ν
High accuracy spot only All spot	-0.03	-0.08	$0.63 \\ 0.64$	$0.24 \\ 0.20$	$0.49 \\ 0.49$
All spot and shiptrack	-0.01	0.09	0.65	0.20	0.69
0.4 -			All spot	racy spot only od shiptrack	

Table 1. Optimal hyperparameters determined for each dataset. Quantities are defined as in the main text.



Ó

Distance, $d(\mathbf{x}, \mathbf{x}')$, km

expressions, and our results are not influenced by spectral leakage caused by truncation of 317 a spherical harmonic expansion at some particular degree. 318

Finally, to obtain a spectrum for any random realisation of \mathbf{c} , we follow Hoggard et 319 al. (2016) and define the power at degree l to be 320

$$p_l = \sum_{m=-l}^{l} c_{lm}^2 \,. \tag{19}$$

5000

10000

Since the coefficients c_{lm} are Gaussian-distributed, p_l follows a generalised χ^2 -distribution. 321 It should be noted that this has a number of counter-intuitive properties, including the fact 322 that the expected value of the power is typically higher than the power of the expected set 323 of coefficients. 324

4 Results and Discussion 325

Covariance, k(x, x') 0.2

0.0 + -10000

-5000

326

4.1 Dataset properties and models of residual topography

We begin by considering the 1160 high-accuracy spot measurements only. We deter-327 mine the hyperparameters μ_0 , σ_1 , σ_2 and ν that maximise the inherent likelihood of this 328 dataset (subject to the assumptions underpinning our approach), as discussed in Section 3.1. 329 Values are given in Table 1; in particular, we find the characteristic length-scale over which 330 each data-point is informative to be around 1500 km. The resulting covariance function 331 $k(\mathbf{x}, \mathbf{x}')$ is shown in Fig. 2. 332

We then condition the Gaussian Process on the observed data, to obtain a model for 333 residual topography. The posterior mean elevation for this high-accuracy dataset is mapped 334

in Fig. 3(a), with uncertainties shown in Fig. 3(b). We observe a spatial pattern of residual 335 topography that is consistent with that reported by Davies et al. (2019), dominated by 336 broad topographic highs within the Pacific, African and North Atlantic regions, separated 337 by a band of topographic lows extending from Antarctica, through the Americas to the 338 Arctic, broadening beneath the Eurasian continent and extending south of Australia. Un-339 surprisingly, given the sparse and uneven data coverage, this model has large uncertainties 340 in many regions of the globe, including continental interiors and substantial sections of the 341 oceans. 342

In any Bayesian approach, the posterior distribution arises as a compromise between information obtained from the data, and information imposed by the prior distribution. To help quantify where the dataset is informative, it is useful to measure the difference between prior and posterior distributions at any point in our model. To do this, we employ a common similarity measure for probability distributions, the Kullback-Leibler divergence (Kullback & Leibler, 1951),

$$D_{KL}(h(\mathbf{x}) \| \tilde{h}(\mathbf{x})) = \int_{-\infty}^{\infty} \rho_h(u) \log \frac{\rho_{\tilde{h}}(u)}{\rho_h(u)} \,\mathrm{d}u$$
(20)

where ρ_h and $\rho_{\tilde{h}}$ represent the probability density functions associated with the distribu-349 tions $h(\mathbf{x})$ and $\tilde{h}(\mathbf{x})$, respectively. Since the natural logarithm is used, information gain is 350 expressed in units of 'nats' (compare 'bits' for base-2 information measures). Where the 351 Kullback-Leibler divergence is low, our posterior distribution remains rather similar to our 352 prior distribution, and so our inference has not taught us much. This is mapped for the 353 high-accuracy spot measurements in Fig. 3(c), and in many regions — particularly conti-354 nental interiors, the Pacific Ocean and the Southern Ocean — there is little information 355 available from data. This must be borne in mind when interpreting the residual topography 356 maps. 357

We repeat this analysis, extending the dataset to include the additional 870 spot 358 measurements for which crustal corrections are not available. As discussed in Section 3.1, we 359 increase the uncertainties associated with these data points by an amount Δ , and determine 360 $\Delta = 0.08 \,\mathrm{km}$ as part of the hyperparameter optimisation process. This is somewhat less 361 than the correction adopted by Hoggard et al. (2017) (who proposed $\Delta = 0.2$ km, based on 362 the median crustal correction applied to the high-accuracy points). Other hyperparameter 363 values are as listed in Table 1: these are generally similar to those obtained using the 364 high-accuracy points, with the characteristic length-scale determined to be ~ 1300 km. The 365 resulting covariance function is depicted in Fig. 2, and is consistent with that obtained using 366 the high-accuracy data alone. 367

Conditioning the Gaussian Process on the observed data, we obtain the residual topography model shown in Fig. 4. Again, this is generally similar to that obtained using the high-accuracy data, although the additional constraints reveal new features in some regions (e.g. the Western Pacific, where some shorter wavelength structure is apparent). Again, we report uncertainties and the information gain between prior and posterior, and these should be considered carefully before any detailed interpretation of model features.

Finally, we extend the dataset once more, incorporating the 20767 measurements de-374 rived from shiptrack bathymetry. Again, to reflect the fact that no crustal corrections are 375 applied to these data points, we increase measurement uncertainties by the amount Δ . We 376 redetermine a value for this using the expanded dataset, obtaining $\Delta = 0.09$ km, and other 377 hyperparameters as given in Table 1. Given the significant expansion of the dataset, these 378 are remarkably consistent with those determined from the spot data alone. Only the order 379 parameter ν changes appreciably. However, Fig. 2 illustrates that this has minimal effect 380 on the form of the covariance function. 381

A residual topography model derived from the full dataset is shown in Fig. 5. Unsurprisingly, given the ten-fold increase in the number of measurements used to construct the model, we see fine-scale detail that is not visible using the spot data alone. Uncertainties

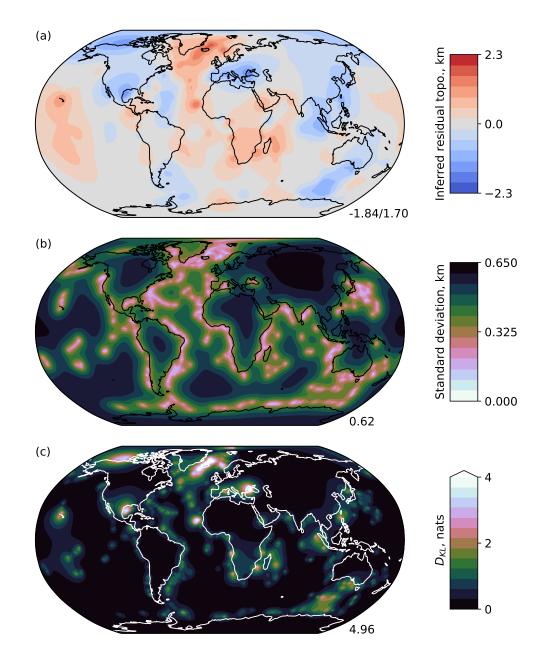


Figure 3. Residual topography determined from high-accuracy spot measurements only. Panel (a) shows the posterior mean point-wise estimate of residual topography, $\tilde{\mu}(\mathbf{x})$, with (b) mapping the associated standard deviation, $\sqrt{\tilde{k}(\mathbf{x}, \mathbf{x})}$. In (c), we show the point-wise information gain from prior to posterior, $D_{KL}(\mathbf{x})$, in units of 'nats' (see text). Colour scales are chosen for consistency with subsequent figures; the maximum values for each map are given in the lower-right corner. Note that uncertainties grow rapidly away from measurement locations, and there are large regions (including all continental interiors) where the spot data is uninformative.

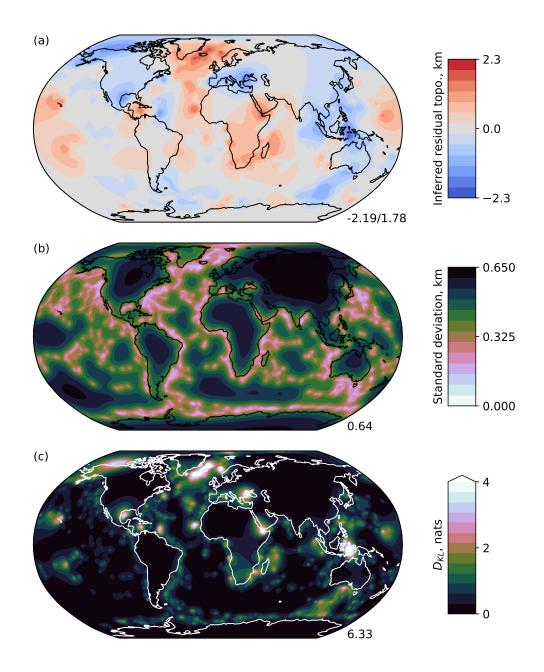


Figure 4. Residual topography determined from spot data, including the 1160 high-accuracy points and a further 870 measurements lacking a crustal correction, for which an additional uncertainty has been estimated. As in Fig. 3, we show (a) the mean residual topography, (b) the standard deviation in this estimate, and (c) the pointwise information gain.

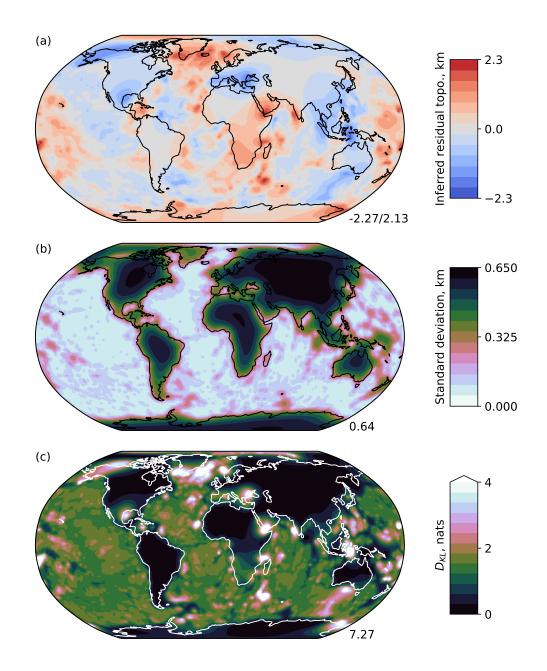


Figure 5. Residual topography determined from all available data, including both spot- and shiptrack-derived measurements. As in Figs 3 and 4, we show (a) the mean residual topography, (b) the standard deviation in this estimate, and (c) the pointwise information gain. The use of shiptrack-derived data substantially expands the area in which the model is informative.

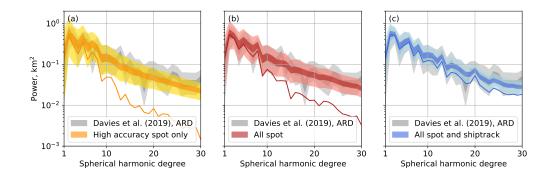


Figure 6. The power spectrum of residual topography: (a) spectrum obtained using highaccuracy spot measurements only; (b) spectrum obtained using all spot measurements; and (c) spectrum obtained using spot and shiptrack data. All panels also show results from Davies et al. (2019) (derived from spot and shiptrack data) for comparison. Solid lines represent the power spectrum of the mean spherical harmonic coefficients; coloured bands represent the ranges spanned by the central 50% and 99% of spectra computed for 100,000 random samples from the distribution of spherical harmonic coefficients. In general, the results obtained in this paper are in agreement with those of Davies et al. (2019).

are markedly reduced throughout the oceans, particularly in the North Atlantic. As expected, the data is informative throughout the oceans, but the lack of onshore data results
 in significant uncertainties within continental interiors.

388 4.2 Power spectra

Having obtained residual topography models based on the three data subsets, we next 389 express each in terms of spherical harmonics and compute power spectra. These are shown in 390 Fig. 6, presented in a format consistent with Hoggard et al. (2016) and Davies et al. (2019), 391 allowing for straightforward comparison. To represent the (non-Gaussian) uncertainties as-392 sociated with these spectra, we generate power spectra for 100,000 random samples from the 393 posterior residual topography models, and Fig. 6 depicts the ranges spanned by the central 394 99% and 50% of samples. An alternative presentation of the same data is shown in Fig. 7, 395 allowing direct comparison between the three data sets at each spherical harmonic degree, 396 and some representative numerical values are provided in Table 2. Files containing the ex-397 pansion coefficients for each data set are included in the software repository accompanying 398 this paper; see the 'data availability statement' for details. 399

In general, all three datasets tell a similar story, which is consistent with that reported 400 by Davies et al. (2019) (results also shown in Fig. 6). The spectrum of residual topography 401 is relatively flat, peaking at or around degree-2 (wavelength $\sim 16000 \,\mathrm{km}$), with steadily-402 declining power at shorter length scales. Based on the high-accuracy spot data, the most-403 probable model has degree-2 power 0.53 km², although the data could support power up to 404 $\sim 1.3 \,\mathrm{km}^2$. By degree-10 (wavelength $\sim 4000 \,\mathrm{km}$) and degree-20 ($\sim 2000 \,\mathrm{km}$), the power is 405 likely in the ranges $0.12-0.18 \text{ km}^2$ and $0.04-0.06 \text{ km}^2$, respectively. In general, the additional 406 information available in the expanded datasets (all spot data; all spot and shiptrack data) 407 enables a modest reduction in the spectral uncertainty, but does not substantially alter the 408 most-probable power. Based on our random samples, we find that degrees 1-3 account 409 for about 85% of the total power in the residual topography field below degree-30 (high 410 accuracy spot data only: 85.9%; all spot data: 83.9%; spot and shiptrack data: 84.2%). 411

	MP ^a	Degree 2 $50\%^b$	99% c	MP^{a}	Degree 5 $50\%^b$	99%c	MP ^a	Degree 10 $50\%^b$) 99%c	MP ^a	Degree 20 $50\%^b$	$99\%^c$
High accuracy spot only All spot & shiptrack	$\begin{array}{c c} 0.53 \\ 0.54 \\ 0.52 \\ 0.52 \end{array}$	$\begin{array}{c} 0.46{-}0.76\\ 0.47{-}0.70\\ 0.47{-}0.65\end{array}$	$\begin{array}{c c} 0.17 - 1.32 \\ 0.23 - 1.12 \\ 0.27 - 0.96 \end{array}$	$\begin{array}{c} 0.34 \\ 0.31 \\ 0.29 \end{array}$	$\begin{array}{c} 0.38 - 0.56 \\ 0.32 - 0.46 \\ 0.29 - 0.40 \end{array}$	$\begin{array}{c} 0.19{-}0.89\\ 0.18{-}0.71\\ 0.18{-}0.58\\ 0.18{-}0.58\end{array}$	0.05 0.08 0.12	$\begin{array}{c} 0.12 - 0.18 \\ 0.13 - 0.18 \\ 0.15 - 0.20 \end{array}$	$\begin{array}{c} 0.06 \\ -0.28 \\ 0.07 \\ -0.29 \\ 0.09 \\ -0.28 \end{array}$	0.01 0.01 0.05	$\begin{array}{c} 0.04 \hline 0.06 \\ 0.05 \hline -0.06 \\ 0.07 \hline -0.08 \end{array}$	$\begin{array}{c} 0.03 - 0.08 \\ 0.03 - 0.09 \\ 0.05 - 0.10 \end{array}$
^a Most probable model, as depicted in Figs. $3(a)-5(a)$. ^b Range spanned by central 50% of 10 ⁶ samples from the posterior distribution, i.e. the interquartile range. ^c Range spanned by central 99% of 10 ⁶ samples from the posterior distribution.	s depict al 50% al 99%	ed in Figs. 3 ¹ of 10 ⁶ sample of 10 ⁶ sample	$\begin{array}{c} (a)-5(a).\\ \text{es from the}\\ \text{se from the} \end{array}$	posterio posterio	r distributic r distributic	m, i.e. the in	nterquar	rtile range.				
			Table 3.		al topograph	Residual topography maximum amplitudes, km.	amplituc	łes, km.				
	MP ^a	Full model $50\%^b$	ءا 99% ^c	MP ^a	Degrees $1-3$ $50\%^b$	-3 $99\%^{c}$	MP ^a	Degrees $4-10$ $50\%^{b}$	-10 $99\%^{c}$	MPa	Degrees $11-30$ $50\%^{b}$	$\frac{30}{99\%^c}$
High accuracy spot only All spot All spot and shiptrack	$\begin{array}{ c c c c } 1.84 \\ 2.19 \\ 2.27 \end{array}$	$\begin{array}{c} 2.12{-}2.39\\ 2.27{-}2.54\\ 2.31{-}2.51\end{array}$	$\begin{array}{c} 1.86{-}2.95\\ 2.00{-}3.09\\ 2.12{-}3.02\end{array}$	$\begin{array}{c c} 0.66 \\ 0.69 \\ 0.74 \end{array}$	$\begin{array}{c} 0.68{-}0.85\\ 0.71{-}0.82\\ 0.74{-}0.82\end{array}$	$\begin{array}{c} 0.49{-}1.16\\ 0.57{-}1.01\\ 0.65{-}0.97\end{array}$	$ \begin{array}{c c} 1.24 \\ 1.32 \\ 1.38 \end{array} $	$\frac{1.27 - 1.48}{1.35 - 1.56}$ $1.42 - 1.63$	$\frac{1.05{-}1.87}{1.12{-}1.93}$ $\frac{1.17{-}2.03}{1.17{-}2.03}$	$\begin{array}{c c} 0.68 \\ 0.95 \\ 1.30 \\ \end{array}$	$\frac{1.01{-}1.14}{1.11{-}1.26}$ $1.35{-}1.48$	$\begin{array}{c} 0.89{-}1.41 \\ 0.96{-}1.61 \\ 1.19{-}1.78 \end{array}$
^a Most probable model, as depicted in Figs. $3(a)-5(a)$. ^b Range spanned by central 50% of 10 ⁶ samples from the posterior distribution, i.e. the interquartile range. ^c Range spanned by central 99% of 10 ⁶ samples from the posterior distribution.	- s depict al 50% al 99%	ted in Figs. 3 of 10 ⁶ sampl of 10 ⁶ sampl	(a)-5(a). es from the es from the	posteric	or distributions of the second	on, i.e. the i	nterqua	rtile range.		_		

 Table 2. Residual topography power, km².

-16-

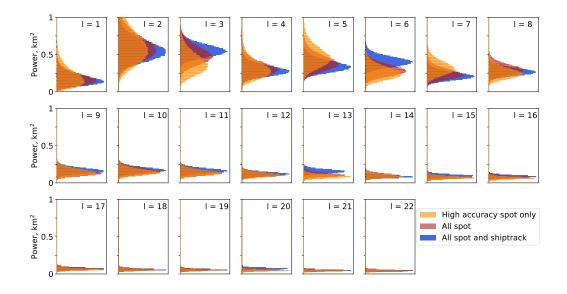


Figure 7. Uncertainties in the power spectrum of residual topography. An alternative representation of the data in Fig. 6: histograms of the power per degree for 100,000 random samples from the posterior distribution of spherical harmonic coefficients obtained using each dataset. Note that the vertical (power) scale here is linear, whereas in Fig. 6 a logarithmic scale is used, to match figures elsewhere in the literature. Beyond l = 22 differences between the two datasets are negligible on the scale of this figure.

By generating random models consistent with the data, we can also quantify the 412 typical height variations associated with residual topography. Some representative figures 413 are given in Table 3. Based on the high-accuracy data, we see that the maximum amplitude 414 of residual topography is likely to be in the range 2.12–2.39 km, with an upper limit of 415 around 3 km. This is consistent with the range of measurements depicted in Fig. 1. When 416 filtered at the longest wavelengths (spherical harmonic degrees 1–3), models consistent with 417 the data typically have maximum amplitudes in the range 680-850 m, with an upper limit of 418 ~ 1.2 km. The expanded datasets reduce these ranges somewhat: the full spot and shiptrack 419 data suggests maximum amplitudes at these wavelengths of 740–820 m, and an upper limit 420 of $\sim 1 \,\mathrm{km}$. 421

422 **4.3** A sy

4.3 A synthetic test

The consistency of results across datasets is a strong indication that we are obtaining 423 models that provide a useful representation of the truth. To further validate our approach, 424 we perform a synthetic test. We take a synthetic residual topography field generated from 425 a global model of instantaneous mantle flow (fully-described in Davies et al. (2019); we here 426 use the simulation 'with shallow structure', which incorporates shallow-density heterogeneity 427 and thermal structure based upon an estimate of lithospheric thickness, and has been shown 428 to have characteristics that are consistent with the observational data). This simulated field 429 is first sampled at the 1160 locations corresponding to our high-accuracy measurements. We 430 use this dataset to construct a global Gaussian Process model, which — as already discussed 431 enables us to generate the mean and standard deviation characterising our prediction 432 for residual topography at any point on Earth's surface. We then sample our synthetic 433 residual topography field at 116062 points drawn from a regular mesh within the oceans, 434 and compare this synthetic ground truth data with our Gaussian Process predictions for 435 these 116062 locations. For each location, we compute the difference between the mean 436

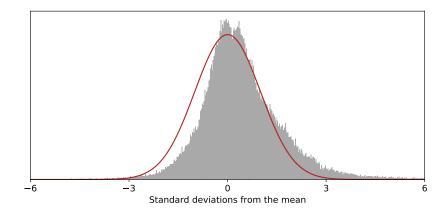


Figure 8. A synthetic test: predictive performance of Gaussian Process residual topography model constructed using simulated data. Output from a global simulation of instantaneous mantle flow from Davies et al. (2019) is used to generate a synthetic residual topography dataset at the same 1160 locations as in our 'high accuracy' dataset. Applying our Gaussian Process approach, we can then predict the residual topography on a regular mesh of 116062 locations throughout the oceans, and associated uncertainties. We then compare these to the 'true' simulation-derived value at each point, expressing any prediction error relative to the uncertainty. In the ideal case, results (grey bars) should follow a standard normal distribution (red line).

predicted topography and the 'true' value, and then divide this by the standard deviation
associated with that prediction. Treating performance at each location as independent, this
quantity should follow a standard normal distribution (i.e., one with zero mean and unit
variance) if our model were perfect.

A histogram of the resulting data is shown in Fig. 8, overlain by a standard normal distribution. As one might anticipate, the two do not match exactly, with the histogram somewhat skewed. Nevertheless, it is clear that the Gaussian Process model provides good predictive performance: the truth lies within one standard deviation of the predicted mean in 65.4% of cases, and within three standard deviations in 96.2%. For a true normal distribution, these figures should be 68.3% and 99.7% respectively.

A number of factors can be identified that contribute to this discrepancy. First, it 447 is not strictly correct to treat the performance at adjacent locations as independent; we 448 ought to account for the fact that they are correlated. However, neglecting this greatly 449 simplifies the analysis, and likely better-represents the manner in which users will approach 450 maps such as those shown in Figs. 3–5. Second, in treating the residual topography field as 451 being representable using a Gaussian Process, we are making a strong assumption about its 452 statistical properties. This is pragmatically justifiable — one cannot analyse any dataset 453 without making assumptions — but is almost certainly imperfect. 454

A third issue may lie within the hyperparameter optimisation procedure. Implicit 455 in this is an assumption that the available data is representative of the field as a whole. 456 Unfortunately, as with most geoscience datasets, our sample locations are not randomly-457 distributed around the globe: they are influenced by a variety of practical, economic and 458 scientific constraints. In the present case, the points within the high-accuracy dataset are 459 governed by the availability of seismic surveys. Inevitably, such surveys are targeted towards 460 particular geological environments — resulting in a bias towards older oceanic crust that 461 abuts continental margins. As a result, there is a significant possibility that our observations 462

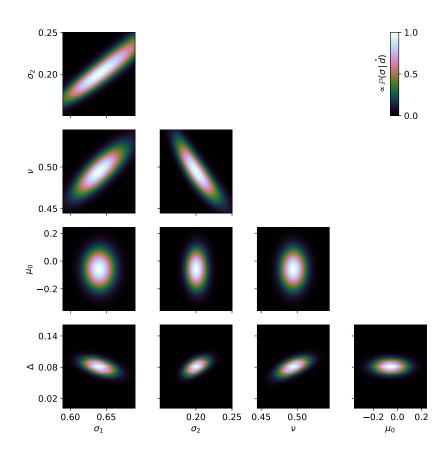


Figure 9. Hyperparameter tradeoffs, computed using all spot data. Plots show the unnormalised likelihood of residual topography measurements, eq. (6), as pairs of hyperparameters are varied in the vicinity of the optimal set (see Table 1); in each plot, all other hyperparameters are fixed to their optimal value. All hyperparameters are fairly tightly-constrained. Relationships between σ_1 , σ_2 , ν and μ_0 computed using the high-accuracy spot data are very similar to those seen here; the parameter Δ is not relevant to that dataset.

present a biased picture of the overall field. This is the case in our synthetic example: 463 although the average global residual topography is zero, the average residual topography at 464 our 1160 sample points is $-422 \,\mathrm{m}$ (globally, the simulation has amplitudes that are somewhat 465 larger than we expect for the Earth, lying in the range -4.95-3.39 km). This sampling bias 466 likely explains (at least partially) the skewness seen in Fig. 8. It is impossible to fully assess 467 whether similar biases are present in the real datasets used elsewhere in this paper (although 468 the consistency across all three datasets suggests that the issue is not as pronounced as in 469 this synthetic test). We stress that any such biases depend heavily on the detail of the 470 residual topography field, and these synthetic results should not be taken to indicate that 471 the models in Figs. 3–5 are similarly skewed towards underestimating topography. 472

473

4.4 Hyperparameter tradeoffs and interpretability

In Fig. 9, we show slices through the hyperparameter likelihood function, computed using all spot data. We see that the maximum-likelihood point is well-constrained, implying that any uncertainty associated with the values presented in Table 1 is small. There are clear correlations: σ_1 is correlated with both σ_2 and ν , while σ_2 is anti-correlated with ν . The mean function hyperparameter, μ_0 , appears to be largely independent of others. Essentially identical patterns are seen with the high-accuracy spot data alone; we use all spot data here since this also illuminates the uncertainty hyperparameter, Δ . This is also seen to have some weak trade-offs with the hyperparameters σ_1 , σ_2 and ν . It is not feasible to make similar plots using the full spot and shiptrack dataset, due to the computational costs involved.

To what extent are these hyperparameters interpretable? In particular, it is notable 484 that the uncertainty correction, Δ , is markedly less than that identified Hoggard et al. 485 (2017), who adopted $\Delta = 0.2$ km. This raises the question: is our value 'more correct'? In addressing this, it is essential to remember that our hyperparameters are optimal in a very 487 specific sense: they maximise eq. 6 for the available data, and are predicated on the various 488 assumptions inherent to our approach. These inevitably represent a simplified version of 489 reality, and one can readily propose alternative assumptions that are — at a minimum — 490 equally plausible. End users must carefully consider whether our choices are appropriate to 491 their application, and thus whether it makes sense to rely on our hyperparameter values or 492 other results. 493

In the present work, we effectively assume that we have no prior knowledge about the 494 hyperparameters, and that all values are equally plausible. If desired, it would be straight-495 forward to modify the hyperparameter determination procedure to incorporate an explicit 496 'hyperprior', expressing a priori expectations or preferences. This is discussed in more de-497 tail in, for example, Valentine and Sambridge (2018). However, for the present problem, we do not see a route to constructing a hyperprior that is objectively more justifiable than 499 the uniform hyperprior implicit in our approach. Nevertheless, we emphasise that the pri-500 mary goal of the present work is the analysis of the high-accuracy spot data, for which the 501 uncertainty hyperparameter Δ is not relevant. 502

503

4.5 The future: obtaining new measurements

Although the first order characteristics of the power spectra obtained here are consis-504 tent across all data sets, there is no doubt that tighter constraints on the exact spatial pat-505 tern, wavelength and amplitude of residual topography require extending the high-accuracy 506 spot measurements across the globe: our knowledge of residual topography is limited by the 507 amount of data available to us. However, obtaining high-quality observations is expensive 508 and time-consuming. A pertinent question is therefore: where would new data be most 509 useful? This question takes us into the field of 'optimal experimental design' (e.g. Curtis, 510 1999), and the Gaussian Process approach allows it to be addressed straightforwardly. 511

As discussed in Section 3.1, our knowledge of residual topography after seeing the data is given by $\tilde{h}(\mathbf{x}) \sim \mathcal{GP}(\tilde{\mu}(\mathbf{x}), \tilde{k}(\mathbf{x}, \mathbf{x}'))$. Suppose we subsequently obtain one more data point at some location \mathbf{x}_0 , measuring value $d_0 \pm \sigma_d$. Incorporating this additional information into our model results in an updated estimate of topography, $\bar{h}(\mathbf{x}) \sim \mathcal{GP}(\bar{\mu}(\mathbf{x}), \bar{k}(\mathbf{x}, \mathbf{x}'))$, where

$$\bar{\mu}(\mathbf{x}) = \tilde{\mu}(\mathbf{x}) + \alpha(\mathbf{x}_0)k(\mathbf{x}, \mathbf{x}_0) \left[d_0 - \tilde{\mu}(\mathbf{x}_0) \right]$$
(21a)

$$\bar{k}(\mathbf{x}, \mathbf{x}') = k(\mathbf{x}, \mathbf{x}') - \alpha(\mathbf{x}_0)k(\mathbf{x}, \mathbf{x}_0)k(\mathbf{x}', \mathbf{x}_0)$$
(21b)

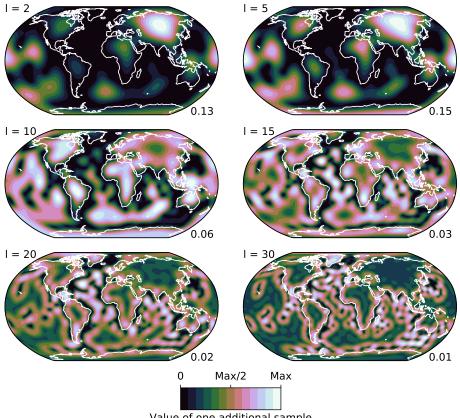
⁵¹⁶ and where we have introduced

$$\alpha(\mathbf{x_0}) = \frac{1}{\tilde{k}(\mathbf{x_0}, \mathbf{x_0}) + \sigma_d^2}.$$
(21c)

⁵¹⁷ We can then express this updated model in terms of spherical harmonics, as before. This ⁵¹⁸ yields a new distribution for the coefficients, $\bar{\mathbf{c}} \sim \mathcal{N}(\bar{\mathbf{y}}, \bar{\mathbf{\Sigma}})$, where

$$\bar{y}_i = y_i + \alpha(\mathbf{x_0}) z_i(\mathbf{x_0}) \left[d_0 - \tilde{\mu}(\mathbf{x_0}) \right]$$
(22a)

 $\bar{\Sigma}_{ij} = \Sigma_{ij} - \alpha(\mathbf{x_0}) z_i(\mathbf{x_0}) z_j(\mathbf{x_0})$ (22b)



Value of one additional sample

Figure 10. Where should data collection efforts be focused? Map of expected information gained (see eq. 24) by collecting one additional sample at that point, broken down by spherical harmonic degree, for the 1160-point high-accuracy spot dataset. Each map is normalised by its maximum value, given in the lower-right corner. Compare with Fig. 3: unsurprisingly, the most-useful places to collect new information correspond to regions of high uncertainty in the recovered topography model.

with 519

$$z_{i}(\mathbf{x}) = f_{l_{i}m_{i}}(\mathbf{x}) - \sum_{pq} f_{l_{i}m_{i}}(\hat{\mathbf{x}}_{p}) \left[\left(\hat{\mathbf{K}} + \mathbf{C}_{\hat{\mathbf{d}}} \right)^{-1} \right]_{pq} k(\mathbf{x}, \hat{\mathbf{x}}_{q}).$$
(22c)

520

Using the Kullback-Leibler divergence (cf. eq. 20), we can now quantify how much the 521 single additional observation has revealed about the spherical harmonic coefficients, given 522 by 523

$$D_{KL} = \frac{1}{2} \left[\operatorname{Tr}(\bar{\boldsymbol{\Sigma}}^{-1} \boldsymbol{\Sigma}) - D + \log \frac{\det \bar{\boldsymbol{\Sigma}}}{\det \boldsymbol{\Sigma}} \right] + \frac{1}{2} \alpha^2(\mathbf{x_0}) \mathbf{z}(\mathbf{x_0})^{\mathrm{T}} \bar{\boldsymbol{\Sigma}}^{-1} \mathbf{z}(\mathbf{x_0}) \left[d - \tilde{\mu}(\mathbf{x_0}) \right]^2, \quad (23)$$

where D represents the dimension of the vector $\bar{\mathbf{c}}$. We therefore see that the amount we learn 524 depends on the value of d: if our additional observation happens to match what we predicted 525 based on the original dataset $(\tilde{\mu}(\mathbf{x}_0))$ we learn less than if the observation reveals unexpected 526 new features. Using $\tilde{h}(\mathbf{x})$, we can determine that the expected value (in a statistical sense) 527 of $(d - \tilde{\mu})^2$ is $\tilde{k}(\mathbf{x_0}, \mathbf{x_0})$ and, hence, we can quantify the expected information gain due to 528

529 an additional observation at \mathbf{x}_0 by

$$G(\mathbf{x_0}) = \frac{1}{2} \left[\operatorname{Tr}(\bar{\boldsymbol{\Sigma}}^{-1} \boldsymbol{\Sigma}) - D + \log \frac{\det \bar{\boldsymbol{\Sigma}}}{\det \boldsymbol{\Sigma}} \right] + \frac{1}{2} \alpha^2(\mathbf{x_0}) \tilde{k}(\mathbf{x_0}, \mathbf{x_0}) \mathbf{z}(\mathbf{x_0})^{\mathrm{T}} \bar{\boldsymbol{\Sigma}}^{-1} \mathbf{z}(\mathbf{x_0}) \,.$$
(24)

This quantity is straightforward to compute, and it is therefore possible to produce maps showing the expected value of one additional sample at any point on the globe. Alternatively, for a more conservative perspective, one might map only the first, parenthesised term in eq. (24): this would represent the minimum information gain from a sample at any point. In particular, one can make such maps considering only the spherical harmonic coefficients corresponding to a particular angular order l, to yield an understanding of how samples in different regions might help constrain different aspects of the residual topography spectrum.

In Fig. 10, we map the expected information gain (i.e., the complete form of $G(x_0)$) 537 based on the high-accuracy spot data. We assume that the one additional data point is 538 measured with an uncertainty of $\sigma_d = 0.1$ km. We clearly see that our knowledge of the 539 low-degree spherical harmonic components can be improved through addition of samples 540 at a relatively small number of locations: continental interiors (especially in Russia), the 541 Southern Ocean, and the South and West Pacific. To constrain shorter wavelengths, our 542 543 analysis indicates that the most useful sampling locations are in regions close to existing data points: this makes sense, as our approach is predicated on collecting only a single 544 sample. 545

546 5 Conclusions

In this paper, we have successfully demonstrated a new approach for obtaining a con-547 tinuous, global model for residual topography given a finite set of point-wise measurements. 548 Our method is based upon the statistical theory of Gaussian Processes (Valentine & Sam-549 bridge, 2020a, 2020b), and has a number of important advantages over those employed 550 in previous studies, particularly: (i) it does not presuppose that the residual topography 551 field is well-represented using a particular set of basis functions, avoiding potential issues 552 such as 'spectral leakage'; (ii) prior information (i.e. the covariance function) is defined 553 within physical space, rather than in an abstract 'model space' where its effects may be 554 poorly-understood; and (iii) the method remains robust for even the smallest datasets. 555

These properties enable us to construct a residual topography model that is based 556 only on the 1160 high-accuracy spot measurements in the database compiled by Hoggard et 557 al. (2017). These points are believed to be particularly robust, since they are obtained via 558 bespoke analysis of seismic and other datasets, and incorporate location-specific information 559 about crustal properties. The resulting model is broadly in agreement with those reported in 560 earlier studies, which have typically relied on a much larger dataset derived from automated 561 processing of shiptrack bathymetry. Applying our method to this larger dataset continues 562 to yield self-consistent results. 563

The global pattern of residual topography is dominated by broad topographic highs 564 in the North Atlantic, African and central Pacific regions, with the largest amplitudes adja-565 cent to Iceland. These are separated by broad topographic lows, extending from Antarctica, 566 through the Americas, beneath the Eurasian continent and to the south of Australia. Based 567 on just the 1160 high-accuracy measurements, the associated spherical harmonic power spec-568 trum peaks at l = 2, with power likely to lie in the range $0.46-0.76 \,\mathrm{km^2}$. This decreases by 569 over an order of magnitude to the range 0.02 km^2 at l = 30. The overall form of the spectrum 570 is consistent across all data subsets analysed. In light of the sensitivity kernels linking topo-571 graphic deflections to density anomalies within the mantle at different depths and spherical 572 harmonic degree (e.g. Richards & Hager, 1984; Colli et al., 2016), and as discussed in detail 573 by Davies et al. (2019), generating this spectrum requires Earth's topographic response to be 574 driven by multi-scale convective processes. The long-wavelength portion of the spectrum is 575

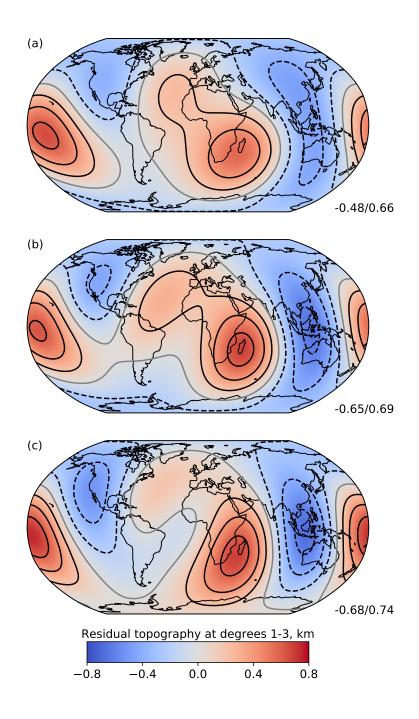


Figure 11. Long-wavelength residual topography. Models (as in Figs. 3–5) filtered to include only spherical harmonic degrees 1–3. We see consistent patterns across all three datasets: (a) high-accuracy spot measurements; (b) all spot measurements; and (c) all spot and shiptrack measurements. Contours are drawn at 200 m intervals, and the amplitude range in each plot is shown, bottom right. Negative contours are dashed. Zero contour is in grey.

primarily attributable to deep mantle flow, while shorter-wavelength features are supported
by flow associated with structural heterogeneity in the shallow mantle and lithosphere.

Given the link between the long-wavelength components of residual topography and 578 deep mantle flow, the patterns and amplitudes associated with low-order spherical harmonics 579 provide an independent constraint on the long-wavelength components of global seismic to-580 mography images and the dynamical interpretation of these images. Accordingly, in Fig. 11, 581 we show our best-fitting models for each data set, filtered to contain only degrees 1–3. The 582 spatial pattern is fairly consistent across all three datasets, with highs centred off South-East 583 Africa that extend into the North Atlantic, and in the central and western Pacific Ocean. As set out in Table 3, models consistent with the observations may have maximum residual 585 topography heights up to around 1.2 km at these wavelengths, although a figure in the range 586 680–850 m is most likely. The residual topography highs strongly resemble the shape and 587 extent of the large low velocity provinces that have been consistently imaged in the deep 588 mantle across a number of global seismic tomography studies (e.g. Becker & Boschi, 2002; 589 Ritsema et al., 2011; Bozdağ et al., 2016), and likely represent concentrations of hot (and 590 potentially chemically-distinct) material that drives upwelling flow towards Earth's surface 591 (e.g. McNamara & Zhong, 2005; Schuberth et al., 2009; Bull et al., 2009; Davies et al., 592 2012; Flament et al., 2013; Davies, Goes, & Sambridge, 2015; Garnero et al., 2016). These 593 long-wavelength residual topography highs are separated by a band of residual topography 594 lows, extending from Antarctica, through the Americas, the Arctic, Eurasia and Australia. 595 These are likely related to downwelling slabs, associated with former subduction zones (e.g. 596 Richards & Engebretson, 1992; Bunge et al., 2002; Flament et al., 2013; Davies, Goes, & 597 Lau, 2015; Rubey et al., 2017). 598

Finally, from a methodological perspective, the Gaussian Process-based approach for 599 inferring global models from point data appears attractive. It is mathematically-elegant and, 600 as we have demonstrated, produces results that are consistent with basis-function expansion 601 strategies. Indeed, as discussed in detail in Valentine and Sambridge (2020b), the two have 602 close links, becoming equivalent under certain conditions as the number of basis functions 603 tends to infinity. We highlight that our approach can be straightforwardly translated to the 604 analysis of other datasets and problems, and suggest that it may prove powerful in cases 605 where models must be constrained using spatially-sparse data. 606

⁶⁰⁷ Data availability statement

The specific data files used for this study are as reported in Davies et al. (2019), and may be obtained from https://github.com/drhodrid/Davies_etal_NGeo_2019_Datasets. Code to analyse this data and produce the results presented here may be obtained from https://github.com/valentineap/DynamicTopographyGP (doi: 10.5281/zenodo.3895317). This repository also includes files containing the spherical harmonic expansion coefficients associated with our mean models.

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