Diagnosing cross-scale kinetic energy exchanges from two submesoscale permitting ocean models.

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Abstract

The upper-ocean at fine-scales (<<100,km) contributes significantly to energy exchanges and dissipation. However, our knowledge of fine-scale motions (in terms of kinetic energy density and transfer) in the real ocean is limited due to lack of sufficient observational datasets at these scales. Kilometric resolving ocean models have been developed in anticipation for the Surface Water and Ocean Topography (SWOT) satellite mission. This mission would provide unprecedented global coverage of energetic oceanic flows down to 10,km. In this study, we investigate the distribution and exchange of energy across different scales down to 10,km based on two state-of-the-art realistic North Atlantic basin simulations with horizontal resolution of $s \ sufficient$ scales undergo both regional and seasonal variability. In particular, this seasonality which is more pronounced at fine-scale is characterized by a significant amount of direct KE cascade in wintertime. In general, we found the forward cascade to be confined to the mixed layer depth while the inverse cascade extends down to about 700m in the interior.

Diagnosing cross-scale kinetic energy exchanges from 1 two submesoscale permitting ocean models. 2

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Key Points:

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• We used two submesoscale permitting ocean models of the North Atlantic Ocean to investigate kinetic energy exchanges at fine-scales. 9

- KE fluxes at fine-scales are strongly impacted by submesoscale turbulence with 10 a stronger forward cascade in winter within the mixed-layer. 11
- Not accounting for ageostrophic motions yields a significant under-estimation of 12 the forward cascade. 13

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14 Abstract

Fine-scale motions (<100 km) contribute significantly to the exchanges and dissipation of 15 kinetic energy in the upper ocean. However, knowledge of ocean kinetic energy at fine-16 scales (in terms of density and transfers) is currently limited due to the lack of sufficient 17 observational datasets at these scales. The sea-surface height measurements of the upcoming 18 SWOT altimeter mission should provide information on kinetic energy exchanges in the 19 upper ocean down to 10-15 km. Numerical ocean models, able to describe ocean dynamics 20 down to ~ 10 km, have been developed in anticipation of the SWOT mission. In this study, 21 we use two state-of-the-art, realistic, North Atlantic simulations, with horizontal resolutions 22 ~ 1.5 km, to investigate the distribution and exchanges of kinetic energy at fine-scales in 23 the open ocean. Our results show that the distribution of kinetic energy at fine-scales 24 approximately follows the predictions of quasi-geostrophic dynamics in summertime but 25 is somewhat consistent with submesoscale fronts-dominated regimes in wintertime. The 26 kinetic energy spectral fluxes are found to exhibit both inverse and forward cascade over 27 the top 1000 m, with a maximum inverse cascade close to the average energy-containing 28 scale. The forward cascade is confined to the ocean surface and shows a strong seasonality, 29 both in magnitude and range of scales affected. Our analysis further indicates that high-30 frequency motions (<1day) play a key role in the forward cascade and that the estimates of 31 the spectral fluxes based on geostrophic velocities fail to capture some quantitative aspects 32 of kinetic energy exchanges across scales. 33

³⁴ Plain Language Summary

The dynamics of oceanic motions with scales <100 km (fine-scales) are currently not well 35 known. This is due to the lack of sufficient observational datasets at these scales in the ocean. 36 There are suggestions from recent studies that this class of motions impacts the distribution 37 and exchanges of kinetic energy in the ocean. To better understand fine-scale motions, 38 the Surface Water and Ocean Topography (SWOT) satellite has been assembled. SWOT 39 is expected for lunch in 2022 and will provide an unprecedented view of the ocean down 40 to a wavelength of 10-15 km. In anticipation for SWOT mission, numerical ocean models 41 capable of resolving fine-scales oceanic motions have been designed and implemented. In 42 this study, we use two of these simulations to investigate how kinetic energy is exchanged 43 between oceanic motions at fine-scales. Our results show that submesoscale turbulence 44 (a class of oceanic turbulence at fine-scale) and high-frequency motions affects the kinetic 45 energy exchanges by providing a route to kinetic energy towards dissipation. Also, we found 46 that kinetic energy exchanges based on the future SWOT dataset might underestimate the 47 true magnitude of the transfer of kinetic energy towards finer scales. 48

49 1 Introduction

The ocean is a turbulent fluid with a broad range of energetic scales, ranging from large \sim 50 O(1000 km) to centimeter scales. The ocean kinetic energy is mostly concentrated in the 51 quasi-geostrophic mesoscale eddy field with scales $\sim O(100 \text{ km})$ (Stammer & Böning, 1992). 52 Due to non-linear interactions among different length scales, energy can be transferred 53 both from large to small (forward, or direct cascade) and from small to large scale (inverse 54 cascade). Understanding the distribution of kinetic energy (KE) and variance across scales 55 in oceanic flows is, therefore, key to our knowledge of ocean circulation (Ferrari & Wunsch, 56 2009).57

To estimate the variance and energy associated with eddy motions at different scales, 58 velocity wavenumber power spectral density has proven to be very efficient (Le Traon et 59 al., 1990, 2008; Fu et al., 2010; Dufau et al., 2016; Uchida et al., 2017). However, spectral 60 density does not indicate the direction of kinetic energy exchanges between the different 61 scales. A better knowledge of cross-scale energy exchanges is gained by looking at the KE 62 cascade due to nonlinearity. This important feature in turbulence study dates back to the 63 work of Charney (1971) and Salmon (1980) on geostrophic turbulence. For stratified rotating 64 quasi-two-dimensional fluid motion, classical geostrophic turbulence theory predicts a direct 65 cascade of energy if the flow is depth-dependent (baroclinic) but an inverse cascade of energy 66 if the flow is depth-independent (barotropic). In particular, for the ocean with a surface 67 intensified stratification, energy from higher baroclinic modes concentrates in the first mode 68 and then converges toward the scale of the Rossby radius of deformation (R_d) (Smith & 69 Vallis, 2002). At R_d , baroclinic energy is converted to barotropic mode via barotropization. 70 At this point, most of the energy near the deformation scale cascade towards larger scales 71 while a small fraction undergoes direct cascades to dissipation (see Figure 1). 72

This prediction of geostrophic turbulence theory has been observed both in numerical 73 simulations and the real ocean but with a little discrepancy. Based on altimeter data, Scott 74 and Wang (2005) showed that an inverse cascade of energy dominates the (Pacific) ocean 75 at scales larger than R_d . So, if one agrees that the altimeter data is reflecting the first 76 baroclinic mode (Smith & Vallis, 2002), then this is in contrast with geostrophic turbulence 77 theory which predicts a forward cascade for a baroclinic flow. Scott and Wang (2005) 78 argued that there might be an inverse cascade associated with the first baroclinic mode and 79 that this would only partially reduce the forward flux of total baroclinic energy. This total 80 energy forward flux is the source of the kinetic energy that arrives near the deformation 81 scale from the large-scale mean flow via baroclinic instability. From this discrepancy, two 82 questions arise. (i) is the inverse cascade seen at the surface due to the barotropic mode? 83 or (ii) is it possible that the baroclinic modes experience an inverse cascade? Scott and 84

Arbic (2007) using a 2-layer model simulation showed that the inverse cascade at the ocean surface is mostly baroclinic with a small contribution from the barotropic mode. The results from Scott and Arbic (2007) are consistent with the proposed modification to geostrophic turbulence by Scott and Wang (2005). More recent literature (Schlösser & Eden, 2007; Sasaki et al., 2017; Tulloch et al., 2011; Aluie et al., 2017; Brüggemann & Eden, 2015; Kjellsson & Zanna, 2017; Khatri et al., 2018) have also shown that an inverse cascade of energy mostly dominates the surface ocean at scales larger than R_d .

In contrast, little is known regarding energy cascade at scales $\langle R_d$, where oceanic 92 motion is dominated by submesoscale motions (< 20 - 50 km). Results from numerical sim-93 ulation and observation have shown an injection of energy in winter at submesoscale (Sasaki 94 et al., 2017). This energy injection is partly responsible for both meso and submesoscale 95 seasonality (Uchida et al., 2017; Capet, Campos, & Paiva, 2008; Sasaki et al., 2014) and has 96 been argued to be associated with mixed layer instability (Callies, Ferrari, et al., 2015; Qiu 97 et al., 2014; Sasaki et al., 2014; Brannigan et al., 2015; Rocha et al., 2016). This seasonality is responsible for the shallowing of KE spectral slope from -3 in summer to -2 in winter and 99 is usually interpreted as a shift from turbulence dominated by interior gradients (Philips 100 regime) to a regime dominated by surface driven turbulence (Charney regime) (Sasaki et 101 al., 2014). Apart from the work of Sasaki et al. (2014, 2017); Schubert et al. (2020), we are 102 unaware of any investigation on the implication of submesoscale seasonality on cross-scale 103 energy exchanges at the basin scale and one of the objective of this study is to investigate 104 how submesoscales modify cross-scale kinetic energy exchanges at fine-scales. 105

Submesoscale resolving ocean models have been developed in anticipation of the Surface 106 Water and Ocean Topography (SWOT) satellite mission (Fu et al., 2010). On a global scale, 107 satellite altimeters remain the primary source of information on the distribution of energy 108 across scales. However, at the moment, the resolution capability of our existing ocean-109 observing satellite altimeters stands at roughly 70 km (Dufau et al., 2016). This limitation 110 undermines our ability to investigate energy exchanges at scales < 100 km. To solve this 111 challenge, SWOT is implemented to provide ten times higher resolution than conventional 112 altimeters, and numerical ocean models have been designed to prepare for SWOT. These 113 state-of-the-art numerical experiments with high-resolution capability, thereby provide an 114 opportunity to study cross-scale energy exchanges down to kilometric scales. 115

In this study, we aim to investigate the distribution and transfer of energy across different scales by using outputs of two submesoscale permitting ocean models of the North Atlantic. In particular, we focus on the seasonality and depth penetration of cross-scale KE variance and transfer with an emphasis on scales < 100 km. This paper is organized as follows; section 2 presents a description of the two numerical simulations. In section 3, we examine the kinetic energy wavenumber spectral density and slope. The KE cascade, its seasonality, and the role of high frequency and ageostrophic motions on the cascade are discussed in section 4. Finally, in section 5, we summarize the findings and discuss the relevance of this work in anticipation for SWOT mission.

¹²⁵ 2 Numerical simulations of the North Atlantic Ocean

In this study, we use numerical outputs from two submesoscale eddy-permitting simulations of the North Atlantic: a NEMO-based simulation with a horizontal resolution of $1/60^{\circ}$ (NATL60) and an HYCOM-based (HYbrid Coordinate Ocean Model) simulation with a horizontal resolution of $1/50^{\circ}$ (HYCOM50).

The NEMO-based NATL60 has a horizontal grid spacing ranging from 1.6 km at 26°N 130 to 0.9 km at 65°N. The initial and open boundary conditions are based on the GLORYS2v3 131 ocean reanalysis with a relaxation zone at the northern boundary for sea-ice concentration 132 and thickness. The model has 300 vertical levels with a resolution of 1 m at the top-most 133 layers. The grid and bathymetry follow Ducousso et al. (2017), while the atmospheric forcing 134 is based on DFS5.2 (Dussin et al., 2018). DFS5.2 forcing is based on ERA-interim reanalysis. 135 The spatial resolution of the atmospheric fields is 0.75 degrees. All variables used to compute 136 turbulent fluxes (air temperature and humidity at 2m, wind velocity components at 10m) 137 are 3-hourly. In order to implicitly adapt lateral viscosity and diffusivity to flow properties, 138 a third-order upwind advection scheme is used for both momentum and tracers in the model 139 simulation. The model was spun-up for six months, and a one-year simulation output from 140 October 2012 to September 2013 is used in this study. The simulation output used in this 141 study is the same as the one used in Amores et al. (2018), Buckingham et al. (2019) and 142 Ajayi et al. (2020). An earlier version of this simulation set-up was used in Ducousso et al. 143 (2017) and Fresnay et al. (2018). 144

NATL60 ocean simulation has been evaluated, using in situ observations in terms of the 145 kinetic energy levels at different wavelengths (see Figure 1 of supporting information docu-146 ment (SI). In terms of the dynamics of the resolved fine-scales in the upper ocean, results 147 from Buckingham et al. (2019) show that the statistics of horizontal velocity tensor predicted 148 by NATL60 agree reasonably well with observation (OSMOSIS datasets). However, their 149 results also show that there is a likelihood of extreme divergent motions in OSMOSIS that 150 is not captured by NATL60. That NATL60 underestimates divergent motions compared to 151 observation isn't that surprising because NATL60 model simulation is without tidal forcing, 152 one of the major sources of wave energy. The model, however, reproduces fairly well other 153 forms of internal gravity waves (see Figure 2 of SI). 154

The HYCOM-based HYCOM50 extends from 28°S to 80°N and has a horizontal grid 155 spacing ranging from 2.25 km at the equator, ~ 1.5 km in the Gulf Stream region, and 1 156 km in the subpolar gyre. As for NATL60, the effective resolution is about 10-15 km. The 157 vertical coordinate is hybrid and consists of 32 layers. The simulation is initialized using 158 potential temperature and salinity from the GDEM climatology and spun up from rest for 159 20 years using climatological atmospheric forcing from ERA-40 (Uppala et al., 2005), with 160 3-hourly wind anomalies from the Fleet Numerical Meteorology and Oceanography Center 161 3-hourly Navy Operational Global Atmospheric Prediction System (NOGAPS) for the year 162 2003. The horizontal viscosity operator is a combination of Laplacian and Biharmonic. The 163 bathymetry is based on the Naval Research Laboratory (NRL) digital bathymetry database. 164 The model configuration and a detailed evaluation of the model results in the Gulf Stream 165 region with observations are documented in Chassignet and Xu (2017). 166

In this study, we use the output from October 2012 to September 2013 for NATL60 167 and year 20 (last year of the simulation) for HYCOM50. Winter and summer correspond 168 to January-February-March (JFM) and July-August-September (JAS), respectively. The 169 summer analysis presented for NATL60 corresponds to that of the year 2013. Since NATL60 170 covers a smaller domain than HYCOM50, we consider the HYCOM50 outputs for the same 171 region covered by NATL60 to have comparable results. To capture regional variability in 172 the distribution of energy across scales, we perform spectral analysis in sub-domains of 173 $14 \, 10^{\circ} \times 10^{\circ}$ boxes across the North Atlantic. We focus specifically on quantifying kinetic 174 energy wavenumber spectral density (Eq. 1) and flux (Eq. 2) using horizontal velocity 175 fields. In equation (1) and equation (2), refers to Fourier transform, * represents the 176 complex conjugate, Re refers to the real part of a complex number and $k = \sqrt{k_x^2 + k_y^2}$. 177 Before performing spectral analysis, the 2D velocity field from each subdomain (box) is 178 detrended in both directions, and a 50% cosine taper window (Tukey windowing) is applied 179 for tapering. An FFT is applied to the tapered data, and a 1D isotropic spectrum is obtained 180 by averaging in the azimuthal direction. Our spectral method is consistent with procedures 181 previously used in Stammer and Böning (1992); Sasaki and Klein (2012); and Chassignet 182 and Xu (2017). 183

$$E(k) = \int_{k}^{k+\delta k} \left[\widehat{\mathbf{u}}^* \cdot \widehat{\mathbf{u}} \right](k) dk \tag{1}$$

$$\Pi(k) = \int_{k}^{k_{s}} -Re\left[\widehat{\mathbf{u}}^{*} \cdot \left(\mathbf{u} \cdot \widehat{\nabla}_{\mathbf{H}} \mathbf{u}\right)\right](k) dk$$
(2)

Both NATL60 and HYCOM50 resolve the first Rossby radius of deformation everywhere
 within the model domain, and these simulations reproduce realistic eddy statistics with levels

of kinetic energy in the range of altimetric observations (Chassignet and Xu (2017), see also Figure 2). A summary of the model parameters is tabulated in Table 1. Both simulations are submesoscale permitting ocean models and a discussion on the ability of these models to resolve the dynamics of submesoscales in terms of eddy length scale, submesoscale energy and their associated seasonality can be found in Ajayi et al. (2020).

We present in Figure 3, the root-mean-square of sea surface height computed from 191 one year daily outputs of NATL60, HYCOM50, compared with AVISO. The AVISO SSH 192 field is derived from observations obtained by altimeter missions and then interpolated 193 onto a 0.25° mercator grid. In this comparison, we have used AVISO mean dynamical 194 topography dataset from October 2012 to September 2013. The SSH fields in Figure 3 195 from NATL60 and HYCOM50 have been resampled onto a 0.25° grid to have comparable 196 results. The SSH rms of the models and AVISO data appears to be consistent in terms 197 of the geographical pattern of energetic oceanic motions except for the differences along 198 the Gulf Stream. The Gulf Stream seems to be more energetic in HYCOM50 compared to 199 NATL60 and AVISO. HYCOM50 has a spin-up of 20 years while NATL60 has a spin-up 200 of 6 months. We hypothesize that HYCOM50 long spin-up allows for the full development 201 of the Gulf Stream energetics and this difference in terms of spin-up could be contributing 202 to differences in the overall energetics of the two simulations. The implication of the short 203 spin-up for NATL60 is more obvious in the time evolution of the domain-averaged kinetic 204 energy (see Figure 3 of SI). The kinetic energy is still increasing with time and is yet to 205 attain equilibrium. A similar curve for HYCOM50 is available in Figure 2 of Chassignet 206 and Xu (2017). The comparison of the KE spectral in summer of the year 2012 vs the year 207 2013¹ (see Figure 4 of SI) further highlights the increase in energy (for NATL60) that is 208 characterized by higher variance and larger eddies at low wavenumbers in year 2013. 209

The two simulations are similar but are not without differences. HYCOM50 appears 210 to be more energetics compared to NATL60. The disparity between the two models' energy 211 level is not the main focus of this paper, but we shall propose a few reasons why the two 212 models could differ in terms of energetic. Firstly, we hypothesize that the eddy structures 213 in NATL60 are not fully developed due to the short spin-up (6 months for NATL60 versus 214 20 years for HYCOM50). The first two years of the HYCOM50 simulation show an increase 215 of total kinetic energy level; see Figure 2 in Chassignet and Xu (2017). Furthermore, 216 the typical scale of eddies are smaller in NATL60 compared to HYCOM50 (Ajayi et al., 217 2020), and this could be a direct consequence of the shorter duration of the model spin-up. 218 Secondly, the question did arise as to whether the coarser vertical resolution in HYCOM50 219

¹ The summer analysis presented in the article corresponds to that of year 2013

(32 hybrid vertical layers versus 300 z-levels in NATL60) could lead to a stronger inverse 220 cascade and hence a higher energy level because of an under-resolved stratification and the 221 depth dependence of flows. A comparison of the vorticity spectral coherence as a function 222 of depth shows that the two simulations are essentially identical in terms of the depth 223 penetration of energetic eddy structures (Ajayi et al., 2020). Furthermore, in section 4, a 224 comparison of the KE spectral flux at depths for the two simulations will show that the 225 HYCOM50 upscale energy flux is not surface intensified and that having only 32 isopycnal 226 vertical levels is not detrimental to the representation of the dynamics in the ocean interior. 227 Thirdly, the choice of sub-grid parameterization is different between the two simulations 228 and could have a substantial effect on how energy is dissipated in each model. 229

Up until recently, most basin-scale numerical models like those used in this study usu-230 ally store simulation outputs in the form of daily averages due to limitations in storage 231 and computational resources. This limitation comes with a caveat. Daily averaging the 232 model outputs suppresses high-frequency motions (> f, where f is the Coriolis frequency). 233 These motions are mostly dominated by ageostrophic motions that include unbalanced sub-234 mesoscales and fast propagating internal gravity waves. There are new shreds of evidence, 235 based on idealized simulations that suggest that high-frequency motions (particularly wave 236 motions) can provide a route to kinetic energy dissipation (Barkan et al., 2017; Rocha et 237 al., 2018). It would be interesting to investigate these new results in realistic simulations 238 like NATL60 and HYCOM50. Unfortunately, NATL60 and HYCOM50 have most of their 239 outputs stored in daily averages except for the last month of simulation for HYCOM50 and 240 surface quantities for NATL60. To that end, the results presented in this study are based 241 on daily averages of velocity fields from the two simulations. We have only used analysis 242 based on hourly output where necessary to illustrate the impacts of high-frequency motions 243 on the kinetic energy distribution and exchanges. 244

245

3 Distribution of Kinetic Energy

In this section, we discuss the variance at different scales of motions by analyzing 246 the kinetic energy wavenumber power spectral density. In general, horizontal wavenumber 247 spectral density exhibits power-law behavior, where the exponent is interpreted in terms of 248 the dynamical processes governing the eddy energy transfer. Existing theoretical frameworks 249 (for horizontal velocity at scale > Rd) predict a spectral slope of -3 and -5/3 for QG and 250 SQG (surface quasi-geostrophic) turbulence respectively. A slope of -2 is also well known for 251 a front dominated flow (Shcherbina et al., 2013; Callies & Ferrari, 2013). Over the years, 252 many research works have tried to establish the accuracy of these predictions by using 253 outputs of realistic ocean models (Sasaki & Klein, 2012; Chassignet & Xu, 2017; Uchida et 254

- al., 2017) and also recently within the context of the real ocean by using altimeter dataset
 (Le Traon et al., 1990; Dufau et al., 2016). Their results have argued for the non-existence
 of a universal wavenumber spectrum (Le Traon et al., 2008) following observed regional
 variability. We shall discuss in section 3.1, the distribution of kinetic energy as predicted
 by NATL60 & HYCOM50 and in section 3.2 we discuss the variability associated with this
 distribution and their associated estimated slope.
- 261

3.1 Spectral Density

- In Figure 4, we present the kinetic energy spectral density as a function of depth for 262 the two simulations. For simplicity, we show this comparison only for Box 8, a box located 263 at the center of the North Atlantic basin. In most of the regions, the peak of the spectral 264 density is around the mesoscale motions (100–500 km). As expected, the energy associated 265 with large-scale motion is relatively higher than that of fine-scales. The peak of the spectral 266 density is preserved with depth, while the variance at all scales decreases with depth. The 267 comparison between the two simulations is illustrated better in Figure 5a where we present 268 the depth-averages of annual KE spectral density for the two simulations in the same region 269 (Box 8). The spectral densities from the two models agree well with an approximate slope of 270 -3, a value that is characteristic for Quasi-Geostrophic (QG) prediction. In this QG regime, 271 submesoscale structures are expected to be weakly energetic while energy is concentrated 272 at the mesoscales. The depth-averages of the winter and summer KE spectral density 273 are presented in Figure 5b. There is a seasonality in the spectral density that is mostly 274 associated with an increase in the variance at submesoscales in winter. The spectral shape 275 in both winter and summer are somewhat QG. We find this surprising given that previous 276 studies (Callies & Ferrari, 2013; Shcherbina et al., 2013; Sasaki et al., 2014) have shown 277 that in the presence of intense submesoscales in winter, KE spectral density is likely to have 278 k^{-2} or $k^{-5/3}$ spectral shape. 279
- As previously highlighted, daily averaging the velocity fields before spectral estima-280 tion could suppress the signature of high-frequency motions (unbalanced submesoscales, 281 ageostrophic wave motions) and this could affect the distribution of energy implied from 282 this sort of spectral analysis. To ascertain this, we compared NATL60 surface KE spectral 283 density for hourly averages versus daily averages in Figure 6a. This comparison is only pos-284 sible for NATL60 because HYCOM50 dataset is stored in daily averages except for the last 285 month of the simulation. In this comparison for NATL60 (Figure 6a), submesoscale motions 286 are more energetic than presented in the daily spectral density. There is a tendency for the 287 distribution (annual average) of energy at submesoscales in the hourly spectral to follow a 288 k^{-2} spectra shape, a characteristics of a regime associated with fronts. 289

The winter versus summer spectral density computed from NATL60 hourly averages 290 of the velocity field (Figure 6b), clearly show how high-frequency motions and energetic 291 submesoscales drive the seasonality of kinetic energy distribution at fine-scales. There is a 292 shift in the spectral slope from -3 in summer to -2 in winter. This shift can be interpreted 293 as a change in dynamics from the interior QG Philips-like regime to a surface intensified 294 Charney-like regime (Sasaki et al., 2014). In this Charney-like regime, submesoscales are 295 associated with large vertical velocity and, in turn, large submesoscale buoyancy fluxes that 296 are suggested to feed mesoscales through an inverse kinetic energy cascade. This result is 297 consistent with the findings of Sasaki et al. (2014); Callies and Ferrari (2013), and it further 298 evaluates the ability of NATL60 to resolve physical processes at fine-scales. In section (4), 299 we shall also discuss how this regime change from summer to winter affects the redistribution 300 of kinetic energy. 301

302

3.2 Spectral Slope

As we have shown in the previous section, a quick way to estimate the wavenumber 303 spectral power law is to compute the 1D wavenumber spectral density then estimate a slope 304 from this spectral by fitting a line to the spectral density curve within a selected wavenumber 305 range. This method is fast and easy to implement and provides a way to investigate regional 306 variability of ocean energetics both at the basin and global scale. For studies on mesoscale 307 energetics using satellite datasets and model outputs, this wavenumber range is mostly 308 within the error limits of the altimeter instrument (~ 70 km) and the horizontal scale of 309 meso/large scale motion (250–300 km). One drawback of this approach is that it does not 310 account for the changes in the scale of average energetic eddy structures with latitude. Scales 311 of motions that are mesoscales in the polar regions could be classified as submesoscales in 312 the tropics. 313

To characterize the spectral signature correctly, several recent studies have tried to propose different approaches to estimate the wavenumber spectral power law. For mesoscale resolving altimetry datasets, Vergara et al. (2019) estimated spectral slope between the peak of the spectral and the minimum of the Rossby radius and the Rhines scale following Eden (2007). A similar approach was presented in Sasaki and Klein (2012), where the authors estimated spectral slope between a fixed wavelength of 30 km (at the lower bound), and a scale that corresponds to the peak of the KE wavenumber spectral.

In order to show how sensitive the estimated slopes are to the selected wavelength range, we present in Figure 7a the average KE wavenumber spectral density and slope for box 3 in March for three different selected wavenumber ranges. The dashed lines with colors red, blue, and black represent the 10–100 km, 10–250 km, and 70–250 km, wavelength,

respectively. The estimated slopes for these three different wavelength ranges have different 325 values, therefore raising the question as to which slope is most representative of the dynamics 326 of this region. We repeat this analysis for all the boxes and present the map in Figure 8. The 327 mismatch is particularly pronounced in the sub-polar region, where the scales of the eddy 328 structures are relatively smaller. The 70–250 km wavelength range is a typical wavelength 329 for estimating spectral slope for satellite datasets because 70 km roughly corresponds to 330 the wavelength where the satellite data becomes noisy. The spectral slope in this range is 331 consistent with the already published work of Dufau et al. (2016) and Chassignet and Xu 332 (2017).333

To avoid the sensitivity of the estimated spectral slope to an a-priori selected wavelength 334 range, we introduce an approach that takes into account the dynamics of the regions and 335 the resolving capability of the model by estimating the spectral slope (Figure 7b) between 336 the energy-containing scale (Kjellsson & Zanna, 2017) and the effective resolution of the 337 model (Soufflet et al., 2016). The energy-containing scale (which represents the scale of the 338 most energetic eddy structure) is estimated from the kinetic energy wavenumber spectral 339 using equation (3) while the effective resolution (a function of the model grid-size) is taken 340 as $5 \times$ the model grid size, which is roughly equal to 10 km for both models. This approach 341 takes into account the scale of the energetic eddy structures within the flow region and also 342 takes into account the geographical variability of this scale, and therefore provides a way to 343 infer dynamical properties of oceanic motions in different regions. 344

$$\lambda_e = \frac{\int \int E(k_x, k_y) dk_x dk_y}{\int \int \sqrt{k_x^2 + k_y^2} E(k_x, k_y) dk_x dk_y}$$
(3)

We apply this technique to the output of both simulations (KE spectral density from 345 daily averages), and we present the estimated spectral slope and the energy-containing scale 346 (integral scale) for all the boxes in the North Atlantic (Figure 9). The estimated integral 347 scale from the wavenumber spectral density represents the averaged scale of energetic struc-348 tures in the selected region. On one hand, this scale varies regionally and fairly follows the 349 variability of the Rossby radius of deformation with latitude, with high values in the south 350 and relatively low values in the north. On the other hand, the estimated slope across the 351 basin is almost uniform and follows the prediction of QG with a slope value $\sim k^{-3}$. This 352 consistency with the QG prediction is observed in both model outputs and also holds in the 353 well known high energetic Gulf stream (box 1) and the low energetic OSMOSIS (box 10) 354 regions. This result form daily fields comes with the caveat of suppressing the impact of 355 high-frequency motions on the estimated slope values. 356

In order to investigate whether accounting for high-frequency motions would affect 357 our estimation of spectral slopes, we estimate the spectral slope (from KE spectral density 358 computed using hourly averages of velocity fields). We do this only for NATL60 because 359 HYCOM50 surface hourly output is available only for one month. The NATL60 spectral 360 slope from hourly spectral density in all the regions (Figure 10) has smaller values than 361 that of daily averages. As previously highlighted in section 3.1, this implies that the actual 362 spectral are shallower in hourly fields (as a result of stronger variance at the fine-scales) 363 compared to daily fields. The impact of intense submesoscale and ageostrophic flows is 364 better illustrated in the winter/summer map of the spectral slope (Figure 11). The slopes 365 in winter have values that are closer to k^{-2} throughout the domain. This indicates that 366 high-frequency motions are quite significant in the North Atlantic ocean basin. In the next 367 section, we shall discuss the impact of different dynamics on the redistribution of kinetic 368 energy. 369

370

4 Kinetic Energy Cascade

In this section, we present and discuss the exchange of energy due to non-linearity across different scales of motion. This exchange is estimated from the horizontal velocity fields using equation (2). A positive flux represents a direct cascade of energy, while a negative value represents an inverse cascade of energy.

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4.1 Annual Averages of Kinetic Energy Cascade

We show in Figure 12 the KE spectral flux computed using one year's daily outputs 376 of surface velocity fields. For simplicity, we show plots for boxes 3, 8, and 11 representing 377 latitudes of 35°N, 45°N, and 55°N, respectively. In all the boxes and both models, the 378 spectral flux is dominated by an inverse cascade of energy at large scales (between 25-50 379 km and 500 km) and a forward cascade of energy below 25-50 km. As observed earlier, 380 the spectral slope from daily averaged fields has a value that is ~ -3 , a characteristic 381 of QG turbulence. The energy exchanges computed using the same data show that the 382 flux is mostly upscale with a little forward flux at fine-scales. While the inverse cascade 383 is a well known phenomenon in geostrophic turbulence, the dynamics responsible for the 384 forward cascade at fine-scales is, however, still a subject for discussion. One would expect a 385 forward flux at fine-scales to be accompanied by a $(k^{-2} \text{ or } k^{-5/3})$ spectral slope. This sort of 386 relationship between spectral density and flux is not observed for estimates coming from the 387 daily fields. This finding is, however not new and agrees with the results of Brüggemann and 388 Eden (2015). The authors show that as soon as ageostrophic dynamics become important 389 in a quasi-QG flow, the flow is no longer restricted to an inverse cascade of energy. Instead, 390

the kinetic energy can proceed toward smaller scales providing a direct route to dissipation. We believe the fine-scale ageostrophic motions resolved by these simulations are significant for the flow to be in a regime that supports a forward cascade of energy.

It is interesting to note that the scale at which the inverse cascade is most intense 394 coincides with the energy-containing scale (dashed line in Figure 12 estimated from the 395 kinetic energy spectral density). This signifies that the inverse cascade is maximum at the 396 scale of the most energetic eddies. Just like the energy-containing scale, the scale of the most 397 intense inverse cascade varies with latitude with relatively smaller values in the sub-polar 398 regions (Figure 13). Also, depending on the region, part of the submesoscale range (0 to 25-399 50 km) falls to the left of the zero-crossing (where the flux changes sign). This implies that 400 submesoscales motions are involved in fluxing energy to large-scale via an inverse cascade 401 of energy. This scale of zero-crossing varies across the basin with higher values in high 402 EKE regions and relatively smaller values in low EKE regions (Figure 13). This regional 403 variability in the value of the zero-crossing shows that at fine-scales, the spatial scale at 404 which energy is injected into the flow is greater in eddy-rich regions, compared to eddy-poor 405 regions. 406

Figure 14 presents the KE spectral flux for Box 8 as a function of depth. The overall 407 shape of the flux is preserved, and the scale at which the inverse cascade is maximum is 408 also consistent with depth. The strength of the inverse cascade decreases with depth, and 409 the direct cascade at high wavenumbers is confined mostly to the surface. In section 3.1, 410 the question was raised as to whether HYCOM50's higher KE, when compared to NATL60, 411 is a consequence of HYCOM50's coarser vertical resolution, which could lead to a surface 412 intensified inverse cascade and hence more energetic surface eddies. However, in Figure 14, 413 we can see that both at the surface and all depth levels, the estimated inverse and direct 414 cascade is stronger in HYCOM50 than in NATL60. This indicates that HYCOM50 upscale 415 energy flux is not surfaced intensified and we can conclude that the disparity between the 416 two models in terms of energy levels is most likely due to differences in the length of the 417 spin-up as well as the choice of sub-grid scale parametrization, and that having only 32 418 isopycnal vertical levels is not detrimental to the representation of the dynamics in the 419 ocean interior. 420

421

4.2 Seasonality of Energy Cascade

In this section, we present the seasonality of the kinetic energy spectral flux by comparing winter (JFM) and summer (JAS) averages. Figure 15 shows the winter cascade (in blue) and the summer cascade (in red). There are two notable differences between the seasons. First, there is a shift in the zero crossings to higher wavenumbers in winter. Second, there

is a stronger forward cascade within the submesoscale range in winter. As highlighted in 426 the preceding section, a zero-crossing at the high wavenumbers partly indicates how much 427 submesoscale motions are involved in feeding large-scale motions via an inverse cascade of 428 energy. So, a shift to higher wavenumbers in wintertime signifies that smaller-scale struc-429 tures are involved in fluxing energy to larger scales. It is noteworthy that the integral scale 430 and scale of the maximum inverse cascade also undergo seasonality. There is a shift in the 431 scale to high wavenumber from winter to summer. This can be interpreted as a reduction 432 in the average size of energetic eddies structures in winter. This sort of seasonality in eddy 433 length-scale is documented in Ajayi et al. (2020). 434

It is interesting to understand the contribution of the different dynamics on the sea-435 sonality of the energy cascade. Recent studies have shown that submesoscales are energetic 436 in wintertime (Mensa et al., 2013; Sasaki et al., 2014; Callies, Flierl, et al., 2015; Rocha et 437 al., 2016), and their emergence is forced by mechanisms such as frontogenesis, wind-induced 438 frontal instabilities, mixed layer instability among many others (Thomas, 2008; McWilliams, 439 2016). Sasaki et al. (2014) argued that submesoscales generated via mixed layer instability 440 could feed large scale motion via an inverse cascade of energy, hence a shift in the zero-441 crossing towards high wavenumbers. More recently, Schubert et al. (2020) used both the 442 coarse-graining approach (Aluie et al., 2017) and spectral analysis to investigate the role 443 of mixed layer baroclinic instabilities on kinetic energy exchanges. Their results show that 444 mesoscale oceanic eddies are strengthened by the absorption of submesoscale mixed layer 445 eddies and that the forward cascade of energy at very fine-scales occur mostly in fronto-446 gentic regions. Following these aforementioned findings, we hypothesize that the increased 447 forward cascade presented in this study, could be associated with frontogenesis and subme-448 soscale frontal instabilities. This seasonality highlights how submesoscale motions modulate 449 the redistribution of energy between scales of motions; hence, the need for climate (ocean) 450 models with submesoscale resolving capability. 451

452

4.3 Impact of High-Frequency Motions on Energy Cascade

In section 3.1, we observed that in the presence of high-frequency motions, the spectral densities (computed from hourly fields) are shallower with an increased variance at finescales compared to daily averaged fields. In this section, we are going to discuss the impact of this increased variance on energy exchanges.

The contribution of high-frequency motions to kinetic energy spectral flux is highlighted in Figure 16a where we show the comparison of the spectral flux computed from daily versus hourly fields. The magnitude of the forward cascade at submesoscales is significantly stronger in hourly spectral flux. That high-frequency motions can provide a pathway to

kinetic energy dissipation is illustrated in these results. The dynamics responsible for this 461 increase in forward cascade are likely due to an energy loss to the generation of gravity 462 waves and unbalanced submesoscales. It is interesting to note that the scale of the most 463 intense inverse cascade remains the same while there is a slight shift in the zero-crossing 464 towards higher wavenumber. This result suggests that the impact of high-frequency motion 465 on energy exchanges is mostly concentrated at fine-scales. As seen in the daily averages, the 466 winter flux is equally stronger in hourly fields (Figure 16a). We believe that this increase 467 is attributed to resolved internal gravity waves and intense (un)balanced submesoscales. 468 However in summer time both daily and hourly flux are identical. 469

470

4.4 Diagnosing Spectral Flux from SWOT

NATL60 and HYCOM50 are submesoscale permitting model simulations that have been 471 created to simulate the scales of motions that we expect SWOT to see from space. SWOT 472 will provide measurements of sea surface heights from which velocities (based on geostrophic 473 approximations) will be inferred. Geostrophically balanced motions dominate the ocean at 474 meso/large-scale, and the inferred geostrophic velocities at this scale mostly reflect the 475 absolute velocity of these large scales motions. However, geostrophy is less accurate for 476 fine-scale motions, particularly at the submesoscales and this remains a challenge due to 477 the projection of ageostrophic motions on SSH field. Recent studies have shown that some 478 classes of non-wave ageostrophic motions could impact the forward cascade of kinetic energy 479 at fine-scales (Capet, McWilliams, et al., 2008). We have equally shown in the previous 480 sections that high-frequency motions in the form of ageostrophic waves and unbalanced 481 submesoscales can contribute significantly to a forward cascade of energy at fine-scales. 482 In light of this, we are not sure if SWOT data can accurately estimate the redistribution 483 of kinetic energy at fine-scales, given that SWOT will provide information down to ~ 15 484 km. We are curious to see if the geostrophically inferred surface velocity would capture the 485 accurate energetics at scales < 50 km where geostrophy is likely to fail. 486

To investigate this, we present in Figure 17 the spectral flux from total velocity and 487 geostrophic velocity for three regions (same as for the previous sections). The geostrophic 488 velocity is estimated from the sea surface height (SSH) using the geostrophic approximation. 489 The strength of the energy cascade differs between the flux computed from the total velocity 490 and that of the geostrophic velocity. This difference is consistent in all three boxes and 491 in the two models. In particular, at the very high wavenumbers, the forward cascade is 492 underestimated in the flux computed from the geostrophic velocity. A possible reason for 493 this mismatch at smaller scales could be explained by the findings of Brüggemann and Eden 494 (2015) that showed that ageostrophic flows at fine scales are an excellent catalyst for energy 495

cascade towards dissipation. Despite the differences in terms of flux magnitude, the overall
shape of the flux is consistent for the two forms of spectral flux. The scale at which the
inverse cascade is maximum is the same irrespective of the type of velocity fields.

⁴⁹⁹ 5 Discussion and Summary

In this study, we presented the analysis of kinetic energy wavenumber spectral density, 500 slope, and flux by using datasets from daily and hourly outputs of two submesoscale permit-501 ting ocean models of the North Atlantic. The analysis presented has shown that in summer, 502 the North Atlantic ocean follows the QG framework (with $\sim k^{-3}$ spectral shape) and in 503 winter, the basin mostly reflects a k^{-2} spectral shape, a characteristic of a front dominated 504 regime. The estimated kinetic energy spectral flux revealed an overall net inverse cascade 505 of energy with a significant direct cascade of energy at high wavenumbers. The spectral 506 flux undergoes a seasonality that is associated with a stronger forward cascade at high 507 wavenumbers in winter. This increased forward cascade in winter is further amplified in the 508 presence of high-frequency motions. The spectral flux as a function of depth reveals that 509 the forward cascade at high wavenumbers is confined to the mixed layer while the inverse 510 cascade dominates the water column down to 700m. We showed that the maximum inverse 511 cascade occurs at a scale that coincides with the energy-containing scale. 512

Until recently, most basin/global scale simulations had their outputs stored in the form 513 of daily averages. Our results show that high-frequency motions (that are only resolved in 514 hourly outputs) affect the distribution and exchanges of kinetic energy. We observed that the 515 difference between the daily and hourly results is mostly in the form of an increased variance 516 and (forward) cascade at fine-scales in favor of hourly fields. High-frequency motions are 517 dominated mainly by ageostrophic motions that include unbalanced submesoscales and fast 518 propagating internal gravity waves. These two classes of motions are out of phase seasonally 519 with stronger submesoscales in winter and stronger internal gravity waves in summer. It is 520 puzzling that while the exchanges of energy at fine-scale are unaffected by high-frequency 521 motions in summertime, the distribution of energy shows the contrary. We observed higher 522 variance at fine-scales in power spectral density estimated from hourly fields. The reason 523 for this disparity between the impact of high-frequency on spectral density and flux in 524 summertime is not apparent but would be an interesting subject to investigate further. 525

The kilometric simulations used in this study have similar horizontal grid spacing but different numerics, sub-grid parameterization, and vertical resolution. Despite these differences, the two simulations agree well on the overall dynamics of the North Atlantic. However, HYCOM50 is more energetic compared to NATL60 both at the surface and in the interior. We found the estimated cascade in HYCOM50 to be of higher magnitude

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compared to NATL60 for both direct and inverse cascade. The difference in energy levels between the two models could be due to the difference in length of spin-up or/and the choice of subgrid-scale parameterization. Initially, we thought that HYCOM50 having just 32 hybrid layers in the vertical, could lead to a more surface intensified energy cascade in HYCOM50 than in NATL60. Nevertheless, this is not the case, because across all scales and at depth, HYCOM50 seems to show stronger energetic compared to NATL60.

NATL60 and HYCOM50 are designed mainly to serve as an observational dataset for the 537 anticipated SWOT mission. SWOT will provide sea surface height, and by using geostrophic 538 approximation, we would obtain the geostrophic velocity, a requisite for computing cross-539 scale energy transfer. At fine-scale, geostrophic approximation is however less accurate. 540 Thus, accurately diagnosing surface velocity from sea surface height at fine-scales remains 541 a challenge, and this has an impact on the estimate of kinetic energy cascade. Our results 542 show that at fine-scales, not accounting for the ageostrophic motions could affect at all 543 scales, the true magnitude of the estimated cascade of kinetic energy. 544

The results presented in this study are based on the output of ocean numerical sim-545 ulations that are forced with realistic atmospheric winds. Recent literature suggests that 546 air-sea coupling at fine-scales could affect the evolution and energetics of oceanic eddies. 547 Renault et al. (2016) using a coupled/uncoupled model of the California Upwelling System 548 argued that the ocean-atmosphere interactions have feedback that acts as an oceanic eddy 549 killer. This feedback deflects energy from the geostrophic current into the atmosphere and 550 dampens geostrophic kinetic energy. A possible future study would be to recompute the 551 (kinetic energy transfer) diagnostics in this study using datasets from an ocean-atmosphere 552 coupled simulation. This sort of analysis would take into account the direct impact of air-sea 553 interaction on the ocean's kinetic energy exchanges. 554

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- ⁵⁶⁷ and SARAL AltiKa altimeter. HYCOM50 and NEMO-NATL60 data can be accessed at
- ⁵⁶⁸ ftp://ftp.hycom.org/pub/xbxu/ATLb0.02/ and http://meom-group.github.io/swot-natl60/access-
- ⁵⁶⁹ data.html, respectively.

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 $\label{eq:Table 1. Table of model parameters for NATL60 and HYCOM50$

| | NATL60 | HYCOM50 |
|----------------------|--------------------------------|------------------------|
| Domain | 26.5N - 65N | 28 - 80N |
| Numerical Code | Nemo v.3.6 | HYCOM |
| Horizontal grid | 1/60: 0.9-1.6 km | 1/50:1.1-2.2 km |
| Vertical coordinate | Z partial cells | Hybrid (Z & isopycnal) |
| Integration period | 6 Months | 20 years |
| Vertical grid | 300 Levels : 1-50 m $$ | 32 Layers |
| Boundary conditions | GLORYS2v3 | GDEM |
| Atmospheric forcing | DFS5.2 | ERA-40 |
| Horizontal Viscosity | Implicit in momentum advection | Laplacian & Biharmonic |



Figure 1. Schematics of kinetic energy spectral flux in the ocean at mid-latitude. Blue : inverse cascade of energy, Red : forward cascade of energy.



Figure 2. Left panel : snapshot of surface currents speed (m/s) on march 1st for NATL60 (a) and HYCOM50 (b). Right panel : surface eddy kinetic energy (cm^2s^{-2}) computed from daily output for NATL60 (c) and HYCOM50 (d).



Figure 3. Standard deviation of sea surface height based on one year datasets for (a) AVISO,
(b) NATL60 and (c) HYCOM50. The SSH values for NATL60 and HYCOM50 were degraded to 0.25° spatial resolution. In this comparison, we have used AVISO mean dynamical topography dataset from October 2012 to September 2013.



Figure 4. One year average of kinetic energy spectral density $(m^2 s^{-2}/cpm)$ for Box 8 computed from horizontal total velocity as a function of depth for (a) NATL60 and (b) HYCOM50.



Figure 5. Kinetic energy spectral density for Box 8 (averaged over 1000 m depth) computed from daily output of horizontal total velocity for NATL60 (thick line) and HYCOM50 (dash line). (a) one year mean (b) winter (blue line) and summer (red line) averages. See Figure 5 in SI for a comparison of the surface vs depth averaged spectral density. A comparison of the spectral density between three different depth levels and the surface is presented in Figure 6 of SI



Figure 6. Comparison between surface kinetic energy spectral density computed from daily averages (thick line) and hourly averages (dash line) of velocity outputs for Box 8. (a) one year mean (b) winter (blue line) and summer (red line) averages. See Figure 7 and 8 in SI for a similar plots for all the boxes.)



Figure 7. (a) Average surface KE spectral density and slope for box 3 (NATL60) in the month of March for three different selected wavenumber ranges. The wavelengths ranges are represented by dashed lines with the color red, blue and black for 10-100 km, 10-250 km and 70-250 km respectively. (b) A schematic to illustrate the proposed dynamical approach to estimate spectral slope. λ_e is the energy-containing scale (which represents the scale of the most energetic eddy structure) and it is estimated from the kinetic energy wavenumber spectral density by using equation (3) while E_r is the effective resolution (a function of the model grid-size) and is taken as 5 × the model grid size. E_r is roughly equally to 10 km for both NATL60 and HYCOM50.



Figure 8. Map of spectral slope estimated from the surface kinetic energy spectral density for three different selected wavenumber ranges. Colour red, blue and black represent 10 - 100 km , 10
- 250 km and 70 - 250 km respectively



Figure 9. Map of spectral slope (blue colour) and energy containing scale, λ_e (black colour) from the surface kinetic energy spectral density. The slope is estimated between the model effective resolution (E_f) and the energy containing scale (λ_e).



Figure 10. Map of spectral slope from the surface kinetic energy spectral density computed from daily versus hourly Fields. The slope is estimated between the model effective resolution (E_f) and the energy containing scale (λ_e) .



Figure 11. Map of spectral slope from NATL60 hourly surface kinetic energy spectral density in winter (JFM) and summer (JAS). The slope is estimated between the model effective resolution (E_f) and the energy containing scale (λ_e) .



Figure 12. One year average of surface kinetic energy spectral flux computed from the daily output of horizontal total velocities. NATL60 (upper panel) and HYCOM50 (lower panel)



Figure 13. A geographical map of the scale of the kinetic energy spectral flux zero-crossing (red) and the most intense inverse cascade (black) estimated from one year average of the surface kinetic energy spectral flux for (a) NATL60 and (b) HYCOM50.



Figure 14. One year average of kinetic energy spectral flux for Box 8 computed from horizontal total velocity as a function of depth for (a) NATL60 and (b) HYCOM50.



Figure 15. Winter (blue line) and summer (red line) average of surface kinetic energy spectral flux computed from daily output of horizontal total velocities. Dash lines represents the energy containing scale. NATL60 (upper panel) and HYCOM50 (lower panel)



Figure 16. Comparison between surface kinetic energy spectral flux computed from daily averages (thick line) and hourly averages (dash line) of velocity outputs for Box 8. (a) one year mean (b) winter (blue line) and summer (red line) averages



Figure 17. Surface kinetic energy spectral flux computed from total velocity (black line) versus geostrophic velocity (green line). NATL60 (upper panel) and HYCOM50 (lower panel)