

# High Frequency (6 Hz) PKPab precursors and their sensitivity to deep Earth heterogeneity

Christoph Sens-Schönfelder<sup>1</sup>, Klaus Dieter Bataille<sup>2</sup>, and Marcelo Bianchi<sup>3</sup>

<sup>1</sup>GeoForschungsZentrum Potsdam

<sup>2</sup>Universidad de Concepcion

<sup>3</sup>Universidade de Sao Paulo

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## Abstract

We present observations on a new precursory phase of seismic waves scattered in the deep Earth. This phase arrives prior to the PKPab wave at epicentral distances larger than  $155^\circ$ , and we call it PKPab precursor. We show that the presence of the PKPab precursor is a necessary consequence of scattering in  $D''$ , which is the commonly accepted cause of the PKPdf precursor at distances smaller than  $145^\circ$ . PKPdf waves that propagate through the inner core should arrive before the PKPab precursor but those, are strongly attenuated in the inner core at frequencies between 4 Hz and 8 Hz used here, making the PKPab precursor the earliest teleseismic signal at distances larger than  $155^\circ$ . Calculated PKPab precursor sensitivity kernel shows that this phase is mostly sensitive to scattering along the closest PKPbc path between source and receiver. It can thus help to constrain the lateral distribution of heterogeneity along  $D''$ .

1           **High Frequency (6 Hz) PKPab precursors and their**  
2                           **sensitivity to deep Earth heterogeneity**

3                           **C. Sens-Schönfelder<sup>1</sup>, K. Bataille<sup>2,1</sup>, M. Bianchi<sup>3</sup>**

4                           <sup>1</sup>GFZ German Research Centre for Geosciences, Telegrafenberg, 14473 Potsdam, Germany

5                           <sup>2</sup>Earth Sciences Department, University of Concepcion, Chile

6                           <sup>3</sup>Instituto de Astronomia, Geofísica e Ciências Atmosféricas, Universidade de So Paulo, Brazil

7           **Key Points:**

- 8           • PKP precursor observed at distance beyond 155 deg  
9           • D" scattering of teleseismic waves at 6Hz  
10          • radiative transfer simulation used to locate regions of heterogeneity

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Corresponding author: Christoph Sens-Schönfelder, [sens-schoenfelder@gfz-potsdam.de](mailto:sens-schoenfelder@gfz-potsdam.de)

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 13 the deep Earth. This phase arrives prior to the *PKPab* wave at epicentral distances larger  
 14 than  $155^\circ$ , and we call it *PKPab* precursor. We show that the presence of the *PKPab*  
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 20 Calculated *PKPab* precursor sensitivity kernel shows that this phase is mostly sensitive  
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## 23 Plain Language Summary

24 A new discovered seismic signal recorded far away from earthquakes, by stations on the  
 25 other side of Earth, will help to study the properties of the core-mantle boundary. We  
 26 use high frequencies at which seismic waves do not propagate through the Earth's in-  
 27 ner core but are instead propagated around it by deflection at heterogeneity located along  
 28 the core-mantle boundary.

## 29 1 Introduction

### 30 1.1 Deep Earth structure

31 The boundary between the core and mantle of the Earth is one fascinating region  
 32 in the deep Earth (Tackley, 2012). Here the solid mantle that consists of silicic miner-  
 33 als is in contact with a liquid mostly consisting of molten iron. The density contrast be-  
 34 tween the core ( $\rho = 9,900\text{kg}/\text{m}^3$ ) and the mantle ( $\rho = 5,800\text{kg}/\text{m}^3$ ) is about twice  
 35 as high as the difference between air and the crust at the Earth's surface, but at the core-  
 36 mantle boundary (CMB) the liquid is heavier, while gravitational acceleration is sim-  
 37 ilar to the Earth surface conditions. At this odd interface, lightweight components of the  
 38 core material, potentially generated by solidification of heavier components, accumulate  
 39 from below (Buffett et al., 2000; O'Rourke & Stevenson, 2016), as well as heavy com-  
 40 ponents of the mantle, accumulate from above. These processes caused significant het-  
 41 erogeneity in the  $D''$ -layer at the base of the mantle.

42 The core-mantle boundary is of significant interest in the dynamics of our planet.  
 43 CMB plays a vital role in two major geodynamic processes as it interfaces the outer core  
 44 that generates Earth's magnetic field and the mantle that hosts plate tectonics. Processes  
 45 and structure of the CMB control plate tectonics engine fueled by the heat from the core.  
 46 The geodynamo depends on continuous convection in the core that is, in turn, also con-  
 47 trolled by heat transfer through the boundary (Olson, 2016; Labrosse, 2014). The CMB  
 48 is believed to be the source region of magmatic plumes that led to episodes of gigantic  
 49 volcanic activity at the surface, accompanied by mass extinction events (Courtilot &  
 50 Renne, 2003).

51 The lowermost 200 km of the mantle form a high complexity zone, the so-called  
 52  $D''$  layer. Images of the  $D''$  have been presented by Global seismic tomography studies  
 53 (Kustowski et al., 2008; Ritsema et al., 2011) while its structure is determined generally  
 54 using top and bottom reflections as well as transmitted and diffracted waves (Wang &  
 55 Wen, 2004; Sun et al., 2013; Frost & Rost, 2014; Shen et al., 2016; Euler & Wysession,  
 56 2017; Hansen et al., 2020) observations. A review of seismic investigations of the lower  
 57 mantle can be found in Lay and Garnero (2011).  $D''$  hosts large low shear velocity provinces

(LLSVP) and ultra low-velocity zones (ULVZ) as reviewed in Yu and Garnero (2018) and McNamara (2019).

Whereas the ULVZ are local features with a lateral extent of 100s of kilometers, the two LLSVP are global features beneath Africa and the Pacific. These regions are associated with large scale material uplift in the global mantle convection. There is no real consensus about the nature of the LLSVP, and potential explanations range from purely thermal anomalies to chemically distinct regions in the lower mantle. From their locations, the LLSVPs are believed to be the hottest regions in the mantle since they match the base of global upwelling. This idea is also confirmed by a large number of hotspots and mantle plumes above them.

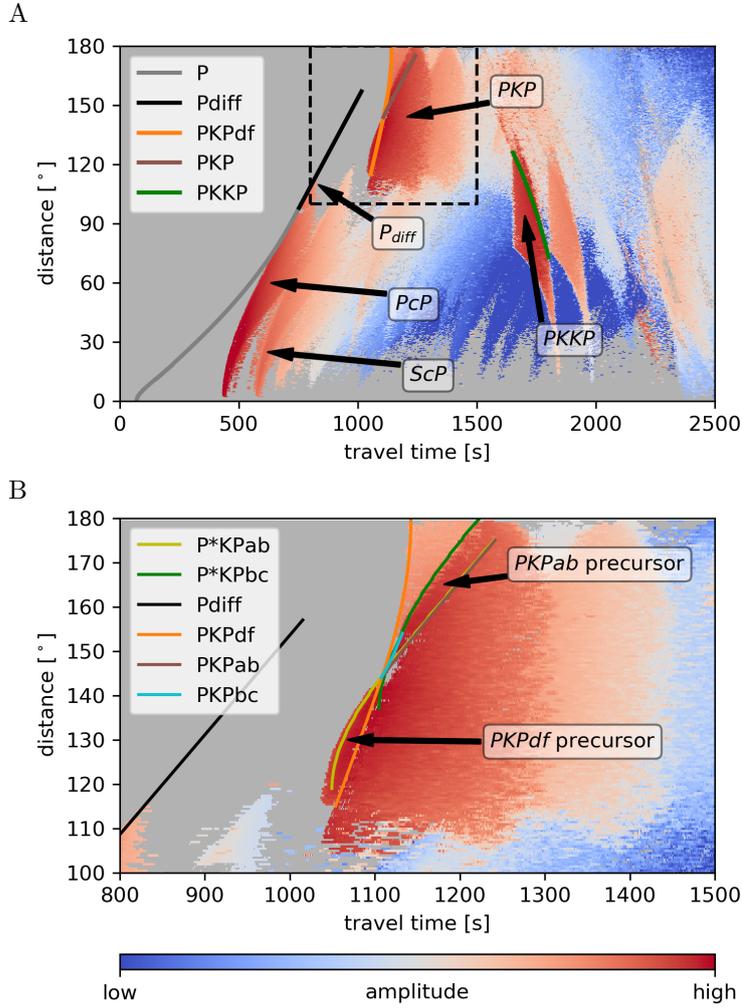
Hypotheses for the origin of the LLSVPs include primordial thermochemical piles of high-density material that accumulated early on in Earth's history and formed a basal *mélange* (Tackley, 2012). Other Hypotheses propose the accumulation of chemical heterogeneity over long geologic timescales through subducted oceanic crust (Li et al., 2014). The presence of post-perovskite (Koelemeijer et al., 2016) best explains the seismic signature of the LLSVP with a reduced shear wave velocity and a normal compressional wave velocity. Estimates of the vertical extent of the LLSVP above the CMB reach up to several 100s of kilometers (McNamara, 2019). The structure at the top of the LLSVP depends on the plumes that rise from the LLSVP. While some geodynamic models predict plumes rising predominantly from the edges of LLSVPs, others predict that smaller plumes may rise from the top of the entire LLSVP area. In fact, the LLSVPs could consist of many thin plumes that focus on large-scale upwelling areas and appear as continuous low velocity features only due to tomographic filtering (Schuberth et al., 2009).

In summary, seismological observations and geodynamic models demonstrated that the lower mantle is a region that might be characterized by chemical heterogeneity but is undoubtedly subject to thermal heterogeneity. Due to viscosity dependence of temperature, the length scale of the thermal heterogeneity can be significantly smaller than what is expected from thermal diffusion, e.g., by the formation of narrow plumes.

## 1.2 Wave scattering in the deep Earth

From geological structures at the surface and the investigation of high-frequency wave scattering in the Earth's crust, we know that geological materials differ not only in their large scale average elastic properties like wave velocity but also, in their small scale internal structure at a length scale below the resolution of seismic imaging (Sato et al., 2012). The statistical properties of these elastic parameter fluctuations are characteristic of the geologic material and can be observed due to the signatures they leave in seismograms. When seismic waves propagate through a heterogeneous medium, the waves are scattered and frequently change direction such that interference generates a complex wavefield (Sato et al., 2012). Scattering in the Earth's crust generates coda waves that follow the arrival of ballistic phases from local or teleseismic earthquakes (Obara & Sato, 1995; Sens-Schönfelder et al., 2009; Gaebler et al., 2015). The envelope of such complex wavefields can be used to investigate the statistical properties of the heterogeneity. The interplay between the length scale of the heterogeneity, wavelength, and intrinsic attenuation of seismic waves causes scattered waves to be best observed at frequencies above 1 Hz. Investigation Earth with scattered waves is different from ballistic waves. Scattered waves do not propagate along deterministic paths predicted by ray theory, but reach the receiver on complicated trajectories that can only be described in a probabilistic sense (Pacheco & Snieder, 2005).

Since the wave velocity at the core is lower than at the mantle, scattering in the deep Earth can cause seismic energy to arrive both at the coda of a ballistic phase and prior to a ballistic phase as a precursor.



**Figure 1.** Increase of seismic intensity due to scattering in a 50 km thick layer above the CMB. Simulations used a 600 km deep P-wave source in the velocity and attenuation model ak135-f (Kennett et al., 1995; Montagner & Kennett, 1996). (A) Arrival times of seismic phases and relevant regions of the time-distance domain that have been investigated for scattering in the deep Earth are indicated. (B) zoom into the time-distance window of PKP waves (dashed box, panel A). Theoretical arrival times for waves scattered at the CMB are indicated and labeled with '\*' indicating the scattering event. The frequently discussed *PKPdf* precursor and the *PKPab* precursor discussed below are labeled.

Figure 1 shows the increase of scattered intensity due to a 50 km thick scattering layer above the CMB simulated with differential radiative transfer simulations as detailed in the supporting information Text S1 which contains additional references to Takeuchi (2016) and Trabant et al. (2012). It shows a number of time-distance windows of the global wave field that have been investigated for waves scattered in the deep Earth. *ScP* and *PcP* top side reflections at the CMB can show precursors that originate by reflections above the CMB as well as coda waves from reverberations in the heterogeneous layer or off great-circle reflections (Wu et al., 2014; Gassner et al., 2015; Shen et al., 2016). Short distance *PKKP* precursors (A. Chang & Cleary, 1978; A. C. Chang & Cleary, 1981; P. S. Earle & Shearer, 1997) also originate from off great-circle bottom side reflections at the CMB (c.f. Figure 1A). *PKP* precursors probe the D'' layer in near-vertical transmission. Scattering of the *PKPab* branch can divert waves in the distance range up to  $145^\circ$  which would not be accessible to *PKPab*, otherwise (Haddon & Cleary, 1974; Hedlin et al., 1997). These waves form *PKPdf* precursors that arrive before the *PKPdf* phase that travels through the inner core (*PKIKP*) and is the earliest phase in the core shadow. This situation provides exceptional conditions for the observation of *PKPdf* precursors (c.f. Fig. 1B). Opportunities to probe the lower mantle by transmission in a near-horizontal direction is provided by *Pdiff* coda (c.f. Fig. 1). While diffraction along the core-mantle boundary vanishes with increasing frequency, at short period *Pdiff* coda waves in the core shadow zone, have been interpreted as a sign of scattering along the CMB (Bataille & Lund, 1996) or, as a signature of scattering throughout the mantle (P. Earle & Shearer, 2001). An overview of the travel time-distance windows in which scattered waves from the deep Earth can be observed, is given in Shearer (2007).

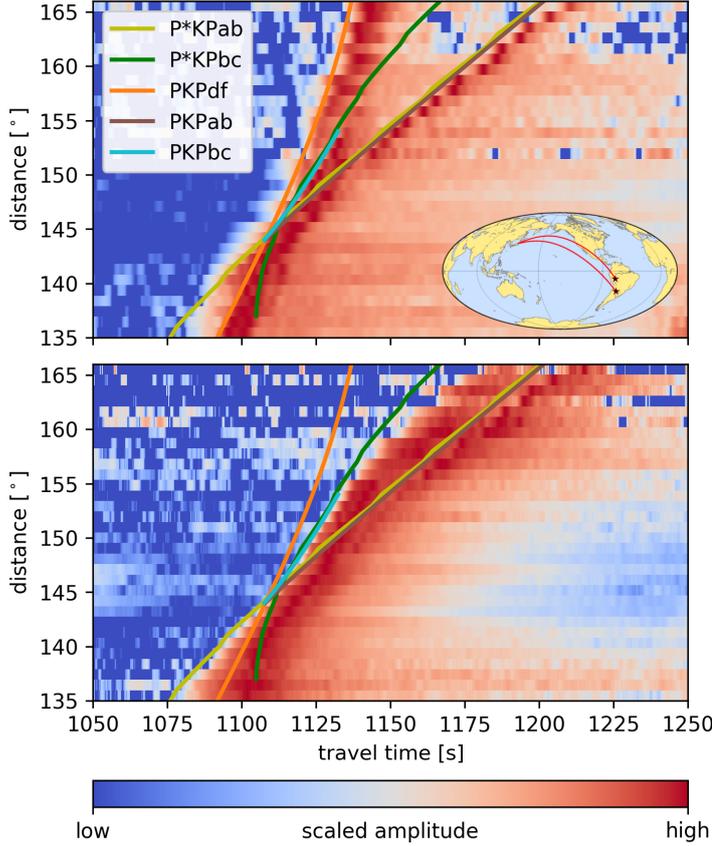
Understanding the origin of such faint signals arriving from the deep Earth provides a powerful tool to investigate the small scale structure of the deep mantle in terms of its statistical properties, i.e., the strength of elastic parameter fluctuations and their size distribution. It can yield valuable information about the distribution of chemical or thermal heterogeneity without the blurring effect of the tomographic filter. It also may help to constrain the depth extent and lateral distribution of features like plume clusters (McNamara, 2019), or accumulations of heterogeneous material in the basal mélange formed from subducted slabs (Tackley, 2012).

## 2 Observation of the *PKPab* precursors

Additionally to the *PKPdf* precursor at  $\Delta < 145^\circ$ , Fig. 1B shows a further arrival of scattered energy at distances  $\Delta > 155^\circ$ . For reasons discussed later, we term this phase *PKPab* precursor. This phase has been discussed sporadically in the literature, and there is no consensus about its origin. Waves propagating through the inner core arrive earlier in this distance range, and it is not clear whether the scattered energy that arrives between the *PKPdf* and *PKPab* should be regarded as a coda of *PKPdf* or as a precursory signal to *PKPab*. In contrast to the *PKPdf* precursor at  $\Delta < 145^\circ$  the *PKPab* precursor at  $\Delta > 155^\circ$  in Fig. 1 might thus be hidden in the *PKPdf* coda depending on the relative strength of both signals.

A possibility to observe the *PKPab* precursor unambiguously is to show its spatial coherency over an extended distance range. To avoid the effect of source-side crustal scattering, we use large deep earthquakes. Since lateral variability of D'' scattering could disturb the spatial coherency when records from different areas are combined, we try to use records from compact regions. Deep sources in South America recorded by the dense Japanese HiNet seismic stations (NIED, 2019; Okada et al., 2004; Obara et al., 2005) offer a perfect source-receiver configuration to observe the desired signals.

Fig. 2 shows HiNet vertical seismogram envelopes from two events stacked in  $1^\circ$  distance bins. The first is a 570 km deep event with Mw 6.8 from January 1st, 2011, in Argentina that covers  $151^\circ < \Delta < 167^\circ$  while the 592 km deep Mw 7.5 Peru event

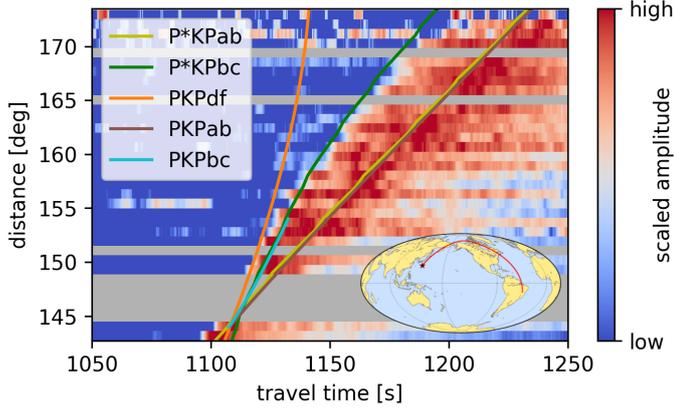


**Figure 2.** Composite image of stacked seismogram envelopes from Argentina and Peru deep earthquakes recorded in Japan. Arrival times of ballistic and scattered core phases are indicated. Top and bottom panels show the 0.35 - 0.7 Hz 4 - 8 Hz frequency bands, respectively. The logarithmic color scale is scaled between maximum and noise level for the individual distance bins. Inset shows the great circle between epicentres (stars) and recording stations.

159 from November 24th, 2015, covers  $135^\circ < \Delta < 152^\circ$ . Data processing for figure Fig. 2  
 160 is described in the supporting information Text S2. Two frequency bands are shown in  
 161 Fig. 2. The low-frequency band between 0.35 Hz and 0.7 Hz shows energetic arrivals fol-  
 162 lowing the *PKPdf* and *PKPab* travel time curves and some energy arriving prior to the  
 163 *PKPdf* below  $145^\circ$  – the known *PKPdf* precursor. Fig. 2B and Figures S1 and S2 (sup-  
 164 porting information) show the same data filtered in the 4-8 Hz frequency band and dif-  
 165 fers significantly from the low-frequency panel.

166 Three main observations can be made: (A) Significant amount of energy travels  
 167 through the entire Earth in the 4-8 Hz band. (B) The *PKPdf* phase is strongly atten-  
 168 uated on the path from Peru to Japan at these high frequencies compared to the *PKPab*  
 169 phase. There is no indication of energy propagating through the inner core in the 4-8 Hz  
 170 band. (C) A distinct increase of energy follows the lines of the earliest possible scattered  
 171 energy arrival from the CMB as indicated by the lines labeled *P\*KPab* and *P\*KPbc*  
 172 in Fig. 2. We call this phase *PKPab* precursor.

173 We would like to emphasize that the presence of the *PKPab* precursor is not due  
 174 to a local effect at the source of the event (Argentina) or local disturbances within the



**Figure 3.** Stacked seismogram envelopes from the Bonin deep earthquakes recorded in Brazil. Arrival times of ballistic and scattered core phases are indicated. Gray intervals represent gaps in the distance coverage of the network. The logarithmic color scale is scaled as in Fig. 2. Inset shows the great circle between epicentre (star) and station network.

175 HiNet. Fig. 3 shows the stacked envelopes of the May 30th, 2015 deep Bonin Islands earth-  
 176 quake (Mw 7.8, depth 677 km) recorded at stations from part of the Brazillian Seismo-  
 177 graphic Network (Bianchi et al., 2018), network codes BL and BR. A clear signal of the  
 178  $PKPab$  precursor following the  $P * KPbc$  arrival time is observed for this wave path,  
 179 too.

### 180 3 Origin of the $PKPab$ precursor

181 The onset of  $PKPab$  precursor emerges at the  $c$ -caustic that connects the  $PKPbc$   
 182 and  $PKiKP$  (inner core reflection) branches with a common slowness. Thus, it seems  
 183 reasonable to assume a relation of the  $PKPab$  precursor to one of these two phases. Pos-  
 184 sible mechanisms could be (A) diffraction of  $PKiKP$  waves along the inner core bound-  
 185 ary (ICB) or the propagation through a heterogeneous waveguide above the ICB, or (B)  
 186 deviation of  $PKPbc$  waves into the shadow of the inner core by scattering in the man-  
 187 tle or outer core. Feasibility to differentiate between these two possibilities is provided  
 188 by the slowness-distance relation of the earliest energy arrival. For mechanism (A), the  
 189 energy diffracted along the ICB should arrive with constant slowness for all distances.  
 190 This should be the slowness of  $PKiKP$  waves at the  $c$ -caustic or a somewhat higher but  
 191 constant slowness if a low-velocity layer is invoked at the ICB. Since the onset of the scat-  
 192 tered energy is clearly curved to higher slowness for increasing distances (Fig. 2C and  
 193 3) the observations do not favor the ICB-diffraction mechanism (A).

194 Mechanism (B) i.e., the deviation of  $PKPbc$  wave direction, would mean that part  
 195 of the  $PKPbc$  wave energy that travels just atop the inner core gets scattered on its path  
 196 through the Earth. Depending on the depth distribution of the heterogeneity that causes  
 197 the scattering, different onset times are possible. However, from the  $PKPdf$  precursor  
 198 at distances  $\Delta < 145^\circ$  it is known that especially the D'' layer above the CMB scat-  
 199 ters wave energy, and is thus a right candidate.

200 Deviating the propagation direction of  $PKPab$  waves at the source (or receiver)  
 201 side to create  $P * KPab$  ( $PKab * P$ ) waves explains the arrival time of the  $PKPdf$  pre-  
 202 cursor energy for  $\Delta < 145^\circ$  (Fig. 2). For  $\Delta > 145^\circ$   $P * KPab$  energy arrives coinci-  
 203 dent with ballistic  $PKPab$ .

204 On the other hand, deviating the propagation direction of *PKPbc* waves at D' can  
 205 shed energy in the distance range  $\Delta < 145^\circ$  that arrives after the *PKPdf* precursor  
 206 and the *PKPdf* wave and is thus hard to observe. For  $\Delta > 155^\circ$  the  $P * KPbc$  en-  
 207 ergy arrives prior to the *PKPab* phase. Since the earlier *PKPdf* arrival is strongly at-  
 208 tenuated in the high frequency, as shown in Fig. 2, the scattered  $P * KPbc$  energy forms  
 209 the first notable arrival.

210 We summarize that (A) scattering of core phases in the lower mantle is a commonly  
 211 accepted process as confirmed for example by the *PKPdf* precursor at  $\Delta < 145^\circ$  and  
 212 (B) in simulations of energy propagation considering a scattering in the lower mantle pre-  
 213 dict the arrival of energy that is in qualitative agreement with the observation of the *PKPab*  
 214 precursor (compare Fig. 1 and 2, 3). These ideas strongly support the hypothesis that  
 215 the observed *PKPab* precursor at  $\Delta > 155^\circ$  is a consequence of the same process that  
 216 causes the well known *PKP* precursor at  $\Delta < 145^\circ$ .

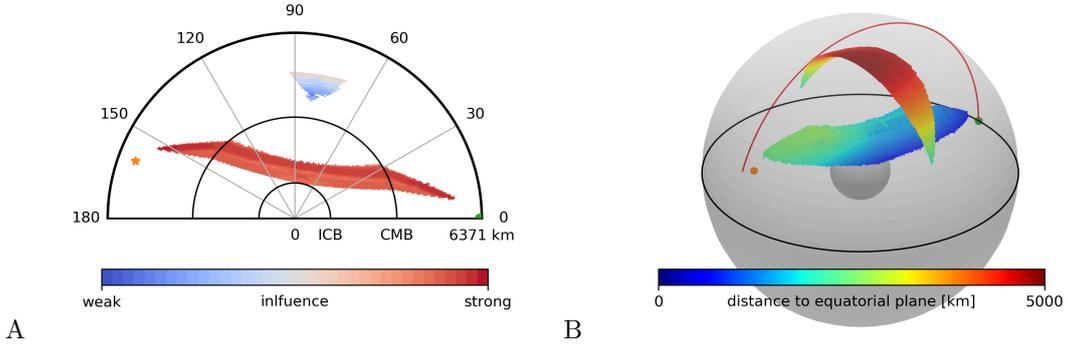
#### 217 4 Local Sensitivity of the *PKPab* Precursor to Scattering

218 Waves scattered in the deep Earth provide means to investigate the structure of  
 219 the lower mantle at a spatial scale below the resolution limits of seismic tomography. The  
 220 *PKPab* precursor offers a new opportunity for this. Here we investigate the spatial sensi-  
 221 tivity of this signal. We use the theory of Margerin et al. (2016) and Zhang et al. (2020)  
 222 to derive an intensity sensitivity kernel, which describes the sensitivity of the seismogram  
 223 envelope to a local increase of scattering strength. We simplify the treatment in three  
 224 ways. (A) Wave propagation through the inner core is blocked. Since we observe that  
 225 *PKPdf* waves vanish in the 4-8 Hz frequency range (cf. Fig. 2, 3, S1), waves that prop-  
 226 agate through the inner core cannot contribute to the scattered arrival either. Scatter-  
 227 ing within the inner core would generate *PKPdf*-coda rather than a separate phase that  
 228 is disconnected from the *PKPdf* arrival. (B) We assume that scattering leading to the  
 229 *PKPab* precursor is isotropic, which simplifies the treatment of scattering angles. In-  
 230 creased probability of forward scattering would reduce the probability of scattering close  
 231 to either source or receiver. (C) The scattering process is restricted to a single scatter-  
 232 ing of P-waves because S-wave propagation is highly unlikely for the short travel time  
 233 of the *PKPab* precursor.

234 Under the assumptions made, we can calculate the volume in which scattering can  
 235 contribute to the observed *PKPab* precursor by convolution of the forward and back-  
 236 ward P-wave intensity. These intensities can be obtained by radiative transfer simula-  
 237 tions, as described in Text S1 for excitation at the location of the earthquake (forward  
 238 simulation) and the location of the receiver (backward simulation). The sensitivity fi-  
 239 nally describes the probability of a wave packet that arrives in a particular time-distance  
 240 window to have traveled from the source to a particular location in space where scat-  
 241 tering occurred, and then continued to the receiver location.

242 Fig. 4A shows a cross-section through the sensitivity kernel in the great circle plane  
 243 for an epicentral distance of  $\Delta = 160^\circ$  and a lapse time of 1155 s, which is within the  
 244 time-distance window of the *PKPab* precursor. It describes the influence of heterogene-  
 245 ity (i.e. the possibility for wave scattering) on the amplitude of the *PKPab* precursor.  
 246 Regions, where this probability is high, have a strong influence on the precursor ampli-  
 247 tude. If this probability is low at some location, the influence is weak because it is un-  
 248 likely that a wave arriving in the time-distance window was scattered there. Zero influ-  
 249 ence means that it is impossible for wave energy to arrive in the time-distance window  
 250 of the *PKPab* precursor even if it is scattered there.

251 High sensitivity is located along the *PKPbc* path through the outer core and lower  
 252 mantle. In this narrow volume, waves are scattered mostly in the forward direction mean-  
 253 ing that small perturbations of the propagation direction of *PKPbc* waves can gener-



**Figure 4.** Volume of sensitivity for an arrival at 1155 s lapse time and epicentral distance of  $\Delta = 160^\circ$ . Orange star and green circle indicate locations of source and receiver, respectively. (A) cross section in the great circle plane of source and receiver with warm colors indicating high sensitivity of the arrival to scattering. CMB and ICB are indicated. (B) 3D representation of the volume of sensitivity with color indicating distance to the equatorial plane. Two distinct regions of sensitivity exist. One is draped on the inner core along the  $PKPbc$  path. Due to the high  $PKP$  amplitudes heterogeneity in this volume has a strong influence on amplitudes of the precursor. Another sickle-shaped region of sensitivity that allows for large deviations from the great circle is formed by scattering of  $P$  waves in the mantle. Heterogeneity in this region, however, has less influence on the amplitude of the  $PKPab$  precursor than in the elongated region that extends through the deep Earth (c.f. Fig. 4A).

254 ate the  $PKPab$  precursor at 1155 s at  $160^\circ$  distance. Due to the high amplitude of the  
 255  $PKP$  phase the influence of this region on the  $PKPab$  precursor amplitude is high as  
 256 indicated by the color in Fig. 4A. Another patch of sensitivity is located in the mid man-  
 257 tle. It indicates that scattering of  $P$ -waves in the mid mantle allows waves to travel around  
 258 the slow outer core and still carry energy to a receiver in the time-distance window of  
 259 the  $PKPab$  precursor. However, considering the smaller amplitudes of the participat-  
 260 ing waves, there is a low probability that the scattering in this region contributes to the  
 261 observed signal – resulting in a weak influence of this region on the precursor.

262 Since scattering allows for off great-circle path propagation, the sensitivity has a  
 263 significant 3D component, as illustrated in Fig. 4B. The volume with the strong influ-  
 264 ence on the  $PKPab$  precursor that extends through the deep Earth is draped on the in-  
 265 ner core and shows small deviations from the great-circle plane. The region of  $P$  wave  
 266 scattering in the mid mantle forms a sickle-shaped volume of sensitivity, perpendicular  
 267 to the great circle plane. Energy in the  $PKPab$  precursor window that were scattered  
 268 in the mantle can, therefore, arrive with significant deviations from the great-circle di-  
 269 rection.

## 270 5 Discussion

271 Using numerical simulations, we show that scattering in the lower mantle results  
 272 in the arrival of scattered energy before the  $PKPab$  at  $\Delta > 155^\circ$ . This energy arrives  
 273 after  $PKPdf$ . In the high-frequency band between 4 and 8 Hz, waves do not propagate  
 274 through the inner core on the path between South America and Japan. This vanishing  
 275 of the  $PKPdf$  energy makes the  $PKPab$  precursor the first notable arrival of the record,

276 which can be readily observed in individual records of deep earthquakes. We speculate  
 277 the *PKPab* precursory signal is also present at lower frequencies where it is masked by  
 278 the earlier *PKPdf* arrival and its coda.

279 The origin of the *PKPab* precursor has been discussed earlier. A number of arti-  
 280 cles discussed the *PKP-C<sub>diff</sub>* phase that should result from the diffraction of com-  
 281 pressional waves around the inner core along the ICB. Nakanishi (1990) present obser-  
 282 vations of 2.5-3.3 Hz *PKP - C<sub>diff</sub>* waves in the distance range  $152^\circ < \Delta < 157^\circ$ .  
 283 From the complex, long-lasting waveforms, their earliest arrival and the high slowness  
 284 Nakanishi (1990) concluded that scattering at the base of the upper mantle around 660 km  
 285 depth is more likely to generate these arrivals than ICB diffraction. Tanaka (2005) in-  
 286 vestigated *PKP-C<sub>diff</sub>* coda using short-period seismic arrays and found slowness rang-  
 287 ing between 1 and  $5s/^\circ$  extending through the whole range covered by *PKPab* and *PKPbc*  
 288 waves. Scattering at the CMB was invoked as an alternative origin of the *PKP-C<sub>diff</sub>*  
 289 coda signal, since the slowness of waves scattered close to the c-caustic is close to that  
 290 of *PKP-C<sub>diff</sub>* waves to be separated by the arrays. These early works are thus in agree-  
 291 ment with our interpretation of the *PKPab* precursor as scattered *PKPbc* with a likely  
 292 location of the scattering close to the CMB.

293 Adam and Romanowicz (2015) report on a scattered phase that arrives 5-20 s af-  
 294 ter the *PKPbc* of *PKPbc<sub>diff</sub>* which they call *M*-phase. Adam and Romanowicz (2015)  
 295 uses coherent stacking of 1 Hz signals within distance ranges up to  $10^\circ$  and concludes  
 296 that the scattered *M*-phase originates at the ICB. Scattering at the CMB is ruled out  
 297 because the *M*-phase appears as an isolated phase in the phase weighted stack with a  
 298 slowness between  $0.7 - 1.6s/^\circ$ . This slowness is too low for *PKPbc* waves scattered at  
 299 the CMB beyond  $160^\circ$  distance. This finding appears to contradict our interpretation.  
 300 However, firstly the 1 Hz frequency range differs from our observation and the argument  
 301 that Adam and Romanowicz (2015) uses to rule out the possibility of *PKPbc* scatter-  
 302 ing close to the CMB is strongly based on the limitation of the slowness range to  $1.6s/^\circ$   
 303 maximum. This constraint is derived under the assumption of distance independent slow-  
 304 ness even though it is not shown that the *M*-phase at  $\Delta > 160^\circ$  has a slowness below  
 305  $1.6s/^\circ$ . The fact that the *M*-phase appears as an isolated phase is enforced by the phase  
 306 weighted stacking and did not exclude the actual presence of an extended wave train orig-  
 307 inating from waves with a significant spread of slowness.

308 Thus, we think that our interpretation of the *PKPab* precursor, as scattered *PKPbc*  
 309 waves is compatible with earlier studies. The discussed *PKP-C<sub>diff</sub>* phase, as well as  
 310 the *M*-phase, may be interpreted as a signal with the same origin. Since the heterogene-  
 311 ity at D'' is widely accepted, it should be taken into account in any interpretation of sig-  
 312 nals that might have passed through D''. In fact, this is true for all investigations of the  
 313 inner core. The difference in coda decay between *PcP* and *PKiKP* coda at small dis-  
 314 tances should not be interpreted without considering the effect of the twofold *PKiKP*  
 315 transmissions through D'' which can significantly alter the shape of the coda.

316 The possibility to observe the *PKPab* precursor at  $\Delta > 155^\circ$  requires strong at-  
 317 tenuation of the earlier arriving *PKPdf* waves that pass through the inner core. Lon-  
 318 gitudinal variations of *PKiKP* vs. *PKPdf* travel time and amplitude differences indi-  
 319 cate hemispherical asymmetry of inner core attenuation (Monnereau et al., 2010). More-  
 320 over, this will likely influence the observability of the *PKPab* precursor. The *PKPab*  
 321 precursor in locations where it can be observed can increase the lateral resolution of *PKP*  
 322 CMB based studies. Combined with *PKPdf*, it allows using earthquakes from a much  
 323 wider distance range.

324 As indicated by the elongated shape of the sensitivity kernel in Fig. 4, the verti-  
 325 cal resolution to determine the location of scattering is relatively poor. Since the required  
 326 deviation of the propagation direction (scattering angle) is small, the scattering can hap-  
 327 pen almost anywhere between the source and receiver. However, it is known from array

analysis of the *PKPdf* precursor (Thomas et al., 1999, e.g.) that the most likely location of scattering is D". The theoretical possibility of propagating seismic energy in the time-distance window of the *PKPab* precursor by *P\*P* scattering in the mid mantle (c.f. Fig. 4) is challenging to test because of the much stronger *PKP* phases. However, for scattering deeper in the mantle, the *P\*P* scattered waves can arrive prior to any scattered core phase and could be used to investigate scattering above D".

## 6 Conclusion

We show that the frequency range for investigation of the deep Earth with teleseismic waves can be extended towards frequencies of several Hertz. The attenuation of high frequency waves in the inner core allows for the observation of scattered *PKPbc* waves as *PKPab* precursor in the shadow of the inner core. Without this attenuation, the *PKPab* precursor would be masked by the *PKPdf* coda. This situation is similar to the *PKPdf* precursor that can only be observed so clearly as the first arriving phase because the low velocity core deviates the *P* phase – thereby creating the (outer) core shadow.

We calculate the sensitivity kernels of the *PKPab* precursor for heterogeneity using elastic radiative transfer simulations. The kernels describe the Earth's region in which scattering would contribute to seismic energies arrival in a given time-distance window. Scattering in D" that causes the *PKPdf* precursor at  $\Delta < 145^\circ$  is also the most likely mechanism causing the *PKPab* precursor at  $\Delta > 155^\circ$ . Combining these sensitivities kernels with observations of scattered energy from *PKPab* and *PKPdf* precursors will improve the imaging and characterization of heterogeneity in the deep Earth .

## Acknowledgments

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# Supporting Information for "High Frequency (6 Hz) PKPab precursors and their sensitivity to deep Earth heterogeneity"

C. Sens-Schönfelder<sup>1</sup>, K. Bataille<sup>2,1</sup>, M. Bianchi<sup>3</sup>

<sup>1</sup>GFZ German Research Centre for Geosciences, Telegrafenberg, 14473 Potsdam, Germany

<sup>2</sup>Earth Sciences Department, University of Concepcion, Chile

<sup>3</sup>Instituto de Astronomia, Geofísica e Ciências Atmosféricas, Universidade de São Paulo, Brazil

## Contents of this file

1. Text S1 to S2
2. Figures S1

**Introduction** In the following we describe the numerical simulations of the deep Earth scattering and of the data processing used to extract the signals of the *PKPab* precursor. We also show the stacked envelopes of the HiNet records for each 1° distance bin.

## Text S1. Differential Monte-Carlo Simulation of Deep Earth Scattering

Radiative Transfer Theory can describe the propagation of scattered seismic waves. We use a version of the elastic simulation code described by Sens-Schönfelder, Margerin, and Campillo (2009) that we adapted to spherical geometry with a 1D velocity and attenuation structure. This code has already been used to model the teleseismic waves by Gaebler,

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Sens-Schönfelder, and Korn (2015). To highlight the effects of localized scattering, we introduce a further conceptual modification that allows us to directly model the *change* of seismic intensity due to the presence of scattering in a specified part of the model. We call this differential modeling. Takeuchi (2016) has used a similar approach.

In the Radiative Transfer approach, the propagation of seismic wave energy is simulated by the number density of a large number of particles (wave packets). The particles propagate through the domain according to the ray theory. Scattering is simulated by discrete scattering events governed by the statistical properties of the medium's heterogeneity. Intrinsic attenuation is accounted for by reducing the weight of the particles.

To simulate the differential intensity, we modify the weights of the particles by an additional factor ( $S$ ). Let us call the region under investigation  $G$  in which the change in scattering properties should be modeled. When particles are launched from the source we set  $S = 0$ . This changes only upon scattering in  $G$ . When a particle is scattered in  $G$  we have to model the increase of scattered intensity as well as the decrease of ballistic intensity. This is done in a probabilistic sense by either changing the direction of the particle and setting  $S = 1$  to simulate the increase of scattered intensity with a probability of 50% or by simply setting  $S = -1$  and keeping the propagation direction to simulate the decrease of ballistic intensity with a probability of 50%. A particle with  $S = -1$  does not interact with the heterogeneity in the  $G$ .

The simulations in this paper use a modified version of the ak135-f model (Kennett et al., 1995; Montagner & Kennett, 1996) obtained from the IRIS DMC Data products (Trabant et al., 2012) with doi:10.17611/DP/9991801. The modification comprised replacing the

shallow partly liquid structure at the Earth's surface with constant structure corresponding to the top side of the discontinuity at 10 km depth.

### **Text S2. Processing of Envelope Stacks**

The high frequency seismograms that we use to observe the scattered wave are affected by local noise, site factors and station sensitivity. To visualize the stacked envelopes, we use the following processing steps.

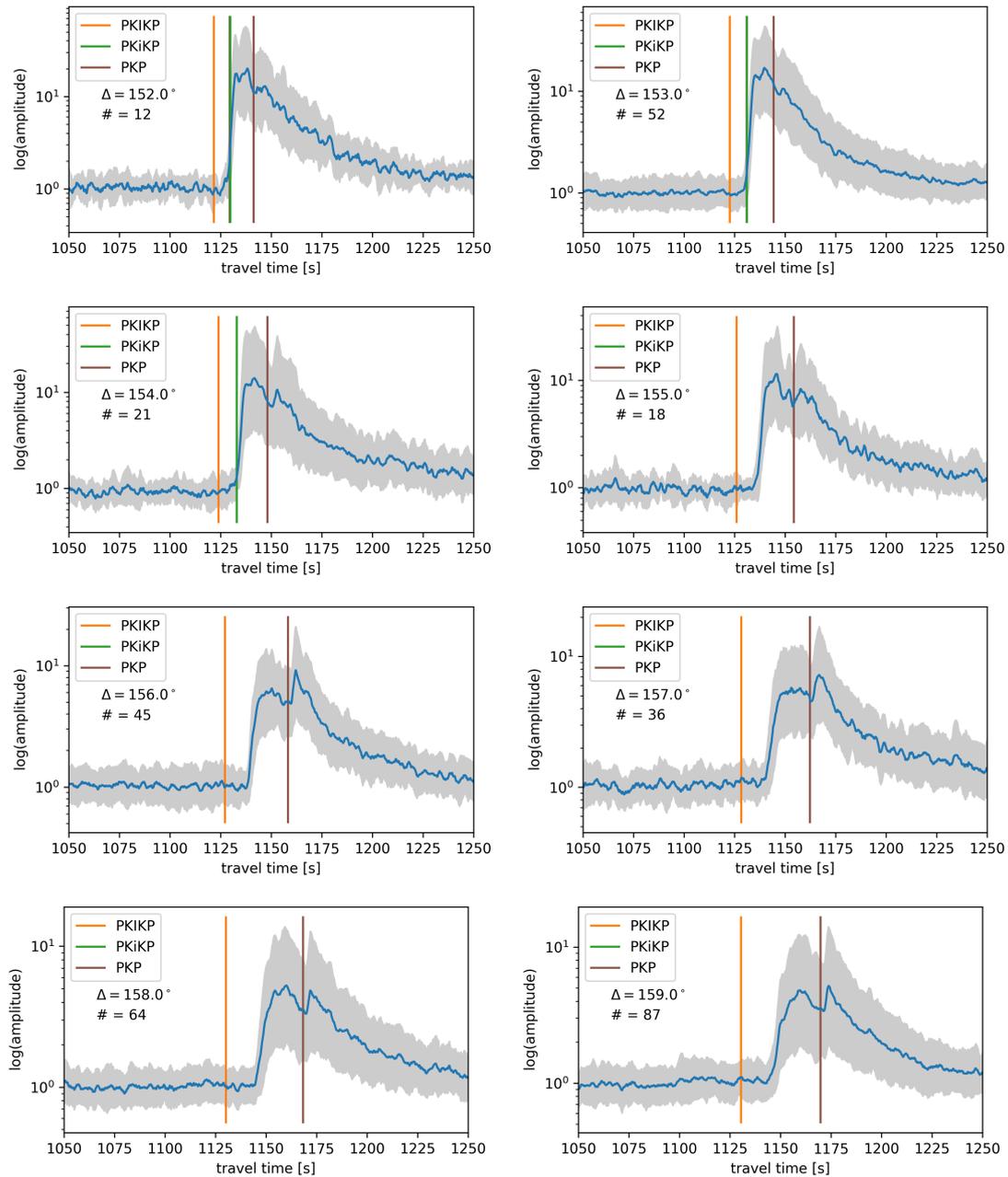
1. data selection
2. filtering
3. envelope calculation using instantaneous amplitude
4. temporal smoothing of logarithmic envelopes
5. alignment to reference phase travel time
6. stacking of logarithmic envelopes in distance bins
7. subtraction of noise level
8. normalization to the maximum amplitude value

### **Figure S1. Data of the *PKPab* precursor**

Here we show the data used to create the color images of the time-distance sections as individual traces, with gray background indicating the pointwise logarithmic standard deviation when different records have been stacked. Figures S1 and S2 show the data from the Jan 1st 2011 Argentina event recorded by the HiNet stations in Japan, processed as described above but without the normalization to the maximum in the last item.

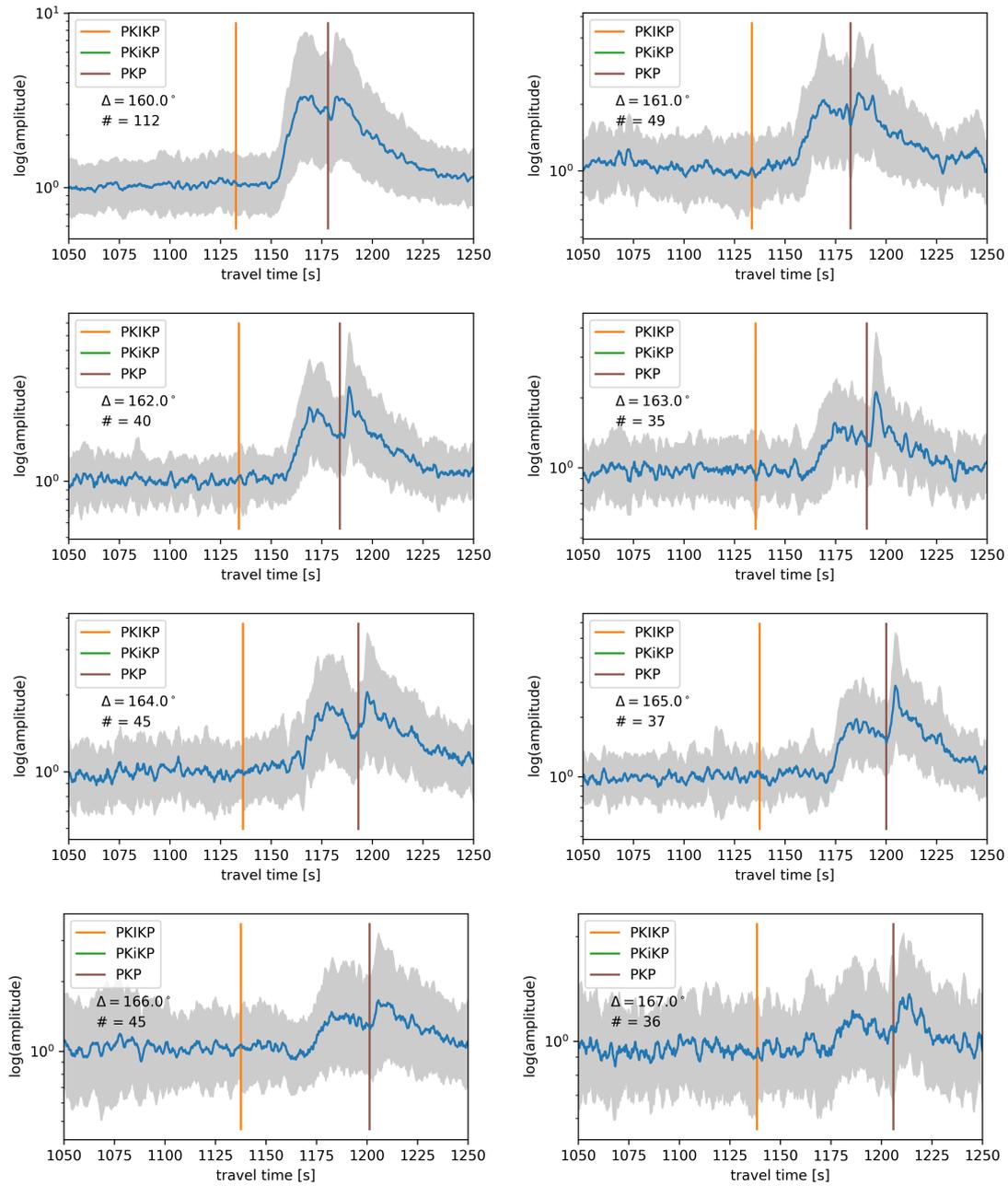
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**Figure S1.** Stacked HiNet records of the Jan 1st 2011 Argentina event for  $1^\circ$  wide bins cantered at the distances given in each panel. The number of stacked records is indicated in each panel.

June 5, 2020, 3:03pm



**Figure S2.** Same as Fig. S1 for further distances.