

Upper Mantle Seismic Anisotropy as a Constraint for Mantle Flow and Continental Dynamics of the North American Plate

Wanying Wang^{1,1} and Thorsten Becker^{1,1}

¹The University of Texas at Austin

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Abstract

The alignment of intrinsically anisotropic olivine crystals under convection is typically invoked as the cause of the bulk of seismic anisotropy inferred from shear-wave splitting (SWS). This provides a means of constraining the interplay between continental dynamics and the deep mantle, in particular for densely instrumented regions such as North America after USArray. There, a comparison of “fast orientations” from SWS with absolute plate motions (APM) suggests that anisotropy is mainly controlled by plate motions. However, large regional misfits and the limited realism of the APM model motivate us to further explore SWS based anisotropy. If SWS is estimated from olivine alignment in mantle circulation instead, plate-driven flow alone produces anisotropy that has large misfits with SWS. The addition of large-scale mantle density anomalies and lateral viscosity variations significantly improves models. Although a strong continental craton is essential, varying its geometry does, however, not improve the plate-scale misfit. Moreover, models based on higher resolution tomography degrade the fit, indicating issues with the flow model assumptions and/or a missing contributions to anisotropy. We thus compute a “lithospheric complement” to achieve a best-fit, joint representation of asthenospheric and frozen-in lithospheric anisotropy. The complement shows coherent structure and regional correlation with independently imaged crustal and upper mantle anisotropy. Dense SWS measurements therefore provide information on depth-dependent anisotropy with implications for tectonics, but much remains to be understood about continental anisotropy and its origin.

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Wanying Wang^{a,b,*}, Thorsten W. Becker^{a,b}

^a*Institute for Geophysics, Jackson School of Geosciences, University of Texas at Austin, Austin, Texas, USA*

^b*Department of Geological Sciences, Jackson School of Geosciences, University of Texas at Austin, Austin, Texas, USA*

Abstract

The alignment of intrinsically anisotropic olivine crystals under convection is typically invoked as the cause of the bulk of seismic anisotropy inferred from shear-wave splitting (SWS). This provides a means of constraining the interplay between continental dynamics and the deep mantle, in particular for densely instrumented regions such as North America after USArray. There, a comparison of “fast orientations” from SWS with absolute plate motions (APM) suggests that anisotropy is mainly controlled by plate motions. However, large regional misfits and the limited realism of the APM model motivate us to further explore SWS based anisotropy. If SWS is estimated from olivine alignment in mantle circulation instead, plate-driven flow alone produces anisotropy that has large misfits with SWS. The addition of large-scale mantle

*Corresponding author

Email address: wanying@utexas.edu (Wanying Wang)

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Keywords: continental dynamics, seismic anisotropy, North American plate

1. Introduction

Upper mantle seismic anisotropy is suggested to be mainly caused by the alignment of olivine aggregates in mantle flow. This is referred to as olivine lattice preferred orientation (LPO), and LPO is expected to align with shear under convection. This relationship provides a link between asthenospheric flow and seismic observations, in particular to study the relationships between surface geology and the underlying mantle dynamics in continental plates (e.g. Silver, 1996; Long and Becker, 2010). In order to obtain information about upper mantle flow, shear-wave splitting (SWS) analysis of teleseismic phases is widely used to infer azimuthal anisotropy. SWS measures the separation of shear waves into two orthogonally polarized pulses upon traversing an anisotropic medium. The polarization plane orientation of the faster shear wave pulse is often called the “fast azimuth”, and is expected to parallel the alignment of the seismically fast [100]-axes of the olivine aggregates and the sense of shear. The delay time between the fast and slow wave arrivals at the surface indicates the anisotropy magnitude accumulated along the path, and by inference, the depth extent or layer thickness of the anisotropic part of the mantle or lithosphere (e.g. Silver, 1996; Savage, 1999). Teleseismic SWS measurements use *SKS*, *SKKS* and *PKS* phases that have nearly vertical ray paths and sample the upper mantle beneath the seismic stations with poor vertical, but good lateral resolution.

Recently, the USArray seismometer deployment during the EarthScope

24 effort provided unprecedented coverage of United States, renewing
25 efforts to investigate mantle dynamics within and underneath the North
26 American plate (e.g. Hongsresawat et al., 2015; Long et al., 2016; Zhou
27 et al., 2018). Here, we compare a range of mantle flow model predictions of
28 upper mantle anisotropy to the observed SWS fast orientations to advance
29 our understanding of North America upper mantle dynamics (Fig. 1).

30 The SWS dataset used in this study is shown in Fig. 1b and newly
31 spans the whole continent at roughly uniform station spacing. The SWS
32 compilation consists of 14,326 splits from the updated compilation of
33 Becker et al. (2012), as well 29,061 standardized splits from Liu et al.
34 (2014), Refayee et al. (2014), and Yang et al. (2016, 2017). We also include
35 results from offshore experiments (Bodmer et al., 2015; Ramsay et al., 2015;
36 Lynner and Bodmer, 2017).

37 As has been discussed earlier based on more limited compilations, the
38 fast SWS orientation within the U.S. are generally E-W to NW-SE (e.g.
39 Silver, 1996), exhibit a circular pattern beneath the Great Basin (e.g. Zandt
40 and Humphreys, 2008; Hongsresawat et al., 2015), and orogen-parallel
41 orientation beneath and around the Appalachians (e.g. Long et al., 2016)
42 (Fig. 1b). While we mainly consider fast azimuths below, we note that the
43 delay times of SWS vary in systematic fashion. Broadly speaking, delay
44 times are larger beneath most of the western U.S. and the south central
45 U.S., and smaller beneath the interior plain, the Appalachians and the
46 southern Great Basin (Fig. 1b).

47 In order to link those observations of azimuthal anisotropy to
48 continental dynamics, we can consider the geological history of the
49 region. In the broadest of strokes, we note that the western U.S. has been
50 tectonically active since the late Mesozoic, from the Laramide orogeny to
51 the ongoing subduction-related orogenesis in the Cascades and extension
52 in the Basin and Range. A relatively thinner lithosphere in the west
53 likely plays a role in suggested scenarios where active mantle flow affects
54 lithospheric deformation beneath the Basin and Range, the Colorado
55 plateau and the Rockies (e.g. Savage and Sheehan, 2000; Karlstrom et al.,
56 2012). Likewise, mantle flow itself may have eroded part of the lithosphere
57 and caused thinning and extension at the Basin and Range (e.g. Lekić and
58 Fischer, 2014). A thinner lithosphere relative to the cratonic eastern U.S.
59 also implies a reduced role of possible shallow, frozen in anisotropy (e.g.
60 Silver, 1996), and perhaps a more readily understandable link between
61 asthenospheric flow and SWS.

62 The area through the central U.S. to the west of the Appalachian
63 mountains is within the extent of the North American Craton, which is
64 part of the oldest lithosphere on Earth that had been stable for over 1.7
65 Ga (e.g. Hoffman, 1989). The lithospheric root beneath the craton extends
66 to over 200 km depth (e.g. Gung et al., 2003; Steinberger and Becker, 2016),
67 and is suggested to have higher viscosity than the surrounding mantle
68 (e.g. Lenardic and Moresi, 1999). Beneath the central U.S., the oldest
69 part of the cratonic region is stable since the Archean, and may preserve

70 relatively larger degrees of shallow, frozen-in anisotropy. On the other
71 hand, the cratonic root may divert upper mantle flow, perturbing flow
72 at the craton's edge and inducing counter flow beneath it, which could
73 possibly strengthen regional lithosphere-asthenosphere coupling (e.g.
74 Silver, 1996; Fouch et al., 2000). This phenomenon is likely important for
75 understanding the details of upper mantle flow dynamics and the origin of
76 azimuthal anisotropy beneath the eastern U.S., which sits atop the cratonic
77 boundary and edge.

78 Tectonic features in the eastern U.S. include the Proterozoic rifting and
79 Paleozoic compressional orogenic events, followed by extensional events
80 in the Mesozoic. Based on SWS splitting and modeling, Fouch et al. (2000)
81 suggested that the observed anisotropy reveals the combined effect from
82 the lithospheric and sublithospheric anisotropy in this region. Small-scale
83 upper mantle density variations and lithospheric thickness variations exist
84 in this region (Fig. 2) and might cause perturbations in anisotropy as well.
85 For example, the northern Appalachian upwelling that can be inferred
86 from slow seismic tomography anomalies (Schmandt and Lin, 2014) might
87 relate to the Great Meteor hot spot track, and possibly indicate convection
88 on relatively small scales in the surrounding mantle (e.g. Schmandt and
89 Lin, 2014; Levin et al., 2018). Lithospheric thickness appears to decrease
90 rapidly from the plateau to the east of the Appalachian, and is suggested
91 to relate to lithospheric weakness from Eocene delamination (e.g. Mazza
92 et al., 2014).

93 Convective flow models should be able to predict the current
94 sublithospheric LPO to match the SWS observations if the models capture
95 the major contributors that affect the present day upper mantle strain (e.g.
96 Long and Becker, 2010). Given the extensive tectonic activity and prior
97 sampling, much of the geodynamic SWS modeling previously focused on
98 the western U.S.. For example, Silver and Holt (2002) jointly interpreted
99 splitting and GPS observations to infer eastward mantle flow. Becker
100 et al. (2006b) computed LPO from mantle flow modeling, and showed that
101 SWS outside the Basin and Range domain could be fit well with relatively
102 simple flow models as long as a downwelling associated with the Farallon
103 slab was included. More recently, Zhou et al. (2018) computed anisotropy
104 from more complex models with lateral viscosity variations (LVVs) and
105 were able to reproduce the circular pattern discussed by Zandt and
106 Humphreys (2008).

107 Given the long geological history of the North American plate, we
108 expect that the lithosphere-asthenosphere system will reflect different
109 contributions to anisotropy. Based on joint surface wave and SWS
110 analysis, Yuan and Romanowicz (2010) suggested layering with various
111 lithospheric azimuthal anisotropy orientations beneath North America,
112 and many authors have made the case that variations in SWS fast
113 orientations with back-azimuth are best explained by a significant
114 lithospheric anisotropy source (Silver, 1996; Savage, 1999).

115 Here, we seek to address azimuthal anisotropy underneath the U.S.,

116 explore which role small-scale lateral variations in density and viscosity
 117 play for predictions of asthenospheric anisotropy, and then return to the
 118 question of lithospheric anisotropy.

119 2. Methods

120 *Mantle flow modeling*

121 This study broadly follows the approach of Becker et al. (2006b)
 122 and Miller and Becker (2012). Under the Boussinesq and infinite
 123 Prandtl number approximations, the conservation equations for mass and
 124 momentum for mantle flow are given by

$$\nabla \cdot u = 0,$$

$$-\nabla p + \nabla \cdot \eta(\nabla u + \nabla^T u) - \delta \rho g \hat{e}_r = 0.$$

125 Here, u is the velocity vector, p is the dynamic pressure, η is the viscosity,
 126 $\delta \rho$ is the density anomaly, g is the gravitational acceleration and $\hat{e}_r$ is the
 127 radial unit vector. We solve the conservation equation using the finite
 128 element software CitcomS (Zhong et al., 2000) in a 3-D spherical domain.
 129 The surface boundary condition of most of our models are prescribed plate
 130 motions in the no-net-rotation (NNR) reference frame (NNR-NUVEL-1,
 131 by Argus and Gordon, 1991). The mechanical boundary condition at the
 132 core-mantle boundary is free-slip. Therefore, the absolute reference frame
 133 of the plate motions is irrelevant for relative velocities, and hence mantle

134 flow predicted anisotropy.

135 Density variations outside continental cratons are assumed to be
136 purely thermal and scaled from seismic tomography anomalies $\ln v_S$ with
137 a simplified scaling of $R = \frac{d \ln \rho}{d \ln V_S}$. To ensure that the system is dynamically
138 consistent, the resulting vigor of density-driven flow is adjusted via R
139 such that when the same density variation is used in a model with
140 free-slip surface boundary conditions, the same RMS surface velocity as
141 for prescribed absolute plate motion (APM) results. The resulting $R = 0.24$
142 is in line with prior work (e.g. Miller and Becker, 2012). Inputs for the
143 density variations come from two models: SMEAN is a composite, global
144 S -wave tomography model (Becker and Boschi, 2002) used for reference
145 (Fig. 2a). In order to capture the possible effect of small scale density
146 anomalies beneath the U.S., we merge the regional tomography model
147 of Schmandt and Lin (2014) with SMEAN to obtain MERGED where the
148 edges of the embedded high resolution region are smoothed (Fig. 2b).

149 Within cratons, where we expect compositional anomalies (e.g. Jordan,
150 1978; Forte and Perry, 2000), we assume the lithosphere to be neutrally
151 buoyant by setting craton-related seismic velocity anomalies to zero. The
152 depth of the cratonic root is suggested to be $\sim 200 - 250$ km, for example
153 by Yuan and Romanowicz (2010) and Gung et al. (2003), and geodynamic
154 inversions (Forte and Perry, 2000). Since the tomography models we use
155 show fast velocity anomalies that extend to ~ 300 km beneath the North
156 American Craton, we use 300 km depth as the extent of the neutrally

157 buoyant zone, for simplicity. The viscosity of the cratonic root is important
158 in maintaining its long term stability. Convection modeling suggests it to
159 be 100 to 1000 times more viscous than the ambient mantle (e.g. Lenardic
160 and Moresi, 1999). Here we assume it to be 10 times more viscous than the
161 continental lithosphere, which is 500 times the regular asthenosphere.

162 Both radial and lateral viscosity variations are considered. The
163 viscosity model is built upon a three layered radial viscosity structure
164 (RVV). The viscosity of each of the 0 – 100 km, 100 – 660 km, and 660 –
165 2891 km layers is 150, 1, and 60 times the reference value. For the 100 –
166 660 km depth range, a temperature dependent lateral viscosity variation
167 (LVV) is applied to the three layered RVV structure, and the viscosity is
168 given by equation: $\eta = \eta_0 \exp E(T - T_{ref})$. In this equation η_0 is from
169 the RVV structure, E scales the effect of temperature dependence with a
170 value of 7, T is the non-dimensional temperature at each point inferred
171 from the tomography models, and T_{ref} is the non-dimensional reference
172 temperature that equals to 0.5. In the upper 300 km, η is then multiplied
173 by a structure dependent viscosity factor to account for the LVVs.

174 The viscosity factor at each of the plate boundaries, the oceanic and
175 continental lithosphere, cratonic keels and oceanic asthenosphere is 0.01,
176 1, 50, 500, 0.01, respectively (cf. Miller and Becker, 2012). Focusing on
177 continental keels underneath the U.S., we test two viscosity structures,
178 models LVV1 and LVV2. The cratonic keel geometry of LVV1 is inferred
179 from global tomography using the approach of Steinberger and Becker

180 (2016) and the model SL2013 (Schaeffer and Lebedev, 2013). In LVV1, the
181 minimum lithospheric thickness is 50 km in both continental and oceanic
182 regions (Fig. 2c). LVV2 is taken from the reference craton model of Miller
183 and Becker (2012) where keel geometry is simpler, and keel depth constant
184 at 300 km (Fig. 2d).

185 *Asthenospheric and lithospheric anisotropy modeling*

186 Based on the mantle circulation models, we then use particle tracking
187 and the D-Rex mineral physics approximation (Kaminski et al., 2004) to
188 compute LPO as the tracers are advected until a logarithmic saturation
189 strain of 0.75 is reached (Becker et al., 2006b; Miller and Becker, 2012).
190 We assume that mantle circulation is stationary over the few Myr that it
191 takes to achieve this strain (cf. Becker et al., 2003, 2006a). Depth-dependent
192 single crystal elasticity constants and Voigt averaging are then used to
193 determine the elasticity tensor $\mathbf{C}$ at 25 km spaced locations underneath
194 each of the stations where SWS is measured.

195 While SKS splitting is well known to not linearly average over $\mathbf{C}$ along
196 the path, such differences are generally limited as long as anisotropy does
197 not vary strongly with depth (e.g. Becker et al., 2012). We conducted tests
198 using the full-waveform approach of Becker et al. (2006b) and found that
199 regionally, details of the SWS predictions were affected. However, our
200 overall conclusions regarding the flow model predictions would be the
201 same.

202 Here, we therefore mainly consider the simplified, depth-averaged
203 tensor approach, computing an average for the 25 to 375 km depth
204 range, but we revisit a two layer case below. Under the tensor-averaging
205 assumption, the Christoffel equation is then solved for the equivalent SWS
206 delay times and fast azimuths using a back-azimuthal average.

207 Upon having predicted the inferred LPO anisotropy caused by mantle
208 flow in the asthenosphere, we compare it with the SWS observation and
209 compute the absolute angular misfit, $\Delta\alpha$, between the two ($\Delta\alpha \in [0^\circ, 90^\circ]$)
210 for a range of flow models. Given the relatively poor overall fit for the
211 study area of those predictions (Fig. 4 to 7) compared to earlier work
212 (e.g. Miller and Becker, 2012), we also explore the possible contributions
213 of the lithosphere more extensively. For this, we assume that there are
214 two anisotropy layers, and the bottom layer is fixed to the flow model
215 predicted anisotropy, which represents the depth averaged asthenospheric
216 anisotropy. Then we invert for the best-fit “lithospheric complement”
217 based on a parameter space exploration and Silver and Savage’s (1994)
218 approach, and find the fast azimuth and delay time of the top layer
219 anisotropy that, results in the best match to the back-azimuthally
220 distributed SWS observations at each station.

221 3. Results

222 *SWS alignment with absolute plate motions*

223 Assuming that plate motions at the surface in some absolute reference
224 frame (APM) are reflective of the orientation of shear between the
225 lithosphere and mantle, APM alignment is a first order test for the origin
226 of anisotropy (Silver, 1996). The SWS fast orientation beneath the U.S. are
227 indeed found to be generally aligned with plate motion directions (e.g.
228 Hongsresawat et al., 2015). Figure 3 substantiates earlier analyses using
229 our denser SWS dataset by comparing it with APM in the NNR reference
230 frame (Argus and Gordon, 1991), and the spreading-aligned reference
231 frame (Becker et al., 2015).

232 On a plate scale, SWS fast axes have NE-SW orientations similar to
233 the NNR APM orientation of North America, especially in the western
234 U.S. (Fig. 3a), leading to a plate-scale mean misfit of $\langle \Delta\alpha \rangle \approx 30^\circ$. The
235 spreading-aligned APM is more similar to the SWS fast orientations and
236 $\langle \Delta\alpha \rangle$ is further reduced by $\sim 5^\circ$ (Fig. 3b). While misfit values thus depend
237 on different APM reference frames (e.g. Becker et al., 2015), similar local
238 misfit fluctuations are observed, and those may be related to mantle flow
239 deviating from implied APM shear. For example, in the southeastern
240 Rockies (Fig. 3), the large angular misfit might relate to local lithospheric
241 thickness variations (Refayee et al., 2014; Hongsresawat et al., 2015).
242 Another significant misfit is found at the eastern U.S. and the southern
243 Appalachian Mountains (Fig. 3), where anisotropy possibly contains a

244 lithospheric frozen-in component (e.g. Levin et al., 2018; Long et al., 2016).

245 *Flow model predictions for SWS*

246 We now test the role of asthenospheric convection other than APM
247 shearing by predicting anisotropy from the sublithospheric flow that is
248 driven by plate motion alone, or in addition by density variations. We
249 investigate the effect of viscosity variations caused by cratons, a weak
250 oceanic asthenosphere layer, and plate boundary weak zones. For each
251 flow model, we explore the flow itself, and compute the misfit between
252 predicted and observed anisotropy, $\Delta\alpha$ (Table 1). We find it helpful to
253 visualize the effect of shearing in flow models by plotting the vector
254 difference $\vec{v}_{shear}$ between the horizontal flow velocities at the surface
255 $\vec{v}_{surface}$ and at a typical, 200 km depth $\vec{v}_{200km}$ with

$$\vec{v}_{shear} = \vec{v}_{200km} - \vec{v}_{surface},$$

256 as such differential velocities can be a rough proxy for LPO alignment.

257 *The effect of plate motion induced circulation with LVVs*

258 Our starting Model 1 only has radial viscosity variations and is purely
259 driven by prescribed plate motions. Given the effects of geometry and
260 return flow, we expect that the induced asthenospheric shearing will be
261 different from the APM model of Fig. 3 even for this simple circulation
262 model (e.g. Long and Becker, 2010), and this is indeed the case.

263 Figure 4a shows that beneath the western and central U.S., the
 264 direction of $\vec{v}_{200km}$ deviates from the plate motion direction due to the
 265 flow perturbation at the Pacific-North America plate boundary. The flow
 266 direction is to the W to SW and the shear direction forms an 110° to
 267 150° angle with the plate motion in this region. While details depend
 268 on the viscosity structure (cf. Becker et al., 2006b), this plate boundary
 269 flow perturbation extends almost throughout half the continent. In the
 270 Eastern U.S., sublithospheric Couette flow (Fig. 5a) is more in line with
 271 APM, such that shear is roughly into the opposite direction (Fig. 4a). The
 272 mis-alignment of shear and plate motion vectors is subdued in the eastern
 273 U.S. but still of order 20° to 30° .

274 Since there are no small-scale flow perturbation or abrupt changes
 275 in viscosity, the orientation of differential velocities of Fig. 4a are
 276 representative of the predicted anisotropy (Fig. 4b). W-E oriented
 277 predicted fast axes fit well with the SWS observation onshore in NW
 278 U.S.. However, there are large misfits with regions of consistent $\sim 90^\circ$
 279 misalignment such as the in the southern Rockies. The overall match
 280 between SWS and predictions is very poor at $\langle \Delta\alpha \rangle \approx 45^\circ$ (which is the
 281 expectation for random). This indicates that plate-induced shear flow
 282 without density anomalies is actually a much worse model in this case
 283 compared to the APM hypothesis of Fig. 3.

284 Based on Model 1, Model 2 adds in weak plate boundaries and strong
 285 cratonic keels from viscosity model LVV1. Comparing Models 1 and 2,

286 changes in horizontal flow mainly occur beneath and around the craton
287 (Figs. 4a and c). The spatial extent of this change is shown in the
288 differential flow velocity profile in Fig. 5d. Due to its high viscosity,
289 the craton maintains and enhances plate-like motion down to ~ 300 km
290 depth, as shown in Fig. 4c, and transfers it to the sublithospheric mantle.
291 The craton also causes minor flow perturbations in the radial direction
292 at the lithospheric thickness discontinuities beneath the Colorado Plateau
293 (Fig. 5d). However, the directional change in flow introduced by the keel
294 is overall small, such that the anisotropy predictions of Models 1 and 2 are
295 fairly similar (Figs. 4b and d). The weak plate boundary effect of Model 2
296 changes the flow and shear direction beneath the Juan de Fuca Plate, for
297 example, slightly reducing $\Delta\alpha$ there (Fig. 4d).

298 Model 3 adds in a 200 km thick oceanic asthenosphere that is 100
299 times weaker than the ambient mantle compared to Model 2 (cf. Becker,
300 2017). Comparing the flow fields in Models 2 and 3, we see significant
301 differences in flow pattern beneath the oceanic plates and adjacent areas
302 (Figs. 4c and e). Differential velocities, $\vec{v}_{200km}$, and the APM within the
303 oceanic region are nearly parallel in Model 2 (Fig. 4c), while in Model 3
304 they are perpendicular within the Pacific and form 40° to 60° angles within
305 the Atlantic domain (Fig. 4e). The flow modification leads to a rotation in
306 predicted anisotropy orientations from NW-SE in Model 2 to W-E in Model
307 3 (Figs. 4b and d).

308 Angular misfits $\Delta\alpha$ in Model 3 are reduced to $< 10^\circ$ in parts of

309 the western and eastern U.S. (Fig. 4f). We see misfit reduction relative
 310 to Model 2 result (Fig. 4d) through most of the study area. Here, the
 311 weak sub-oceanic asthenosphere causes flow directional change to become
 312 more APM parallel than Model 2 through the south central and south
 313 eastern U.S.. This effect, though small, can be seen from the change in
 314 shear direction and magnitude. At greater depth the flow changes to
 315 westward, so the depth-averaged shear vector and predicted fast axes
 316 orient approximately W-E instead of parallel to the plate motion. Overall,
 317 the weak asthenosphere in plate-driven flow models accommodates the
 318 lithospheric shear beneath the Pacific plate, slows down the westward
 319 sublithospheric flow motion beneath the U.S., and amplifies return flow at
 320 400 km depth (Figs. 5c and e). The misfit is overall reduced to $\langle \Delta\alpha \rangle \approx 36^\circ$
 321 for Model 3. These tests suggests that a sub-oceanic viscosity reduction, as
 322 a much larger-scale feature compared to plate boundaries and continental
 323 cratons, can have a major control over the plate-driven shear (cf. Conrad
 324 and Lithgow-Bertelloni, 2006).

325 *The effect of density driven flow*

326 We next investigate the effect of density-driven flow by adding
 327 anomalies inferred from SMEAN and MERGED tomography models to
 328 Model 2, resulting in Models 4 and 5, respectively. The direction of $\vec{v}_{200km}$
 329 in Model 4 changes nearly 180° from the western to central U.S. (Fig.
 330 6a) relative to Model 2, also clearly seen in the flow profile of Fig. 6e.

331 This flow patterns results from an upwelling underneath the western U.S.
332 and a lower-mantle, Farallon-related slab sinker anomaly. Those were
333 earlier shown to lead to strong, APM opposite counter flow underneath
334 the western half of the U.S. (e.g. Becker et al., 2006b), and are here seen to
335 be further modulated by the cratonic keel.

336 Density anomalies from SMEAN as incorporated in Model 4 result
337 in shear and predicted anisotropy fast axes oriented W-E to WSW-ENE
338 beneath the north western U.S. and west central U.S., W-E to WNW-ESE
339 beneath the east central U.S. and north eastern U.S., and SW-NE beneath
340 the south eastern U.S. (Figs. 6a and b). In these regions, the predicted
341 anisotropy fits the SWS observation nearly as well as the APM model
342 (Fig. 3), and the overall misfit is $\langle \Delta\alpha \rangle \approx 32^\circ$. This substantiates that a
343 contribution of density-induced flow to plate-driven shear is needed for
344 an appropriate prediction of LPO anisotropy, and hence a realistic mantle
345 circulation estimate, as has been argued for global models (e.g. Behn et al.,
346 2004; Becker et al., 2015).

347 The flow pattern and predicted anisotropy orientation in Model 5
348 based on MERGED are overall similar to Model 4 (Figs. 6c, d and Fig. 6f),
349 but have, expectedly, more small-scale perturbations due to the higher
350 resolution, regional tomography model of Schmandt and Lin (2014).
351 Those features include the radial flow beneath Yellowstone and Snake
352 River Plain (e.g. Savage and Sheehan, 2000), Salton Trough, northern
353 Great Valley, Rio Grande Rift, New England and central Appalachian

(Fig. 2b). Some of these smaller-scale flow structures inferred from MERGED affect the predicted anisotropy significantly. For example, beneath the northern Great Valley, which corresponds to a $\sim 5^\circ \times 5^\circ$ region with large $\Delta\alpha$ in Model 4, the dense structure that is suggested to be a lithospheric instability (Zandt et al., 2004) changes the predicted anisotropy orientation from nearly N-S in Model 4 to either SW-NE or NW-SE in Model 5, and results in a $\sim 45^\circ$ improvement in $\Delta\alpha$ values locally.

However, on balance, a degradation of the fit to SWS results on the scale of the whole U.S. is seen when the presumably better resolved MERGED tomography is used, with mean misfit increased to $\langle\Delta\alpha\rangle \approx 38^\circ$ (Fig. 6d). This means that asthenospheric flow is sensitively and diagnostically mapped into SWS predictions, but simply adding newer density models to existing flow computations at constant scaling does not provide a more consistent description of mantle dynamics. In fact, the opposite is true.

The effect of different LVVs in density and plate-driven flow models

LVVs were seen to improve the fit of purely plate-driven flow to SWS observations (cf. Figs. 4b and f). Adding density-driven flow on large scales further improved the fit to observation to a level that is comparable to the APM model (Fig. 6b), but not for the smaller-scale anomalies of MERGED (Fig. 6d). We therefore explore the other major contribution to

376 flow besides density, viscosity variations, further.

377 To complement the tests of Figs. 4 and 6 and focus on LVVs specifically,
378 we explore six additional models (Figs. 7 and 8). For the tomography
379 model SMEAN, we build a new reference, Model 6, by prescribing density
380 variations to Model 1. We then build Models 7 and 8 by prescribing LVV
381 models LVV1 and LVV2 to Model 6. Model 7 is different from Model 4
382 because it has the oceanic asthenosphere to allow for full investigation
383 of the LVVs and also to help to distinguish the effect of the craton from
384 the oceanic asthenosphere when comparing to Model 4. Similarly, for
385 tomography model MERGED, we have Model 9 in which there are no
386 LVVs, and Models 10 and 11 that use LVV1 and LVV2.

387 Without LVVs, the anisotropy predicted by Model 6 has W-E
388 orientation beneath the north western U.S. and west central U.S. and fits
389 the observed SWS regionally quite well (Fig. 7a). Other regions have very
390 large angular misfits, raising the average to $\langle \Delta\alpha \rangle \approx 50^\circ$, worse than for
391 pure plate-driven shear (Fig. 4b). Model 9 shows similar patterns (Fig. 7d),
392 besides the southeastern edge of the western U.S., for example.

393 Comparing Model 7 to 6 (Figs. 7a and b) and Model 10 to 9 (Figs. 7d
394 and e), we see that prescribing LVVs in flow models degrades the fit
395 offshore the east coast, improves the fit between predicted and observed
396 anisotropy in the central and eastern U.S., and largely modifies, although
397 does not improve, the predicted anisotropy in the western U.S. (overall
398 drop in $\langle \Delta\alpha \rangle$ is $\approx 6^\circ$ compared to no LVVs). Craton flow modification

399 (Figs. 8b and e) affects regional misfits but does not lead to an overall
400 improvement compared to the best flow model of Fig. 6b. Changing
401 the viscosity structure to the LVV2 models leads to better coupled flow
402 with the plate motion beneath the craton and thus changes the predicted
403 anisotropy orientation in the northern part of the central and eastern U.S.
404 (Figs. 7-c and f). LVV1 and LVV2 have different keel shapes (Figs. 2c
405 and d) and deflect or lead the flow differently. Indeed, comparing the
406 regional mean misfits of Fig. 7b and c, as well as e and f, we can see
407 changes particularly for the MERGED model. However, these effects of
408 anisotropy modification are not overall beneficial, and the mean misfit
409 values for models with the two viscosity structures are comparable.

410 Comparing flow profile residuals of the SMEAN flow models (Figs. 8b
411 and c), they both show better coupled sublithospheric flow velocity
412 beneath south central U.S. to the plate motion. The craton slows down the
413 eastward flow beneath the western U.S., and speeds up the westward flow
414 beneath the eastern U.S., relative to simpler viscosity models. For SMEAN,
415 LVV1 causes more perturbations on the radial direction to the flow
416 beneath it, while LVV2 mainly leads the sublithospheric flow horizontally.
417 The MERGED flow models have similar residual flow pattern overall, but
418 see more variations in magnitude and direction upon adding the cratons
419 (Figs. 8e and f), suggesting that the LVVs can amplify the density variation
420 effects.

421 In summary, we find that the effects of different assumptions on

asthenospheric density anomalies lead to the largest differences in predicted anisotropy. Yet, presumably higher resolution tomography does not improve the fit to SWS observations without additional model adjustments. Lateral viscosity variations help improve the fit when cratons and sub-oceanic viscosity reductions are introduced. Modifying the keel geometries between models LVV1 and LVV2 does improve the fit to SWS locally, but none of the modified LVV models we considered can make up for the degradation of fit observed for MERGED compared to SMEAN density anomalies.

Model Number	Viscosity Structures	Density Variations ($\delta\rho$)	Average Misfit ($^{\circ}$)
1	no LVVs	no $\delta\rho$	44.6 $^{\circ}$
2	cratons and plate boundaries in LVV1	no $\delta\rho$	45.0 $^{\circ}$
3	cratons, plate boundaries and oceanic asthenosphere in LVV1	no $\delta\rho$	35.5 $^{\circ}$
4	cratons and plate boundaries in LVV1	SMEAN	32.5 $^{\circ}$
5	cratons and plate boundaries in LVV1	MERGED	38.2 $^{\circ}$
6	no LVVs	SMEAN	49.9 $^{\circ}$
7	all structures in LVV1	SMEAN	40.5 $^{\circ}$
8	all structures in LVV2	SMEAN	40.1 $^{\circ}$
9	no LVVs	MERGED	45.1 $^{\circ}$
10	all structures in LVV1	MERGED	41.8 $^{\circ}$
11	all structures in LVV2	MERGED	40.9 $^{\circ}$

Table 1: Summary of the main information of all flow models discussed in this paper, through Model 1 to 11. Column 2 and 3 list the corresponding viscosity structure and density variation model of each flow model. Column 4 lists the average angular misfit between model predicted anisotropy and SWS observation. The surface boundary condition is prescribed APM for all flow models.

431 4. Discussion

432 *Sensitivity of mantle flow modeling*

433 We confirm that lateral viscosity variations can play an important
434 role in controlling upper mantle flow underneath continental regions
435 (e.g. Fouch et al., 2000; Miller and Becker, 2012). Plate-motion
436 induced mantle-flow model predictions of SWS observations of azimuthal
437 anisotropy are much improved when LVVs are added (Figs. 4b and f). This
438 improvement is mainly due to the implementation of a strong cratonic
439 keel and a weak oceanic asthenosphere which lead to enhancement and
440 reduction of the coupling between plate motions and sublithospheric
441 mantle, respectively (e.g. Conrad and Lithgow-Bertelloni, 2006; Becker,
442 2017). However, for purely plate-driven flow, the addition of a stiff
443 craton does not cause significant regional flow deflection in lateral or
444 radial directions, unlike what might be expected given experiments using
445 simpler geometries (e.g. Fouch et al., 2000). Moreover, the fit to SWS
446 of plate-driven flow is worse than the likely unphysical assumption of
447 alignment with APM motions.

448 In models that also include the effect of mantle density anomalies
449 for flow, in contrast, the craton amplifies the small-scale radial flow and
450 causes more significant lateral deflection and strong downward deflection
451 on scales that are relevant for regional anisotropy. In conjunction, the
452 effects of density-driven flow and lateral viscosity variations are reflected
453 in anisotropy, and SWS observations therefore do appear diagnostic of

454 both density and viscosity anomalies on scales of 100s of km.

455 SWS and flow dynamic studies have, of course, long suggested the
456 importance of density anomalies for North American plate dynamics,
457 for example related to the Juan de Fuca and Farallon slabs (e.g. Becker
458 et al., 2006b; Zandt and Humphreys, 2008), possible mantle drips (e.g.
459 West et al., 2009) and mantle upwellings (e.g. Savage and Sheehan,
460 2000). Such anomalies should be better captured by the MERGED model
461 based on regionally improved tomography, which makes it interesting
462 that the addition of smaller-scale mantle structure actually leads to
463 a worsening of the misfit between model predictions and azimuthal
464 anisotropy observations (Figs. 6b and d). This was unexpected given prior
465 successes of the general modeling approach.

466 Let us assume that structural models from seismology have in fact
467 improved thanks to USArray, and that the most fundamental assumptions
468 for our approach hold, i.e. that upper mantle anisotropy is at least partially
469 caused by LPO alignment under asthenospheric mantle flow, and that
470 mantle flow can be estimated with mantle circulation models (e.g. Long
471 and Becker, 2010). There are then several possible, not mutually exclusive,
472 reasons for why our best circulation-based model is one that is based on
473 plate-driven flow, the SMEAN large-scale mantle density anomalies, and
474 simple LVVs.

475 First, given the sensitivity of LPO predictions to details of the
476 LVVs, different keel structures, non-linear rheology, variations in volatile

477 content, or additional compositional dependence of viscosity may all
478 lead to lateral viscosity variations that counterbalance the detrimental
479 effects of adding small-scale density structure of MERGED. A formal
480 inversion for these variations is possible, but none of our forward tests
481 (most not shown) trying different LVV structures have led to plate-scale
482 improvement in mean misfit. Figure 7 illustrates the sorts of variations
483 in LPO predictions one might expect. These effects are in line with
484 arguments about local effects, e.g. of drips and the like, but we leave
485 the exploration of more complex mantle LVV models that could possibly
486 reconcile the predictions for later. The general applicability of such
487 optimized models will also be questionable should the LVVs not be based
488 on some additional, general physical relationship not explored here.

489 Second, our scaling between seismic tomography and density
490 anomalies might be wrong, and this is clearly the case in principle,
491 given the highly simplified nature of our linear, depth-independent
492 scaling. Besides temperature, other properties, especially compositional
493 heterogeneity and anelasticity, can also affect seismic wave velocity (e.g.
494 Forte and Perry, 2000; Cammarano et al., 2003). This might be of
495 particular importance for the high resolution tomography model, which
496 might demand lateral variations in the scaling factor. We expect that
497 cratonic regions of the continental lithosphere may be neutrally buoyant
498 (“isopycnic”, Jordan, 1978) which is why we corrected for this effect in a
499 coarse fashion in our mantle flow models. The isopycnic assumption is not

500 expected to be perfectly true at all depths, nor is the extent of cratons or the
501 thickness of the lithosphere well constrained (e.g. Lekić and Fischer, 2014;
502 Steinberger and Becker, 2016). We therefore cannot rule out that more
503 sophisticated models including a wider range of compositional anomalies
504 would lead to better predictions of LPO based anisotropy using the high
505 resolution tomography models such as MERGED.

506 However, we conducted a range of tests where we varied the R scaling
507 step wise from zero to its reference value, and found that the signal
508 inherent in MERGED leads to a degradation of the fit compared to SMEAN
509 as soon as the density effects are felt by mantle flow. This implies that
510 compositional anomalies would have to cancel out much of the signal seen
511 in MERGED compared to SMEAN to at least not degrade the fit. This is
512 possible, but would also question the general interpretations of seismic
513 tomography for regional tectonics.

514 Third, time-dependence of mantle convection, and in particular
515 changes in plate motions, may complicate the interpretation of LPO based
516 anisotropy even for the relatively short time-scales needed to saturate
517 fabrics within the asthenosphere (e.g. Kaminski et al., 2004; Becker et al.,
518 2006a). On global scales, Becker et al. (2003) showed that this effect was
519 detectable, but seismological models did not allow determining which
520 models were better within uncertainty.

521 Regionally, the story may be different, and Zhou et al. (2018) explored
522 such effects for the western U.S. in detail. The authors pointed out the

523 importance of the Juan de Fuca slab and a hot mantle anomaly beneath the
524 western U.S. for the formation of the circular anisotropy beneath the Great
525 Basin. However, the anisotropy adjacent to that pattern was not well fit,
526 implying similarly mixed results in terms of a comprehensive explanation
527 of SWS observations.

528 There are thus at least three plausible reasons why a purely
529 asthenospheric origin of anisotropy appears to be a moderately successful
530 explanation of the large-scale SWS signal for the U.S. at best. In the
531 remainder, we will instead assume, for the sake of argument, that
532 our computations are in fact very good predictions of asthenospheric
533 anisotropy, so good that we can ask about a missing lithospheric
534 component needed to fit SWS observations.

535 *The lithospheric complement*

536 A lithospheric, frozen-in origin of anisotropy has long been discussed
537 for the shallow oceanic lithosphere, as well as the bulk of the thicker
538 and petrologically more heterogeneous continental lithosphere (e.g. Silver,
539 1996). Assuming that the difference between the SWS observations and
540 flow predictions of LPO anisotropy arises entirely from the lithospheric
541 component, we can augment a flow model with its corresponding
542 lithospheric complement that would be needed to achieve a (near) perfect
543 fit to *SKS* splitting.

544 Figure 9 shows results for the lithospheric complements for the

545 best performing LPO based on flow models, Model 4 (SMEAN) and 5
 546 (MERGED). The lithospheric complement is found by fitting individual
 547 splits from Liu et al. (2014) with a two-layer model, in which the bottom
 548 layer is fixed to the flow predicted anisotropy. The values of the apparent
 549 splitting parameters from the hypothetical two-layer anisotropy and the
 550 average of the *SKS* splits are similar (Fig. S1), with angular difference
 551 of $\Delta\alpha \lesssim 5^\circ$, which would be within the typical "error" of SWS estimates.
 552 For stations where the flow predicted anisotropy has similar orientation
 553 to the bottom layer from an independent two-layer inversion of *SKS*
 554 splits, the hypothetical lithospheric anisotropy is also similar to the top
 555 layer from the independent inversion. This suggests the validity of this
 556 approach for studying multi-layer anisotropy. Besides the two layer
 557 parameter space exploration approach, we also explore a simple method
 558 of matching SWS by inverting for the best-fit thickness and anisotropy
 559 orientation of a lithospheric layer that consists of frozen-in anisotropy
 560 represented by a single elastic tensor (supp. mat.). Using this method, the
 561 inferred lithospheric complement has similar orientation with our current
 562 approach, but the delay times are less realistic (Fig. S2). We leave the
 563 exploration of back-azimuthal dependence of *SKS* splitting for a future
 564 joint analysis with surface-wave depth-dependent anisotropy.

565 As Figs. 9a and b show, the patterns of the best fit lithospheric
 566 complement are fairly smooth over much of the study area. This might be
 567 expected from the spatial heterogeneity of SWS and seismic tomography,

568 but implies that there could be a relation with a deterministic tectonic or
569 convective process. The lithospheric complement is different for the two
570 flow models in detail, but there are also consistent features. That said, the
571 connection of the lithospheric complement's azimuthal alignment patterns
572 to geological history is not immediately apparent, at least to us.

573 However, we can check if the features of the complement are at least
574 consistent with other possibly related observations. To this end, we
575 visually compare the complements with an azimuthal anisotropy model
576 inferred from 16 s period Rayleigh waves by Lin and Schmandt (2014)
577 (Fig. 9e). While mainly sensitive to the uppermost crust, the anisotropy
578 orientations appear related to tectonic regions, such as the Great Basin,
579 the Rockies and the Precambrian Rift Margin (Lin and Schmandt, 2014).
580 Without going to details of the relationship between crustal anisotropy
581 and tectonics, we note that there are fairly good correlations in orientations
582 between our lithospheric complement and the crustal anisotropy along the
583 west coast of the U.S., beneath the Columbia Plateau, the southern Basin
584 and Range, south of the Colorado plateau, Texas and the southern Coastal
585 Plain (Figs. 9a and b). Beneath the eastern U.S., Model 5's lithospheric
586 complement matches the crustal model while Model 4's does not.

587 To expand this comparison to the uppermost mantle, we further
588 compare the Model 5 lithospheric complement with the Pn anisotropy
589 model by Buehler and Shearer (2017) (Figs. 9f). This model provides
590 information beneath the Moho. In this model, the NE-SW oriented

591 orogeny parallel anisotropy beneath the Appalachian mountain and east
592 central U.S. only exist in the central region (Fig. 9f). In other regions
593 of the eastern U.S., the anisotropy is E-W, which might relate to plate
594 motion (Buehler and Shearer, 2017). If this is the case, we would expect
595 orogeny parallel anisotropy at shallow depths, and more plate motion
596 parallel anisotropy beneath. This is true when we look at the Model 5
597 results, where the flow model predicted anisotropy parallels the plate
598 motion (Fig. 6d), and the lithospheric complement parallels the orogeny
599 (Fig. 9b). Since the *SKS* splits have a more dominant orogeny parallel
600 pattern compared to the uppermost mantle anisotropy, there might be
601 a significant crustal contribution in the SWS observation at the eastern
602 and east-central U.S., which partly explains the misfit we observed when
603 comparing the flow predicted anisotropy to SWS in this region.

604 To investigate the anisotropy at different depths in the eastern and
605 east-central U.S., we compare our lithospheric complement with the
606 regional model by Deschamps et al. (2008) (Fig. 9g). The lithospheric
607 complement of Model 5 (Fig. 9c) has similar patterns with the Rayleigh
608 wave anisotropy at periods < 60 s, which approximately shows the
609 lithosphere. The longer period (160 s) Rayleigh wave anisotropy, however,
610 does not match the lithospheric complement, but matches the flow
611 predicted anisotropy in the same region (Fig. 9d). This depth constraint
612 of anisotropy further suggests that the actual lithospheric anisotropy is
613 reasonably estimated by the lithospheric complement, and the lithosphere

614 has notable contribution to the SWS observation, at least in the eastern and
615 east-central U.S..

616 Good correlation between lithospheric complement and the
617 lithospheric anisotropy, and between flow model prediction and the
618 sublithospheric anisotropy in the eastern U.S. indicate the possibility that
619 MERGED predicts the sublithospheric anisotropy better than SMEAN
620 even if the asthenospheric LPO alone leads to a poor fit. This substantiates
621 importance of understanding lithospheric anisotropy, and may help to
622 resolve the connection between small-scale mantle structures and the
623 upper mantle anisotropy formation and SWS observation.

624 New insights into continental dynamics may yet be revealed by
625 modeling anisotropy due to mantle flow. However, the answer might
626 at least regionally have to involve more detailed study of the lithosphere
627 and longer-term geological history. Such future work should be especially
628 promising once noise and ballistic surface wave inferences for crustal and
629 mantle anisotropy are adequately incorporated.

630 **5. Conclusions**

631 Azimuthal anisotropy in the upper mantle as seen by shear wave
632 splitting throughout the U.S. and offshore portions of the North American
633 plate can be modeled by mantle circulation models. These models
634 allow exploring the effect of density anomalies and viscosity variations
635 within the asthenosphere, which strongly affect predictions when acting

636 together. Large-scale flow models lead to misfits that are comparable
637 to the absolute plate motion alignment hypothesis for the study region.
638 This confirms the general validity of the approach, but smaller-scale
639 density anomalies of modern, EarthScope era tomography degrade the
640 fit, and none of the viscosity models we considered can make up for
641 it. “Lithospheric complements” can be estimated from the best flow
642 model based anisotropy, and those match independent estimates of crustal
643 anisotropy. This indicates promising avenues forward, but much is still to
644 be learned about the link between seismic anisotropy and mantle flow and
645 continental dynamics.

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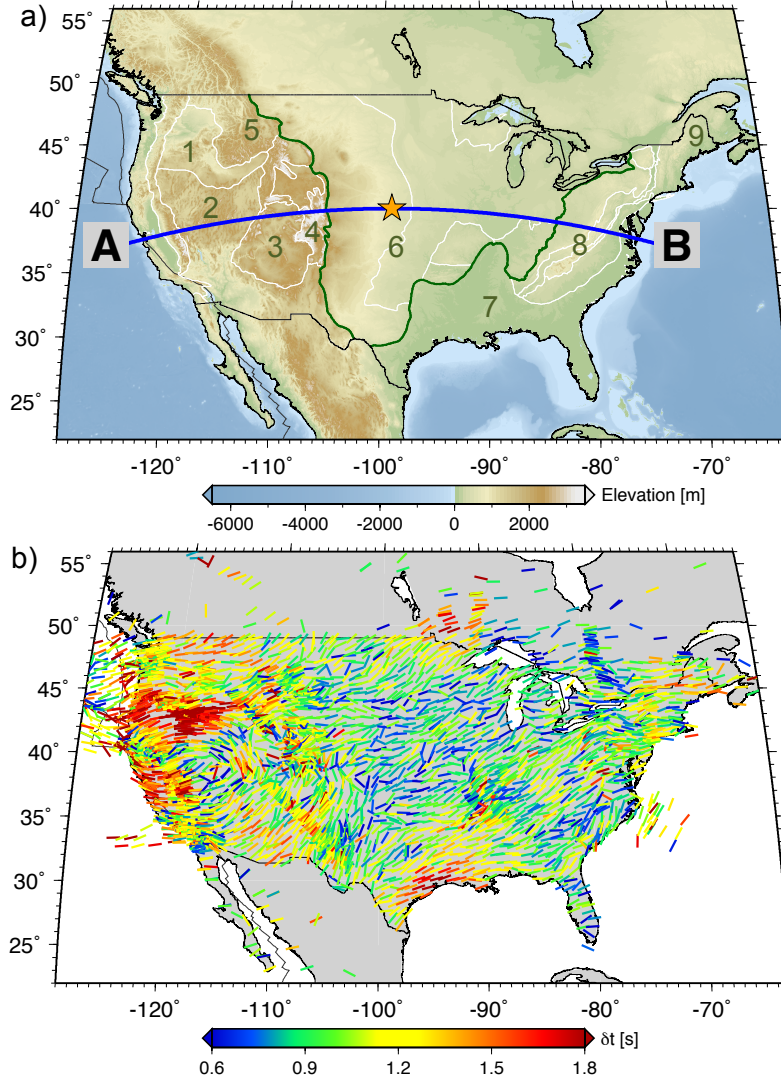


Figure 1: Study area showing topography and physio-graphic regions (a) and station-averaged shear-wave splitting measurements (b). In a), elevation is shown in the background; green lines are the orographic boundaries here used to define the western, central and eastern U.S.; white lines are the boundaries of different physio-graphic regions; blue profile shows the location of the cross section of the flow profiles discussed below. Main physio-graphic regions that are discussed in this paper are marked with numbers, they are: 1) Columbia Plateau, 2) Basin and Range, 3) Colorado Plateau, 4) Southern Rocky Mountains, 5) Northern Rocky Mountains, 6) Interior Plains, 7) Coastal Plains, 8) Appalachian Mountain Range, and 9) New England province. In b), fast orientation and delay times (δt) of the SWS measurement compilation are shown by stick orientation and color, respectively.

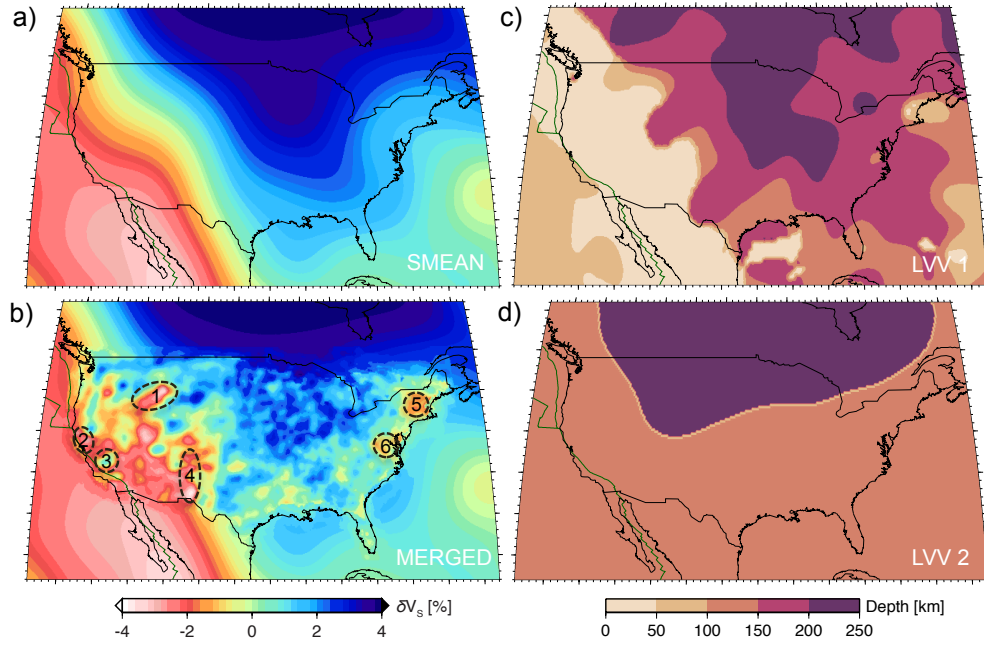


Figure 2: Tomography (a and b) and lithospheric thickness (c and d) models. Colors in a) and b) indicate S wave velocity anomalies (δV_s) for the SMEAN (Becker and Boschi, 2002) and MERGED (cf. Schmandt and Lin, 2014) tomography models, respectively, at 200 km depth. In plot b), the outlined features are upper mantle anomalies that are discussed in the result section, they are: 1). Yellowstone and Snake River Plain, 2). Salton Trough, 3). northern Great Valley, 4). Rio Grande Rift, 5). New England, and 6) central Appalachian. Colors in c) and d) show the inferred depth of the lithosphere in viscosity models LVV1 and LVV2, respectively.

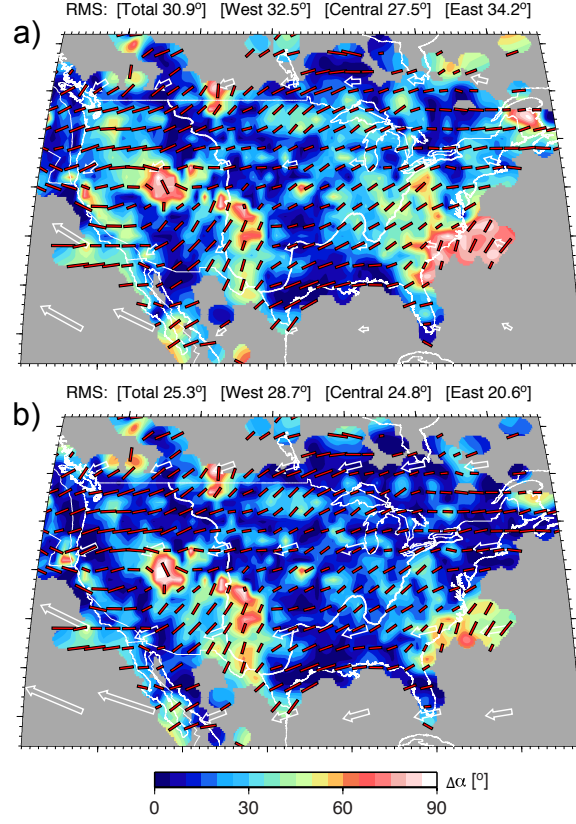


Figure 3: Absolute angular misfit ($\Delta\alpha$) between *SKS* splits and absolute plate motion (APM) orientations in the no-net-rotation (NNR) reference frame of Argus and Gordon (1991) (a), and in the spreading-aligned reference frame of Becker et al. (2015) (b). $2^\circ \times 2^\circ$ grid averaged *SKS* splits (based on Fig. 1b) are shown by red sticks. APM motions are indicated by white, open vectors. Background color indicates the value of $\Delta\alpha$, and title shows the map-wide and regional means of $\Delta\alpha$ for the sub domains indicated by heavy white lines, $\langle\Delta\alpha\rangle$.

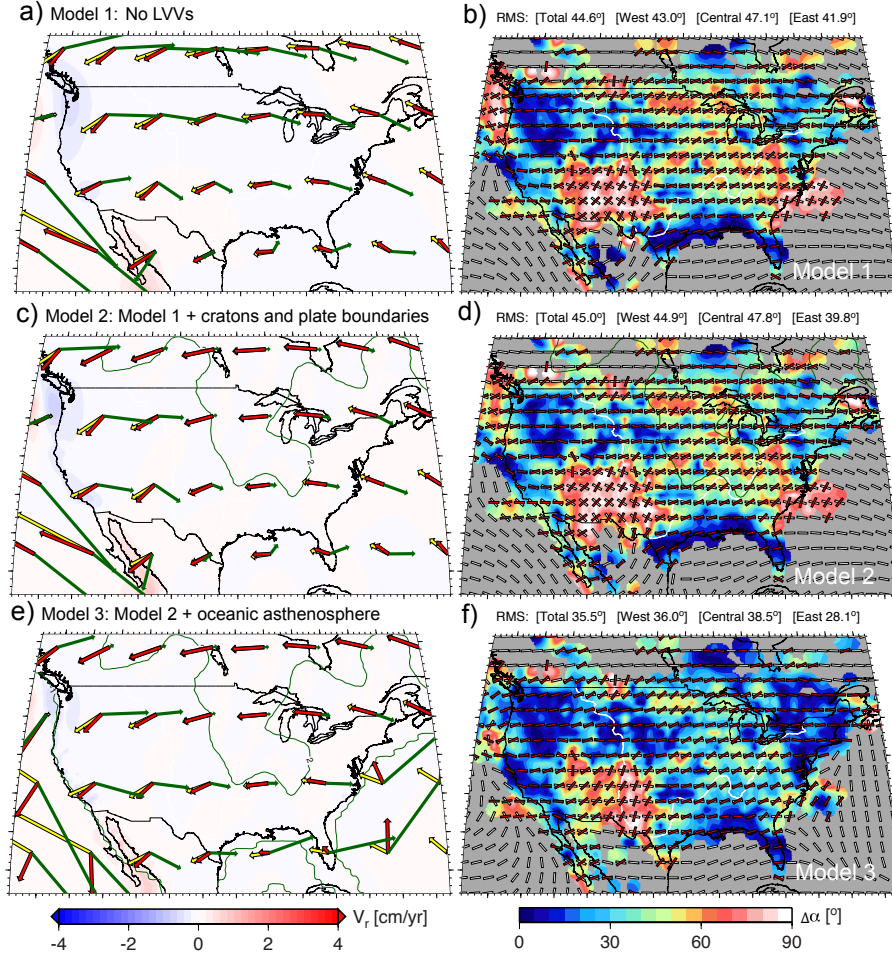


Figure 4: Upper mantle flow (a, c, and e) and the resulting angular misfit ($\Delta\alpha$) between *SKS* splits and flow-model predicted anisotropy (b, d, and f) of Models 1 (plate-induced shear, a and b), 2 (added cratons and weak zones, c and d), and 3 (added oceanic weak asthenosphere, e and f). In plot a), c), and e), radial flow is shown in background coloring (upwelling positive); surface velocities, flow at 200 km depth, and their vector difference (amplified by 5) are indicated by yellow, red and green vectors, respectively. In plot b), d), and f), $\Delta\alpha$ is shown in the background; SWS observed and flow model predicted SWS fast orientations depicted by red and white vectors, respectively. Title for b), d), and f) shows mean angular misfits as in Fig. 2.

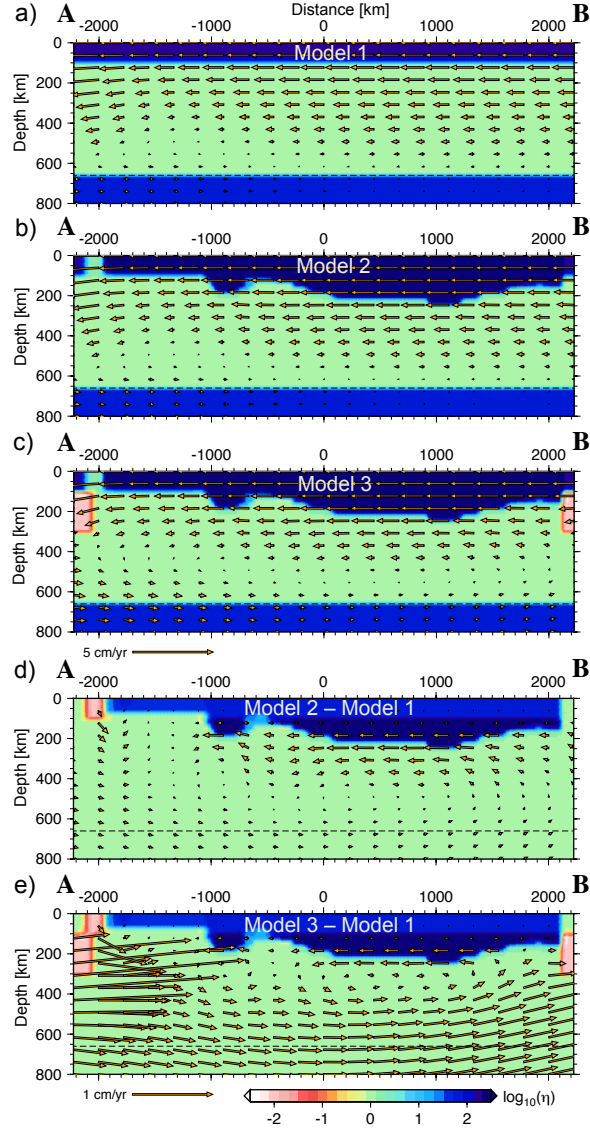


Figure 5: Cross-section of mantle flow for Models 1 (a), 2 (b), and 3 (c) along the profile shown in Fig. 1a. Background color shows the decadic logarithm of the upper mantle normalized viscosity, and orange vectors show flow velocity with the length scale shown beneath the bottom left corner of plot c). Sub-plots d) and e) show the differences in flow field and viscosity between Models 1 and 2 (d), and between Models 1 and 3 (e). Length scale of the differential flow vectors is shown beneath plot e).

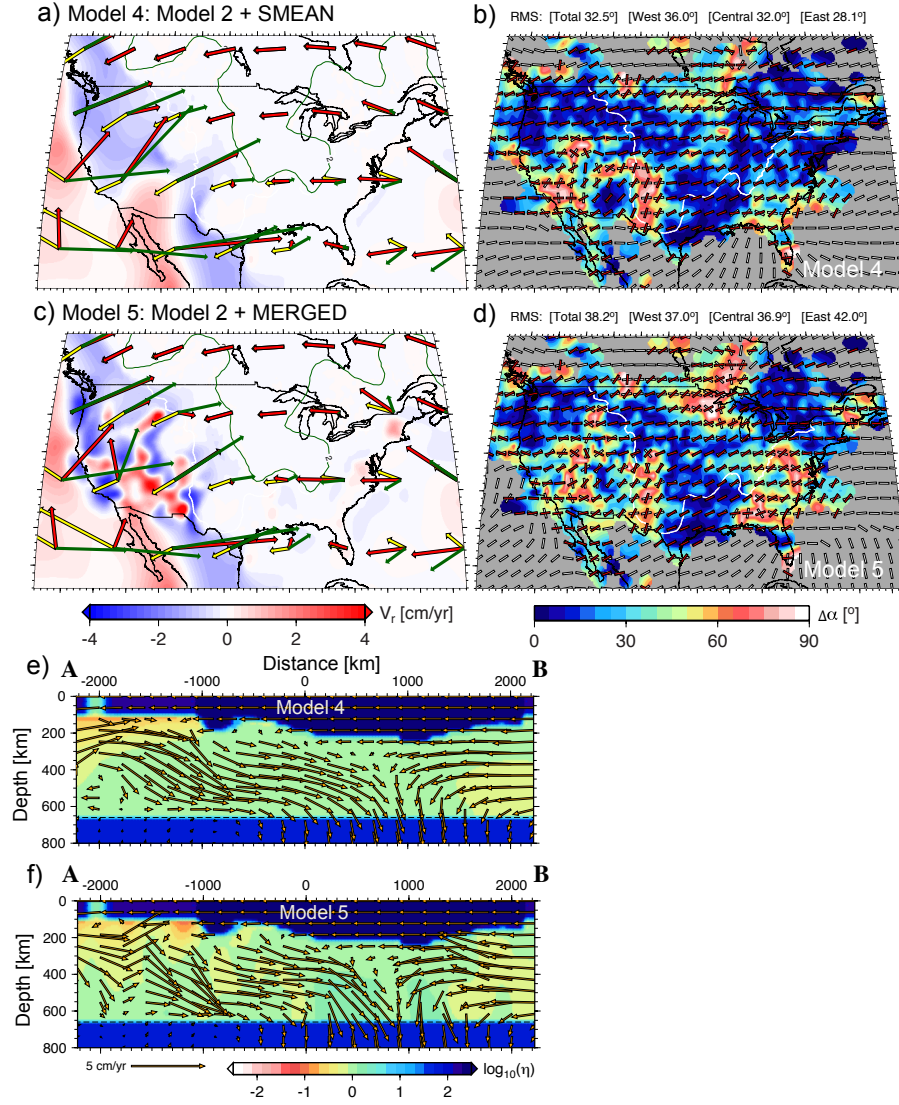


Figure 6: Effect of density anomalies. Flow field at 200 km depth (a and c), the resulting $\Delta\alpha$ between *SKS* splits and flow predicted anisotropy (b and d) of Model 4 (SMEAN density driven flow, a and b) and 5 (MERGED density, c and d), and velocity and viscosity profiles for Models 4 (e) and 5 (f). Scale of velocity vector length is shown beneath plot f). See Fig. 4 and Fig. 5 for details.

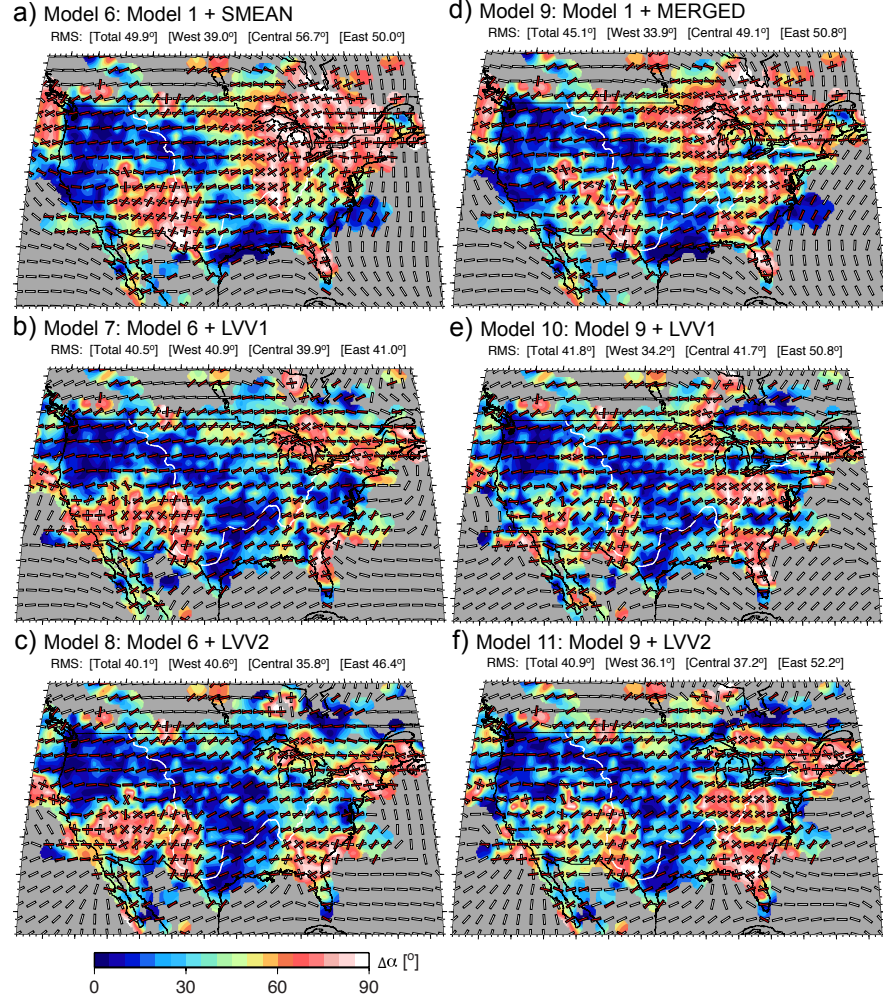


Figure 7: Effect of lateral viscosity variations. $\Delta\alpha$ between *SKS* splits and flow-model predicted anisotropy of Models 6 (a), 7 (b), 8 (c), 9 (d), 10 (e), and 11 (f). See Fig. 4 for details.

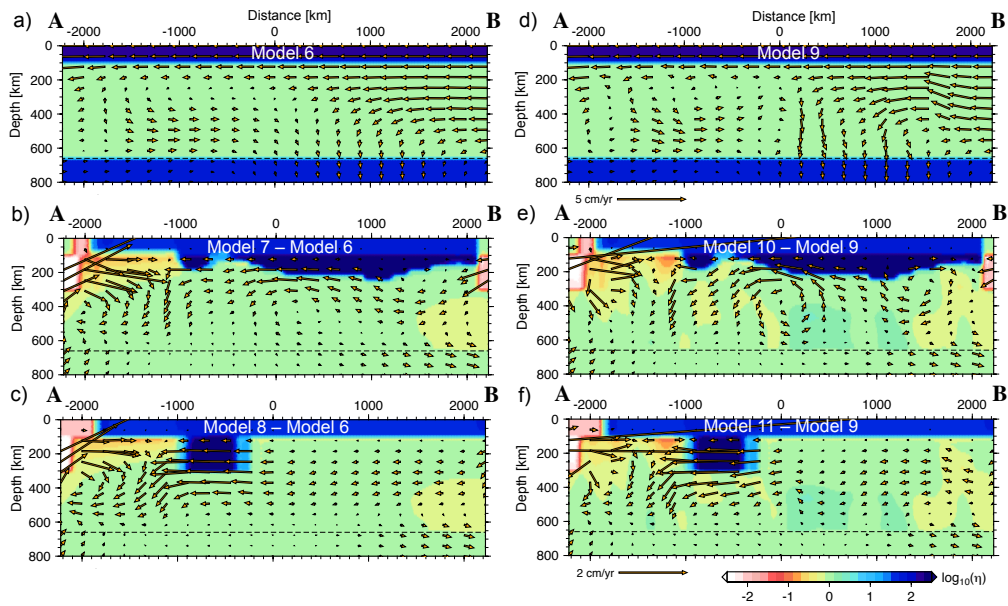


Figure 8: Flow profiles of Models 6 (a) and 9 (f) with velocity vector length scale shown beneath plot f). Plots c), e), h) and j) show the differences in velocities and viscosity between Models 7 and 6 (c), Models 8 and 6 (e), Models 10 and 9 (h), and Models 11 and 9 (f). Differential velocity vector length scale is shown beneath plot j). See Fig. 5 for details.

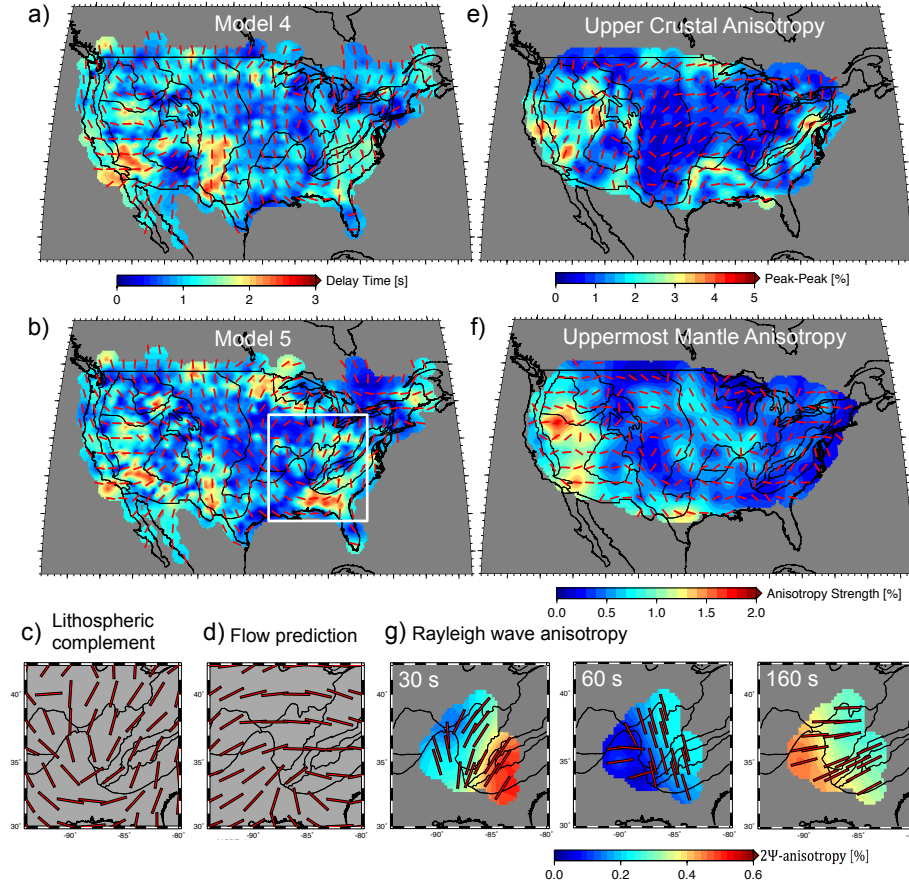


Figure 9: Inferred lithospheric anisotropy resulting from matching Model 4 (a) and 5 (b) flow-predicted anisotropy to the SWS dataset. c) shows the Model 5 lithospheric anisotropy orientation within the white rectangular box in b), and d) shows the Model 5 flow predicted anisotropy within the same region. e) shows crustal anisotropy from Lin and Schmandt (2014). f) shows the uppermost mantle anisotropy from Buehler and Shearer (2017). g) shows regional Rayleigh wave anisotropy model from Deschamps et al. (2008), with each of the subplots showing different Rayleigh wave periods as noted on the upper left corner. In all plots, the red vector shows fast orientation. The background color shows the delay time in a) and b), shows the peak to peak amplitude in e), anisotropy magnitude in f), and shows 2ψ anisotropy magnitude in g).

Supplementary Material

Wanying Wang, Thorsten W. Becker

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1 Here, we discuss additional material for the lithospheric complement
2 computation discussed in the main text. Whenever we estimate average
3 asthenospheric anisotropy, we compute an arithmetic average of tensors
4 and then use the Christoffel matrix approach and that mean tensor to get
5 average SKS. To match average splitting, we align and scale a lithospheric
6 tensor before averaging by means of a parameter space search. When
7 accounting for back-azimuth dependence, we use a parameter space
8 search and Silver and Savage's (1994) approach.

9 To illustrate that our approach of inverting for best-fit lithospheric
10 anisotropy layer gives reliable results, we picked 11 stations throughout
11 the study area, and show in Fig. S1 that the apparent splitting parameters
12 from the resulting hypothetical two-layer anisotropy are similar to the
13 station averaged splitting parameters from SWS. Moreover, the two layer
14 model from the constrained inversion where we fix the asthenospheric
15 (lower) layer to that expected from an average of our flow model predicted

16 LPO is usually similar to that of a station measurement only based two
17 layer model.

18 We do not expect the resulting lithospheric complement to be exactly
19 the same as the top layer from SWS two-layer inversion. However, when
20 the flow predicted anisotropy has similar orientation with the bottom
21 layer from the two-layer inversion of SWS, the hypothetical lithospheric
22 anisotropy is also similar to the top layer of the SWS two-layer inversion
23 (Fig. S1), suggesting the validity of this approach.

24 Figs. S2a and b show a comparison between the SWS top layer
25 anisotropy and our hypothetical lithospheric complement. The SWS top
26 layer anisotropy has a wider range of azimuthal orientations and delay
27 times (Fig. S2a), while the hypothetical lithospheric complement shows
28 more smooth and regionally consistent anisotropy patterns, and less
29 variation in delay times. Although differences exist, we see consistency
30 between the two top layer anisotropy estimations in the central U.S. where
31 there is relatively small delay time, in the Great Basin where there is the
32 near circular anisotropy pattern, in the southern Basin and Range where
33 there is the NE-SW anisotropy pattern, and in the eastern U.S. where there
34 is the Appalachian Mountain parallel anisotropy pattern.

35 We also explore a simple method of matching the SWS observation
36 by inverting for the best-fit thickness and anisotropy orientation of a
37 lithospheric layer that consists of frozen-in anisotropy represented by a
38 single elastic tensor. The tensor used here is an averaged single-crystal

39 tensor with 70% olivine and 30% pyroxene from Estey and Douglas
40 (1986). Instead of using a two layer fit and accounting for back-azimuth
41 dependence of splits, we then only seek to match the average fast
42 axes by means of a simple averaging approach. Using this simple
43 method, the resulting hypothetical lithospheric anisotropy show similar
44 orientations to our two layer inversion approach, but the delay time is
45 overall larger (Fig. S2b and c). With the result from this, we can try fit
46 the two-layer anisotropy from this simple method for $\pi/2$ backazimuth
47 distribution, and compare it to the fit from our two layer inversion
48 approach of computing the lithospheric complement. In Fig. S1, the
49 current lithospheric complement approach (solid curve) fits well with the
50 *SKS* splits, while the simple method (dashed curve) fitting is off at some
51 stations.

52 **References**

- 53 Estey, L. H., Douglas, B. J., 1986. Upper mantle anisotropy: A preliminary
54 model. *J. Geophys. Res.* 91 (B11), 11393.
- 55 Silver, P. G., Savage, M. K., 1994. The Interpretation of Shear-Wave
56 Splitting Parameters In the Presence of Two Anisotropic Layers.
57 *Geophys. J. Int.* 119 (3), 949–963.

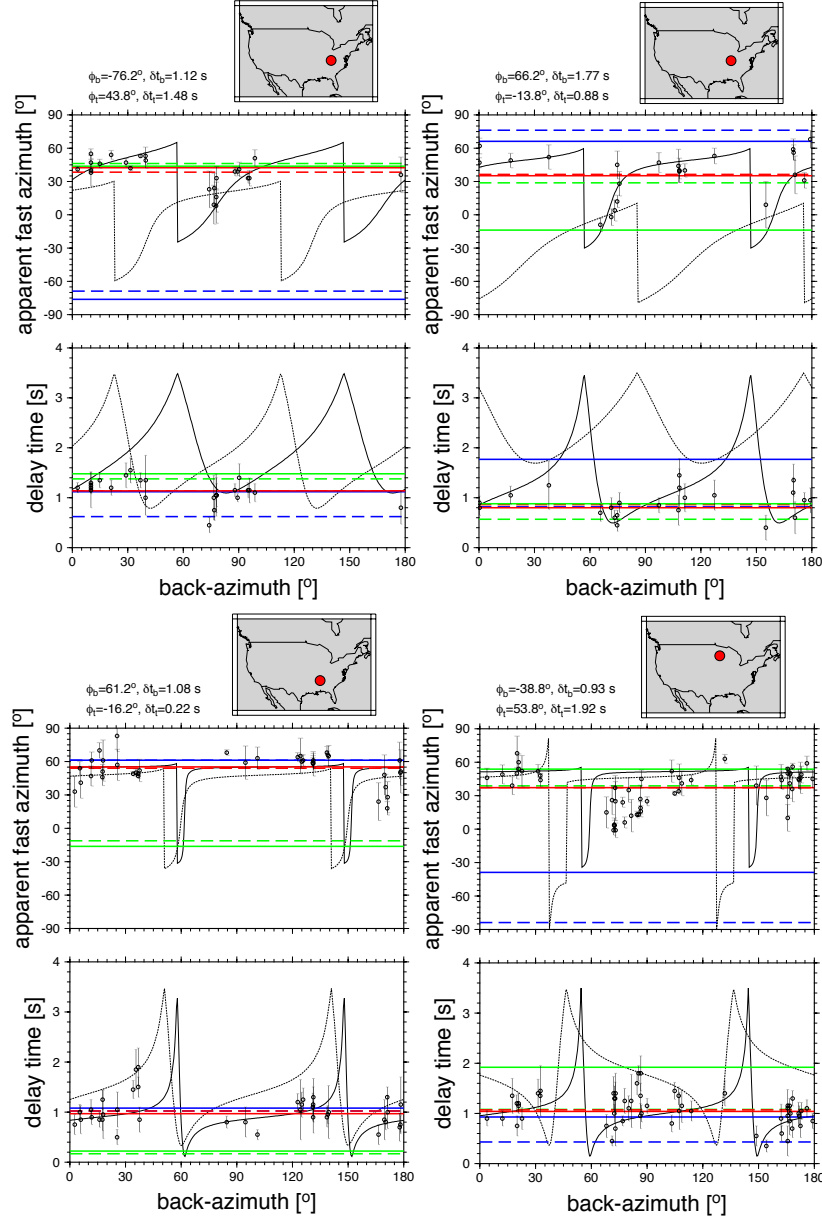


Figure S1: Examples of statistics at stations showing the shear wave splitting (SWS) two-layer inversion results, and the flow model (Model 5) hypothetical two layer anisotropy (top layer: flow model predicted anisotropy, and bottom layer: the hypothetical lithospheric complement). Caption continues on next page.

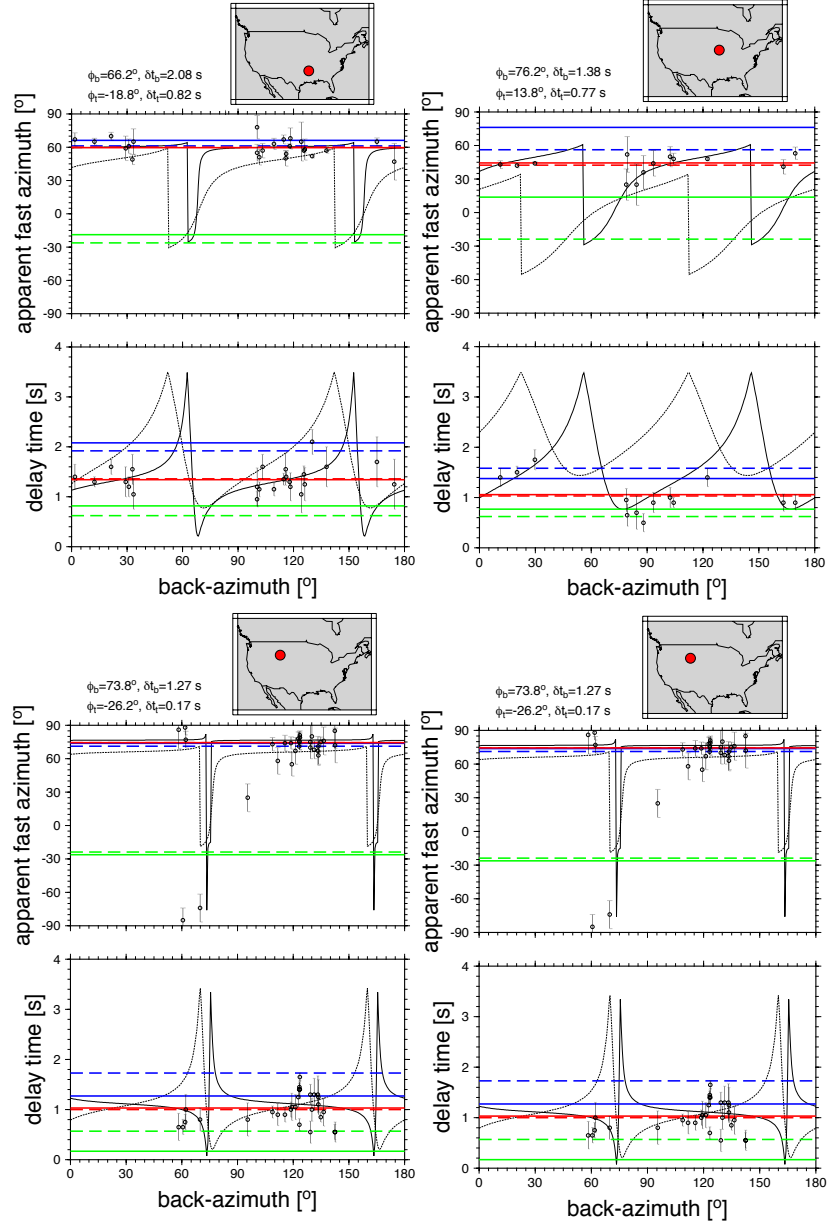


Figure S1: Continues from last page. The average values of the SWS top layer, bottom layer, and station average are shown by the solid blue line, solid green line, and solid red line, respectively. The average values of the top layer, bottom layer, and station average from the hypothetical two layer anisotropy are shown by the dashed blue line, dashed green line, and dashed red line, respectively. Caption continues on next page.

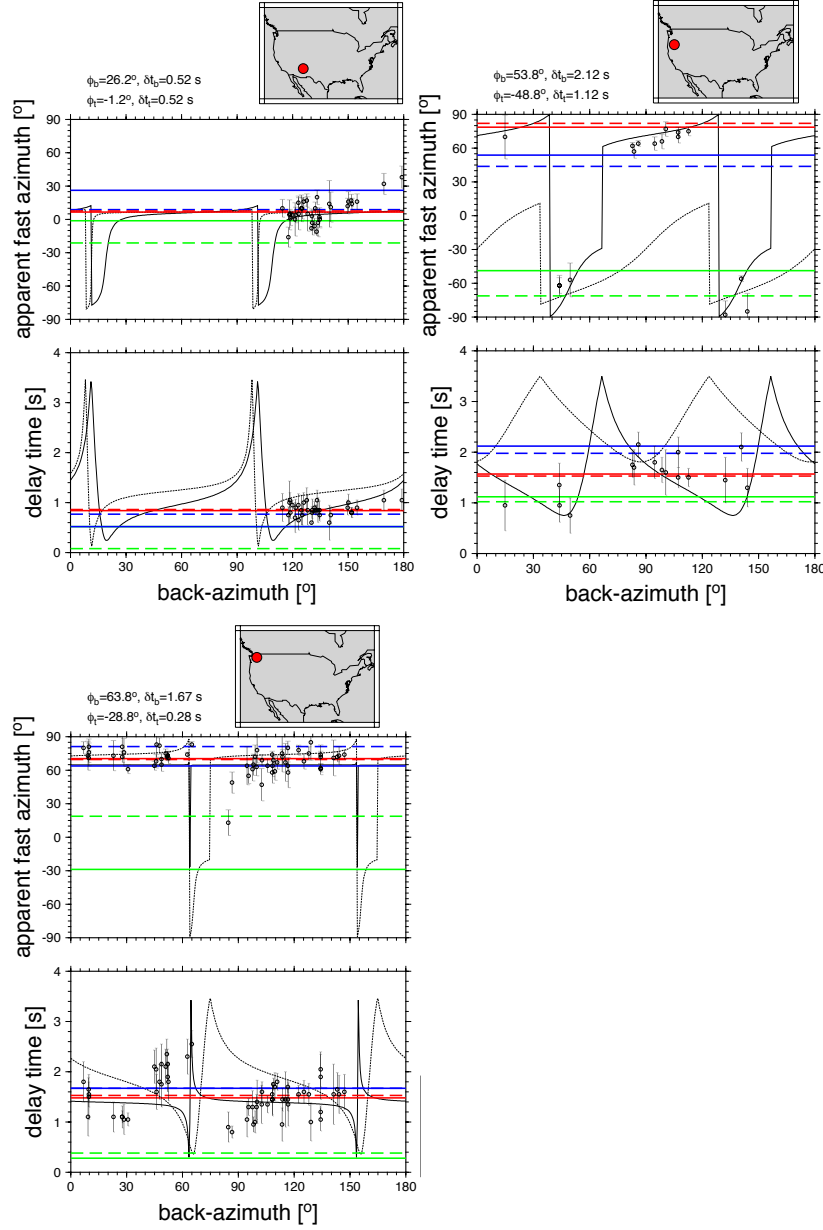


Figure S1: Continues from last page. Splitting parameters from SWS are shown by the black circles, and the vertical lines through the circles show estimated errors. The solid black curve shows the two-layer fitting of the hypothetical two layer anisotropy. The dashed black curve shows the two layer fitting for the simple method of computing lithospheric complement. The location of each station is shown by red dot on the map.

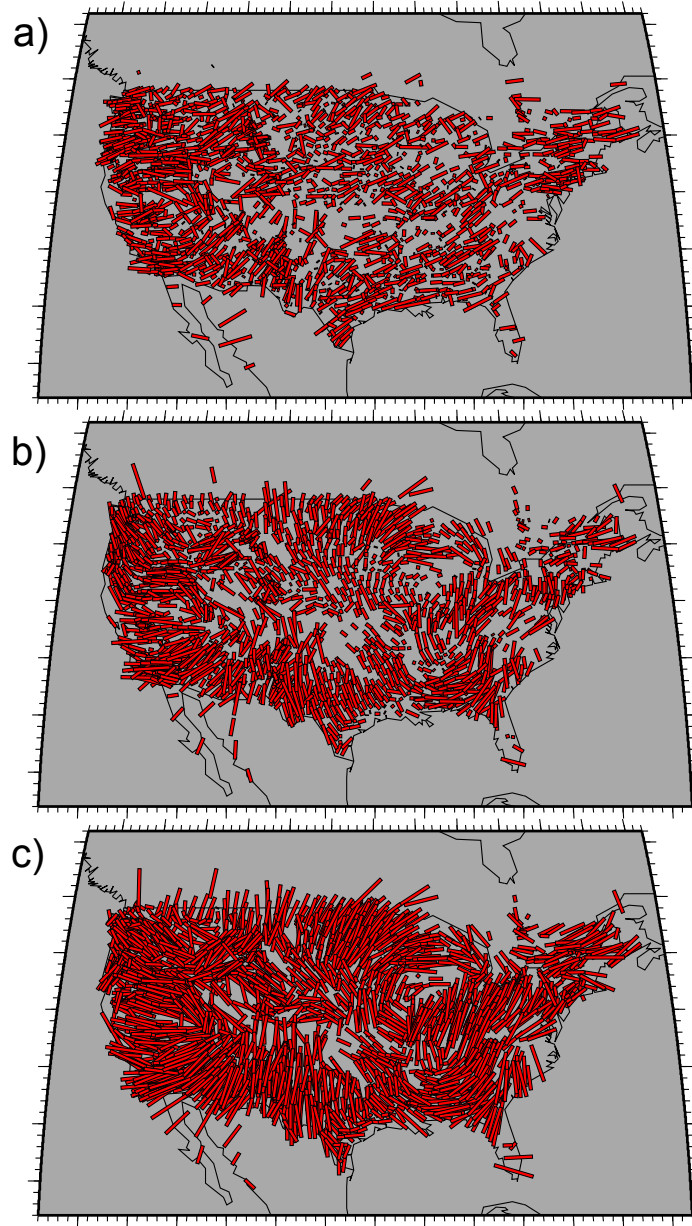


Figure S2: The top-layer anisotropy from a) station by station SWS two-layer inversion, b) method used for the main text of computing the lithospheric complement, and c) the simple method of computing lithospheric complement based on modeling average angles.