# Cluster Curlometry Limitations in the Ring Current Region

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#### Abstract

During its ongoing mission, the Cluster II constellation has provided the first small-scale multipoint measurements of the space environment, and dramatically advanced scientific understanding in numerous regimes. One such region is the Earth's inner magnetospheric ring current, which could now be computed using the curl of the magnetic field over a spacecraft tetrahedron instead of via plasma moments. While this produced the first 3D current estimates, it also produced different results from prior ring current studies with differing magnitudes and correlations with storm indices/local times. In this analysis, we revisit Cluster ring current data via curlometry, and conduct additional quantitative sensitivity simulations using actual spacecraft position data. During the orbits that observed ring current structure, tetrahedron shape and linearity assumptions can create large errors up to 100% of physical current magnitude in curlometer output that contradict accepted estimated quality parameters. These false currents are directly related to the structure of the current environment, and cannot be distinguished from the actual currents without additional limiting assumptions. The trustworthiness of curlometer output in the ring current is therefore dependent on the linearity of the magnetic structure relative to the tetrahedron orientation, which requires additional characterization. The Cluster curlometer output in the ring current is then explored in light of these new uncertainties, with the computed current magnitude and direction both potentially impacted by the production of false currents.

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9	Key Points:
10 11 12 13 14 15 16	<ul> <li>Examination of Cluster curlometer output shows strong evidence of contamination by linearization errors in the ring current region.</li> <li>False currents are computed by the curlometer when a large tetrahedron dimension aligns with the magnetic gradient.</li> <li>The curlometer technique has highly constrained utility in the inner magnetosphere that requires careful consideration of limitations.</li> </ul>
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#### 24 Abstract

During its ongoing mission, the Cluster II constellation has provided the first small-scale 25 multipoint measurements of the space environment, and dramatically advanced scientific 26 understanding in numerous regimes. One such region is the Earth's inner magnetospheric ring 27 current, which could now be computed using the curl of the magnetic field over a spacecraft 28 tetrahedron instead of via plasma moments. While this produced the first 3D current estimates, it 29 also produced different results from prior ring current studies with differing magnitudes and 30 correlations with storm indices/local times. In this analysis, we revisit Cluster ring current data via 31 32 curlometry, and conduct additional quantitative sensitivity simulations using actual spacecraft position data. During the orbits that observed ring current structure, tetrahedron shape and linearity 33 assumptions can create large errors up to 100% of physical current magnitude in curlometer output 34 that contradict accepted estimated quality parameters. These false currents are directly related to 35 the structure of the current environment, and cannot be distinguished from the actual currents 36 without additional limiting assumptions. The trustworthiness of curlometer output in the ring 37 38 current is therefore dependent on the linearity of the magnetic structure relative to the tetrahedron orientation, which requires additional characterization. The Cluster curlometer output in the ring 39 current is then explored in light of these new uncertainties, with the computed current magnitude 40 and direction both potentially impacted by the production of false currents. 41

#### 42 Plain Language Summary

The ring current is a structure in near-Earth space that causes magnetic changes on the 43 Earth's surface and is very important for plasma transport as well. Previously, it has been measured 44 45 using single spacecraft by sampling the electrically charged particles. However, the Cluster mission allowed the ring current to be calculated in greater detail by observing magnetic fields, 46 which are easier to measure. This technique was dubbed the 'curlometer,' and has applications in 47 many regions. For the ring current region in particular, the curlometer technique produces very 48 49 different ring current strengths than the particle-measurement methods. To resolve this, we reevaluated the curlometer technique and tested it against simulated data for which the current can 50 51 be analytically calculated, showing larger uncertainty than previously thought. Thus, the 52 curlometer has limitations for using the Cluster mission to measure the ring current.

#### 53 **1. Introduction**

The Cluster II multi-spacecraft mission, launched in 2000, has been a resounding success 54 in probing a multitude of environments by collecting data in varying three-dimensional 55 configurations. By providing multipoint measurements, Cluster benefits from improved 56 calibration and novel techniques that cannot be implemented by a single spacecraft (Paschmann & 57 Daly, 2000). One such technique has been named the 'curlometer,' because it uses the four 58 spacecraft as vertices of a tetrahedron to compute the linearized gradient of magnetic field and find 59 the average current density in the tetrahedron volume (Dunlop et al., 1988). Although requiring 60 61 limitations and assumptions, the curlometer technique has been applied to many regions containing magnetic structure that has a larger characteristic scale than the tetrahedron (Dunlop et al., 2016). 62 Thus, we have gained additional insight into the three-dimensional current structure of features 63 from the cusp to the magnetotail (e.g., Henderson et al., 2008; Dunlop et al., 2015; Petrukovich et 64 65 al., 2014).

In the ring current region of the inner magnetosphere (~3-7 R<sub>E</sub>) (e.g., see reviews and 66 references by Ganushkina et al., 2018, and Dandouras et al., 2018), the curlometer technique has 67 68 been applied to provide a better estimate of ring current magnitudes critical for understanding 69 storm indices and current closure, also qualitatively capturing field-aligned currents in this region (Vallat et al., 2005; Zhang et al., 2011). Curlometry offers an alternative to computing current from 70 plasma pressure moments, which have higher uncertainty through methodology and 71 72 instrumentation (Dandouras & Barthe, 2011). Pressure moments can also only vield a single 73 component of current orthogonal to the plane defined by the spacecraft trajectory and local magnetic field, whereas curlometry provides the full current vector. 74

75 Currents computed from Cluster via curlometry in the ring current region suggest a westward current that varies from near zero to a few tens of  $nA/m^2$ . Vallat et al. (2005) found that 76 77 the magnitude had no correlation with geomagnetic activity indicated by the Disturbance Storm 78 Time (Dst) index (see Figure 16 and Section 7, Vallat et al., 2005), and using subsequent orbits, other papers (Zhang et al., 2011; Grimald et al., 2012; Shen et al., 2014) have built statistical 79 80 compilations for the ring current using different thresholds for storm/quiet delineation. However, these results are contrary to plasma moment and single spacecraft magnetometer current 81 82 calculations, which suggest currents approaching 10 nA/m<sup>2</sup> occur only near storm-times, and are

consistently weaker than this value with lower activity (e.g., Lui et al., 1992; Greenspan & 83 Hamilton, 2000; Jorgensen et al., 2004; Le et al., 2004). Furthermore, time series or radial plots of 84 curlometer output (i.e., Vallat et al., 2005) often show structure that is inconsistent with 85 expectations; namely, near-constant current magnitude instead of an inverse relation with L-shell 86 (assuming the pressure peak is radially inward of perigee). These discrepancies urge a thorough 87 reanalysis of Cluster data to account for these observational differences and to probe the underlying 88 assumptions of current calculation (Liemohn et al., 2016). Herein, an approach to curlometry that 89 is mindful of limitations caused by tetrahedron geometry and alignment is developed to 90 characterize the curlometer technique's ability to reproduce known current systems. We assess the 91 Cluster tetrahedron configurations near perigee relative to imposed idealized current sheets, 92 examining the resulting current densities obtained from linear curlometry calculations. Through 93 94 simulation, the conditions that cause the curlometer to produce false currents are quantitatively characterized, adding a better understanding of the uncertainties associated with these ring current 95 96 measurements. Through simulation, the conditions that cause the curlometer to produce false currents are quantitatively characterized, adding a better understanding of the uncertainties 97 98 associated with these ring current measurements.

#### 2. Methodology 99

#### **2.1. Implementation of the Curlometer Technique** 100

The curlometer technique has been well-documented in numerous papers (e.g., Dunlop et 101 102 al., 1988; Robert et al., 1998; Vallat et al., 2005); thus, only a brief discussion is provided here. 103 According to the Maxwell-Ampere Law, and assuming stationarity (removal of time dependence 104 term):

105 
$$\mu_0 \vec{J} = \nabla \times \vec{B} \tag{1}$$

which can be rewritten with respect to a reference magnetic field vector at a reference location:

107 
$$\vec{J} \cdot \left( \left( \vec{r}_i - \vec{r}_{ref} \right) \times \left( \vec{r}_j - \vec{r}_{ref} \right) \right) = \frac{1}{\mu_0} \left( \left( \vec{B}_i - \vec{B}_{ref} \right) \cdot \left( \vec{r}_j - \vec{r}_{ref} \right) - \left( \vec{B}_j - \vec{B}_{ref} \right) \cdot \left( \vec{r}_i - \vec{r}_{ref} \right) \right)$$
(2)

With four spacecraft, the curl of the magnetic field can be computed by cyclically differencing 108 over each face of the tetrahedron to find the three-dimensional linear gradients and summing 109 (Equation 2), yielding the average current density within the tetrahedron volume. Note that in 110

Equation 2, 'r' denotes the spacecraft position vector, the 'ref' subscript refers to the reference spacecraft, and the 'i' and 'j' subscripts are iterated through non-reference spacecraft pairs. A more detailed treatment of the technique can be found in Dunlop et al. (1988) or Middleton and Masson (2016).

For this study, the curlometer computation was performed by modifying a Python script 115 provided by the Cluster Science Archive (CSA) (http://www.cosmos.esa.int/web/csa/multi-116 spacecraft/). The script was thoroughly tested for correct methodology: first with sample data 117 provided by CSA, then by alternating reference spacecraft and perturbing parameters, and finally 118 119 by 'flying' the constellation through simulated linear current environments. In each case, the output was as expected, demonstrating independence of reference spacecraft choice and correctly 120 capturing the simulated currents. This analysis followed the work of Robert et al. (1998) and 121 independently confirmed the results of that study. The code was further verified by replacing 122 123 magnetic field data with constant field, correctly producing no current, with an idealized dipole field, producing small current ( $< 2 \text{ nA/m}^2$ ) due to nonlinear magnetic gradients, and with the 124 125 International Geomagnetic Reference Field (IGRF) (Alken et al., 2021), which also created only small current outputs (< 3 nA/m<sup>2</sup> with 2002 tetrahedra) as expected (Dunlop et al., 2020). Thus, 126 the curlometer script was confirmed to function correctly in a variety of environments. A sample 127 of these tests for different tetrahedra is provided in supplemental material. Extensive testing 128 provided the necessary foundation for later conclusions by definitively verifying the curlometer 129 computation. 130

131 Magnetic field data from the Fluxgate Magnetometer (FGM) instruments on each spacecraft were obtained in spin resolution along with spacecraft ephemeris data. The resolution 132 of the magnetometer measurement at the observed field strength is roughly 0.125 nT (Balogh et 133 al., 1997): the effect of this resolution was determined to be no more than +/-2 nA/m<sup>2</sup> in curlometer 134 output in agreement with the analysis done by Vallat et al. (2005), and a sample is provided in 135 supplemental material. For this present study, temporal data resolution was experimentally 136 determined to have only a small effect on current values, within the ranges of available datasets; 137 therefore, spin resolution (~4 s cadence) sufficiently captures the scale of desired features. All 138 139 values were converted from Geocentric Solar Ecliptic (GSE) to Solar Magnetic (SM) coordinates 140 using the SpacePy Python package (Morley et al., 2011) and tested to ensure they retained their magnitudes. Note that westward azimuthal current is defined to be positive due to convention in 141

prior studies. IGRF values were then subtracted from the magnetic field data to remove as much
nonlinear magnetic gradient as possible and therefore allow a more robust linearity assumption
and curlometry result (Dunlop et al., 2016; Dunlop et al., 2020). This has a significant effect on
current densities for spacecraft separations above 200 km.

#### 146 **2.2. Ring Current Event Selection**

147 During the period of study for this analysis, the Cluster spacecraft operated in highlyelliptical orbits with perigee of ~4 RE, apogee of ~20 RE, inclination of ~90°, and period of ~57 148 149 hours. The spacecraft flew in various evolving formations that had interspacecraft separations less than 1000 km. The Cluster tetrahedron was most regular within the ring current region for select 150 151 dates from 2001-2004, which is critical for accurate curlometer output. Using prior studies for guidance in event selection (Vallat et al., 2005; Zhang et al., 2011; Shen et al., 2014), six hours of 152 data surrounding every perigee pass individually underwent visual inspection. Each pass was run 153 through the curlometer script and examined. Perigee passes were removed if data was missing 154 from anywhere in the expected ring current region, or if there were obvious errors in the data (e.g., 155 discontinuities, asymptotes). This filtering of missing data therefore removed passes that 156 experienced eclipses. 157

Figure 1 presents an example of the Cluster data examined for this study, specifically from 158 the 9 May 2002 perigee pass, when the tetrahedron shape was within the nominal guidelines for 159 yielding reasonable current densities from the curlometer technique. The tetrahedron characteristic 160 size was just under 200 kilometers, with the largest separation at about 250 kilometers. This pass 161 has a minimum radial distance of 4.3 R<sub>E</sub> and a magnetic local time near 20:00. The timeseries 162 shown is from south to north, maintaining the magnetic local time near 20:00 with only small 163 deviations from meridional motion. The magnetic equator is at approximately 18:20 UTC, with 164 Southern Hemisphere observations prior to that time and Northern Hemisphere observations 165 subsequently. The top two panels show the current density computed from the curlomter Python 166 167 script, first in SM coordinates (Figure 1a) and then just the azimuthal component (Figure 1b). In Figure 1a, the total current magnitude is plotted in the black line, while colored lines are Cartesian 168 169 current components. Figure 1c shows the interspacecraft differences in magnetic field magnitude used in the curlometer calculation. Spacecraft geometric parameters, tetrathedron quality, and 170

radial current are also plotted for reference in the supplemental material. The vertical red dashedlines roughly outline the region of interest, as marked by smooth curlometer output.

Several features are readily seen in Figure 1. First, there is a smooth section of current 173 around the magnetic equator, surrounded by intervals with highly fluctuating currents. Second, the 174 magnetic field differences do not show a drastic change at the variable/smooth boundary (the red-175 dashed vertical lines). That is, the current densities come from rather small deviations in magnetic 176 field differences, sometimes hardly noticeable on the plotted scales. The perturbations in the 177 178 magnetic field that cause large oscillations in the current are very small, less than 0.1 nT, and the 179 differences across the tetrahedron (Figure 1c) that are used for computing current are far smaller than the typical field strength at this distance of about 400 nT. Also note an apparent symmetry in 180 181 the azimuthal current along the trajectory, as expected.



flying south to north at approximately 20:00 magnetic local time. **a**) Curlometer current components in SM coordinates. The black line is the total current magnitude. **b**) Azimuthal current component  $(J_{\phi})$  in SM coordinates. **c**) Inter-spacecraft magnetic field differences. Dashed vertical lines delineate the region of interest. L-shell values are provided every 30 minutes.

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Similarly to the example in Figure 1, all perigee passes were examined to capture and identify the region of interest as the area with smooth current profile. Outside the region of smooth current signatures (outside red dashed lines), curlometry yields wildly oscillating current components as a result of the complex structure in these regions below the tetrahedron scale size (Vallat et al., 2005). Conversely, the presumed ring current section of the orbit is marked by a 188 smooth timeseries indicating a homogeneous structure. After identification, the set of all cases was189 then used to demonstrate the uncertainties detailed herein.

#### 190 2.3. Nonlinear Current Simulations and Tetrahedron Shape

191 To investigate the performance of the curlometer in different environments, simplified simulations were used to probe for possible error. When testing curlometry through simulated 192 193 magnetic fields, deviation from linear magnetic gradient causes errors in the output current. Robert et al. (1998) suggests that this effect is <10% for a regular tetrahedron with elongation and 194 195 planarity both below 0.8 (see Paschmann & Daly (1998), Chapter 13, for detailed treatment of tetrahedron geometric parameters). Defined as the ratio between magnetic divergence and 196 197 magnetic curl, the quality parameter Q = |div(B)|/|curl(B)| has also been used to determine regions where the curlometer has better performance, with Q = 0.5 as the standard threshold (e.g., Dunlop 198 et al., 1988, Vallat et al., 2005, Zhang et al., 2011). However, a comparison between idealized 199 current as input and curlometer output has not been conducted using actual spacecraft position 200 data. This process is distantly similar to that of Dunlop et. al (2002), but uses a more idealized 201 situation and a more focused region of interest. 202

To represent the simulated currents, an infinite planar current sheet was constructed with 203 thickness scaled to the selected perigee pass and quadratic variation from 0 nA/m<sup>2</sup> at the 204 boundaries to 10 nA/m<sup>2</sup> at the center; the simulated current sheet was then offset to the center of 205 the selected spacecraft flight track. Although greatly oversimplified, this model creates gradients 206 consistent with structural understanding of the magnetosphere (eg. Le et al., 2004) and of similar 207 magnitude as proposed by Vallat et al. (2005). Using the Biot-Savart law, the magnetic field 208 vectors were computed at each spacecraft location, then passed through the curlometer script. A 209 simplified visualization is provided in Figure 2, depicting the spacecraft trajectory, variation of 210 current, and magnetic field for a single case. Figure 2a shows the current density constructed for 211 the simulated system in blue; the corresponding magnetic field generated by this current is in 212 213 Figure 2b in black. Finally, Figure 2c shows another representation of the current density as blue shading, magnetic field as black arrows, and a cartoon Cluster constellation flying through the 214 215 simulated environment from bottom to top in red.



**Figure 2.** Simulated Current System. **a)** Slice through simulated current sheet, in  $nA/m^2$ , of the J<sub>y</sub> current component (blue). This plot is the current experienced by a Cluster spacecraft as it flies through the simulated environment. **b)** Magnetic field corresponding to the simulated current (black). These values replace the measured magnetic field from the Cluster spacecraft for simulation purposes. **c)** Alternate visualization of simulated environment. Black arrows denote magnetic field from (b). Blue shading indicates quadratically-varying current density from (a). Simplified spacecraft flight track and tetrahedron evolution are shown in umber.

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Because the Cluster tetrahedron had some degree of irregularity in shape near perigee on all orbits, the simulated current sheets were iterated through all orientations and current directions. In other words, the simulated current sheets were constructed in each component direction,  $J_x$ ,  $J_y$ , and  $J_z$ , and also with the current sheet normal direction varied for each case. Additionally, for each current sheet orientation, the tetrahedron was rotated about the barycenter by way of individual spacecraft positions and reanalyzed to deduce the combined effect of current direction and tetrahedron aspect. Thus, the possible geometric combinations of the simulated current andtetrahedron were completely explored for the ring current environment as observed by Cluster.

#### 226 **3. Results**

227 As discussed in the methodology section above, idealized current configurations were used to replace the observed Cluster magnetometer data with simulated data, for which the 228 229 corresponding current density is known. Figure 3 shows the results of an idealized current sheet in the  $J_v$  direction centered in the z = 0 plane. The current density varies quadratically from J = 0 at 230 the boundaries to a maximum in the center of 10 nA/m<sup>2</sup> with a total thickness of 5 R<sub>E</sub>. This 231 thickness was based on the observed ring current thickness of the selected event. The tetrahedron 232 233 was 'flown through' the simulation using actual position data from 10 October 2003 (other events provided in supplemental materials) and replacing the magnetic field with values computed from 234 235 the simulated current sheet using the Biot-Savart law (Figures 3a-3c). Figure 3a shows the 236 calculated currents from the original unrotated tetrahedron, with the imposed currents represented by dashed lines and the calculated currents represented by solid lines. The green lines are the  $J_{y}$ 237 component, which is the imposed component that is expected to be captured. The blue lines are 238 239 the  $J_z$  component, which is set to 0 in the imposed current and expected to be computed as 0 as 240 well. Figure 3b contains the corresponding quality parameter Q. The boundaries of the threshold Q = 0.5 are marked with vertical red dashed lines. In Figure 3c, the locations of all spacecraft 241 relative to the constellation barycenter are plotted as the vertices of the tetrahedron, to visualize 242 the constellation shape. Figure 3c also contains arrows to show relevant vector directions, 243 244 including current direction (black arrow), magnetic gradient (blue arrow), and longest planar axis (red arrow). Note that the planarity vector shows the direction of the largest semiaxis, instead of 245 the planarity normal direction. The tetrahedron was then rotated about its barycenter around the x-246 axis by 90° to produce Figures 3d-3f. Figure 3d shows the currents as in Figure 3a, Figure 3e shows 247 248 the Q calculation as in Figure 3b, and Figure 3f shows the constellation as in Figure 3c, following the 90° rotation. 249

It is immediately apparent that, despite the simulated current being purely in the  $J_y$ direction, a large false current is output by the curlometer technique because of the non-zero  $J_z$ current. In this instance, the false current grew as large as 5 nA/m<sup>2</sup> in the  $J_z$  direction (solid blue line, Figure 3a). The unaltered tetrahedron orientation produced nearly the greatest false currents

of any rotation direction. Furthermore, the large false currents occur well below the Q < 0.5 quality 254 255 standard in Figure 3b, with significant stature even at more stringent thresholds. In other words, 256 the false currents occur between the red dashed lines denoting acceptable curlometer quality. Rotating the tetrahedron about its barycenter, however, produced different results (Figure 3d). 257 Instead of a large false current, the same tetrahedron parameters with a quarter rotation captured 258 the currents remarkably well with little deviation between dashed and solid lines (Figure 3d), and 259 Q is much lower throughout the pass (Figure 3e). All else equal, rotation dramatically changed the 260 output by altering the planar direction with respect to the magnetic field gradient. When rotated 261 farther, currents appear again, related to the orientation of magnetic gradient with respect to planar 262 direction. An animation of rotating the tetrahedron with corresponding curlometer output is 263 provided in supplemental materials. Larger false currents are produced when the largest planar 264 semiaxis is more parallel to the magnetic gradient; this can be seen by comparing the red and blue 265 arrows in Figures 3c and 3f. During analysis, the tetrahedron was rotated independently and in 266 combination of all three axes, with only the extrema of false currents highlighted in the figure. 267 Rotation about other axes produced smaller false currents, so the uncertainty contributions from 268 269 those axes are smaller. False currents are directly related to the length of the tetrahedron semiaxis parallel to the magnetic gradient; therefore, some rotations will roughly preserve that quantity. 270



**Figure 3.** Quadratically varying infinite current sheet centered on the z=0 plane using the orbit spatial parameters from 10 October 2003. From unaltered flight track position data: **a**) Curlometer output (bold) and imposed simulation current (dashed). **b**) Quality parameter Q. **c**) Tetrahedron shape and environment vectors. For **d**) - **f**), the current sheet remained the same, but the tetrahedron was rotated about the barycenter by 90°, pivoting clockwise around the x-axis. Red dashed lines denote the region with Q < 0.5. In (c) and (f), the black arrow 'J' depicts the direction of simulated current. The blue arrow ' $\nabla$ B' shows the gradient of magnetic field. The red arrow 'P' shows the orientation of the longest tetrahedron semiaxis.

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Using simulated currents and tetrahedron rotations like those that produced Figure 3, the curlometer calculation yielded large non-physical currents in complementary components with different simulated current directions and tetrahedron rotation axes as well. Despite this example emphasizing the production of  $J_z$ , gradients in other directions do impact the azimuthal current in the same way. Thus, without a priori knowledge of the environment, false currents may appear in any direction depending on the gradients present. These false currents are in the direction of the magnetic gradient associated with the magnetospheric current sheet (like the spurious  $J_z$  in Figure 3a), but the magnitude was highly dependent on tetrahedron orientation and shape parameters. Even in regions where Q < 0.5 for the curlometer computation, false currents could be as much as 100% of the imposed current at a given time (5 nA/m<sup>2</sup> of  $J_z$  created by 5 nA/m<sup>2</sup> of  $J_y$ ), with significant implications that ring current curlometry can be highly inaccurate despite accepted data filtering methods.

Rotating the tetrahedron about the barycenter provided additional insight into the 284 relationship between shape, orientation, and false current. Elongation was consistently near 0.8 for 285 286 all passes and timestamps, but planarity evolved quickly from 0.4 to near unity in the few hours surrounding perigee. Thus, planarity was observed to have the larger effect of the shape parameters 287 288 in these flight tracks, and maximum false current was produced by the largest planarity semiaxis that was close to being parallel to  $\nabla B$  and therefore experiencing the largest linearization error. In 289 290 cases with multiple gradients, false and physical currents added linearly through the curlometer, requiring understanding of physical currents to deduce the false components. However, false 291 292 current was also strongly related to nonlinearity in magnetic gradient, so without quantitative 293 knowledge of this it is impossible to correct in-situ data for these differences. Numeric correction of curlometer output requires additional study and technique development and will be severely 294 295 limited by enhanced stationarity assumptions.

#### **4. Application to Cluster Data**

297 Considering the production of false currents in the simulated environment, this section 298 provides an initial view of the Cluster ring current observations in the context of increased 299 uncertainty. The application of the curlometer technique to this dataset produces current structures 300 that are consistent with nonlinearity current artifacts as seen in the simulated currents.

**301 4.1. Representative Curlometry** 

Previously in this study and others (e.g., Vallat et al., 2005, Zhang et al., 2011), the quantity Q = div(B)/curl(B) was used as a quality flag, with values less than 0.5 considered acceptable uncertainty. However, simulated currents can produce false results with the same order of magnitude as the imposed currents even when Q is below the 0.5 threshold (as seen in Figure 3), depending on tetrahedron orientation. Orbits in the Cluster dataset each only contain a small

duration under this threshold and are likely to contain larger magnitude errors outside this range. 307 While the latitudinal extent of the ring current can be established with reasonable confidence via 308 plasma data and regularity of current profile, information regarding the magnitude and direction 309 of the current may be grossly inaccurate. Examining tetrahedron orientation through simulated 310 current sheets produced new maximum uncertainty estimates that exceed the elongation-planarity 311 312 plots constructed by Robert et al. (1998), which suggest current magnitude errors rarely exceed a mere 10% when both elongation and planarity are below 0.8 (see, for example, Figures 16.7 and 313 314 16.8 in Robert et al., 1998).

New selection criteria for suitable ring current data must examine the curlometer output in 315 316 conjunction with the quality parameter Q and a priori knowledge of the magnetospheric environment to judge data validity. Only data with a maximum Q of 0.5 should be considered in 317 climatological studies of the ring current because instances above this threshold almost certainly 318 have large errors in both magnitude and direction. Lowering the threshold further increases 319 confidence, but at the expense of sample size. To effectively filter curlometry output, a more 320 complete analysis of tetrahedron and environmental parameters needs to be constructed, both 321 322 through higher-degree estimates of nonlinear magnetic field gradient and the orientation with 323 respect to the tetrahedron. Curlometer uncertainties stem not only from tetrahedron geometry, but from the geometry in combination with the magnetic environment. Then, additional quality 324 standards can be developed to increase trustworthiness of current calculation at the expense of the 325 resolution. 326

### 327 4.2. Single-Event Current Features

328 Figure 4 provides a representative sample of a single perigee pass, from 10 October 2003. 329 This is the actual curlometer calculation from magnetometer observations, in contrast to the simulation in Figure 3. While the observed current structures varied widely across the total dataset, 330 this event provides a typical case that shows common trends and avoids extremes. It is important 331 to note that the tetrahedron characteristic size varies greatly between data from 2001, 2002, and 332 2003-04 flight configurations. Thus, the example chosen in Figure 4 represents the most 333 observations but not necessarily the best. As in Figure 1, the spacecraft are traversing the inner 334 magnetosphere from the Southern Hemisphere to the Northern Hemisphere with only small 335 meridional deviation. The magnetic equator is crossed near 13:00 UTC, and this pass is located at 336

local time 09:40. Figure 4a shows curlometer output in SM Cartesian current components, with 337 the components clearly diverging later in the timeseries. The total current magnitude is plotted in 338 black. Figure 4b displays the azimuthal current component computed from the Cartesian 339 components. These current values are then compared to the quality parameter Q = div(B)/curl(B)340 in Figure 4c and the tetrahedron geometry in the form of elongation (red) and planarity (blue) in 341 Figure 4d. Note that the geometrics are plotted on separate y-axes in Figure 4d to capture detail. 342 The plot limit times have been chosen to fully enclose spacecraft perigee by several hours and with 343 the "smooth" sections used as a means for identifying the region of interest. 344

345 For this inner-magnetospheric traversal, a few key features immediately stand out. The Cartesian current components quickly diverge near the end of the selected data (Figure 4a), which 346 in turn causes the azimuthal current component to develop a strong negative trend (Figure 4b). 347 Figure 4b shows an environment where the azimuthal current is westward in the Southern 348 349 Hemisphere before decreasing to 0 just north of the magnetic equator and becoming eastward into the Northern Hemisphere. This is clearly nonphysical behavior, and cannot be trusted to accurately 350 351 represent the ring current. Given the south-to-north trajectory, the azimuthal current should vary with radial distance/L-shell, not with latitude as seen here. Instead, the curlometer-derived current 352 topology in Figure 4 looks very similar to the simulated environment in Figure 3, suggesting false 353 currents are to blame for the divergence. The large uncertainty in the data is corroborated by the 354 calculation of Q, which is only below the standard threshold of 0.5 in the beginning of the flight 355 track (Figure 4c). However, even during periods of relatively high confidence, the currents are still 356 diving towards negative values. A potential explanation for the diverging current lies in the 357 tetrahedron geometry (Figure 4d), which has a high elongation (E > 0.8) and an increasing 358 planarity (from P < 0.4 to P = 1) where low values are more regular and desirable. These extreme 359 geometric factors produce false currents as in Section 3. 360



**Figure 4.** 10 October 2003 Perigee Pass. The satellite constellation is flying south to north at approximately 09:40 magnetic local time. **a)** Cartesian current components in SM coordinates, showing clear divergence after perigee. Total current magnitude is shown in black. **b)** Azimuthal current density  $J_{\phi}$ . **c)** Quality parameter Q, satisfying the Q < 0.5 threshold until just after 13:00. This is not centered on the ring current region. **d)** Tetrahedron shape parameters elongation and planarity. Note that planarity is steadily increasing to unity throughout the ring current traversal, and that the axes scale differently.

In all orbits where Cluster achieved relatively regular tetrahedra when traversing the ring 362 current region, curlometry results varied greatly in structure and magnitude. Nearly all passes had 363 364 azimuthal current components in a westward direction with a magnitude peaking at or below 10 nA/m<sup>2</sup>. Many perigee passes also featured similar azimuthal current profiles to the one on 10 365 October 2003 in Figure 4. This is inconsistent with our understanding of physical current structures 366 (e.g., Ganushkina et al., 2018), and calls the curlometer result for these diverging regions into 367 serious doubt. Data from 2003-2004 is especially susceptible to the decrease and reversal of 368 azimuthal current in similar fashion to Figure 4b, especially in regions where the tetrahedron 369 evolves to higher planarity. While the azimuthal current never becomes negative while Q < 0.5, it 370 does begin to decrease towards zero within this criterion. 371

372 Curlometer output also showed large cross-hemisphere field-aligned currents (FACs), 373 around 10 nA/m<sup>2</sup> on 10 October 2003. This can be seen in Figure 4a using the  $J_z$  component as a 374 proxy for FACs due to magnetic field lines primarily along that axis. The persistence of FACs in 375 the curlometer output for all events will be discussed in more detail in the next section.

#### **376 4.2. Cluster-Derived Ring Current Environment**

377 If the simulated current environments and the false currents produced in those simulations have any bearing on the actual ring current system, similar effects should appear in curlometer 378 379 output using the original spacecraft magnetic field observations. To establish a view of the ring current at all local times, it is important to have a large and representative sample of measurements 380 381 throughout the whole precession of the Cluster orbit. This aim would be hindered by the extensive data filtering required to produce meaningful curlometer results, removing any non-physical 382 383 results or using an advanced algorithm to determine magnetic gradient nonlinearity. Thus, the holistic plots here are subject only to the restriction on quality parameter Q, and investigated for 384 signs of false current presence despite an accepted quality. 385

All perigee passes are considered in Figure 5. Taking data with Q < 0.5 and with L < 7 R<sub>E</sub>, events were sorted into "disturbed-time" and "quiet-time" categories delineated by Dst = -25 nT. This threshold was set to have a very clear set of "quiet-time" observations, despite placing some nearly-quiet observations into the "disturbed" category. Figure 5a provides curlometer output for disturbed magnetospheric conditions (red data) and for quiet conditions (black data) as a function of L-shell. While disturbed events can show wild variation and sparse data coverage, quiet events provide insight into the radial ring current distribution. The spread of quiet data only is visualized via boxplot in Figure 5b, binned by L-shell in 0.2 R<sub>E</sub> increments. Figure 5c shows the number of observed current values in each bin. The majority of ring current observations with Q < 0.5 lie Earthward of L = 5, so conclusions in this region are more robust.



**Figure 5.** Ring Current Statistical Radial Profile. **a)** Azimuthal current as a function of L-shell. Quiet magnetosphere conditions are shown in black, and disturbed conditions are shown in red. **b)** Quiet-time current distributions. Box heights show the median (blue) and interquartile range (IQR), box widths represent current binning with respect to L value at 0.2 L per bin, and vertical bars indicate values up to 1.5\*IQR. **c)** Observation count per data bin.

The ring current as measured by Cluster between 2001 and 2004 was largely below 10 397  $nA/m^2$  in magnitude, with a marked decrease towards  $L = 4 R_E$ . This suggests that the eastward 398 399 reversal of the ring current usually lies inside the Cluster orbit and is not observed, although there is less data in this region, and agrees with prior studies (Shen et al., 2014; Vallat et al., 2005). 400 However, at all values of L, the median of ring current magnitudes was below 7.5  $nA/m^2$ , and the 401 third quartile was below 10  $nA/m^2$ , as seen in Figure 5b. This better agrees with plasma pressure 402 calculations but contradicts earlier curlometry, such as the results of Vallat et al. (2005). 403 Additionally, this could still be impacted by up to  $2 \text{ nA/m}^2$  of uncertainty from magnetometer 404 resolution, and by the production of false currents. For disturbed magnetospheric conditions, less 405 data were available and the scatterplot reveals higher spread (red data, Figure 5a). The linearization 406 assumption becomes even less robust during geomagnetic activity, because the extent of via viable 407 408 disturbed-time data is much smaller than quiet-time data (the limiting factor is usually that Q < 0.5for a shorter duration). Nevertheless, the radial profiles of azimuthal current shown in Figure 5 do 409 410 not show any characteristic signs of contamination by false currents.

411 Figure 6 visualizes a subset of the full current vectors in 3D space. The data were subjected to the same quality constraint Q < 0.5 as before and were then binned and averaged in 15-minute 412 intervals for plotting clarity. The plot contains both disturbed and quiet-time data, and all local 413 414 times. In this figure the  $J_z$  component is strong in most cases, and almost always negative. In fact, 415 the dominant current at most latitudes is southward field-aligned current as indicated by the arrows generally pointing southward. The largest FACs are in excess of 20 nA/m<sup>2</sup> in perigee passes from 416 2002, and FACs in most other orbits were southward at  $\sim 5 \text{ nA/m}^2$  and did not vary significantly 417 between hemispheres. Prior curlometer analyses do not focus on FACs because they are 418 understood to be poorly represented. However, the topological identification of FACs has 419 precedent (e.g., Vallat et al., 2005, Zhang et al., 2011). This is the first report of such large currents 420 in such consistent direction, and should be viewed with extreme caution. The FACs in the inner 421 magnetosphere are not expected to be so persistently in the same cross-equatorial direction at all 422 latitudes and local times. Thus, the large field-aligned component provides evidence that the 423 curlometer output may not be a good representation of the actual currents. A consistent current of 424 425 this magnitude and direction is characteristic of false currents produced by the curlometer technique under unfavorable magnetic environments. Because of the interconnected nature of 426 curlometer output components, these FACs also call the azimuthal currents computed by 427

428 curlometry into question. This ubiquity and consistency of a southward FAC, regardless of 429 geomagnetic activity, latitude, or local time, is characteristic of false currents.



430

For further visualization of these non-physical curlometer outputs, currents with Q < 0.5431 inside L = 7 (same criteria as Figure 5) were averaged for each perigee pass; the results are 432 provided in Figure 7. Figure 7a shows the full horizontal current vectors, and Figure 7b shows just 433 434 the azimuthal component. Note the large radial components of current at all local times for nearly all events in Figure 7a; this is not expected quiescent ring current structure (e.g., see review and 435 references by Ganushkina et al., 2018), nor has it been reported in any previous study. 436 Consideration of just the pure azimuthal component (Figure 7b) provides a misleading picture 437 438 because it omits the strong Earthward components, as well as the z-component as seen in Figure 6. Providing the full current vector or combining azimuthal and radial components adds additional 439

- 440 evidence of the presence of false currents by allowing the full complexity of the output currents
- that lack a plausible physical explanation.



**Figure 7.** Current vectors averaged by orbit over Q < 0.5 for L < 7. **a)** Full current vectors in the x-y plane. Each vector is averaged about the magnetic equator for clarity. Note that this is a top-down view of Figure 6. **b)** Azimuthal current components of the vectors in (a).

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While the above estimates provide methodically-consistent ring current computation, each 443 perigee pass can also be assessed individually to account for signs of nonphysical output in the 444 computed currents. These include the divergent current components as displayed in Figure 4a. 445 Data from 2001 should be discarded because it does not usually show a clear transition to the ring 446 current region, and because the tetrahedron characteristic size is larger than 1000 km. Additionally, 447 448 data from the 2003-2004 orbits can sometimes be noisy or have large cross-equatorial trends that should be eliminated. Finally, the case from 18 March 2002 should be excluded from holistic 449 analysis because of the uniqueness of the observed current structure. Despite the focus on this 450 event in Vallat et al. (2005), the physical implications for plasma populations and magnetic 451 topology necessitated by a constant and strong ring current through a large swath of the orbit 452 require special case study. However, aggressive removal of non-ideal perigee passes does not alter 453 454 the persistent large FACs show in Figure 6 and Figure 7, nor the radial profile in Figure 5. This suggests a systematic presence of curlometer errors that are appearing in all cases, regardless of 455 456 the available filtering criteria.

Considering these structures and the likely reasons for the non-physical currents, Cluster 457 curlometer output is limited for assessing the inner magnetospheric ring current density. There is 458 459 more information contained in analysis of individual passes than in views of the system as a whole, and qualitative trends provide the most confidence in these cases. It is sufficient to note that the 460 ring current is a highly-variable structure, in both magnitude and extent, and therefore small 461 variations in the environment can lead to drastic changes. Without a definitive way to remove the 462 463 full 3D effects of false current generation, however, quantitative analysis of ring current Cluster data has significant uncertainty that precludes definitive determination of current densities and 464 orientations. 465

#### 466 **5. Discussion**

The curlometer technique has produced unexpected results in the discussed applications: simulated current environments, single perigee passes, and holistic studies of all available events. In the simulated current environments, the curlometer calculation reproduced the imposed current with high fidelity in regular tetrahedra. The best results occur where nonlinearities were constrained or eliminated (Dunlop et al., 2020). However, using actual spacecraft position data

and a nonlinear magnetic environment, false currents were detected by the curlometer that were 472 not imposed in the simulation. These false currents were produced while the quality parameter O 473 474 was within standard thresholds. Standing alone, this directly challenges the efficacy of Q but does not imply anything about the ring current itself. However, single perigee passes also displayed 475 unique current structures. The strong hemispheric asymmetry of the current throughout the perigee 476 passes, especially in the 2003 data where current components remained continuous without sudden 477 spikes but diverged later in the flight track, casts doubt on the validity of the observed structure. 478 These structures are contrary to plasma organization by magnetic field lines, which are oriented 479 cross-hemisphere, and disagree with a pressure peak at low latitude. Although much of the 480 divergent current region lies outside of Q < 0.5 and is therefore untrustworthy, the trends begin 481 well within the filtered data. The diverging currents increase with tetrahedron irregularity, 482 483 suggesting a causal relationship. In light of the simulated false currents, which also produce this divergence with tetrahedron irregularity, the single perigee passes all seem to contain some 484 485 combination of poor tetrahedron quality or strong magnetic gradient that produces poor current estimates. 486

Radial and 3D plots of full current vectors combined from all perigee passes analyzed by 487 488 curlometry are similarly unexpected. Currents are dominated by a large field-aligned component in both hemispheres at all local times that is southward and cross-equatorial. There is also a 489 significant radial component Earthward at all local times. The most valuable product is azimuthal 490 current as a function of L-shell, which produces a consistent value for quiet-time ring current below 491 492 10 nA/m<sup>2</sup> and hints at a current reversal to the eastward ring current within L = 4. Although less steep than anticipated, the curlometer does detect a radial current peak. Attempts to clarify and 493 sharpen these analyses by restricting data to stable, trustworthy, or expected structure for 494 individual perigee passes using established quality indices were unsuccessful in refining the 495 dataset. Therefore, the ring current using Cluster data must stand as presented here, including the 496 unusual features and limitations, and the confounding presence of false currents that greatly 497 498 increase uncertainty.

Revisitation of Cluster ring current computation has called into question the efficacy of the curlometer technique in this region due to increased uncertainty. Individual time series plots show clear errors in current magnitude and direction, especially when the tetrahedron is distorted. Furthermore, extensive simulation shows that the parameter Q does not always accurately represent the current estimate quality, and at the accepted threshold of Q < 0.5 can produce false currents as large as 100% of the actual current in complementary components. Lowering the threshold reduces the sample size to a highly-restricted domain that obscures sought trends. Previous current estimates of ~20 nA/m<sup>2</sup> are nevertheless not upheld by this study, with quiet cases overwhelmingly below 10 nA/m<sup>2</sup>.

The discrepancy between legacy plasma moment estimates of ring current strength and distribution and Cluster curlometer values has narrowed significantly, with a newfound magnitude median peaking at just 7 nA/m<sup>2</sup>. The new analysis presented above finds that the ring current is weaker during the 2002-2004 period than some previous curlometer estimates. Climatologies using these data should nevertheless be extremely cautious of the quality of current estimates, and effectively removing unsatisfactory cases reduces the sample size to preclude judgment on largescale trends without removing the effects of false currents.

#### 515 **6.** Conclusions

This study conducted a systematic assessment of current calculations using the curlometry 516 technique in the inner magnetosphere, using the specific satellite alignments of the Cluster mission 517 when it had a tetrahedral configuration at perigee. It was found that curlometry sometimes yields 518 excellent reproductions of imposed currents, but other times produces large false currents due to 519 520 the linearization within the calculation. The appearance of false currents is directly related to the orientation of the tetrahedron relative to the imposed current. Specifically, it was determined that 521 522 larger false currents are produced when the largest planar semiaxis of the tetrahedron is more 523 parallel to the local gradient of the magnetic field. These false currents can appear even when the 524 elongation and planarity parameters signify only limited tetrahedron irregularity, and can be quite large even when the Q factor is below the nominally acceptable level of 0.5. 525

Keeping these limitations in mind, a statistical compilation of inner magnetospheric currents was then calculated. Although the new median ring current density of 7 nA/m<sup>2</sup> is in better agreement with other computation methods, there are still uncertainties associated with this measurement. The instrument magnetic field resolution alone introduces +/- 1.5 nA/m<sup>2</sup> of uncertainty (Vallat et al., 2005), and one must also consider the presence of false currents as an artifact of the nonlinearity. Even at Q < 0.5, these uncertainties can be quite large or even dominate the observed current structure, and are difficult to remove because of the complex dependence on tetrahedron orientation relative to the unknown currents. Furthermore, the interdependent nature of the curlometer current components necessitates the consideration of the large radial and fieldaligned currents, which raise additional concerns due to unexpected structure.

Thus, any inner magnetospheric curlometry analysis that relies on Cluster ring current data, including magnitudes and climatologies, should be viewed with full knowledge of the caveats and limitations described herein. Although shortcomings in curlometry are discussed in detail here, the extent of this analysis is restricted to Cluster spacecraft within the specified orbits and around perigee. The use of curlometry elsewhere should be examined for similar effects, but the uncertainties in the ring current region cannot be directly applied to other regions because of the strong dependence on magnetic field nonlinearity and tetrahedron shape/orientation.

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Supporting Information for

## **Cluster Curlometry Limitations in the Ring Current Region**

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Figures S1 to S20

## Additional Supporting Information (Files uploaded separately)

Captions for Movies S1 to S2

### Introduction

The plots and animations herein supplement the material in the article by providing enhanced visualization and additional method verification. The first set of figures verify the correct operation of the curlometer script used for this study. Be careful to note that both vertical and horizontal axes are scaled to show detail and are not necessarily the same for comparable plots. These are followed by plots to be compared directly with selected figures from previous studies, with detailed descriptions of the differences in computation and generation. Finally, two animations are provided for cases involving the tetrahedron rotation and the generation of false currents. These animations add clarity to a complex process that is difficult to represent in stationary plots. The animation files are in the GIF format.



**Figure S1.** 9 May 2002 perigee pass with magnetometer data replaced by a simple approximation of Earth's dipole field.



**Figure S2.** Curlometer currents output from the 9 May 2002 perigee pass with magnetometer data replaced by a simple approximation of Earth's dipole field.



**Figure S3**. 10 October 2003 perigee pass with magnetometer data replaced by a simple approximation of Earth's dipole field.



**Figure S4.** Curlometer currents output from the 10 October 2003 perigee pass with magnetometer data replaced by a simple approximation of Earth's dipole field.



**Figure S5.** 9 May 2002 perigee pass with magnetometer data replaced by the International Geomagnetic Reference Field (IGRF) magnetic environment.



**Figure S6.** Curlometer currents output from the 9 May 2002 perigee pass with magnetometer data replaced by the IGRF magnetic environment.



**Figure S7.** 10 October 2003 perigee pass with magnetometer data replaced by the IGRF magnetic environment.



**Figure S8.** Curlometer currents output from the 10 October 2003 perigee pass with magnetometer data replaced by the IGRF magnetic environment.



**Figure S9.** 9 May 2002 perigee pass with magnetometer data replaced by a simulated infinite current sheet (slab) in the z = 0 plane. The simulated slab current is constant over a finite thickness larger than the tetrahedron size, and tapers quadratically to either side.



**Figure S10.** Curlometer currents output from the 9 May 2002 perigee pass with magnetometer data replaced by a simulated infinite current sheet in the z = 0 plane. The simulated slab current is constant over a finite thickness larger than the tetrahedron size, and tapers quadratically to either side. Note that the black current captures the imposed simulation current almost perfectly at 10 nA/m<sup>2</sup>.



**Figure S11.** 10 October 2003 perigee pass with magnetometer data replaced by a simulated infinite current sheet (slab) in the z = 0 plane. The simulated slab current is constant over a finite thickness larger than the tetrahedron size, and tapers quadratically to either side.



**Figure S12.** Curlometer currents output from the 10 October 2003 perigee pass with magnetometer data replaced by a simulated infinite current sheet in the z = 0 plane. The simulated slab current is constant over a finite thickness larger than the tetrahedron size, and tapers quadratically to either side. Note that the black current captures the imposed simulation current almost perfectly at 10 nA/m<sup>2</sup>. Also note that, while in the sharp gradients to either side, the curlometer technique becomes inaccurate in all components.



Figure S13. Investigation of instrument sensitivity for Cluster tetrahedron. Using the 9 May 2002 perigee pass, the magnetometer resolution of 0.125 nT was subtracted uniformly in time from the Cluster 1 spacecraft, chosen as the most irregular in the direction of the gradient for largest effect. The curlometer output of the modified magnetic data showed a decrease of about 2 nT in the azimuthal current. Compare Figure S13 to Figure S14, Figure S15, and Figure 1 in the paper.



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**Figure S14.** Investigation of instrument sensitivity for Cluster tetrahedron. Using the 9 May 2002 perigee pass, the magnetometer resolution of 0.125 nT was added uniformly in time from the Cluster 1 spacecraft, chosen as the most irregular in the direction of the gradient for largest effect. The curlometer output of the modified magnetic data showed an increase of about 1 nT in the azimuthal current. Compare Figure S14 to Figure S13, Figure S15, and Figure 1 in the paper.





**Figure S15.** Full parameter consideration for the 9 May 2002 perigee pass. The panels from top to bottom depict:

- Cartesian and total currents in SM coordinates
- Azimuthal and radial current components in SM/local cylindrical coordinates
- Tetrahedron quality parameter Q

- Tetrahedron geometry factors elongation (E) and planarity (P)
  Magnetic field differences between spacecraft pairs



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**Figure S16.** Curlometer currents output from the 18 March 2002 perigee pass. This figure is constructed in the same manner as Figure 9 in Vallat et al. (2005). Note that this figure has the signs of some components reversed to match the convention used in that paper.



**Figure S17.** Curlometer currents output from the 13 April 2002 perigee pass. This figure is constructed in the same manner as Figure 2 in Vallat et al. (2005). Note that this figure has the signs of some components reversed to match the convention used in that paper.





**Figure S18.** Curlometer currents output from the 20 April 2002 perigee pass. This figure is constructed in the same manner as Figure 14 in Vallat et al. (2005). Note that this figure has the signs of some components reversed to match the convention used in that paper.



**Figure S19.** Curlometer currents output from the 6 February 2004 perigee pass. This figure is constructed in the same manner as Figure 2 in Zhang et al. (2011).



**Figure S20.** Curlometer currents output from the 19 March 2001 perigee pass. This figure is constructed in the same manner as Figure 9 in Shen et al. (2014).

**Movie S1.** Tetrahedron rotation for the 10 October 2003 perigee pass. This animation shows rotation of all spacecraft together around the tetrahedron barycenter, and the effect on curlometer output of a simulated current environment. Note that, depending on orientation with respect to a stationary current structure and magnetic topology, the exact same tetrahedron shape and size can produce very different current outputs. Two still frames from this animation are provided in the paper as Figure 4, to highlight the extrema of the animation.

**Movie S2.** Tetrahedron rotation for the 18 March 2002 perigee pass. This animation shows rotation of all spacecraft together around the tetrahedron barycenter, and the effect on curlometer output of a simulated current environment. Contrary to Movie S1, this event is from 2002, which has a much smaller tetrahedron size. Thus, in the exact same simulated current sheet, the smaller tetrahedron does not produce a large false current at any rotation. However, events from this timeframe make up a small subset of the existing curlometer analysis in the literature.