Geophysical evidence for crustal and mantle weak zones controlling intra-plate seismicity - the 2017 Botswana earthquake sequence

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Abstract

Large earthquakes away from plate boundaries pose a significant threat to human lives and infrastructure, but such events typically occur on previously unknown faults. Most cases of intra-plate seismicity result from compression related to farfield plate boundary stresses. The April 2017 Mw 6.5 earthquake in central Botswana, and subsequent events, occurred in a region with no previously known large earthquakes, occurred away from major present day tectonic activity, and accommodate extension rather than compression. Here, we present results from an integrated geophysical study that suggests the recent earthquakes may be a sign of future activity, controlled by the collocation of a weak upper mantle and weak crustal structure, between otherwise strong Precambrian blocks. Magnetotelluric data highlights Proterozoic continent accretion structure within the region, and shows that recent extension and seismicity occurred along ancient thrust faults within the crust. Our seismic velocity and resistivity models suggest a weak zone in the uppermost mantle, that does not persist to greater depths, and is therefore unlikely to represent mantle upwelling. The Botswana events may therefore be indicative of top-down extension as a response to large scale extensional forces.

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12 Abstract

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Large earthquakes away from plate boundaries pose a significant threat to human lives and infrastructure, but such events typically occur on previously unknown faults. Most cases of intra-plate seismicity result from compression related to far-field plate boundary stresses. The April 2017 M_W 6.5 earthquake in central Botswana, and subsequent events, occurred in a region with no previously known large earthquakes, occurred away from major present day tectonic activity, and accommodate extension rather than compression. Here, we present results from an integrated geophysical study that suggests the recent earthquakes may be a sign of future activity, controlled by the collocation of a weak upper mantle and weak crustal structure, between otherwise strong Precambrian blocks. Magnetotelluric data highlights Proterozoic continent accretion structure within the region, and shows that recent extension and seismicity occurred along ancient thrust faults within the crust. Our seismic velocity and resistivity models suggest a weak zone in the uppermost mantle, that does not persist to greater depths, and is therefore unlikely to represent mantle upwelling. The Botswana events may therefore be indicative of top-down extension as a response to large scale extensional forces.

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13 1. Introduction

The 3rd April 2017 Botswana earthquake (moment magnitude, M_W 6.5) 14 was the largest event on the African continent outside the East African Rift 15 System (EARS) for over 80 years (cf [1]) (Figure 1). It was part of a sequence 16 of 15 events with magnitudes up to M_W 5. That sequence lasted for 4 months 17 following the main event, with the final event occurring 200 km away on 12th 18 August 2017. Intra-plate earthquakes require sufficient stress to build-up, with 19 most events attributed to far-field effects of deformation at plate boundaries [2]. 20 Given that stable continental lithosphere is rigid and strong, these stresses can 21 be transferred over long distances [3]. Such a model of earthquake generation 22 is compatible with thrusting or strike-slip mechanisms and general horizontal 23 contraction, which is observed in most intra-plate events [4]. 24

The Botswana event is different in this respect as its focal mechanism indi-25 cates normal faulting and extension. However, such a mechanism is consistent 26 with stress and stress-gradient calculations for southern Africa [5, 6] which indi-27 cate large scale extension. Controls on the precise locations of intra-plate events 28 remain debated. Tectonic controls, such as gradients in lithospheric thickness 29 and the presence of weak zones may facilitate movement [7]. For example, 30 earthquakes in the New Madrid Seismic Zone, USA, are thought to be asso-31 ciated with pre-existing faults and possibly a weak mantle below [8]. Recent 32 work suggests that these large scale factors determine the style of faulting and 33 the stress release, while transient events, possibly due to non-tectonic forces, 34 are responsible for triggering the earthquake [9]. For the Botswana earthquake 35 it has been hypothesized that fluid leaks from the upper mantle have triggered 36 the event [10]. 37

Although Botswana does not show any strong earthquakes in instrumented history, the clustering of small magnitude events in the Okavango delta in northern Botswana has led to speculation of an incipient rift [13] - the Okavango Rift Zone (ORZ) [14] - which may represent the southwestern continuation of the EARS [15]. Most of the current understanding of the deeper crustal and upper



Figure 1: Map of southern and central Africa. The study area is shown with a blue rectangle and we show the location and focal mechanism of the 03/04/2017 earthquake. Black lines show plate boundaries whereas the dashed line indicates the proposed south-western continuation of the East African Rift System (EARS) to the Okavango Rift Zone (ORZ). Colored boxes show all seismic events in the USGS earthquake catalogue with moment magnitude > 5, where color and size indicate magnitude. Black bars indicate maximum horizontal stress direction from the World Stress Map[11]. Blue arrows show inferred plate motion with respect to the Nubian plate[12] and the yellow star marks the Euler pole for the Somalia-Nubia plate motion. The locations of the SAMTEX MT sites are shown in grey and red circles, where the red ones are modelled in this paper.

⁴³ mantle structures in the region is based on a profile of magnetotelluric (MT)
⁴⁴ data to the east [16], and inversion of receiver functions [17], shear wave split⁴⁵ ting analysis [18], and seismic tomography [15] along the profile of the SAFARI
⁴⁶ experiment. Potential field data have been used to identify the boundaries be⁴⁷ tween different tectonic units [14, 19], but significant uncertainty remains as to
⁴⁸ the location and nature of those units (compare [14, 16, 18, 20, 21]).

49 2. Inverting magnetotelluric and surface wave data

Here we present 3D models of the lithosphere south of the ORZ and centered 50 on the April 2017 event, based on MT measurements (e.g. [22]) and regional 51 surface wave data. We use magnetotelluric transfer functions from 81 stations in 52 the vicinity of the hypocentre (Figure 1) from the publicly available SAMTEX 53 dataset[23], and invert them using the 3D inversion methodology described in 54 [24] that includes correction for static distortion [25]. We select data at 24 pe-55 riods between 1 s and 600 s corresponding to depths between 5 km and 80km 56 as confirmed by sensitivity tests (see supplementary material). We start the 57 inversion with a high smoothing regularization term to recover the broad con-58 ductivity structure and successively lower the weight of the regularization until 59 we achieve an adequate fit to the observed data. The final inversion model 60 (Figures 2 and 3) explains the data to a RMS of 1.3 assuming an error floor of 61 2% of the Berdichevsky invariant of the impedances. This choice of error floor 62 down-weights small diagonal elements in the inversion, but has the advantage 63 of making the misfit rotationally invariant. Still we observe an excellent fit to 64 all elements of the impedance tensor. 65

Figure 4 shows a representative selection of magnetotelluric data and the associated model fits. Stations 24 and 25 (top row in Figure 4) are located closest to the epicentre of the 3rd April event and show excellent fit for all components across the whole frequency range. Station 5 (bottom left in Figure 4) is located close to the 12th August event and shows a good fit for all components and frequencies. Station 79 (bottom right in Figure 4) is the site with highest



Figure 2: Horizontal slices through our preferred resistivity model. The magnetotelluric stations used in the inversion are marked as black squares. We show the location and focal mechanism of the April recent earthquake as well as the location of seismicity in the area (blue dots) and the August 2017 magnitude 5 event (red dot). Solid black lines show the boundaries of tectonic units with additional crustal units from [26] (dashed lines). Thick black bars show the direction of maximum horizontal stress from the world stress map[11].



Figure 3: Cut through the preferred resistivity model viewed from the South-West. MT measurement site locations are marked by black squares and the location of the hypocentres are marked by dots. The main April 2017 event is marked in red and subsequent events in white. For the main event we plot the preferred fault plane from the moment tensor solution and D-InSAR modelling [27] as a transparent plane. The white lines mark the boundaries of major crustal units [26]

RMS misfit of all sites considered in the inversion. Note that the apparent resistivities of the two off-diagonal components of impedance differ by two orders of magnitude, an indication of strong static distortion. Despite this, we achieve a reasonable fit to the observed data even though some of the more subtle features are not reproduced by the model.

The surface wave inversion uses a two stage approach to generate the to-77 mographic models similar to the methodology outlined in [28]. The Rayleigh 78 wave portion of the seismograms (periods of 50 - 120 s) are inverted to find the 79 average 1D shear (S_v) velocity structure between source and receiver. In the ap-80 proach taken here, for each waveform inversion four different starting models are 81 used that incorporate prior information on crustal, and long wavelength man-82 tle structure (see e.g., [28]). A particular advantage of incorporating the prior 83 mantle structure in the starting models is that for the upper mantle there are 84 significant differences between the general structure of oceans and continents. 85 Using a 1D radially averaged starting model (such as PREM or ak135) will limit 86 the recovery of the amplitude of anomalies beneath the different regions due to 87 the necessary regularisation in the waveform inversion. 88

The resulting 1D velocity models are then combined to produce tomographic 89 images, as a series of depth slices at 25 km intervals, of the lateral velocity vari-٩n ations within the upper mantle. To improve the reliability of these tomographic 91 models, data from closely adjacent paths are clustered, this has the benefit of 92 limiting the impact of 1D models that are not consistent with adjacent results, 93 and somewhat downweighting areas that would be dominated by path coverage 94 in one particular direction. For the recovery of the variations in velocity there 95 are two steps in the inversion. Initially, a strongly damped inversion using over 96 > 45,000 1D models, is performed to recover the longest wavelength structure. 97 Subsequently, the tomographic model is updated through an inversion using 98 a parameterisation with knot points at 3-degree intervals. This intermediate 99 stage provides good recovery of structures such as the mid ocean ridges and 100 subduction zones, and therefore minimises the potential for these velocity fea-101 tures to be smeared into the final model. For the final inversion focused on 102



Figure 4: The fit of the final conductivity model for four selected MT sites. Stations 24 and 25 (top row) are the two stations closest to the main event. Station 5 (bottom left) is the MT site closest to the August 12th event. Station 79 (bottom right) has the highest misfit of all inverted MT stations. These sites are marked by red squares in Figure 2.

southern and east Africa a subset of paths is included, quantile-quantile plots 103 are used to remove outliers (further limiting the impact of data that cannot be 104 fit in the inversion procedure), and almost 19,500 paths are incorporated into 105 the tomography (see Figure 5). The final models for each depth slice are chosen 106 based on the trade off between data fit and a model norm regularisation. In this 107 approach to regularisation, the specific choice of damping has an impact on the 108 amplitude of the velocity anomalies, however the spatial location of variations in 109 velocity in the resulting models remain consistent. Checkerboard tests illustrate 110 that the path coverage in the region is sufficient to recover structures around 111 300 km in diameter with limited smearing (see Supplementary Information for 112 associated figures). 113



Figure 5: Events (red circles), stations (triangles) and path coverage (gray lines) used to construct the seismic surface wave model. The light gray paths show the coverage for the large scale model which is used as starting model for the regional model shown here. Dark gray paths indicate coverage for the final regional model.

114 **3. Crustal structure**

Given the heterogeneous data coverage with the MT sites located on profiles along accessible roads, we focus our discussion of the resistivity model on structures close to these profiles. Our sensitivity tests demonstrate though that we have some sensitivity to off-profile structures (see Supplementary material). We estimate that we can recover structures up 1.5 skin depths in lateral direction and blank the areas in Figure 2 where we do not have sensitivity.

At depths between 15 km and 40 km, most of the significant conductors 121 (resistivity $\rho < 25 \ \Omega m$) terminate at geological boundaries (Figure 2). This is 122 particularly evident for anomalies A, B, and C, but other conductive structures 123 also show the same pattern. Resistive structure D, which emerges at a depth of 124 30-40 km, is bound on both sides by the inferred boundaries of the Proterozoic-125 age Limpopo Belt. We note that the hypocentres of all seismic events in the 126 region are located at the boundary of conductive structures (see Figures 3 and 127 6).128

Conductors in the middle and lower crust can have a variety of origins depending on the geological setting. In strongly tectonically active areas they have been interpreted as accumulations of melt [e.g. 29]. This requires an unusually hot crust and thus can be ruled out in a stable continental setting. In such regions, enhanced conductivities at depths between 10 and 30 km are typically attributed to relatively small amounts of saline fluids [e.g. 30] or interconnected graphite and, to a lesser degree, sulphides [e.g. 31].

Regardless of which of these processes are considered, they all require the 136 conductive phase to be interconnected over distances of several kilometres in or-137 der to cause an observable increase in conductivity. For graphite and sulphides, 138 the simplest geological process to achieve such interconnectivity is deformation 139 along shear zones creating thin boundary films [32]. Consequently many con-140 ductivity anomalies in the middle to lower crust have been interpreted as signs 141 of significant deformation, particularly when there is strong variation in depth 142 to the conductor [16, 31, 33, 34, 35]. Where fluids are considered to be the 143

cause of enhanced conductivity, they are often thought to be trapped under 144 an impermeable layer in the middle crust [36]. Large faults can breach such a 145 seal and allow fluids to migrate upwards. This explanation has been invoked 146 to explain the observed conductivities of major active fault systems such as the 147 San Andreas Fault (SAF) [37, 38]. At the SAF, a deep (30-60 km) conductor is 148 interpreted as a fluid reservoir that feeds a more shallow fault related fracture 149 zone imaged as a narrow vertical conductor. Similar images and interpreta-150 tions have been obtained in other active fault zones, e.g. the North Anatolian 151 fault and the Niigata-Kobe Tectonic Zone in Japan [38]. In regions without sig-152 nificant ongoing tectonic activities, interconnected graphite is usually favoured 153 as an explanation for fault related conductivity as fluids migrate upwards over 154 geological time scales [33] and the deep conductor found in active regions ap-155 pears to be missing. However, fluids can assist in transporting graphite during 156 deformation and contribute to the formation of connected films [33]. 157

Of particular interest for our study is the conductive structure associated 158 with the hypocentre of the main event. Figure 3 shows a 3D cutout view of the 159 preferred resistivity model together with the preferred fault plane solution based 160 on the moment tensor, and the Differential Interferometric Synthetic Aperture 161 Radar modelling of [27]. The inferred fault plane coincides with a significant 162 change in depth of the crustal conductor in this area. In the foot wall on the 163 western side its top is located at a depth of 14 km, whereas in the hanging wall to 164 the east, the conductor reaches the surface. Sensitivity tests (see supplementary 165 material) demonstrate that we have good resolution to the depth of the deep 166 conductor, and that the top on the eastern side cannot be located deeper than 167 7 km. 168

Considering the above discussion of causes for high conductivity in fault zones, these structures could be a direct expression of fault related deformation or could be an originally continuous structure that has been displaced by movement on the fault. Given the spacing between the MT sites (20 km), we cannot directly image the fault zone, which is at most hundreds of meters wide. Instead we image the effect of the fault on the surrounding structures. Based



Figure 6: Magnified view of the model shown in Figure 2 around the earthquake sequence. The magnetotelluric stations in the area are marked as black squares. We show the location and focal mechanism of the April recent earthquake as well as the locations of seismicity in the area from the USGS catalogue (blue dots), the precursor locations determined in [10] (blue squares).

on the published estimates of the geometry of the fault for this event [27, 10], it 175 is unlikely that we are imaging fluid pathways or shear signatures caused by the 176 currently active fault. Instead it is more plausible that the two conductors were 177 originally at the same depth and subsequently displaced by movement along the 178 fault. However, the sense of motion necessary to produce such a displacement is 179 opposite to the observed current fault motion. Thus our preferred interpretation 180 is that the earthquake reactivated an existing thrust fault associated with the 181 deformation associated with the collision of the Kaapvaal and Zimbabwe Cra-182 tons. This interpretation is consistent with other observations [27] and similar 183 interpretations have been made for other paleo-faults [34]. 184

The reactivation of an existing fault fits well with other studies of intra-plate earthquake nucleation [8]. However, the question remains to which degree the mid-crustal event corresponds also to deeper regional structure? In particular, can we identify a fluid reservoir that corroborates the hypothesis that this event
was triggered by fluid released from the manite [10]?

¹⁹⁰ 4. Upper mantle structure

In the context of deeper regional structure, Figure 7 shows the S_v velocity 191 for the region of southern Africa at depths of 75 km and 175 km (top row) 192 and for the study area (bottom row), together with heat flow measurements 193 [39], the directions of maximum horizontal stress [11] and relative plate motion 194 from GPS data [12]. At 75 km depth the areas of the Kaapvaal and Zimbabwe 195 cratons are clearly marked by high velocities ($v_s > 4.6 \text{ km/s}$), as expected for 196 cold cratonic mantle. Similar fast velocities are observed beneath other areas of 197 Archean age, e.g., the Tanzanian Craton, and fragments of the Congo Craton 198 such as the Kazai shield. In the vicinity of the Botswanan earthquake we ob-199 serve a low velocity structure ($v_s \approx 4.4 \text{ km/s}$) at 75 km trending NW-SE and 200 with a velocity minimum in the region of the earthquake. In contrast, at 175 km 201 depth, fast velocities ($v_s > 4.6 \text{ km/s}$) typical of thick continental lithosphere are 202 observed across a broader region of much of southeastern Botswana consistent 203 with features observed in global tomographic models [40]. While low velocity 204 zones in the upper mantle can represent zones of high temperature, and po-205 tentially partial melting, the underlying faster velocities make this explanation 206 untenable. Although the heat flow measurements are moderately high (40-60 207 mW/m^2) away from the Kaapvaal and Zimbabwe Cratons [39, 41], the spatial 208 variability and lack of correlation with velocities at 175 km depth, suggest a pre-209 dominate crustal control on heat flow rather than variations due to lithospheric 210 thickness. 211

Examining our resistivity model between 40 and 75 km (Figure 2), we see that the deep parts are generally resistive at depth with most parts exceeding resistivities of 500 Ω m. Based on sensitivity tests (Supplementary material), we conclude that our data do not indicate a significant difference in resistivity between the Limpopo Belt and the surrounding Cratons at this depth and assume



Figure 7: Horizontal slices through our regional surface wave model at depths of 75 km and 175 km, respectively. The top row shows the wider southern African context, while the bottom row shows the region around the earthquake. In addition to stress orientations [11] (black bars), we also show the movement relative to the Nubian plate [12] (blue arrow) and heat flow measurements [39] (coloured dots) in the area. The seismic stations in the plotted area are shown as red triangles.

values of 200 $-1000 \ \Omega m$ as representative. Similar resistivities at these depths have also been observed in studies of the surrounding areas [16].

Dry Archean lithospheric mantle is expected to show resistivies in excess of 219 $10,000 \ \Omega m$ based on laboratory experiments within the typical compositional 220 variations between Lherzolite and Harzburgite [23]. Such high resistivities are 221 observed at the cores of the Kaapvaal Craton [42] and Congo Craton [16] at 222 depths between 100-200 km, and in parts of the Slave Craton [43]. These high 223 resistivity areas also show S-wave velocities exceeding 4.6 km/s in the seismic 224 velocity model as expected for old lithosphere. The resistivity values we observe 225 cannot be explained by a dry mantle, but match the range of resistivities of 500-226 2000 Ω m estimated at this depth for typical mantle compositions with a water 227 content of 150 ppm [44]. Such a water content agrees well with the estimated 228 average water content of the lithospheric mantle [45]. Calculations of S-wave 229 velocity for a range of compositions and temperature profiles predict values 230 in excess of 4.5 km/s at a depth of 80 km [46] which matches the values we 231 observe towards the south, in the Kaapvaal Craton, but is significantly higher 232 than the velocities recovered around the epicentre. So, while the resistivity 233 model indicates a relatively homogeneous, normal lithospheric mantle structure, 234 the seismic model requires a strong change in physical properties between the 235 cratons in the south and the region of the epicentre. 236

237 5. Discussion and conclusions

It has been suggested that the event was triggered by fluid release from the 238 mantle bringing a critically loaded fault network to failure [10]. The crustal 239 structure in our resistivity model is compatible with such a scenario. As ex-240 plained above, we cannot directly image the fault zone as this would require 241 denser site spacing near the fault and higher frequency data than what is cur-242 243 rently available. The two displaced conductive structures could be fluid related although this would require some form of seal to prevent those fluids from mi-244 grating upwards. For this reason we consider an explanation in terms of graphite 245

²⁴⁶ more likely. Even if the high conductivity in the crust is at least partially caused ²⁴⁷ by saline fluids, these structures cannot be the source for the fluid pulse that ²⁴⁸ triggered the event, as the epicentre is located below these conductors at the ²⁴⁹ transition to more resistive material.

A major region of elevated fluid content in the mantle would manifest itself 250 as a region of high conductivity [37]. We do not see such structures in our model. 251 In fact, the lack of strong variation in resistivity in the upper mantle underneath 252 the study area, suggests a homogeneous thermal structure and water content 253 as these are the two major controlling factors on resistivity in the nominally 254 anhydrous minerals (NAMs) of the lithosphere [47]. Therefore, either the source 255 region of the fluids is spatially restricted (less than a few kilometres in diameter), 256 the fluids are derived from moderate amounts of ambient water in the mantle or 257 another triggering mechanism is responsible. Based on our results, we cannot 258 distinguish between these alternatives. Thus, although our model does not show 259 the expected features of a mantle fluid reservoir, we cannot refute the hypothesis 260 put forward by [10]. 261

We will now focus the discussion on the potential origins of the low velocity 262 zone at 80 km depth. Variations in temperature or water content would result 263 in observable resistivity variations [47]. We can therefore exclude these two 264 variables as an explanation for the low velocities. Furthermore, both have a 265 similar effect on resistivity and seismic velocity and thus an increase in temper-266 ature accompanied by a decrease in water content or vice versa is not feasible 267 either. This leaves two possible explanations for a decrease in velocity that is 268 not accompanied by a change in resistivity: i) Variations in mantle composition 269 and ii) variations in grain size of olivine. A bulk compositional change compat-270 ible with our observations would have to maintain iron content (or equivalently 271 Magnesium number: Mg#) as variations in Mg# have observable effects on 272 conductivity [23]. 273

Compositional explanations for low velocities in the uppermost lithosphere
have been discussed previously. [48] suggested qualitatively that paragasitic amphiboles could contribute to lowering velocities in central Australia, in a region

of thick lithosphere, but noted that this would require a complicated layered 277 structure with no clear mechanism of formation. The presence of chrome, thus 278 lowering the depth of the spinel transition, has been invoked as a possible expla-279 nation for the velocity gradients seen in Precambrian lithosphere of a number 280 of areas [49]. Modelling of phase velocity profiles for cratonic regions also indi-281 cated that models of constant composition have a systematic variation from the 282 seismic data [50] and further studies using these data indicate that a metaso-283 matic component (water or carbonate fluids) improve the fit to the seismological 284 observations [51]. However, the velocity variations observed in our study region 28 have larger variations than those modelled in [51]. 286

The idea of enhanced concentrations of amphibole, has been revisited, and 287 invoked to explain low velocities at a similar depth range and magnitude in-288 ferred from S-receiver functions [52]. The electrical resistivity of amphiboles 289 at upper mantle conditions is currently unclear, but laboratory measurements 290 under lower crustal conditions suggest a significant decrease in resistivity from 291 amphibole enrichment [53]. We therefore cannot rule out amphibole as a source 292 of the observed low velocities, but consider the high concentrations ($\sim 20\%$) in-293 voked by [52] to explain a similar magnitude low velocity anomaly improbable. 294

Variations in grain size have been shown to affect seismic velocities in the 205 mantle and a reduction in size from approximately 1 cm below the cratons 296 to several millimeters below the mobile belt is sufficient to explain the lower 297 seismic velocities below the Limpopo mobile belt [54]. Such sizes are consistent 298 with estimated values in undeformed cratonic lithosphere and deformed mobile 299 belts, respectively [55]. Electrical resistivity shows dependence on grain size 300 for sizes below 1 mm, but for the range of sizes considered here is negligible 301 [47]. Deformation can result in a grain size reduction in the upper mantle 302 that can persist for several hundred million years [56]. Thus we consider a 303 reduced grain size below the Limpopo belt the most likely explanation for our 304 observations. Interestingly, our two most likely explanations, reduced grain size 305 and amphibole enrichment are typically observed in samples from mantle shear 306 zones[57]. Furthermore, a reduced grain size results in a reduced viscosity [55] 307

indicating that the low velocity zone underneath the Botswana earthquake isan expression of a weak mantle.

Our combined magnetotelluric and seismic study demonstrates that the re-310 cent Botswana earthquake sequence reactivated previous faults in the area. For 311 the main event, this reactivation occurs in the opposite sense to the original 312 fault movement. All events occur above a region of low velocities and relatively 313 high resistivities in the upper-most mantle that we interpret as a region of re-314 duced grain size and thus weaker material compared to its surroundings. The 315 observed extensional fault movement is compatible with the ambient stress pat-316 tern in southern Africa. Our results can neither confirm nor refute the proposed 317 triggering of the event by mantle derived fluids. We do however see signs of a rhe-318 ologically weak upper mantle. The lack of a significant deep lithospheric thermal 319 anomaly then suggests that this process is initiated from the top, through inter-320 action of the ambient stress field with ancient structures, rather than through 321 thermal weakening from below. 322

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